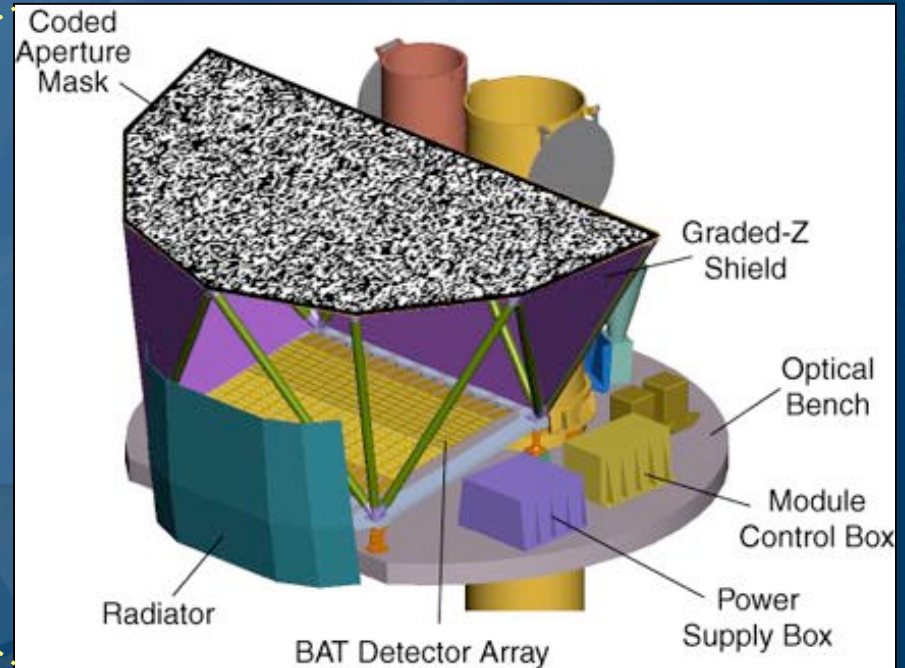
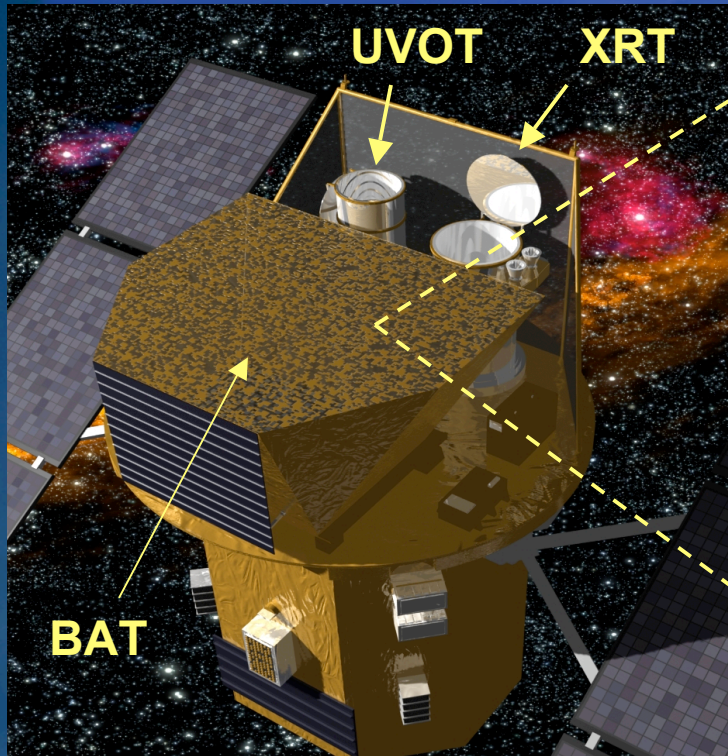

Energy Calibration Status of Swift/BAT

T. Sakamoto
on behalf of BAT team

Outline

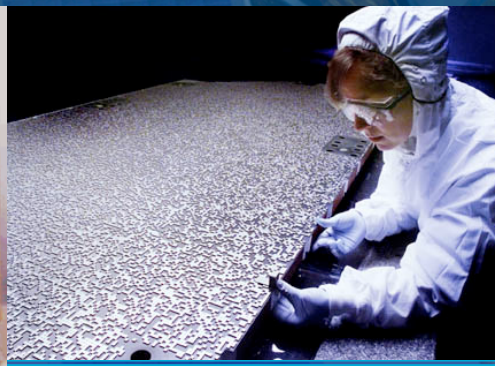
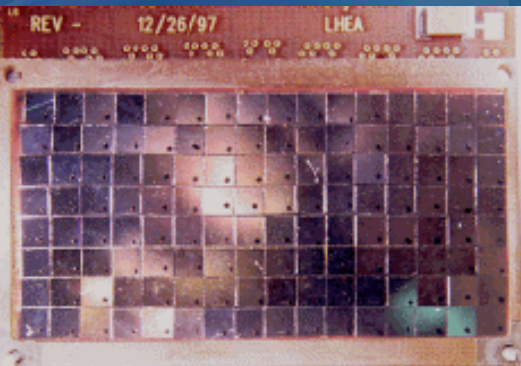
- Update on BAT energy response
- Cross-calibration work using GRBs:
Konus-Wind, Suzaku/WAM, and Swift/BAT
- Gain change in BAT (work in progress)

Burst Alert Telescope (BAT)



One sandwich (128 CdZnTe)

Coded Mask



Energy range	15-150 keV
Area	5240 cm ²
Detector	CdTeZn
Size of Det	4 x 4 x 2 mm ³
# of Det	32768
Energy resolution	~4 keV@60 keV
FOV	120° x 90°
Position Resolution	<3'

2009. 4. 27.

