



Thermal SNR Working Group: Report from this IACHEC

*Paul Plucinsky on behalf of the IACHEC
Thermal SNR Working Group*



Thermal SNR Working Group

One of the “Standard candle” working groups.

This presentation is a summary report of this group’s work:

| | |
|----------------|---|
| XMM-Newton | Andy Pollock & <i>Matteo Guainazzi</i> (ESAC) |
| Chandra HETG | Dan Dewey (MIT) |
| XMM-Newton MOS | <i>Steve Sembay</i> (Leicester) |
| XMM-Newton pn | Frank Haberl (MPE) |
| Chandra ACIS | <i>Jenny Posson-Brown</i> , Joe DePasquale, & <i>Paul Plucinsky</i> (SAO) |
| Suzaku XIS | <i>Eric Miller</i> (MIT) |
| Swift XRT | <i>Andrew Beardmore</i> (Leicester) |
| Models | Adam Foster & Randall Smith (SAO) |

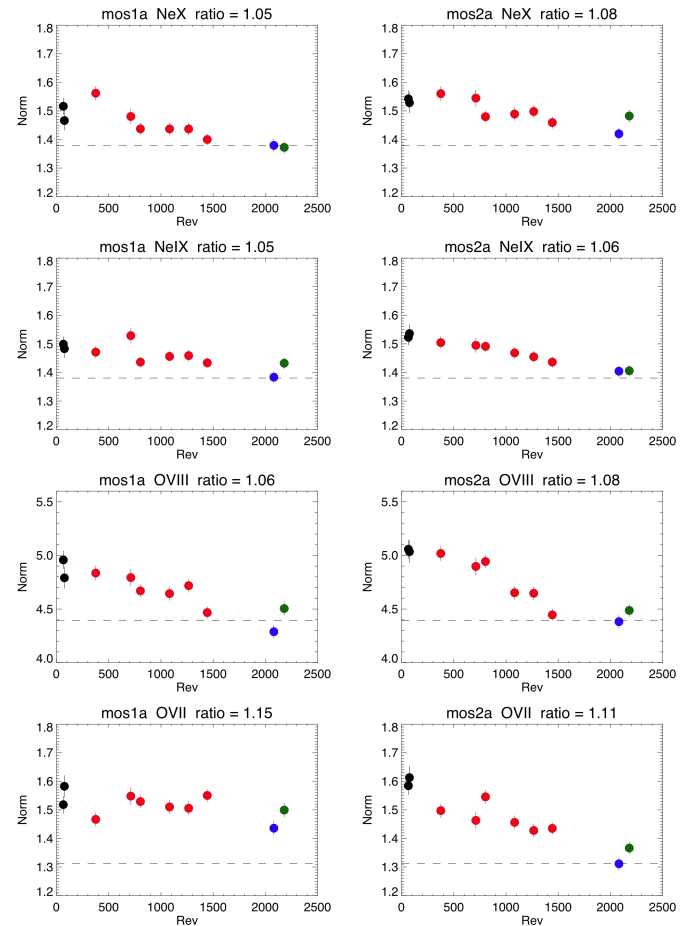


E0102: How to Compare the Instruments ?

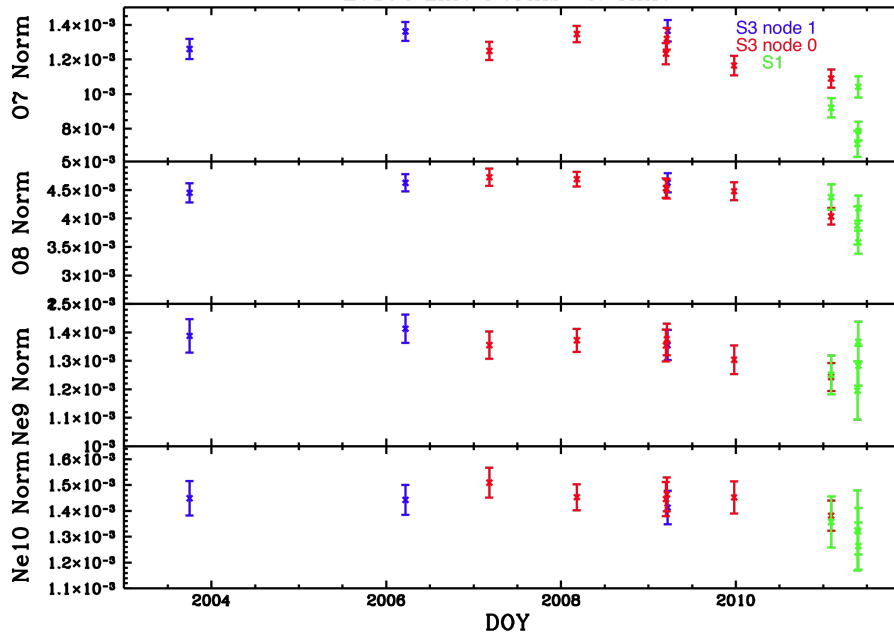
The major challenge over the last few years has been characterizing *the time-dependent* performance of the various instruments. How do we perform a meaningful comparison ??

MOS:

ACIS:



E0102 Line Norms vs. Time





E0102: How to Compare the Instruments ?

Steve Sembay had the brilliant suggestion:

Don't Do It !!!

Organization of an A&A Paper on E0102:

- ① Compare line normalizations/effective areas for the early mission data sets used in the 2008 SPIE paper with the calibration available today.
- ② Each instrument will write its own section on its time dependence.



