

# RXTE/PCA Calibration Status

Keith Jahoda

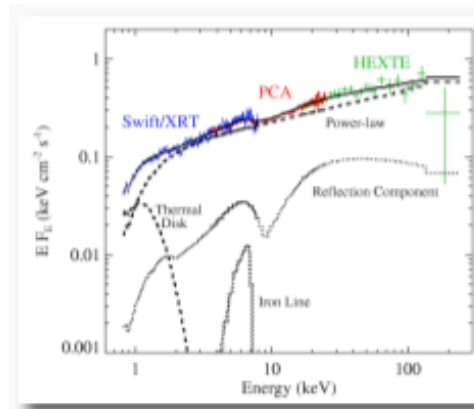
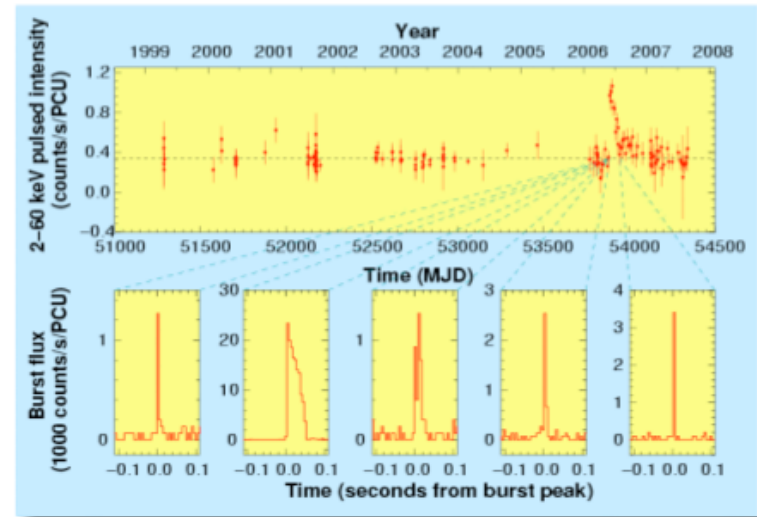
IACHEC, Schloss Ringberg 2008

# Outline

- PCA is a mature instrument undergoing a calibration update
  - Launched December 30, 1995
  - Operations currently approved through Feb 2009
- Previous presentations concentrate on our calibration successes
- This talk concentrates on what is left undone, and which may be left to the archival users
- Calibration makes substantial use of the power-law approximation to the Crab. (tomorrow we may learn what is wrong with this)

# RXTE/PCA characteristics

- Large area:  $\sim 7000 \text{ cm}^2$
- Dynamic range:
  - $\sim 10^{-11}$  to  $10^{-7} \text{ erg/s/cm}^2$
  - $\sim \mu\text{-sec}$  to years
- Flexible, “all” sky pointing enables multiwavelength campaigns



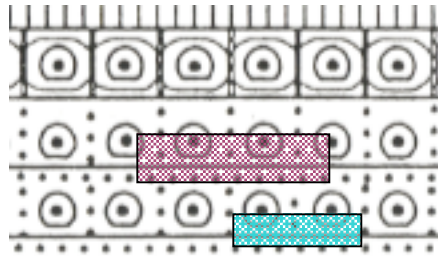
Smeared reflection features (in the hard state) with [Swift](#).

PSR 1846-0248, a rotation powered pulsar that shows magnetar behaviour. (Gavriil et al. 2008) - an example of an investigation that uses dynamic range in timing, scheduling flexibility

Simultaneous observations of Galactic Black Hole (GRO J1650-500)