4th IACHEC Meeting

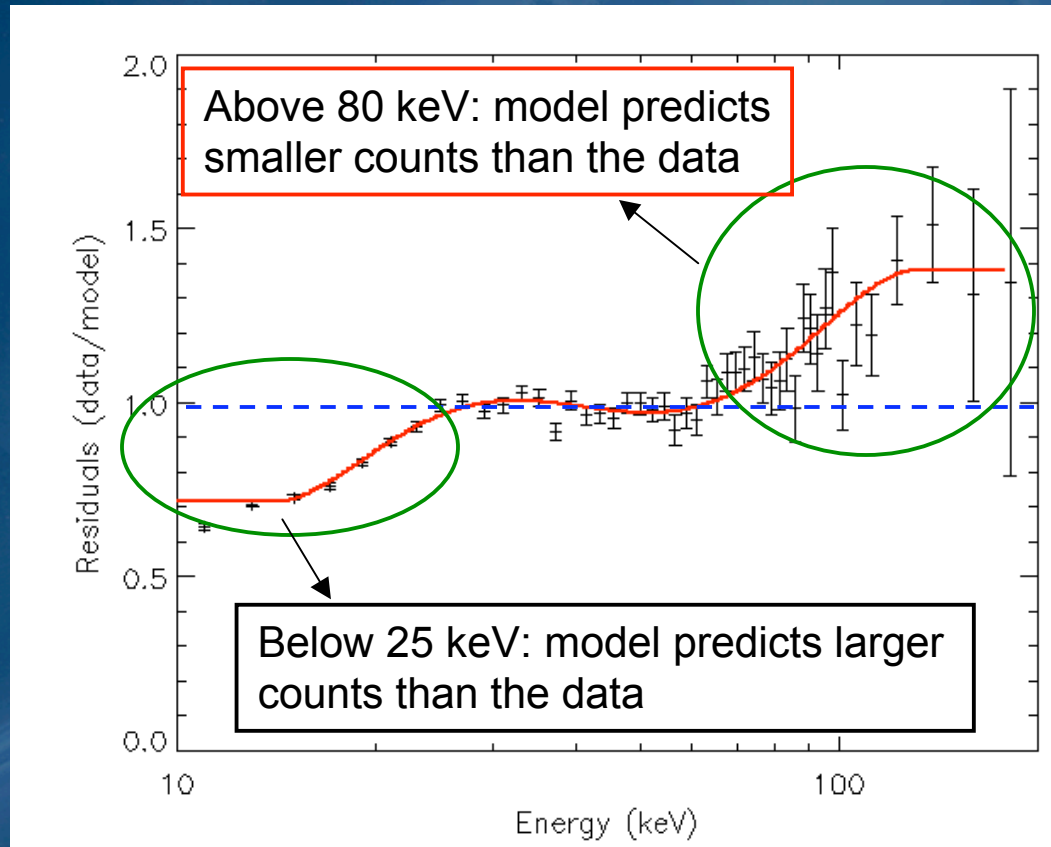
Swift

Update on BAT energy response

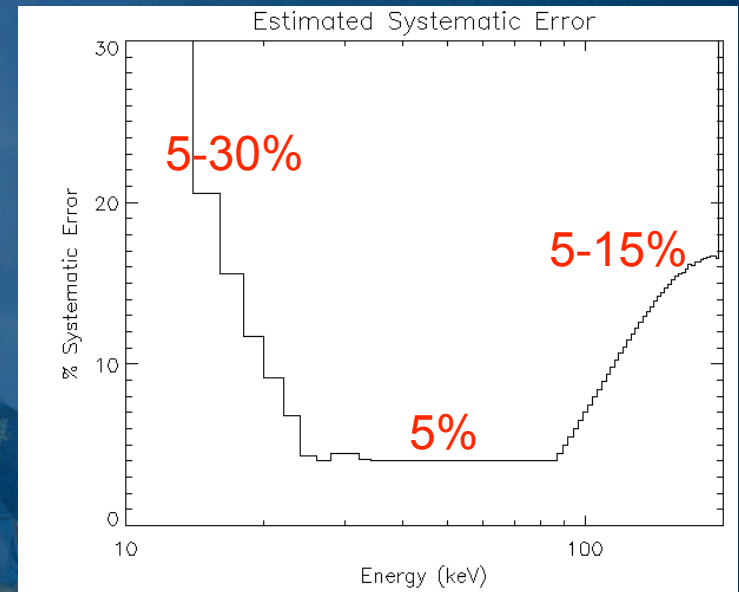
Fix in high-energy (>80 keV) part of the response

Previous energy response

Residuals (data/model) vs. Energy



Systematic error (batphasyserr)

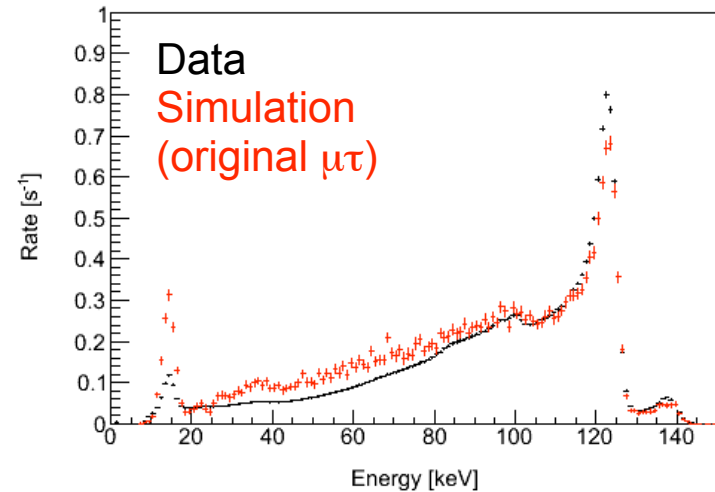
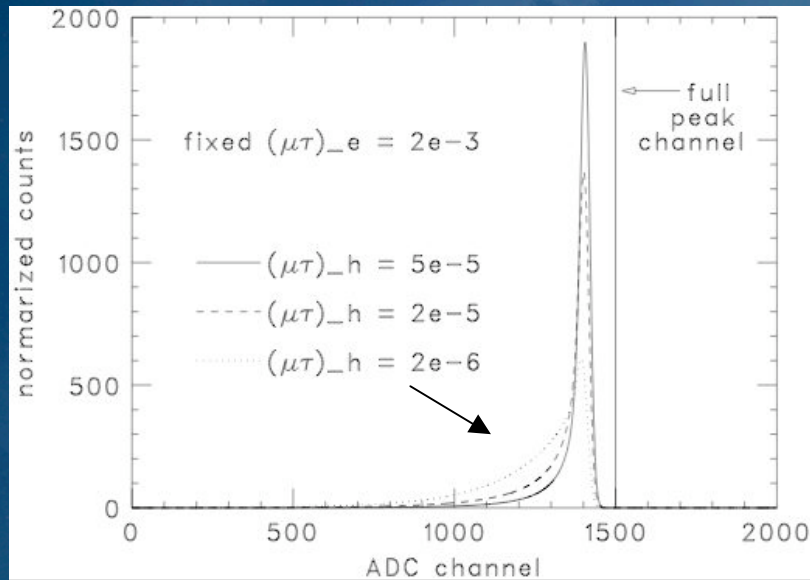


Two corrections in DRMs:

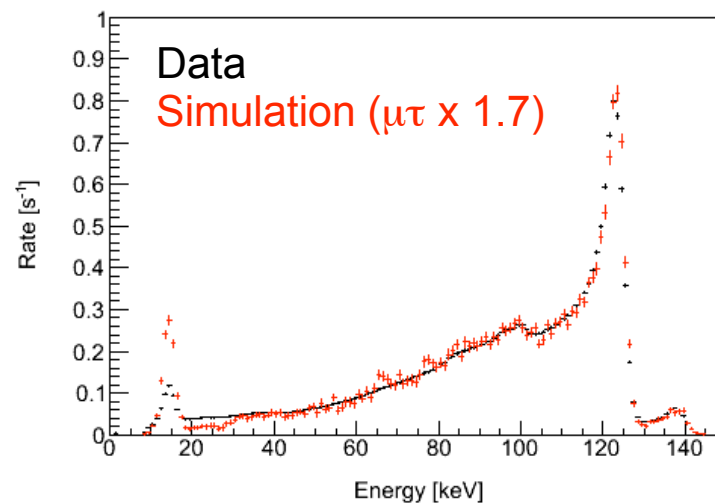
1. Correction at < 20 keV (Correct for unaccounted passive absorption in BAT field of view),
2. Correction at > 80 keV (adjustment in effective area).

Modification in $\mu\tau$ measurement (G. Sato)

Effect of $(\mu\tau)_{\text{hole}}$



↓ $(\mu\tau) \times 1.7$



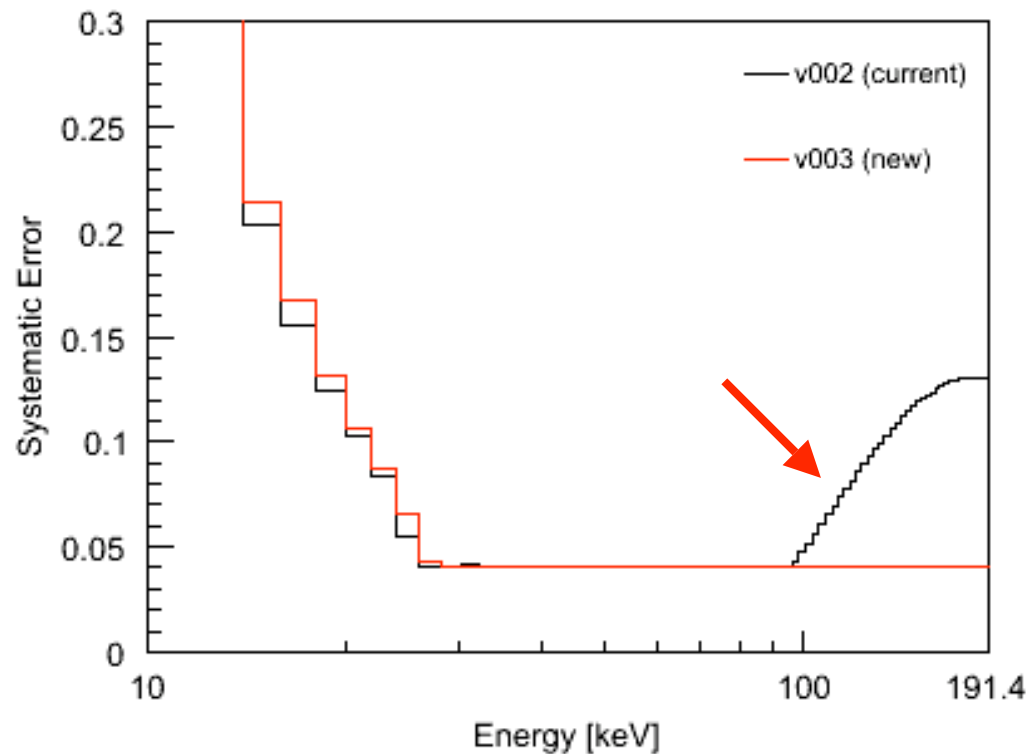
- $\mu\tau$ determination: ^{57}Co source
 - No scatter due to the source packaging
- MC simulation taking into account the material of the packaging
 - Large excess due to the scattering of the source packaging