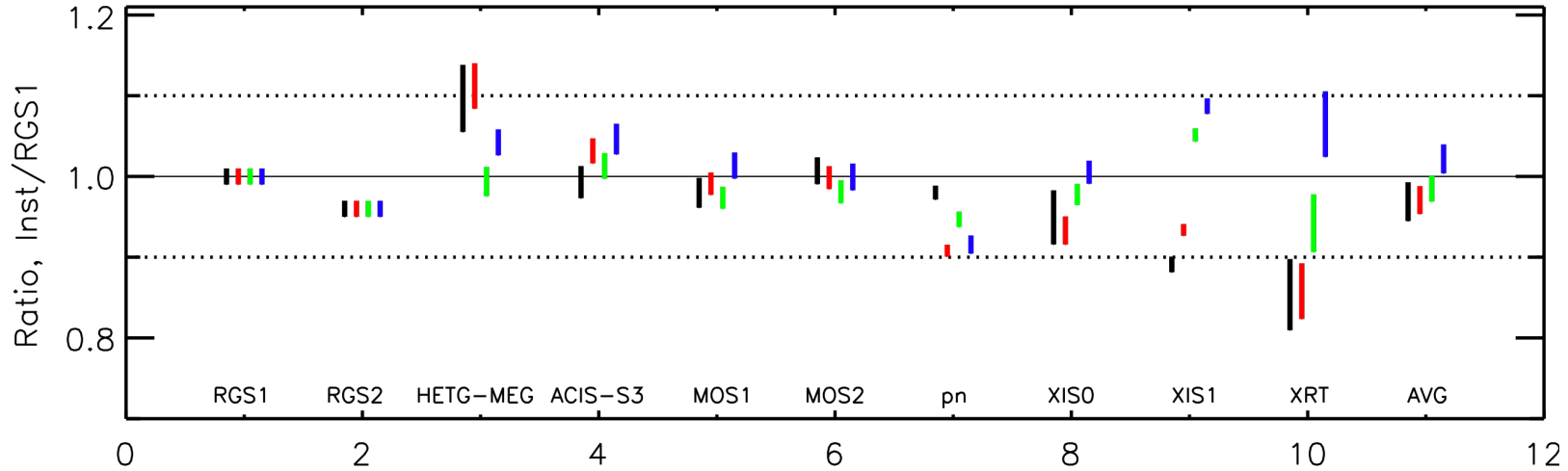
Chandra X-Ray Observatory

CXC

Comparison of Fluxes for Bright Line Complexes:

- 2008 SPIE results updated in 2010

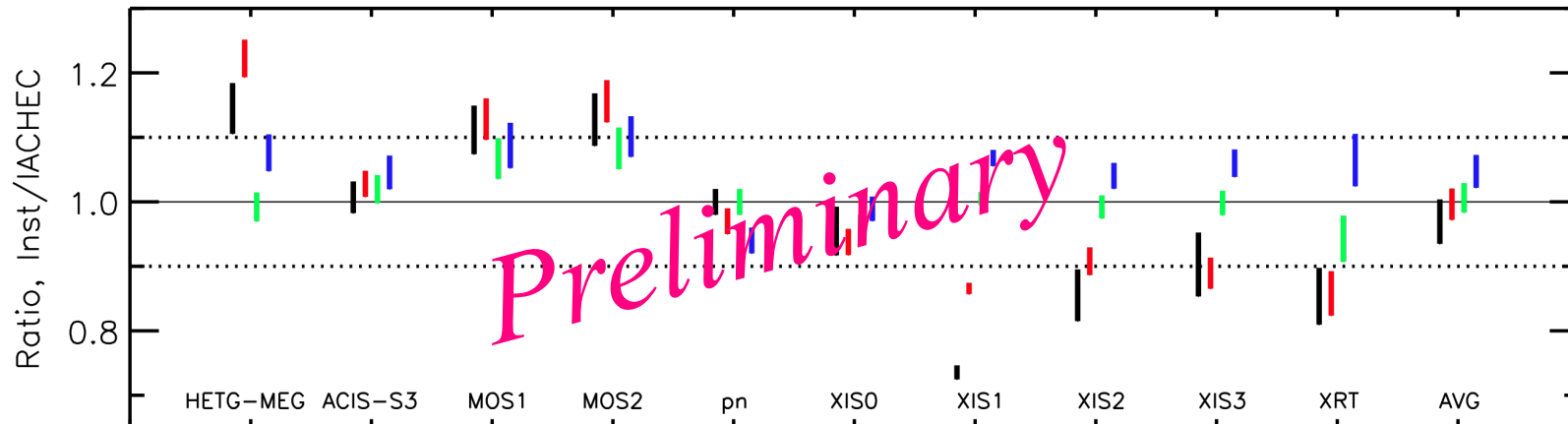
DePasquale(SAO)



OVII black **OVIII red** **NeIX green** **NeX blue**

Posson-Brown(SAO)

