## Status of the eROSITA Mission



#### eROSITA Collaboration

#### **Core Institutes (DLR funding):**

MPE, Garching/D
Universität Erlangen-Nürnberg/D
IAAT (Universät Tübingen)/D
SB (Universität Hamburg)/D
Astrophysikalisches Institut Potsdam/D

#### **Associated Institutes:**

MPA, Garching/D
IKI, Moscow/Ru
USM (Universität München)/D
AIA (Universität Bonn)/D

#### **Industry:**

Media Lario/I Mirrors, Mandrels
Kayser-Threde/D Mirror Structures
Carl Zeiss/D ABRIXAS-Mandrels
Invent/D Telescope Structure

pnSensor/D CCDs

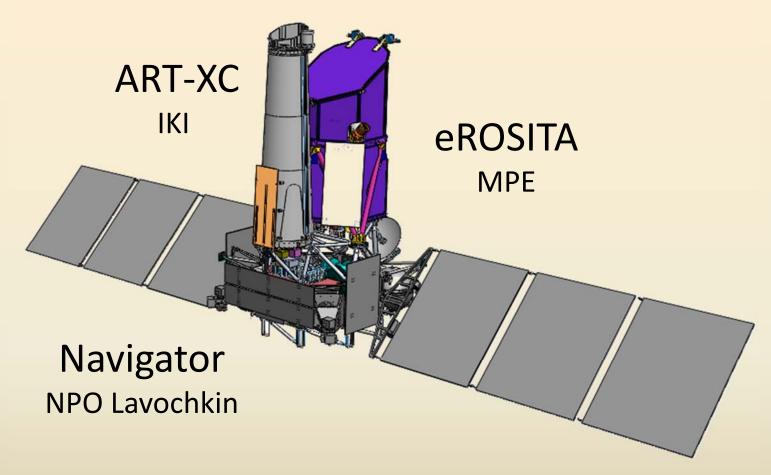
IberEspacio/E Heatpipes
RUAG/A Mechanisms

HPS/D,P MLI + many small companies

MPE: Scientific Lead Institute, Project Managment

Instrument Design, Manufacturing, Integration & Test Data Handling & Processing, Archive etc.

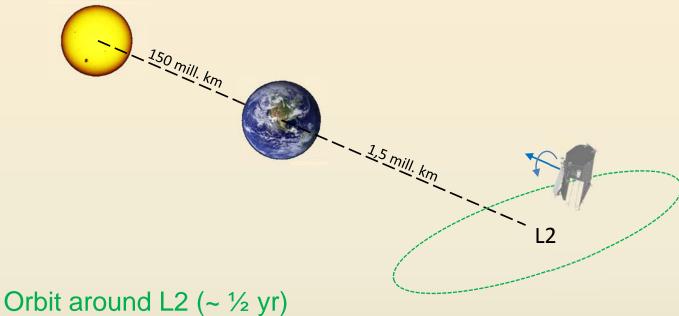
### SRG



Launch from Baikonur with Zenit-Fregat, 2014

Zenit-Fregat-Navigator = standard configuration, also used for Elektro-L (Jan 2011) and Radiastron (Jul 2011).