$\times 1.7 \mu\tau$ is needed to reproduce the ground calibration data

Fix in response and systematic error

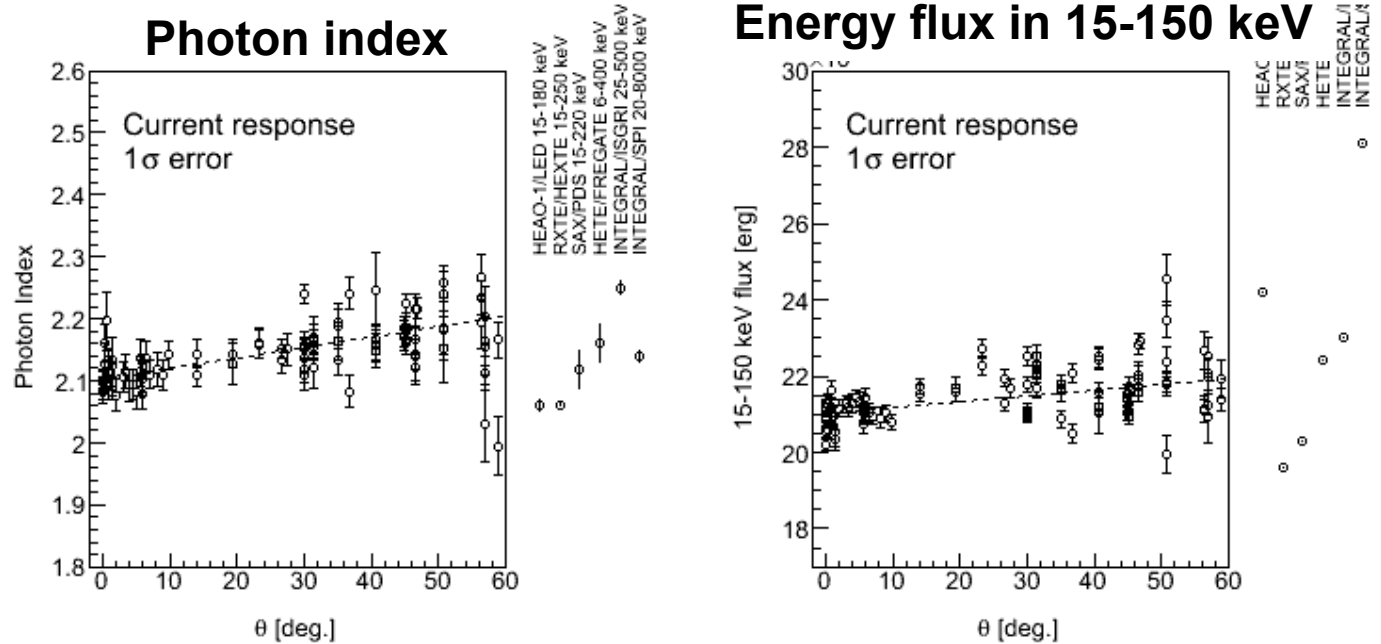
Fix in BAT energy response

- Multiplied by 1.7 of current $\mu\tau$ table (CALDB)
- Major update in systematic error vectors > 80 keV (CALDB)
 - Factor of 2-3 smaller at high energies!

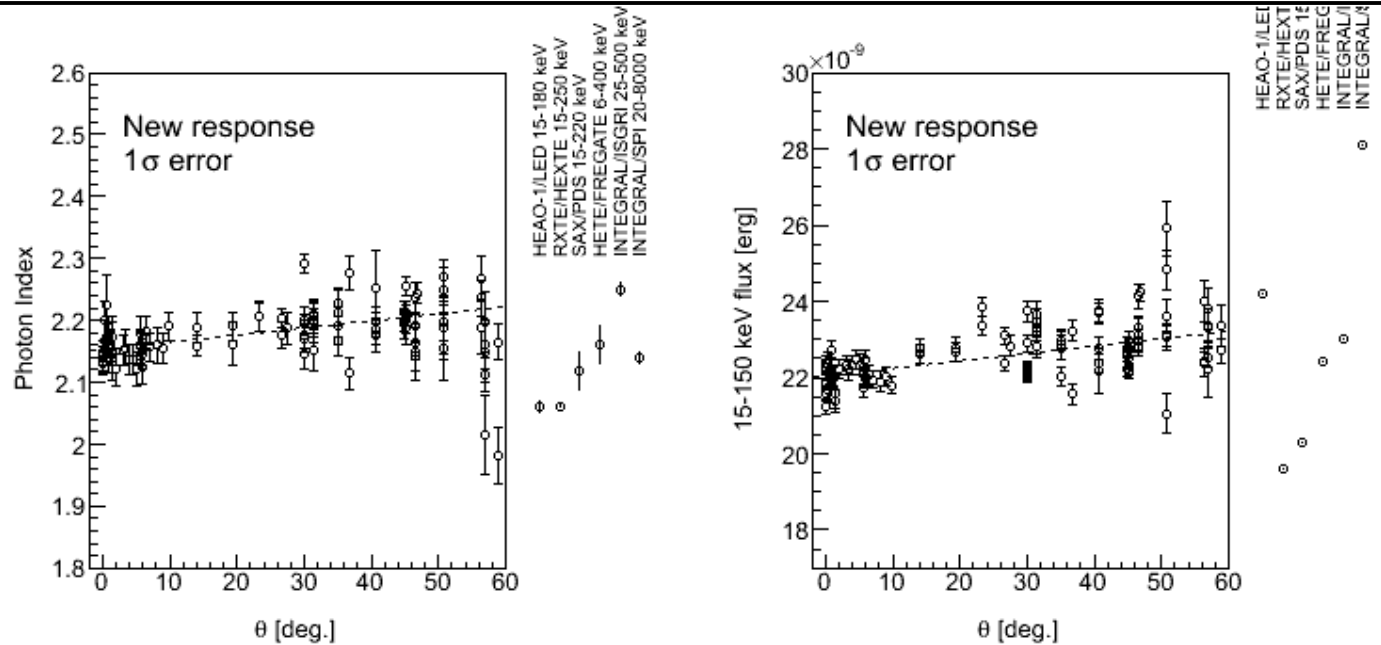


Crab re-analysis

Previous response



New response



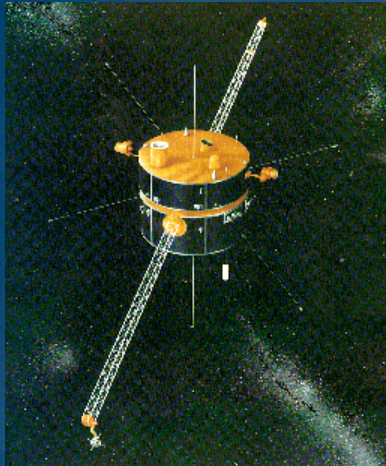
Summary

- High energy part (>80 keV) of the BAT energy response has been updated by increasing the values of $\mu\tau$.
- No correction at the high energy part anymore.
- Systematic error vector >80 keV has been reduced.
- This new response is available at the standard BAT ftools/CALDB.