IACHEC 7, 28 Mar 2012



Preliminary



Chandra X-Ray Observatory

CXC

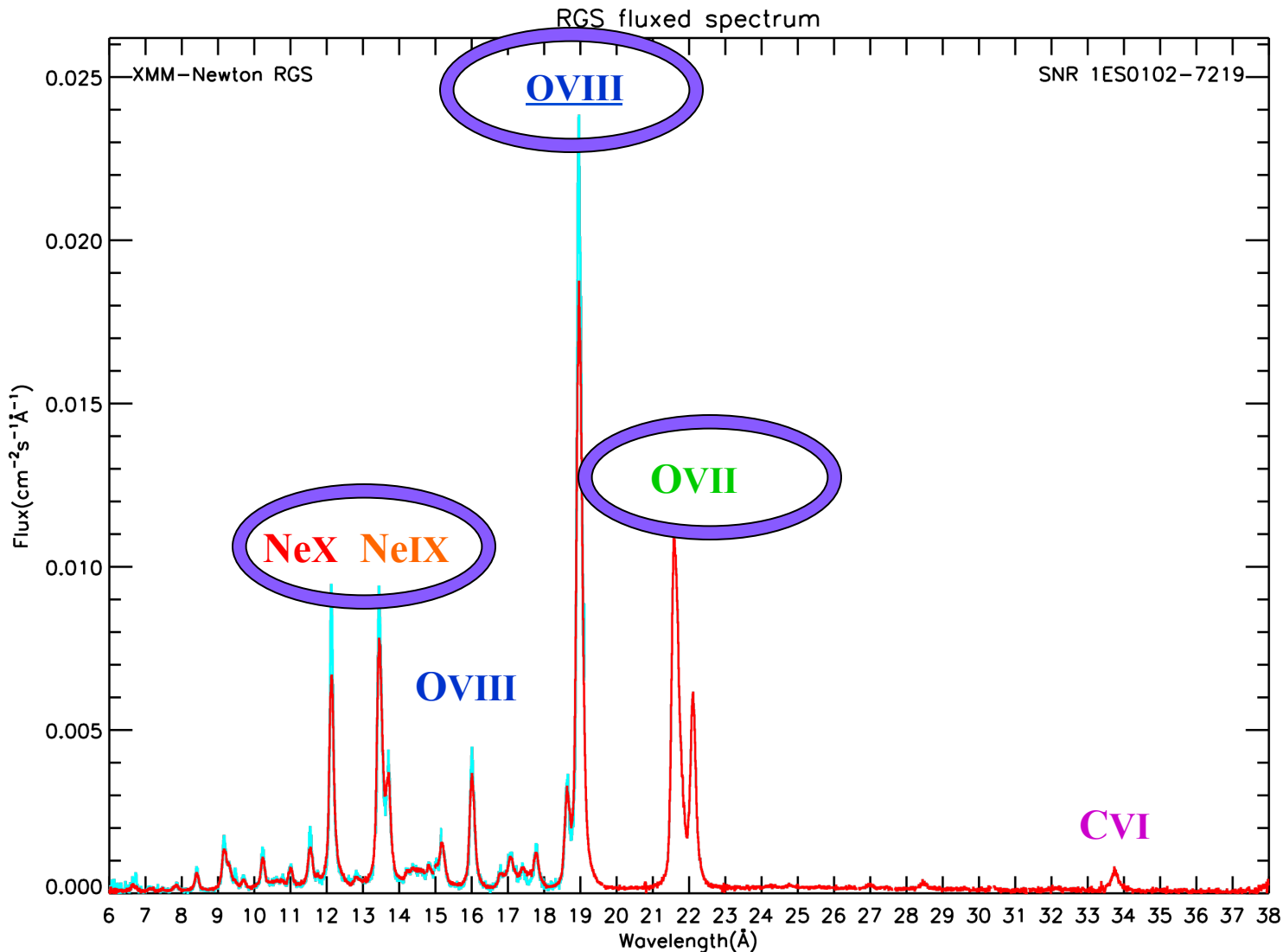
XMM-Newton RGS Spectrum of E0102:

*Pollock
(ESAC)*

Relatively simple spectrum dominated by O & Ne,

little or no Fe emission

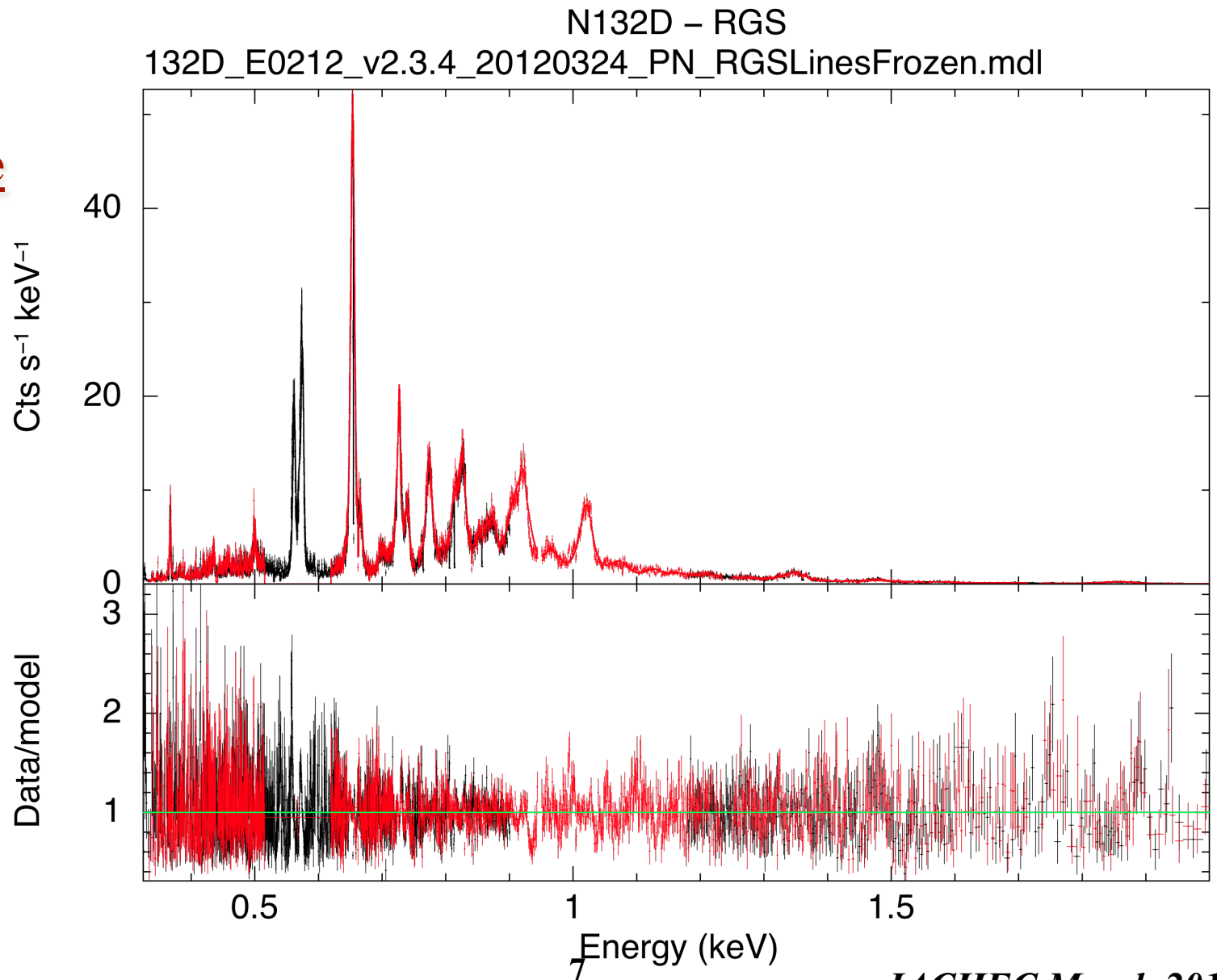
In most applications, line energies & widths are frozen, only 4 normalizations are allowed to vary **OVII**, **OVIII**, **NeIX**, & **NeX**