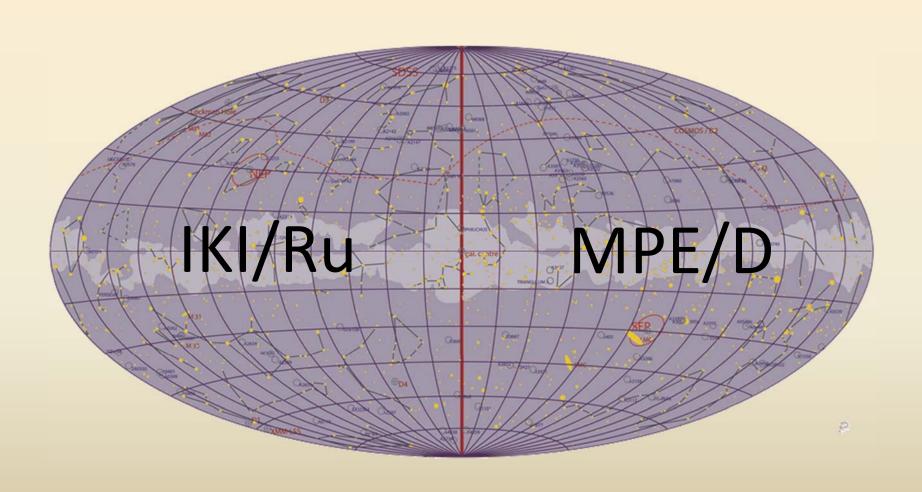
### Mission Profile



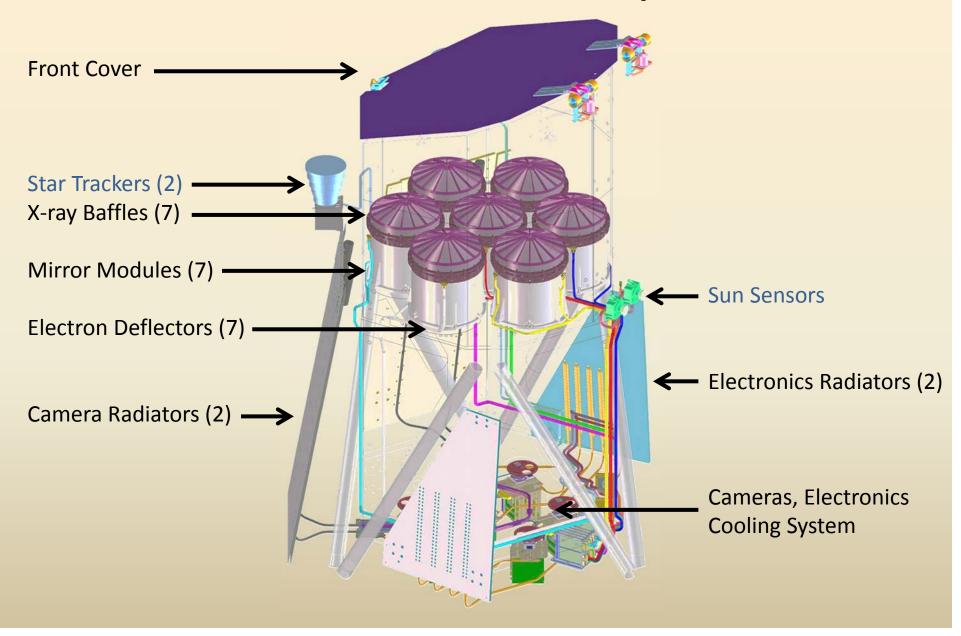
Continuously rotating during Survey, ~ 4 hours / revolution