Cross-calibration using GRBs:

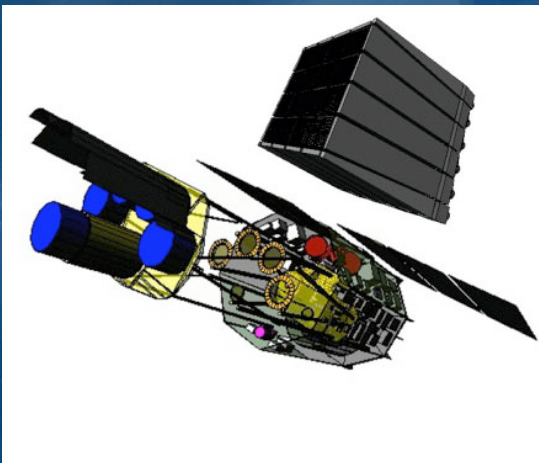
Konus-Wind, Suzaku/WAM and Swift/BAT

Konus-Wind, Suzaku/WAM and Swift/BAT



Konus-Wind

- L1 point (low background)
- Two NaI(Tl): 13 cm (d) x 7.5 cm (h)
- 10-750 keV and 0.2-10 MeV (63 channels)
- Time variable spectral intervals (64 ms- 8.192 s)
- >20 keV is calibrated



Suzaku-WAM

- BGO (400 cm²@1 MeV)
- 4 detectors
- 50 keV - 5 MeV (55 channels)
- 0.5 s spectra (1 s spectra)
- >120 keV is calibrated