N132D:
Brightest
SNR in the
LMC

*Guainazzi,
Stuhlinger,
& Pollock
(ESAC)*





N132D: Development of IACHEC model

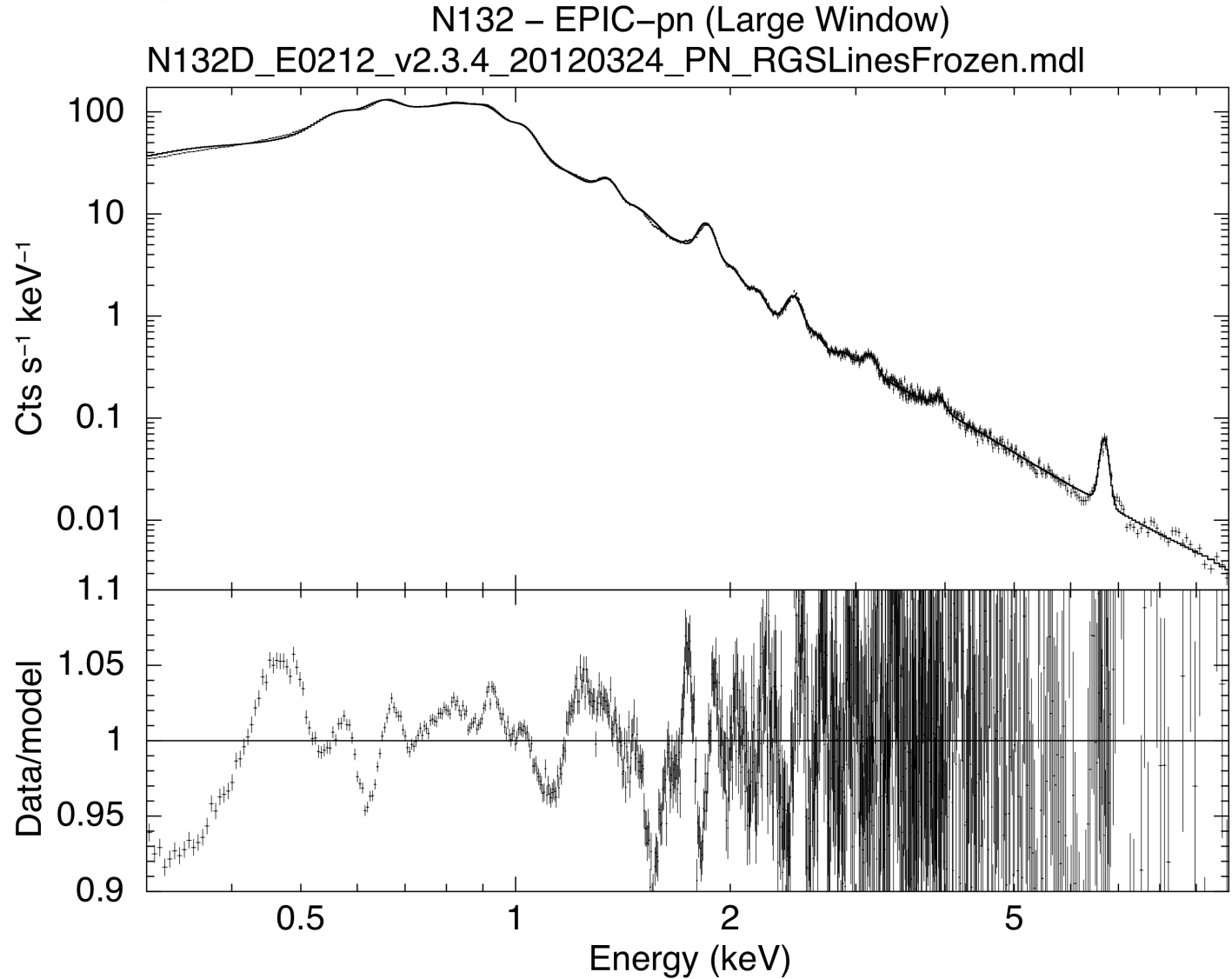
Guainazzi, Stuhlinger, & Pollock (ESAC)