4 years all-sky survey, 3 years pointing

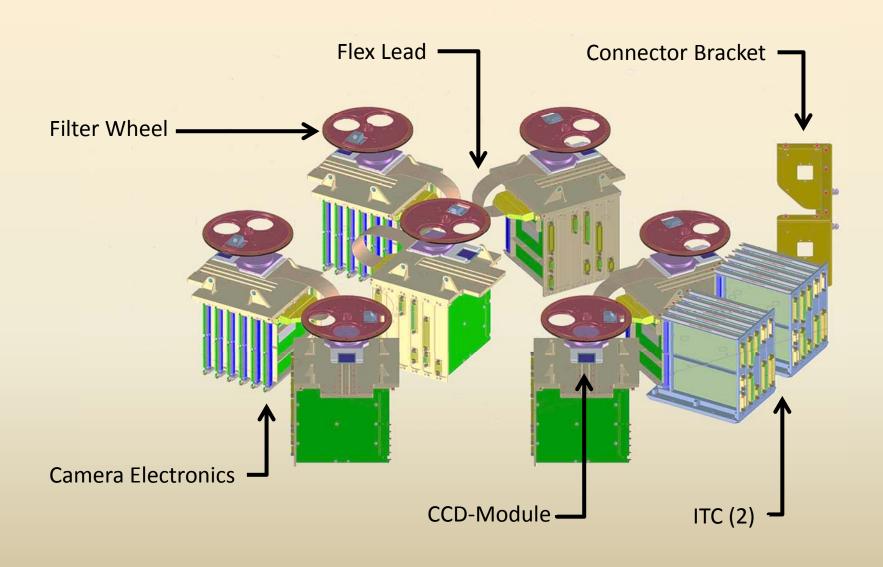
# **Sky Division**



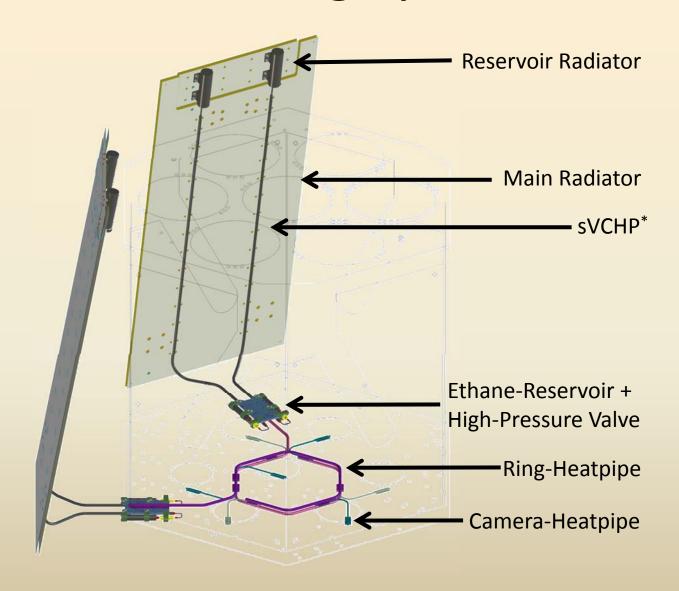
## Scheme: Telescope



### Scheme: Cameras



## Scheme: Cooling System



### **HW-Status: Overview**

	QM, EM, STM, TM, DM	FM-readiness		
Structure	Qualified (Vb)	100%		
Mirror Modules	Qualified (Vb, TV, X) 82%			
CCD-Cameras	Qualified (Vb, TV, X, P)	100%		
Cooling System	Qualified (Vb, TV, C)	100% (parts)		
X-ray Baffles	Qualified (Vb, TV, X)	30%		
Electron Deflectors	Qualified (Vb, TV, ED)	100%		
Filterwheels	Qualified (Vb, AC, TV) 100% (par			
Electronics	Qualified (Vb, TV) TM ready (del. to NPOL) EM in test	80% (parts, ITAR)		
PFM-Telescope	Qualified (MP, Vb, TV, AC,	PS), waiting for TV		
GSE	ready			

Vb = vibration, TV = thermal vacuum, AC = acoustic noise, X = x-ray, P = proton radiation damage, C = cooling, ED = electron deflection, MP = mass properties, PS = pyro shock

#### The MPE PANTER X-ray Test Facility

#### · The X-ray test Facility

- Is located in Neuried, south west of Munich
- 120 m X-ray beamline, 1 m diameter
- 12 m instrument chamber, 3.5 m diameter
- Large cleanroom for handling X-ray optics
- Movable 10 m extension with 0.25 m diameter and 3 m instrument chamber, 1.2 m diameter

#### Detectors

- PSPC: Position Sensitive Proportional Counter
- TROPIC: Single photon counting CCD camera
- PIXI: Integrating CCD Camera

#### X-ray Sources

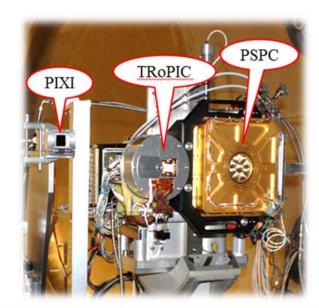
- Multi target electron impact source
- · Seifert closed X-ray source
- X-ray Monochromators