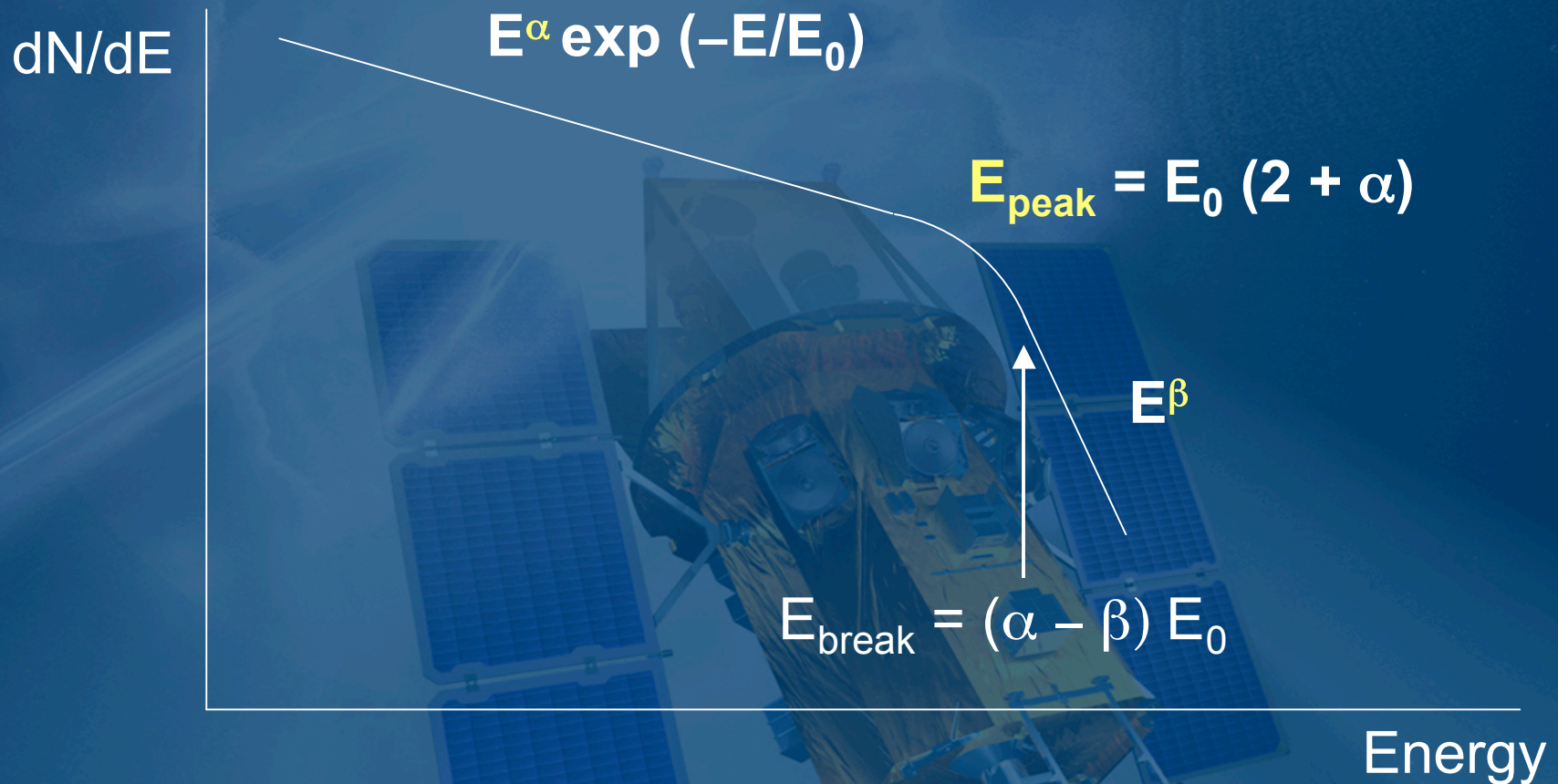
Swift-BAT



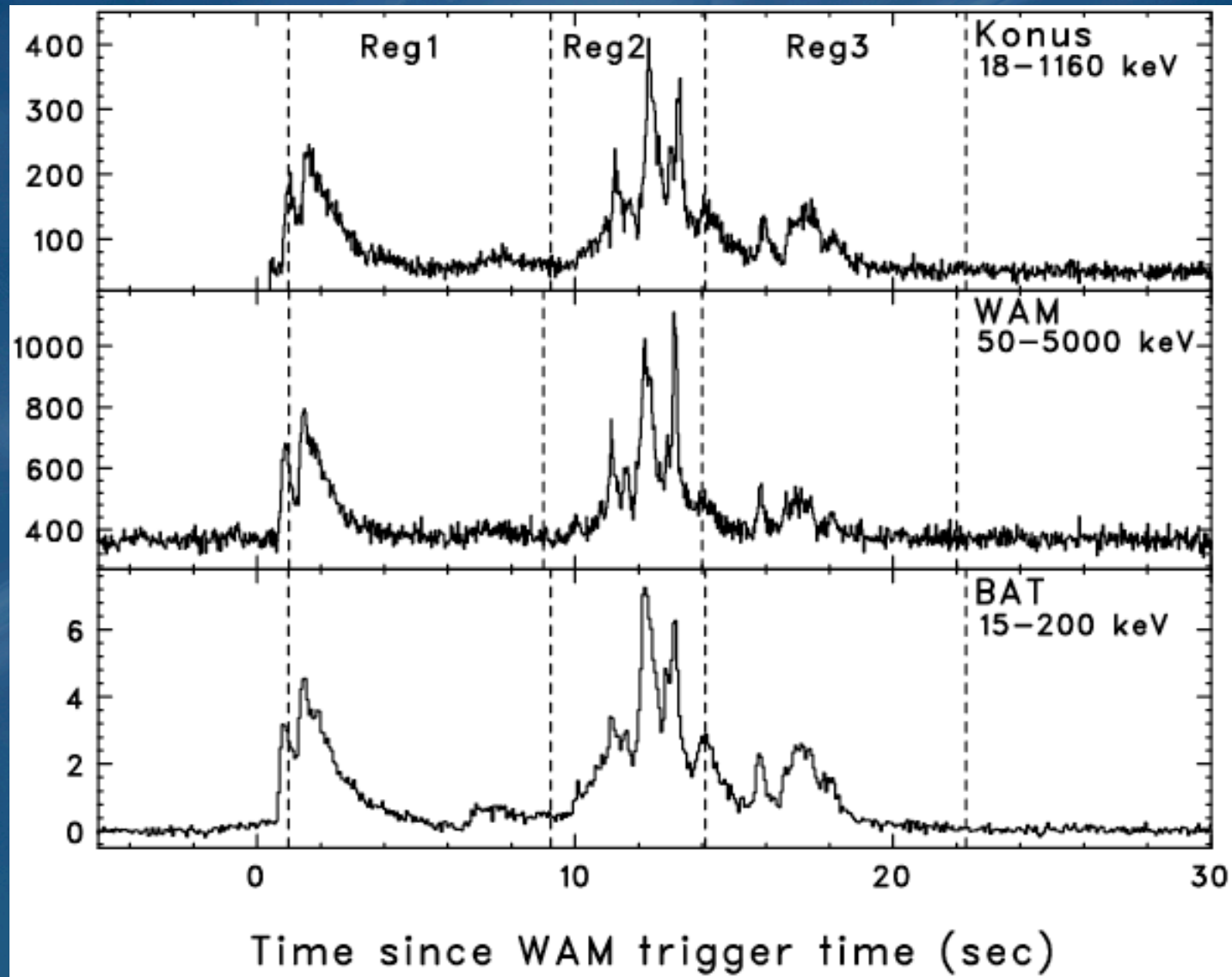
Empirical spectral models of GRBs

Band function

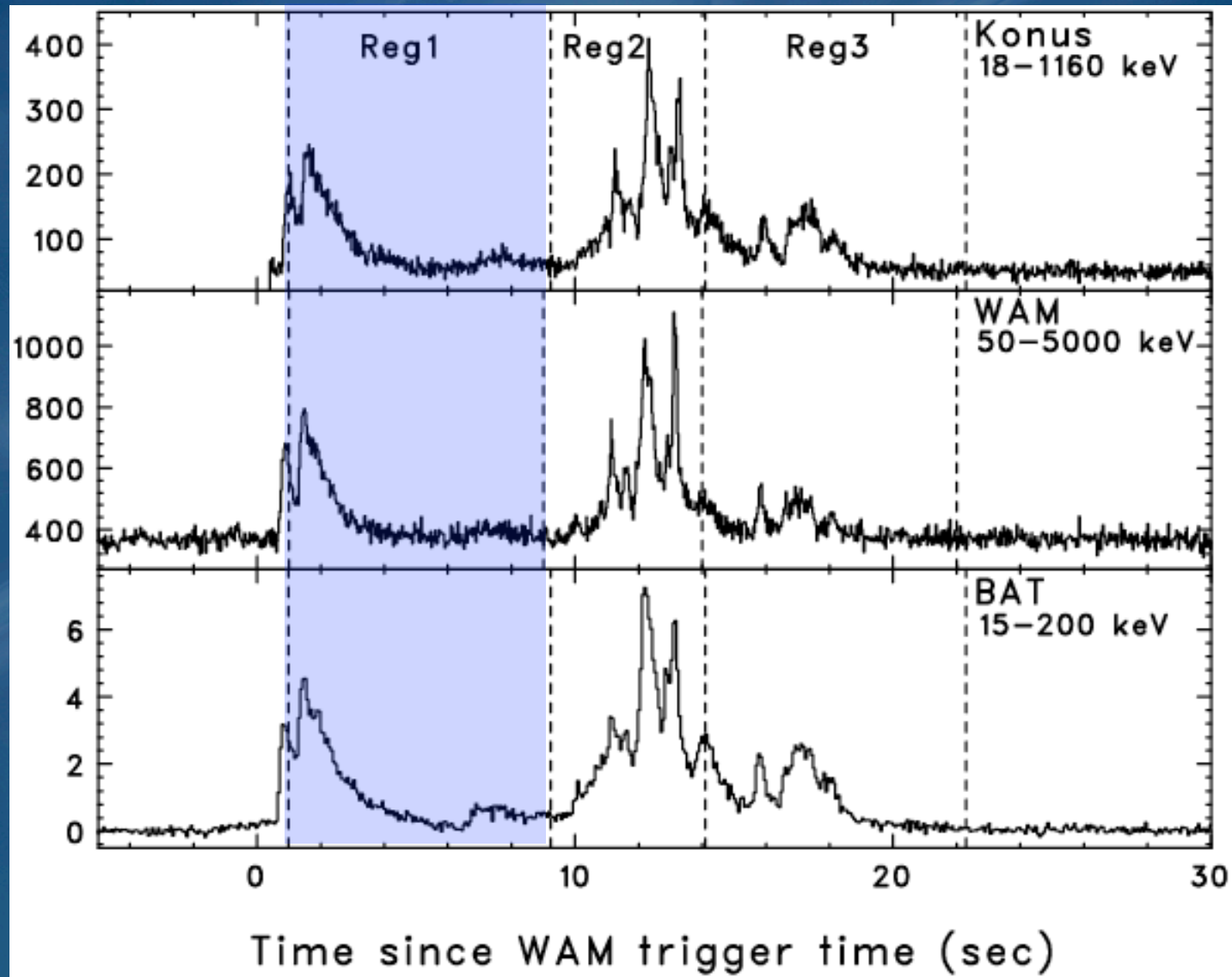
(Band et al. 1993)