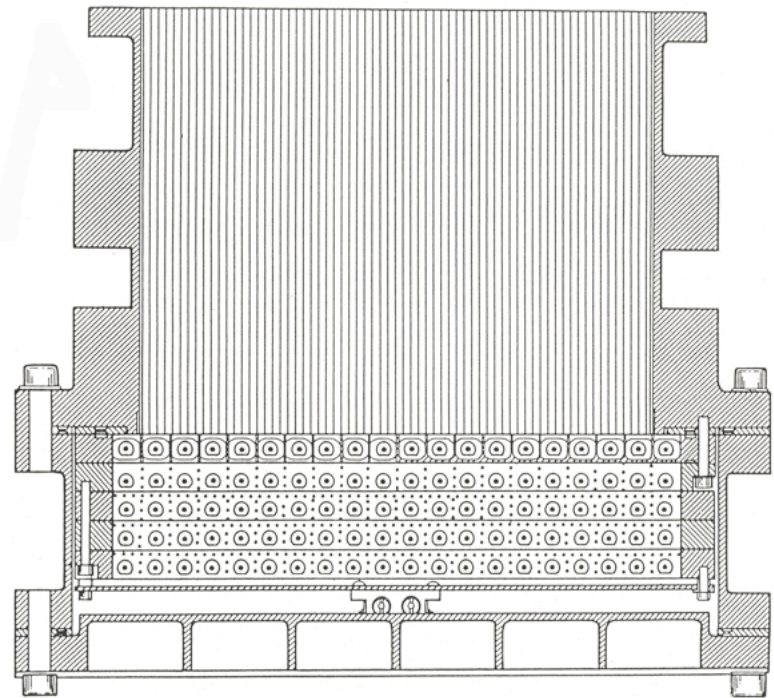
# Energy response

Matrix generator assumes slab elements



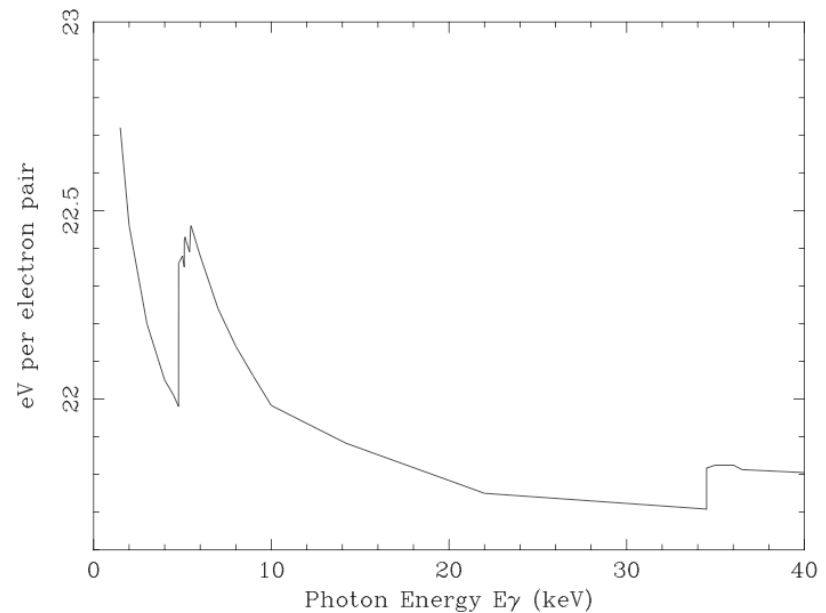
Interior window (25  $\mu$  mylar) spans  
~1.3 cm x 10 cm, supports 0.05 or 1  
atm.

Boundary between layers is formed by  
4 wire ground plane in each 1.3 cm  
cell

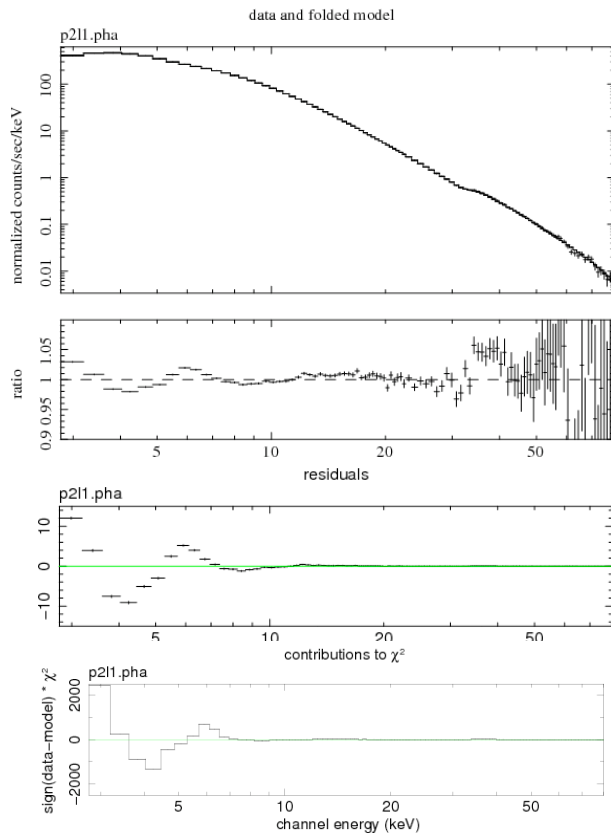


# Energy response

- Discontinuities in energy scale at L-edge are  $\sim$ half of best atomic theory.
  - Pcarmf reduces “contrast” in eV/ion pair by  $\sim 0.5$
- $\Delta E$  (L3,L2,L1) = 0.085, 0.032, 0.012 keV
- $\Delta E$  (K) = 0.18 keV