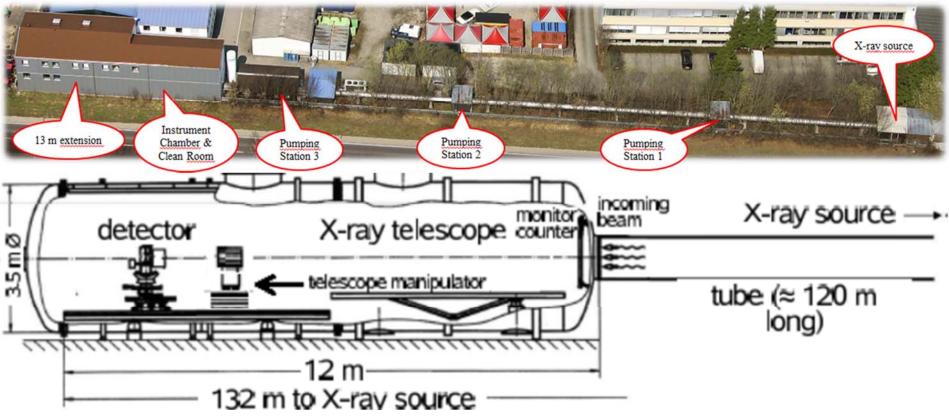
#### Beam Monitors

- · Proportional counters
- Drift Chamber

#### Manipulators

- · X, Y, Z translation stages
- Tip-tilt, rotation stages





### eROSITA Calibration at PANTER

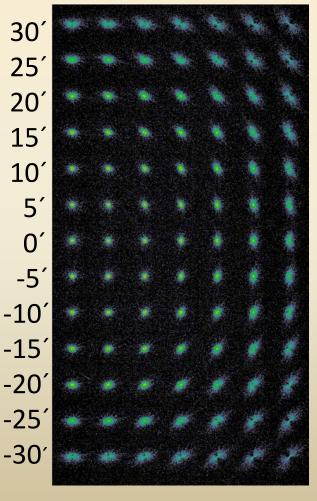
- HEW & Effective area
  - on-axis 0.27 8.04 keV
  - off axis 1.49 8.04 keV
- Out of focus images
- Focal plane Mapping

	FM1	FM2	FM3	FM4	FM5	FM6	FM7	FM8
X1-ACC	$\checkmark$	✓	X	X	X	X	X	X
X2-BAF	$\checkmark$	X	X	X	X	X	X	X
X3-PVI	$\checkmark$	-	-	-	-	-	-	-
X4-PTV	X	-	-	-	-	-	-	-
X5-TEL	X	-	-	-	-	-	-	-
X6-CAL	X	Х	Х	Х	Х	Х	X	Х

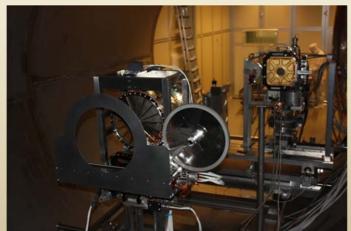
## Status: Mirror System

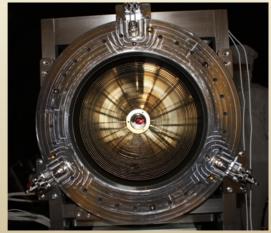
### FM2 off-axis measured PSF over the field of view