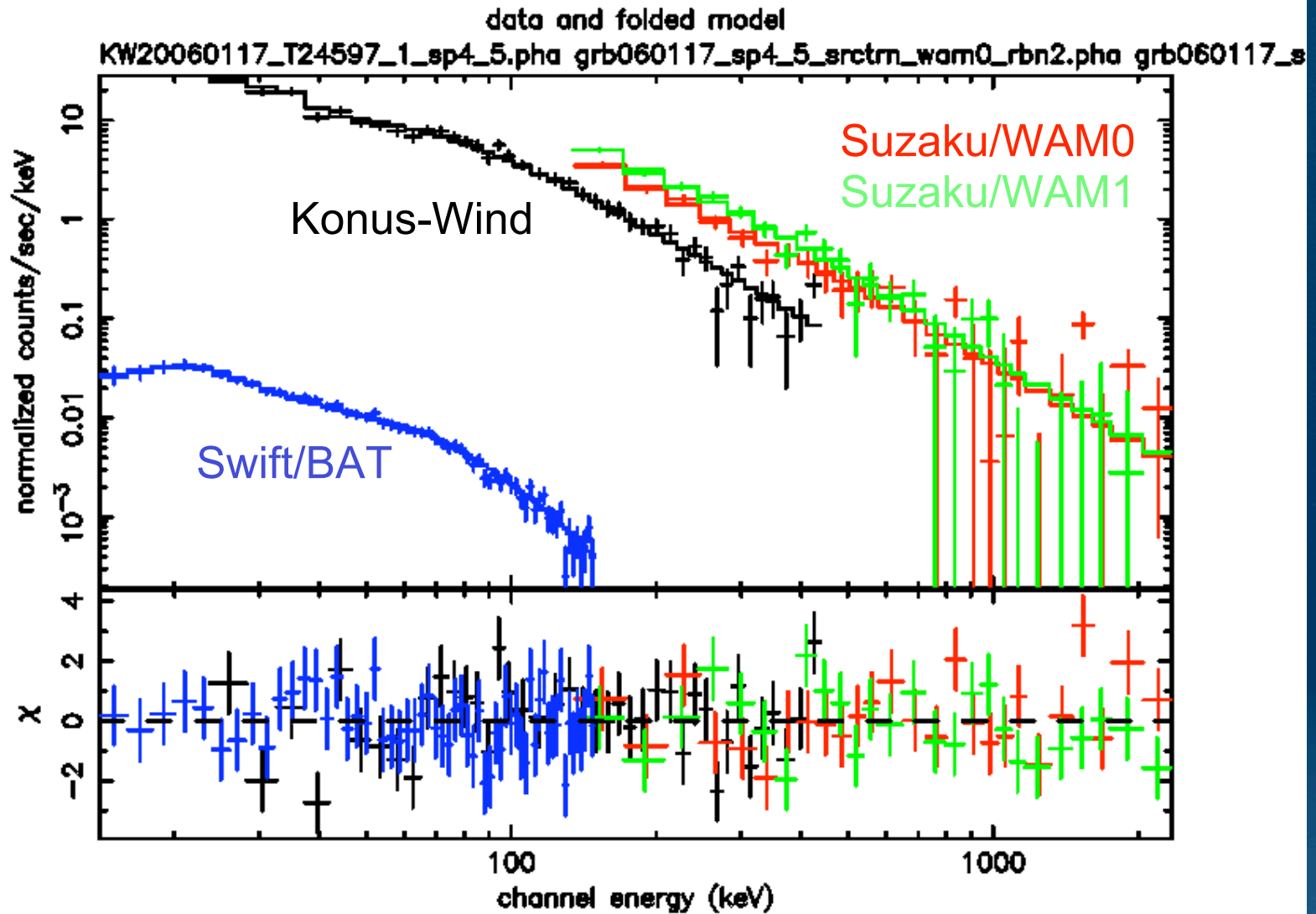
Example: GRB 060117



Example: GRB 060117

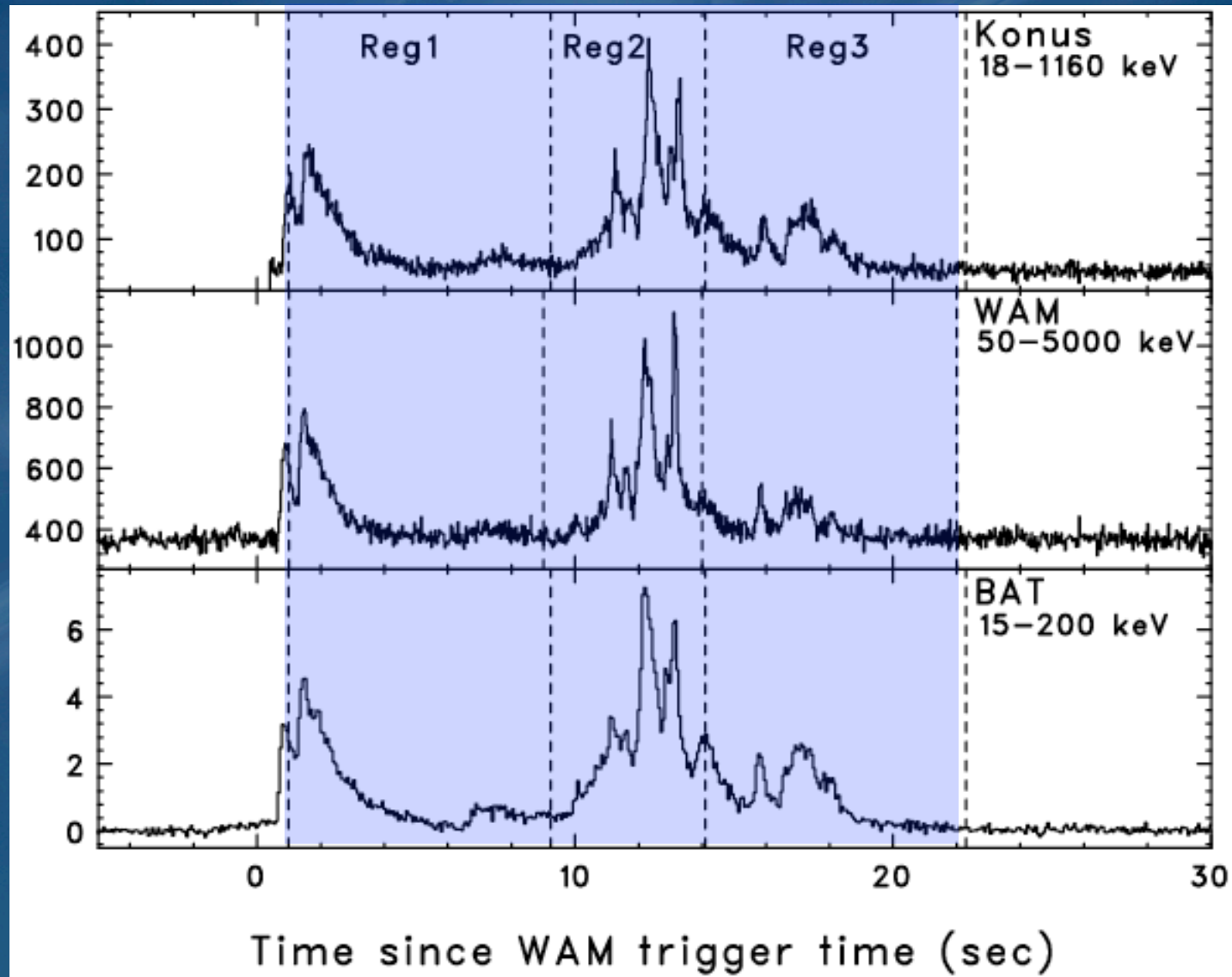


Example: GRB 060117

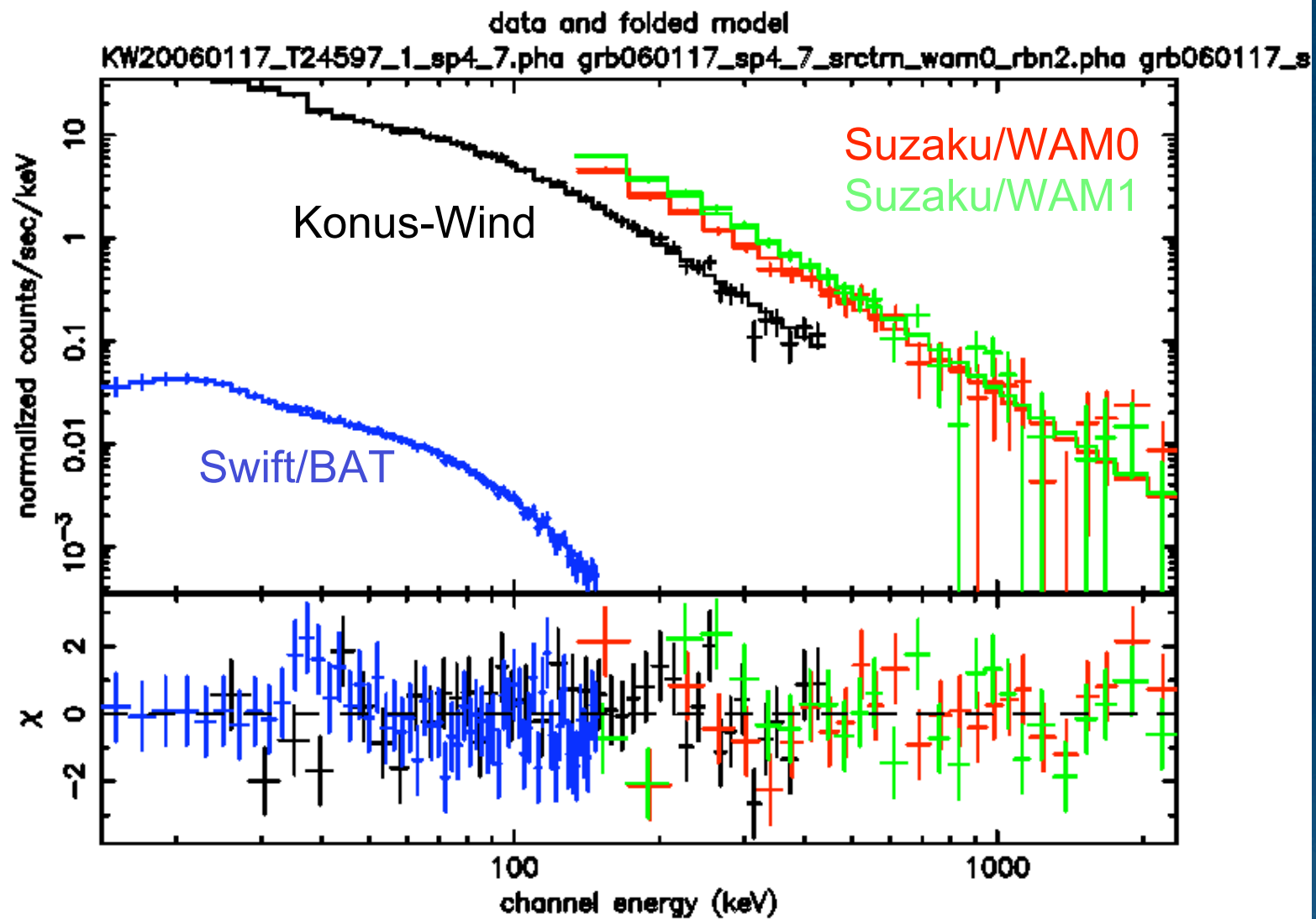


takanori 16-Dec-2008 11:52

Example: GRB 060117



Example: GRB 060117



takanori 16-Dec-2008 14:35

Spectral parameters

Low energy
photon index

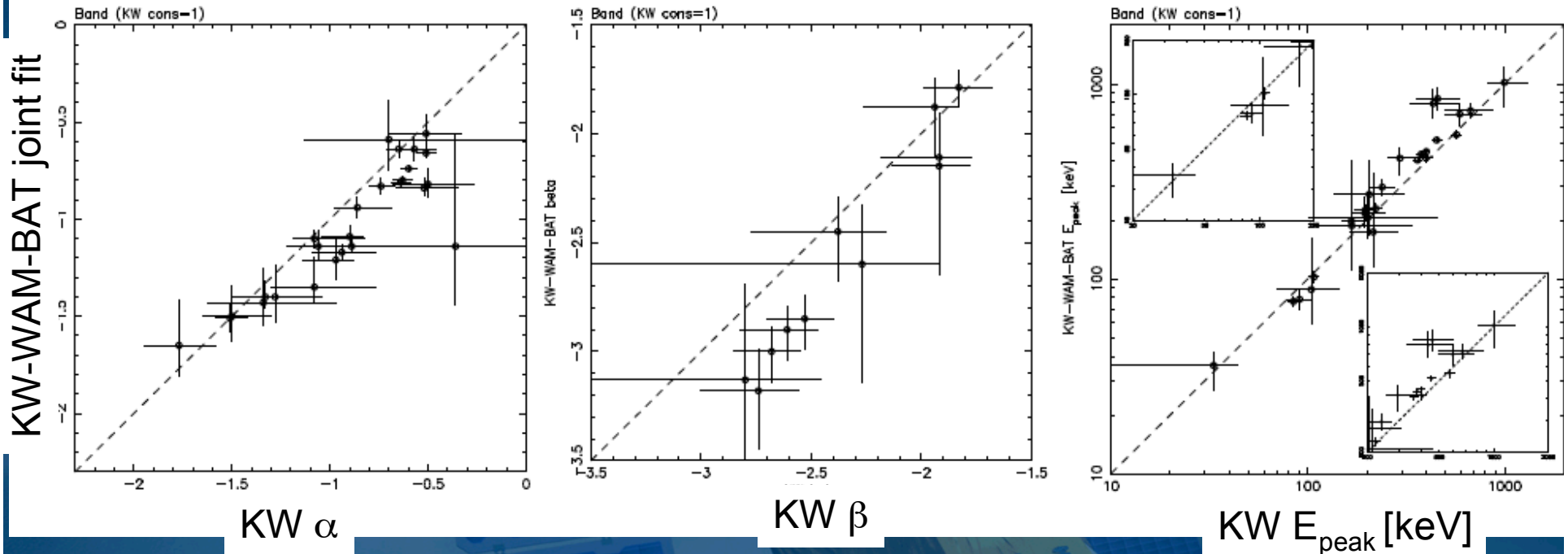
α

High energy
Photon index