- Fit the EPIC-pn spectrum in the 2-10 keV band
 - ◆ Continuum+ Σ unresolved Gaussian lines+power-law (high-energy background)
- Freeze all the parameters in #1, and fit the RGS spectrum
 - ◆ Another continuum+photoelectric absorption+ Σ unresolved Gaussian lines
 - ◆ We try to add full “atomic series” of lines (e.g.: OVIII, NeIX, FeXX ...) rather than being simply driven by the statistics.
 - ◆ Lines are identified and their energy frozen to the value in ATOMDB (via the XSPEC `identify` command)
 - ◆ Once the lines are identified, a fit cycle is run on the whole model to determine the best-fit line width for each series. The width of each line in each series is calculated rescaling linearly with energy the width of its first line
 - ◆ This is: N132D_E0212_v2.3_20120324_RGS.mdl
 - ◆ This model does not necessarily work well above 2 keV
- Freeze the energy and the width of the Gaussian components in the RGS energy bandpass in #2. and fit the EPIC-pn spectrum again
 - This is: N132D_E0212_v2.3_20120324_PN.mdl
- Free the energy, the widths *and the normalizations* of the Gaussian components in the RGS energy bandpass in #2, and fit the EPIC-pn spectrum again
 - This is: N132D_E0212_v2.3_20120324_PN_RGSLinesFrozen.mdl