0′ 5′ 10′ 15′ 20′ 25′ 30′

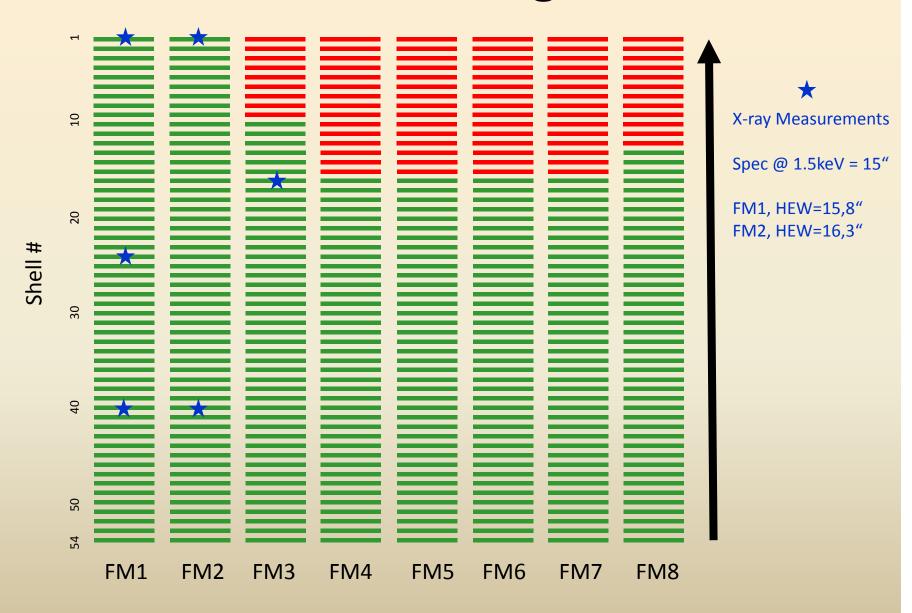


- DM-Mirror delivered 12/2009
  - Qualification tests (TV, Vib, X-ray, early 2010)
  - Together with X-ray Baffle, late 2011
  - Part of QM-Telescope tests (2012)
- FM-Program
  - All mechanical parts (spider, MIS) ready since 2010
  - FM Shell integration started in 2011
  - Witness X-ray measurement: in spec (HEW = 15")
  - FMs completed, delivered, and tested in PANTER
  - FM-1 (12/2012-01/2013)
  - FM-2 (03/2013)





## Mirror Shell Integration



## Status: X-ray Baffle

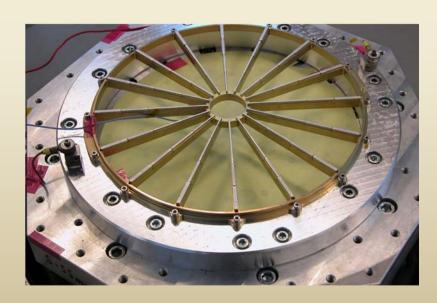
- 7 Baffles, each with 54 invar cylinders in front of Mirror Modules
  - Process qualified (cutting, bending, welding, integration)
  - Accuracy < 50μm
  - Qualification (Vibration, Thermal Vacuum, X-ray tests)
- FM-Program
  - FM-1 and FM-3 completed
  - Integration of FM-4 and FM-5 started
  - Integration of FM-2 on hold (separate tests for removing damaged shell)





### Status: Electron Deflector

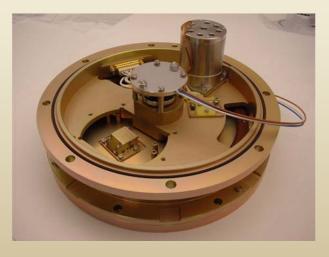
- 7 spider wheels with permanent magnets behind Mirror Modules
  - Qualified (Vibration, Thermal Vacuum, Electron Deflection)
- FM-Program
  - FM 1-6 manufacturing completed
  - FM 7-8 waiting for magnets bonding
  - FM 1,2 acceptance tested (Vib, outgassing, TV)
  - FM 3-6 outgassed, TV-tested

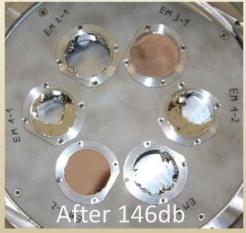




### Status: Filterwheel

- 7 motor driven wheels in front of cameras
  - 4 positions: Open, Closed, Filter, Fe<sup>55</sup> calibration source
  - Qualified (Vibration, Thermal Vacuum, Acoustic Noise)
  - Lifetime test OK, (3 months, vacuum, 3000 cycles)
  - Purge test OK (4 months, 5x nominal flow)
- FM-Program
  - All parts (motors!) in stock
  - Purchasing of Fe<sup>55</sup> targets (1Mbeq) as late as possible

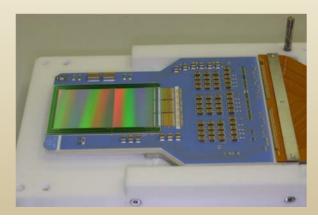




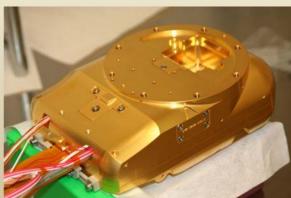
But after a slight redesign: Everything OK!