# Energy response



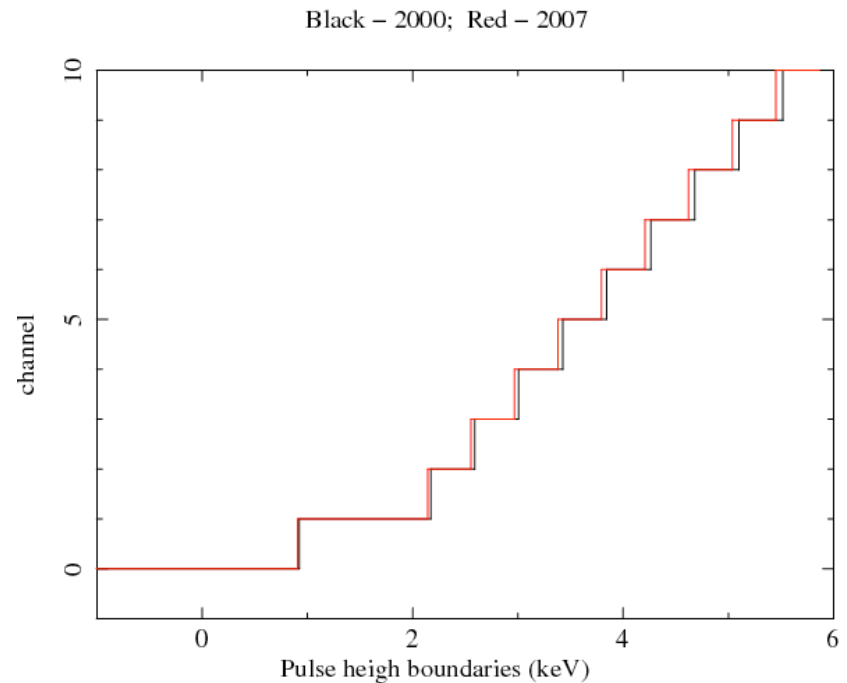
Power law fits to Crab show percent level deviations in “ratio”; “resid” and “chi” can be large (and are exposure dependent)

- fitting a gaussian (in “absorption”) near 4.1 keV typically produces an equivalent width  $\sim 50$  eV

At Xenon-L edge, there are discontinuous changes in quantum efficiency, partial charge collection, and energy to channel relationship.