β

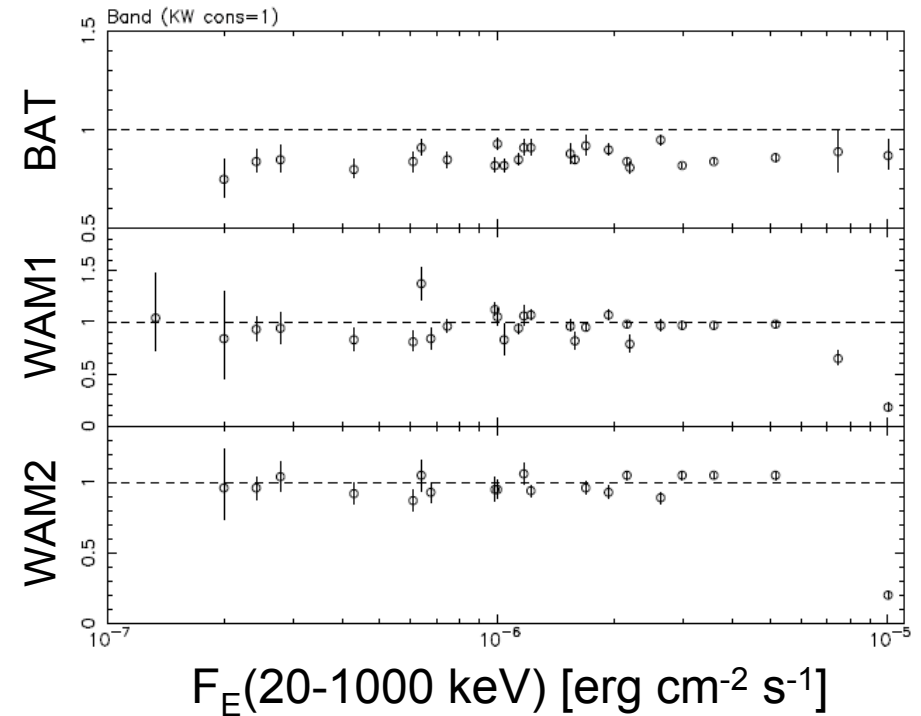
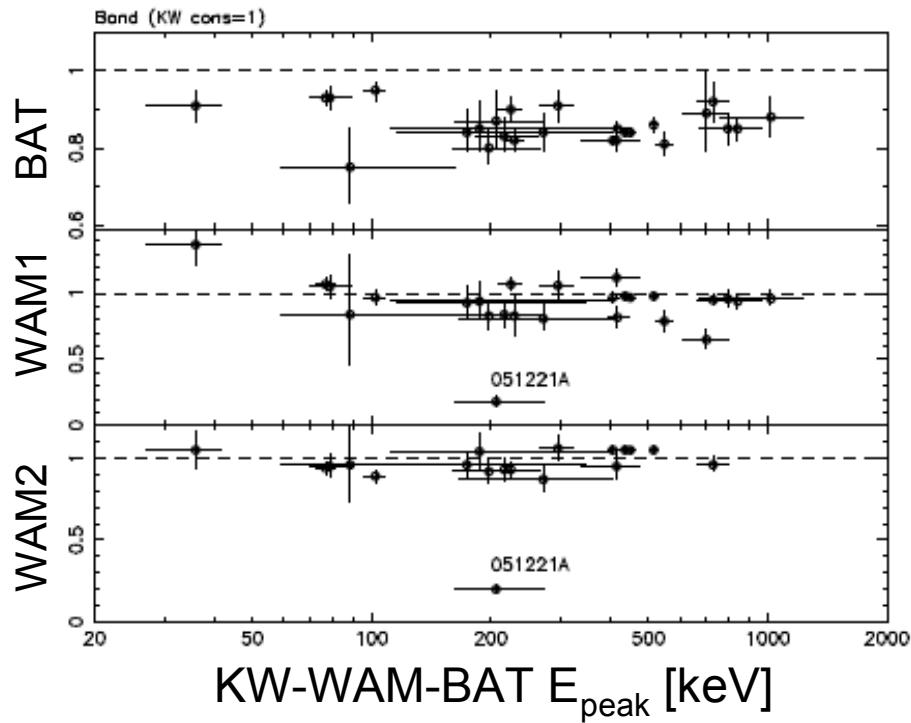
Peak energy
in νF_ν

E_{peak}



- Systematically steeper (softer) in the joint fits
- Systematically higher E_{peak} ($E_{\text{peak}} > 200$ keV) in the joint fits

Normalization factor



- WAM constant factor agrees to KW within 20% range (0.8-1.2).
- BAT constant factor is systematically smaller than KW by 10-20%.