#### Status: Camera

- 7 pnCCDs on front-end ceramic PCBs, cooled to -95°C
  - Prototype camera in operation since early 2007 ("TRoPIC" in PANTER)
  - eROSITA-CCDs in stock since early 2010
  - EM-Camera in X-ray test since late 2010
  - Proton radiation damage tests in 2010 and 2011
- FM-Program
  - All CCDs and CAMEXs manufactured, screened, tested
  - All PCBs manufactured, mounting pending
  - Mechanical parts:
    - All cold casings ready
    - Proton shields pending





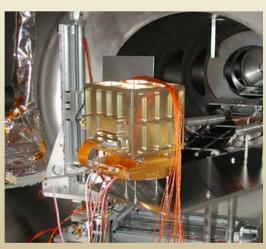


### Status: Electronics

- 7 Camera Electronics Boxes + 2 Interface & Thermal Controllers
  - Qualified (Vibration, Thermal Vacuum)
  - All PCBs designed, layouted, manufactured (EM)
  - EM in test
- FM-Program
  - 80% of components in stock, but still some ITAR issues.





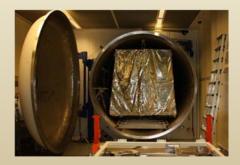


### Status: Thermal

- Cooling of cameras
  - 2 Radiators, 4 sVCHPs, 2 "Ring-HPs", 7 Camera-HPs
  - Qualified (Vibration, Thermal Balance Test in PANTER)
- Cooling of electronics
  - 2 Radiators, 9 standard ammonia heatpipes
  - Qualified parts (HPs), SS-Qualification at QM-Tests
- Thermal system of telescope
  - MLI insulation
  - Thermal baffle in front of mirror system
  - Heating of mirror modules
  - Partly qualified, SS-qualification at QM-tests







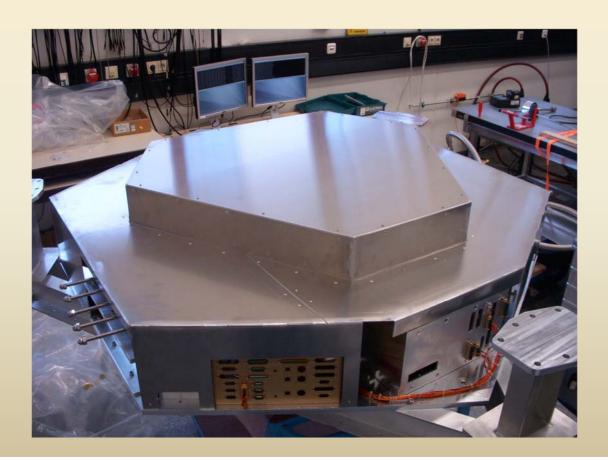
### Status: Telescope Structure



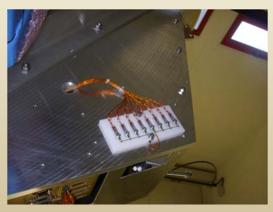
- CFRP Structure
  - Hexapod
  - Optical Bench
  - Sunshield
  - Cover
- Complete (FM)
- Qualified
  - Vibration
  - Acoustic Noise
  - Cover mechansim
  - Space Simulation pending
- Refurbishment after test campaign

# Status: Technological Model

- Model for testing mechanical and electrical interface to S/C
  - Just arrived in Moscow for testing







# QM-Test: Packing







# QM-Test: Transportation to IABG





# Cleanliness







## QM-Test: Mass Properties

		eRosita QM			
Mass [kg]	m	810.3			
CoG [mm] (w.r.t. the origin of ordinates)	х	-34.1			
	Υ	+2.5			
	Z	+1260.4			
MoI [kgm²] (w.r.t. axes passing through CoG parallel to axis system)	I <sub>xx</sub>	619.8			
	lyy	621.0			
	Izz	237.9			