# Energy response

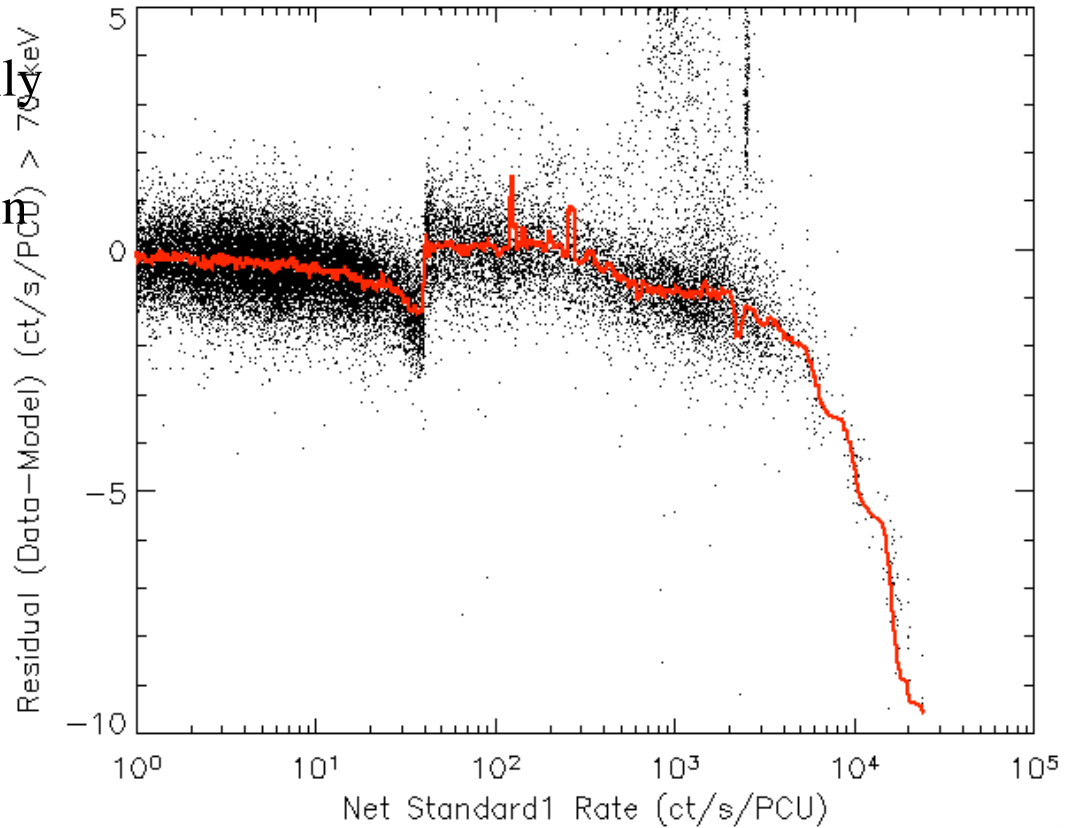
- Gradual shift in energy scale, combined with difficulties in low energy incomplete charge collection are a challenge
  - Fixed channel selection includes lower energies as time goes on; leads to higher reduced chi-square
- Self veto probabilities have no energy dependence



# PCA Background

Net rate above 70 keV typically expected to be near 0

Over-subtraction occurs in both faint and bright background models





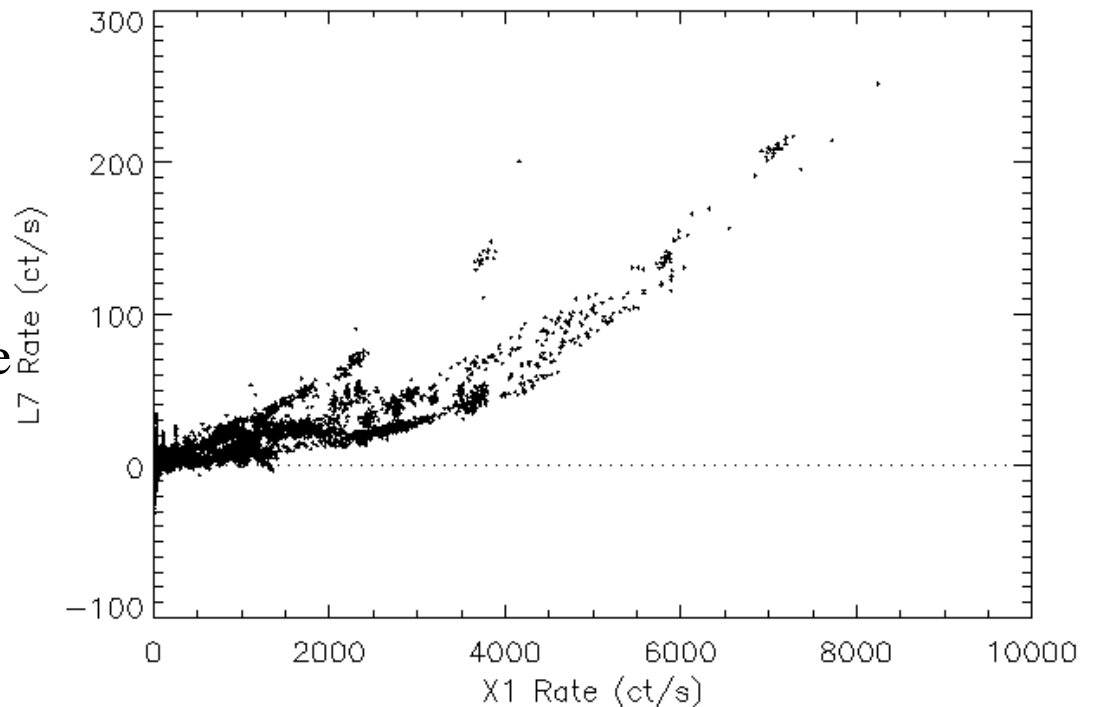
# PCA Background

Occultations allow measurements of L7 vs source rate

- source contributes to L7
- Dependence is probably spectral, and certainly more complex than just rate