Summary

- Cross-calibration work among Konus-Wind, Suzaku/WAM, and Swift/BAT has been performed using the simultaneously observed GRBs in 2005-2007.
- Systematically softer (steeper) low energy photon index α and higher E_{peak} energy are obtained based on the joint fits comparing to those of the KW fit.
- Constant factor of WAM based on the joint fits is consistent with the KW fits within 10-20% range. However, the constant factor of BAT is systematically smaller by 10-20%.

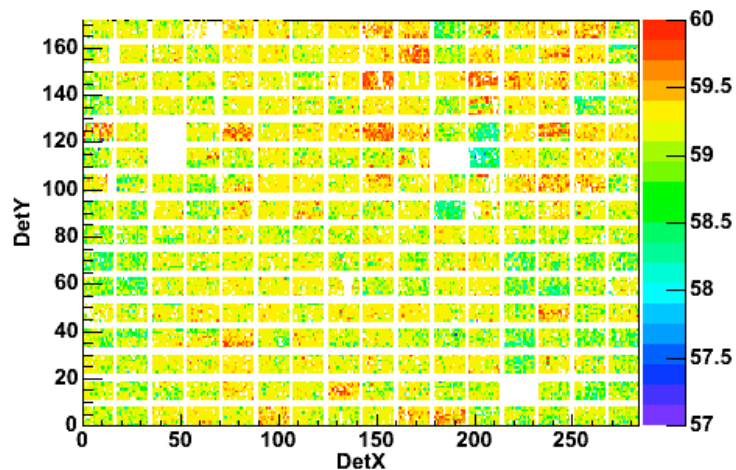


Gain change in BAT (work in progress)

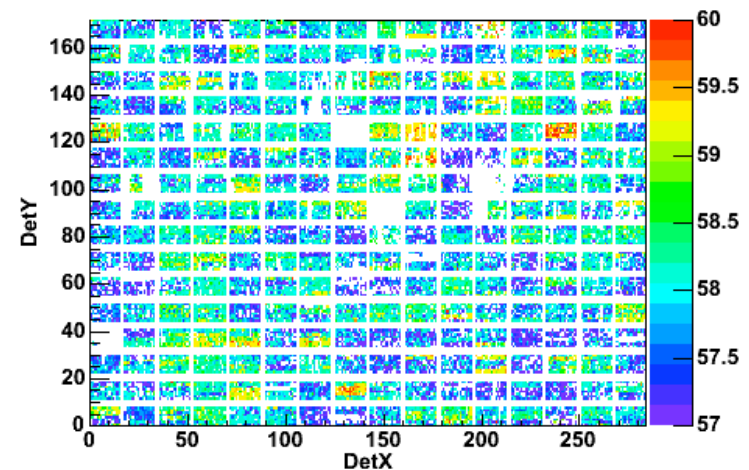
Peak Shift to lower energy

^{241}Am 59.5 keV line

2005 doy 1-10



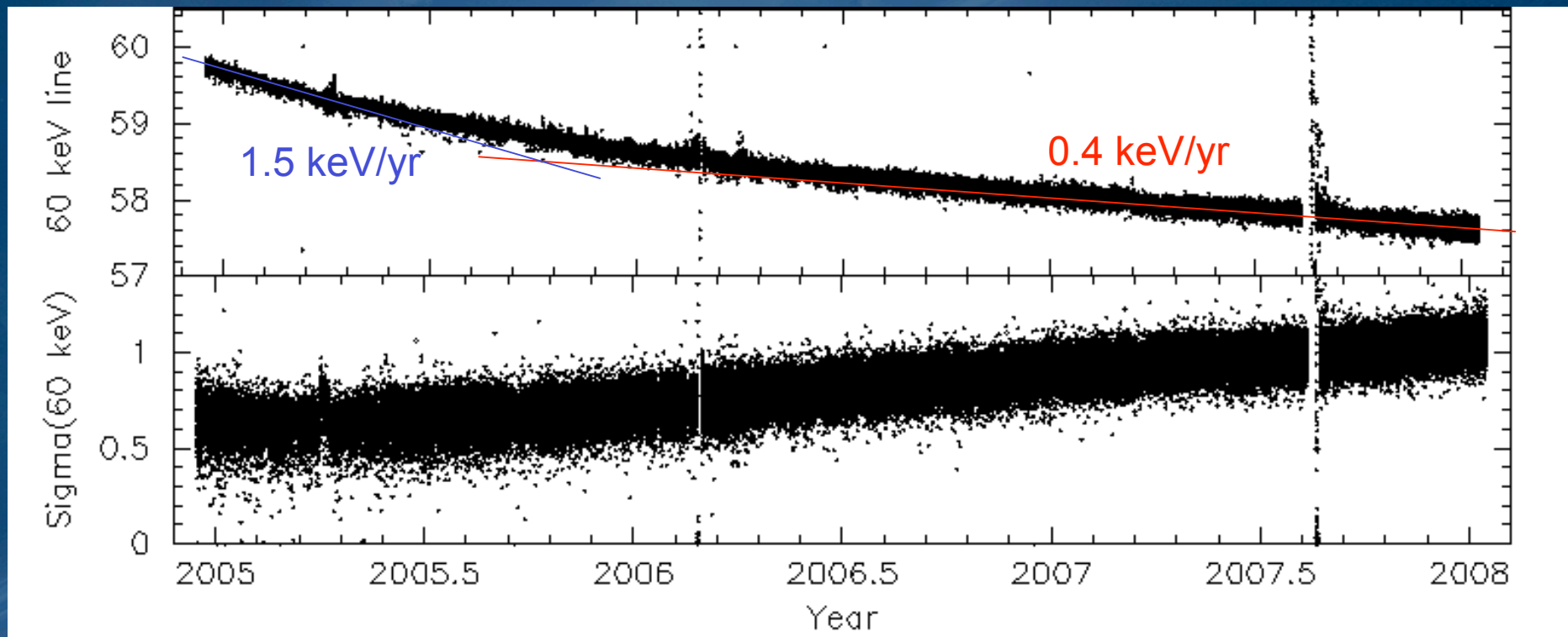
2006 doy 152-161



- BAT measurement of 59.5 keV line has been shifted for 1-2 keV in ~ 1.5 years.

Long term trend of 59.5 keV line

^{241}Am spectrum: array averaged



- 1.5 (early) - 0.4 (current) keV/yr shift in 59.5 keV line (array averaged).
- The pulse width seems getting wider.

Summary

- Peak shift of 59.5 keV line to low energy has been seen (^{241}Am source).
- Shift rate is 1.5 keV/yr (early) - 0.4 keV/yr (current).
- We believe the shift is due to intrinsic properties of CZT detector rather than electronics (still investigating...).
- We are working on updating the software/CALDB to correct this effect.