- Mass: includes startrackers and sunsensors
- Possibly 20kg reduction due to actual mirror weight







## QM-Test: Acoustic Noise

• Qualification Level 143db, OK







# QM-Test: Pyro Shock





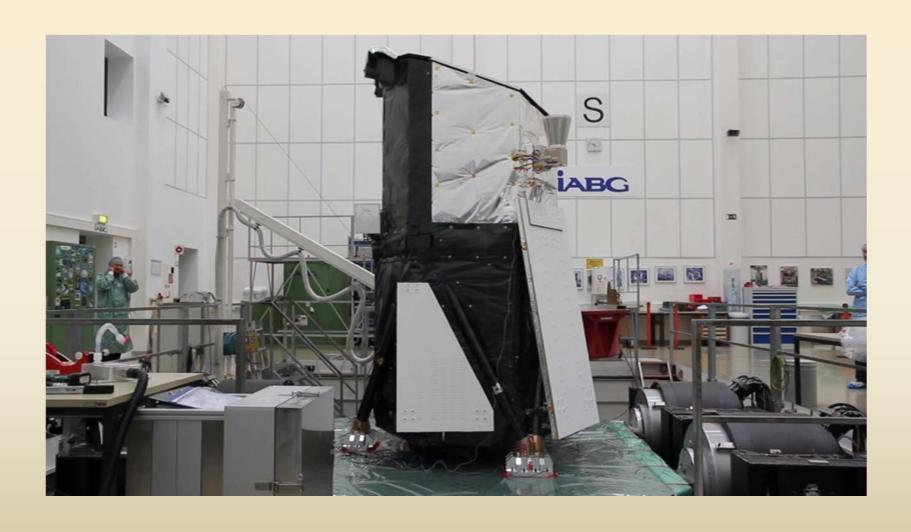
# QM-Test: Vibration

• Qualification level, all three axes, OK





# QM-Test: Vibration

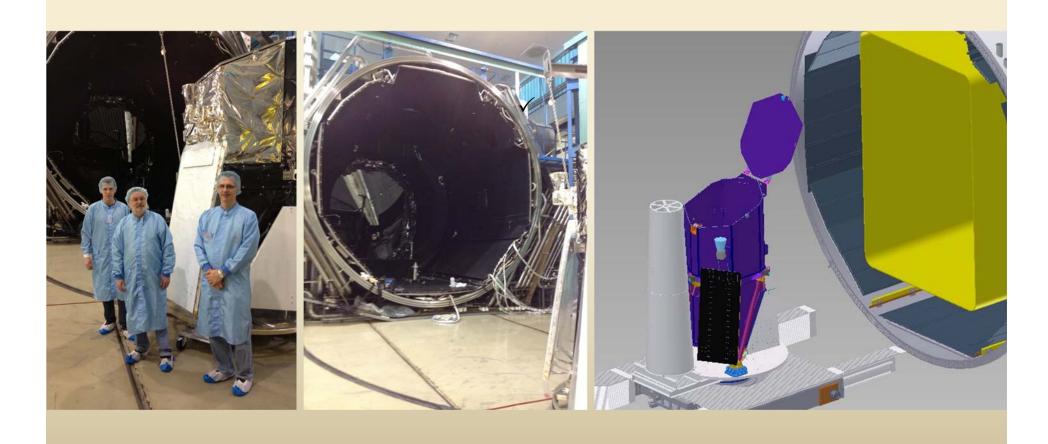


# QM-Test: Cover



# QM-Test: Space Simulation

- Performed in beginning of January
   1 dummy startracker replaced by SED26
  - ART-XC dummy installed



### Summary

- Assembly and completion and testing of most of the components of eROSITA are schedule.
- All environmental and space-simulations tests have been
- As most components come in multiples of at least 7 we have still many calibration campaigns ahead of us for both the mirrors and detectors.
- An End-to-End test at PANTER of the complete telescope is planned for early 2014
- The launch is still scheduled for end of 2014
- Next talk will be on eROSITA in-flight calibration