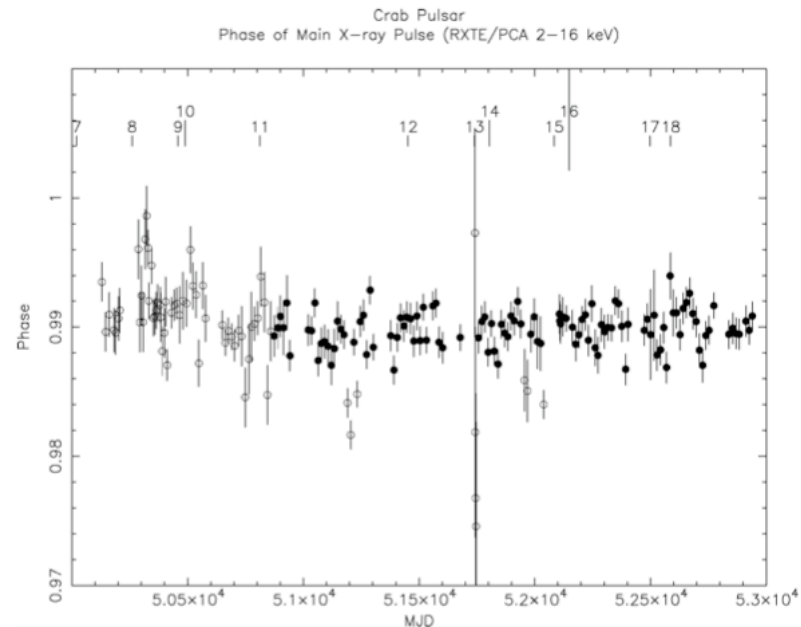
-Additional issues

- Statistical error in L7 (typ 10% in 16 sec) introduces excess variance on short time scales
- Tools report expected poisson error in bkg; should report systematic error in



# Absolute Timing Calibration

- Micro-second time tags
- Few micro-second uncertainty
- Statistical error in comparison of X-ray/radio phase for Crab now smaller than scatter



# Power Spectra

Assuming paralyzable deadtime

$$P_d(f) = P_1 - P_2 \cos\left(\frac{\pi f}{f_{Nyq}}\right)$$

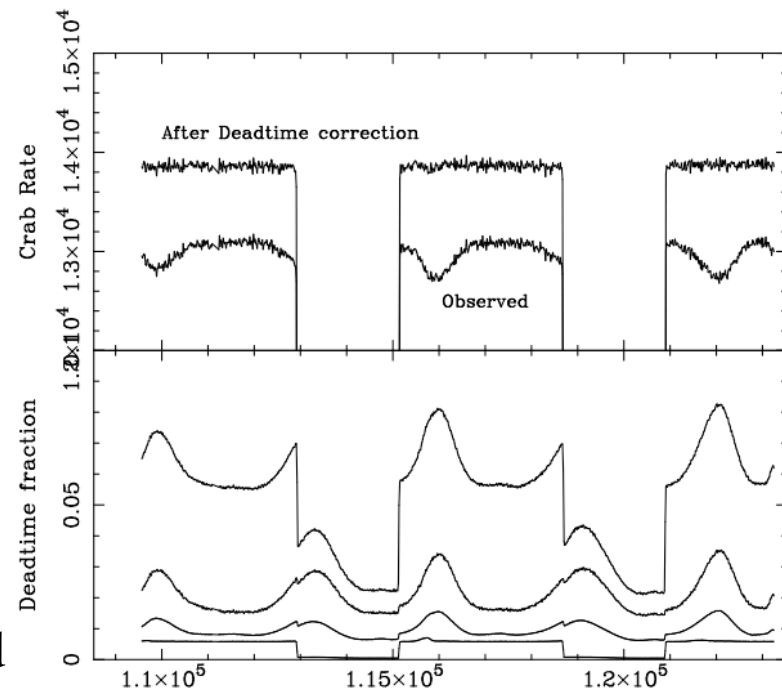
$$P_1 = 2\left[1 - 2r_0 t_d \left(1 - \frac{t_d}{2t_b}\right)\right] \text{ and}$$

$$P_2 = 2r_0 t_d \frac{N-1}{N} \left(\frac{t_d}{t_b}\right)$$

$$P_{vle}(f) = 2r_{vle} r_0 \tau^2 \left(\frac{\sin \pi \tau f}{\pi \tau f}\right)^2$$