N132D: Comparison of IACHEC model to pn data

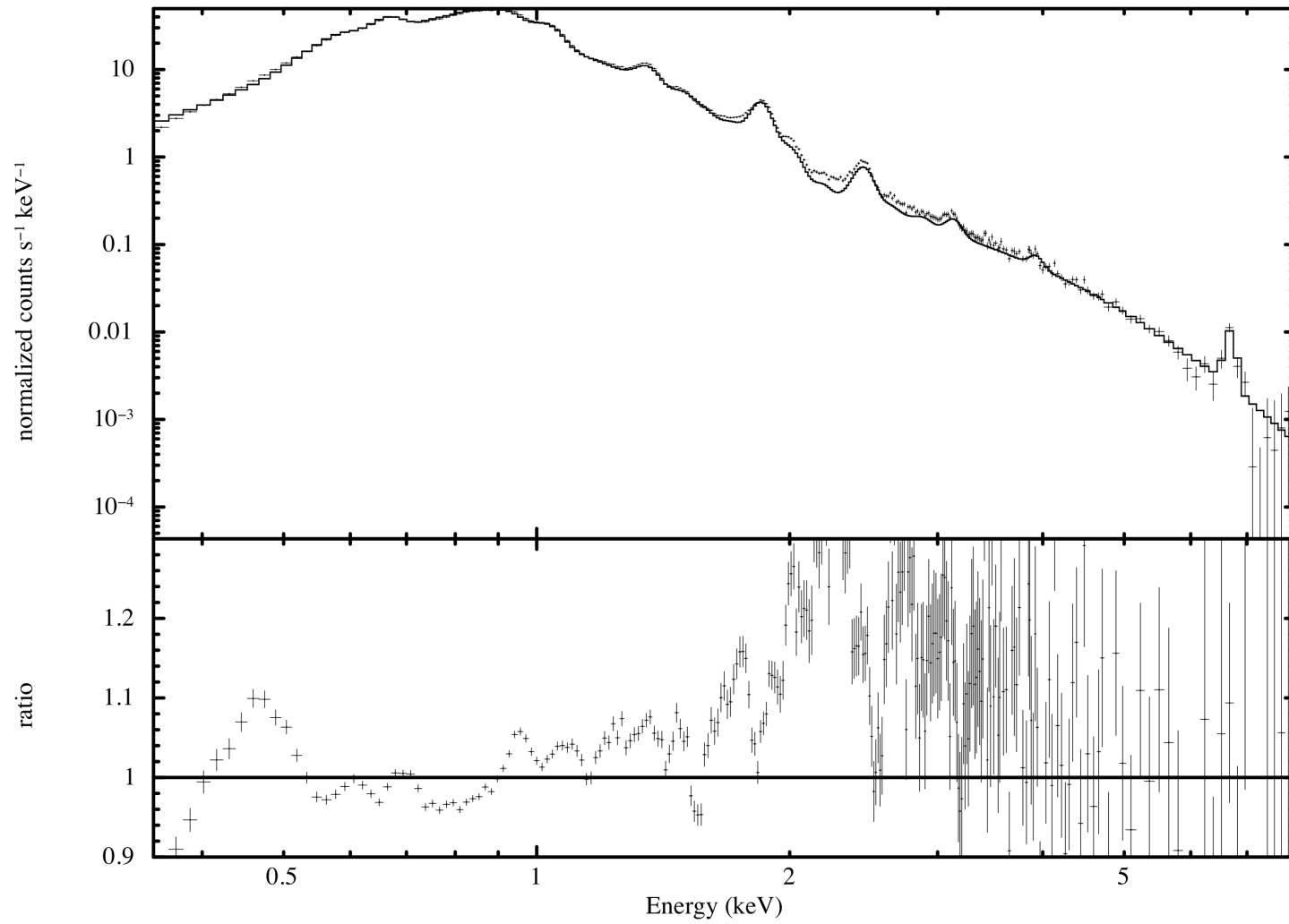
*Guainazzi
(ESAC)*





N132D: Comparison of IACHEC model to ACIS data

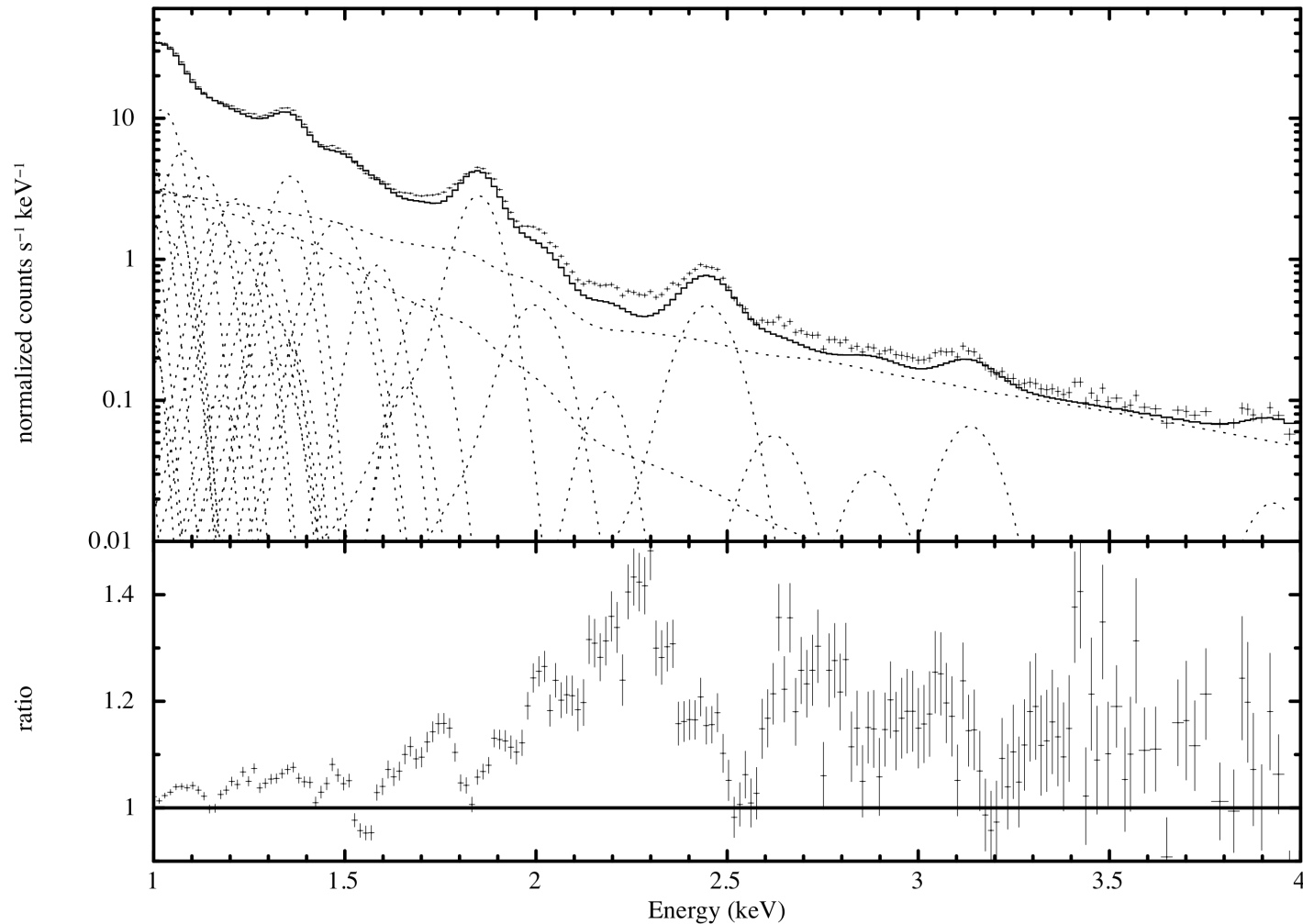
N132D, ACIS S3 spectrum 89 ks, XMM RGS+pn ver 2.3.3 model
Cstat=6382.8, 643 DOF