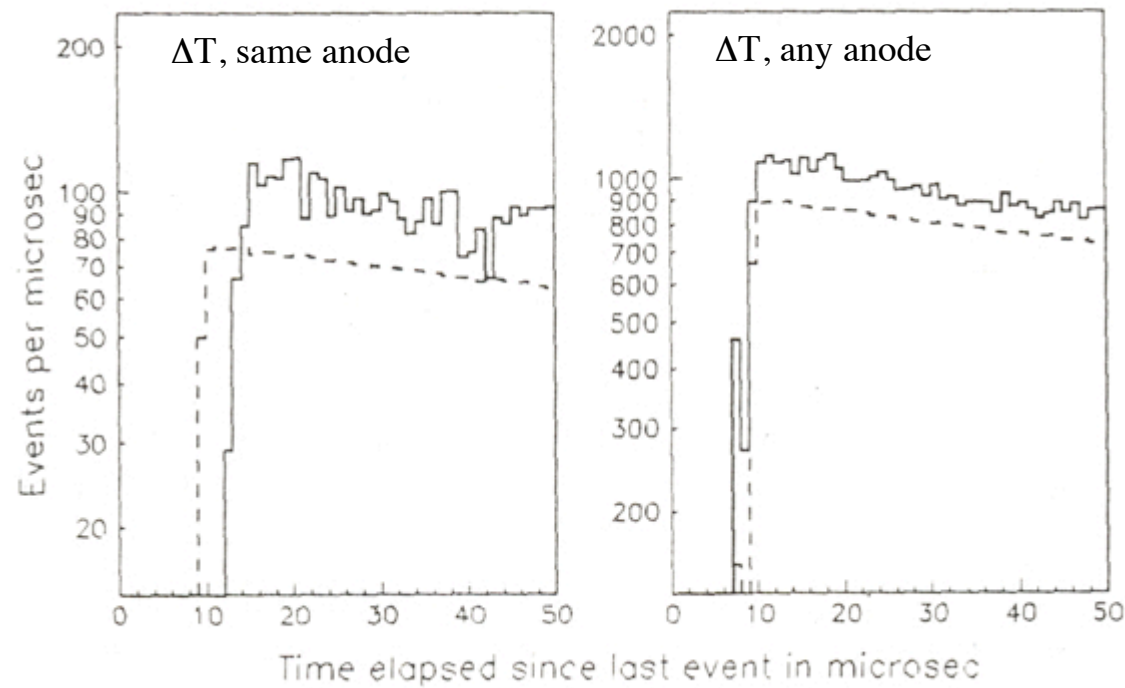
Fitting  $t_d, \tau$  to power spectra gives 8.8, 138  $\mu\text{sec}$  (Wei, 2006, MIT thesis)

But “flattening” the Crab light curve requires 10, 170  $\mu\text{sec}$  (Jahoda et al., 2006, ApJS)



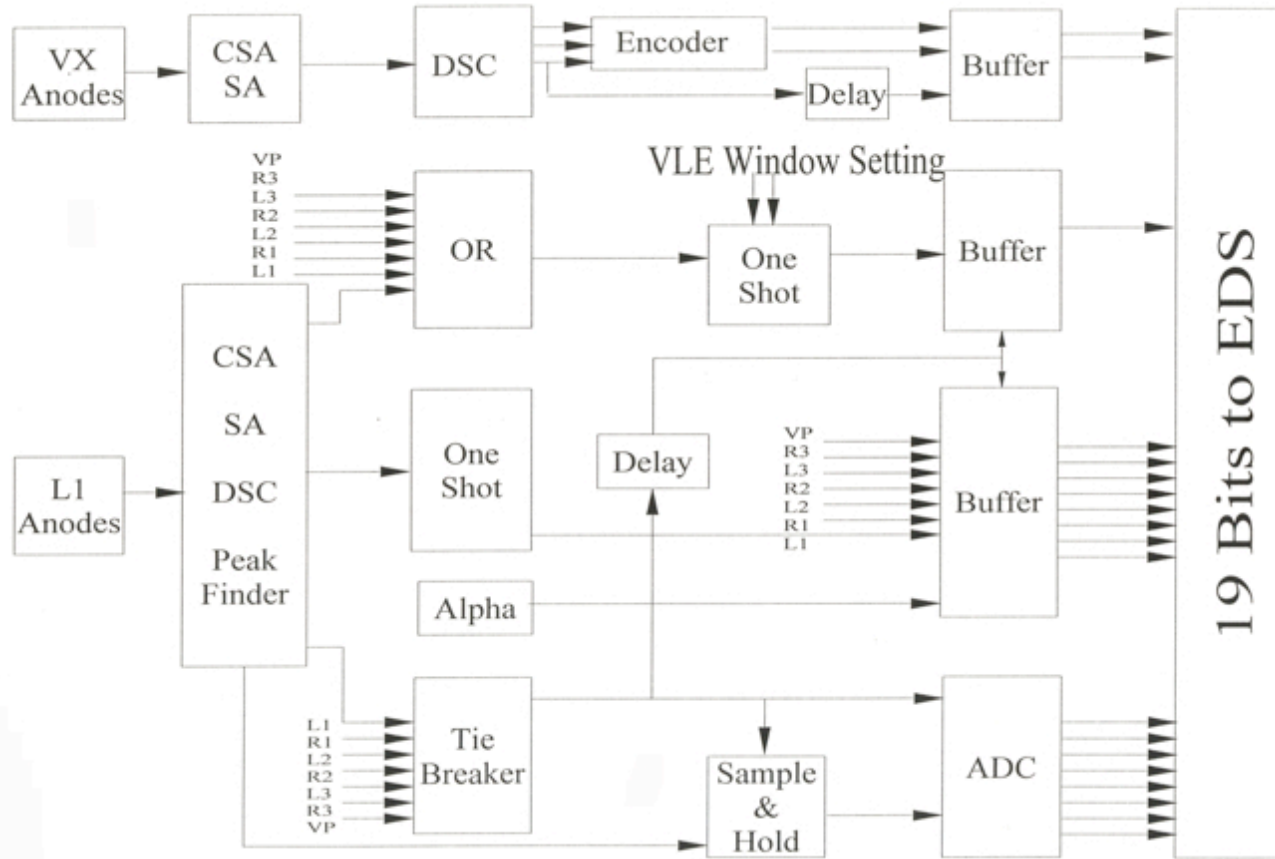
Suggests that the deadtime is complicated