N132D: Comparison of IACHEC model to ACIS data

N132D, ACIS S3 spectrum 89 ks, XMM RGS+pn ver 2.3.3 model
Cstat=6382.8, 643 DOF

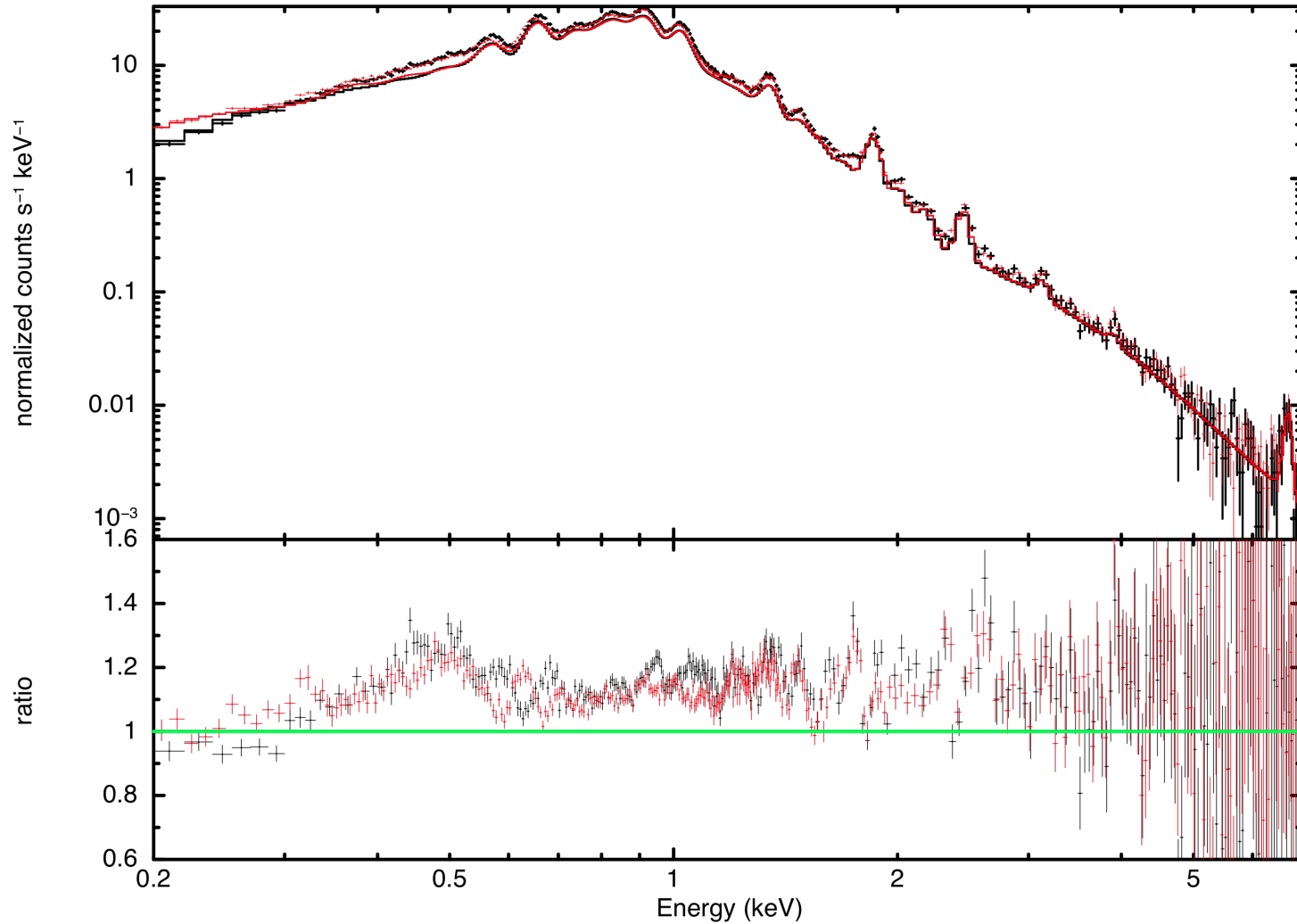




N132D: Comparison of IACHEC model to MOS data

N132D - MOS1

Sembay
(Leicester)