# Energy Dependence of dead time



Dashed - 6 keV  
Solid - 22 keV

# Timing Logic

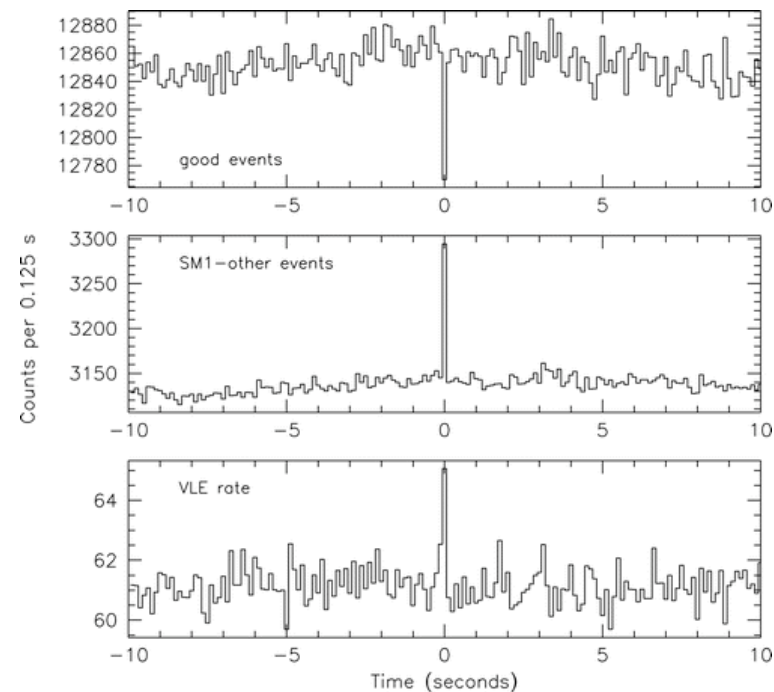
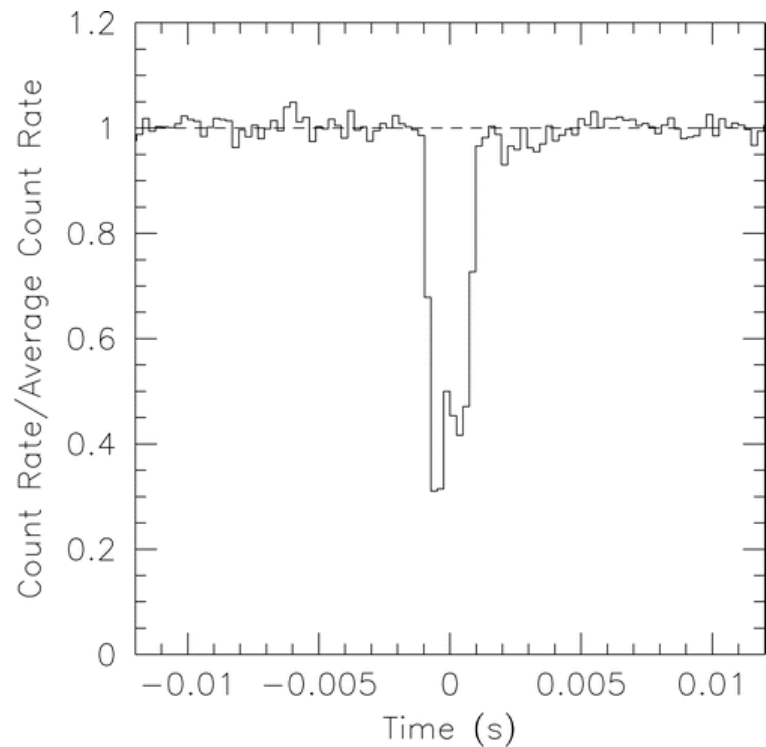


Paralyzable deadtime

Non-Paralyzable deadtime

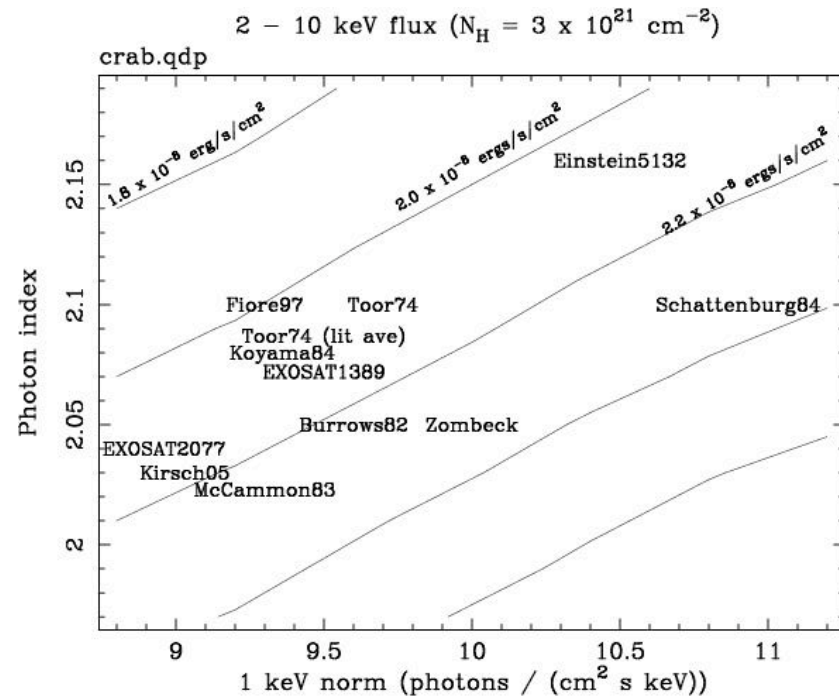
# Dips and TNO

Sco X-1 is in ecliptic, has many counts per msec, and might be occulted by km sized Trans Neptunian Objects (TNO). Chang et al. (astro-ph/0701851) find dips (!) but the dips appear to have in instrumental origin :(



# Collimator transmission

- Absolute area normalized to “Crab” (Zombeck value, which is high)
- Flux in 2-10 keV
- Crab flux plot
- No energy dependence in arf, likely a slight over simplification



Dead-Time of VLE Events