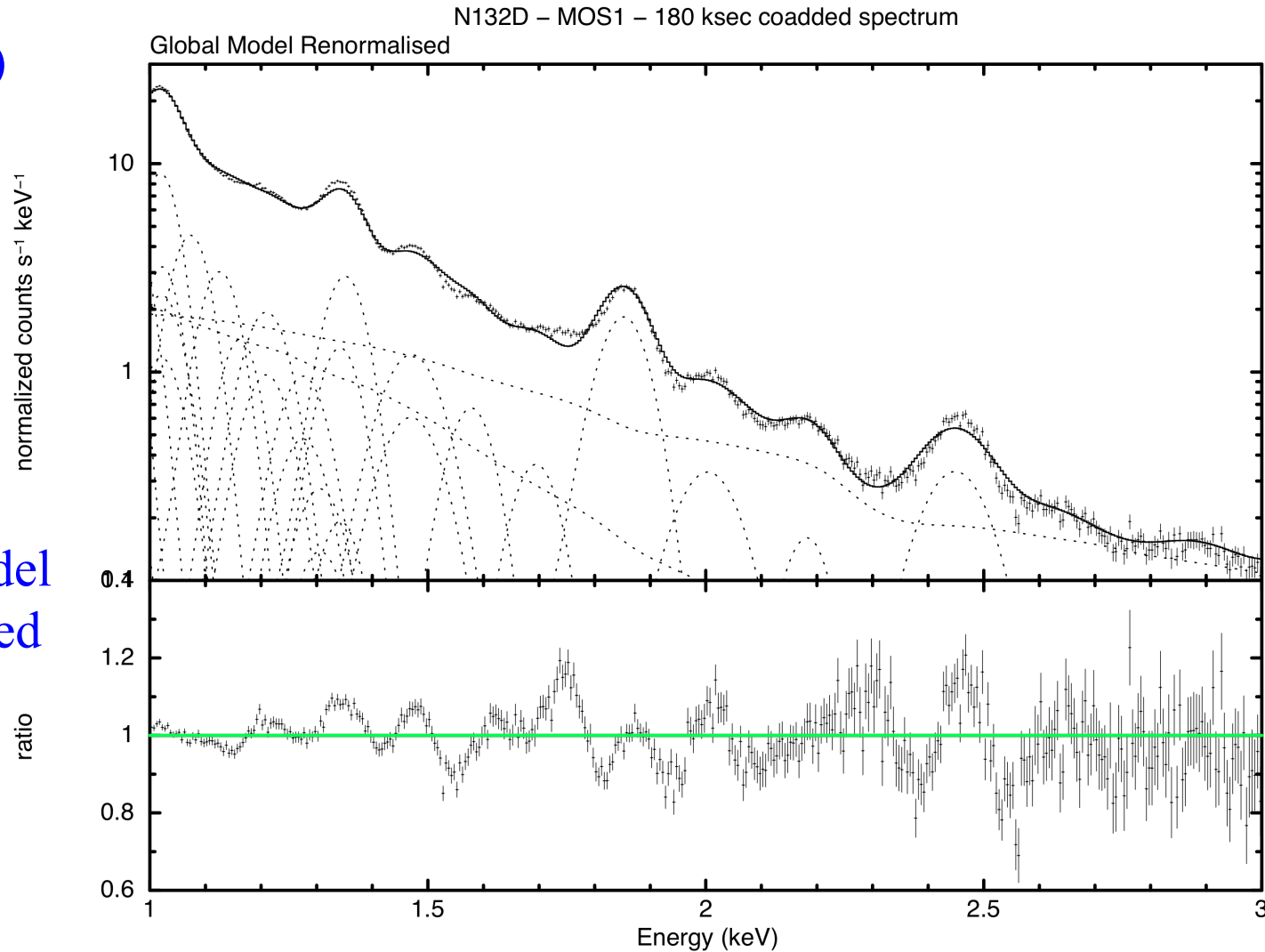




N132D: Comparison of IACHEC model to MOS data

Sembay
(Leicester)

Global model
renormalized

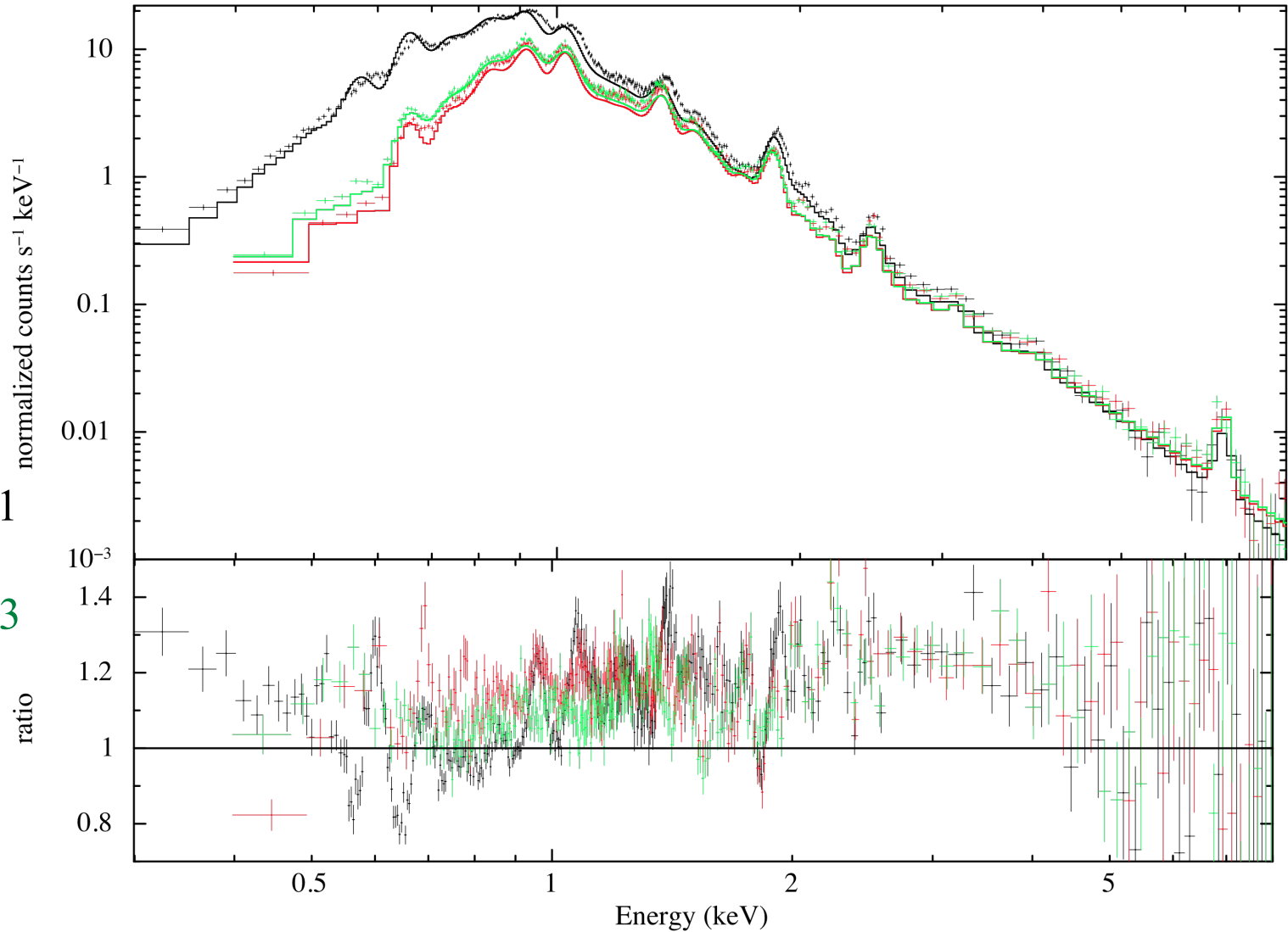




N132D: Comparison of IACHEC model to XIS data

Miller
(MIT)

black = XIS1
red = XIS0
green = XIS3





Summary

1 E0102-7219:

- We need to decide what we want to publish (if anything) on the multiple measurements over the course of the mission

N132D:

- We want to develop a standard IACHEC model of N132D that we can use for calibration purposes

Fitting Methodology:

- We want adopt the approach of using unbinned spectra, modeling the background, & using the C statistic
- We believe the IACHEC should take the lead in encouraging the User community to adopt this approach