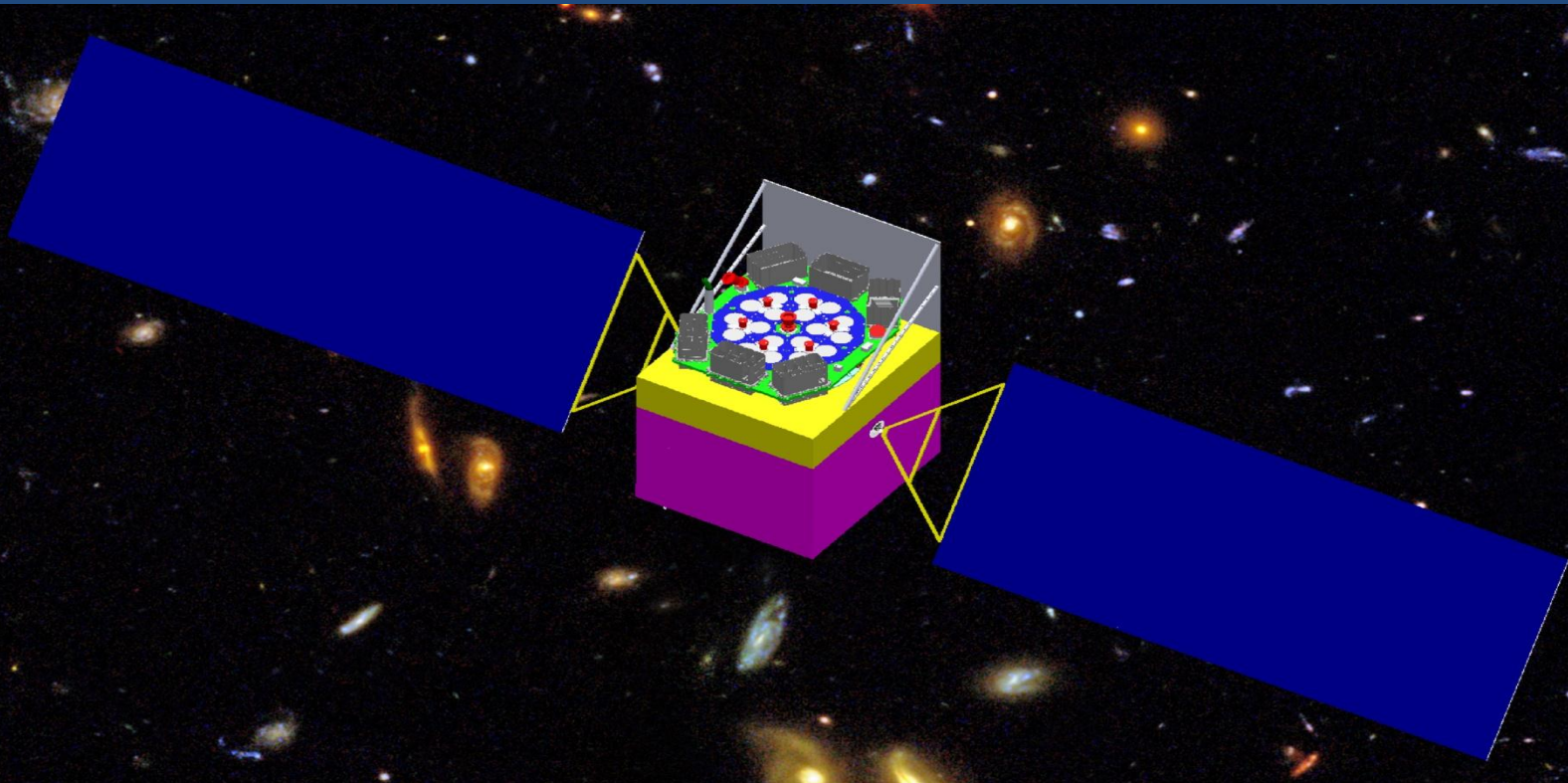


Production of HXMT calibration Files

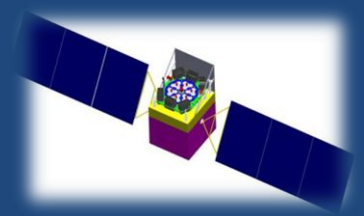


8th IACHEC meeting

LI Xiao-bo

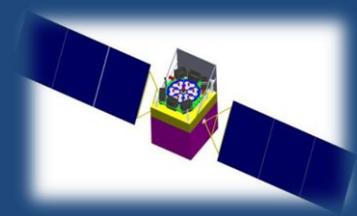
Institute of High Energy Physics, CAS

Outline

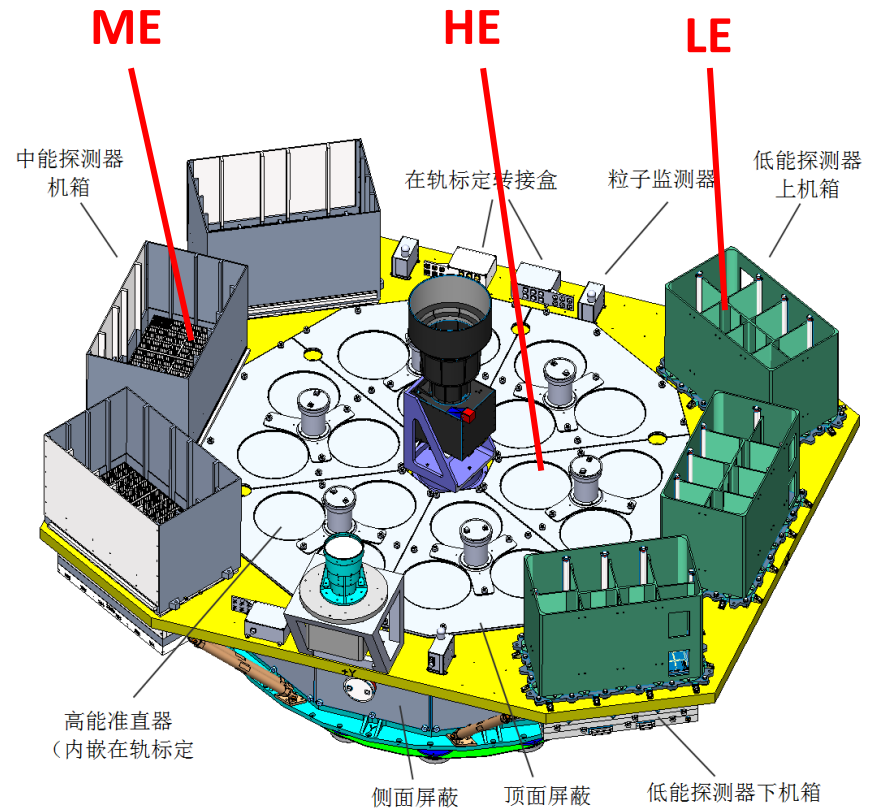


- 1. Hard X-ray Modulation Telescope(HXMT)**
- 2. Flow of calibration data files**
- 3. Calibration parameters for energy response matrix**
 - Energy scale
 - Energy resolution
 - Efficiency
- 4. Monte carlo simulation based on Geant4**
- 5. Performance of HXMT detectors in electronics test**
- 6. Summary**

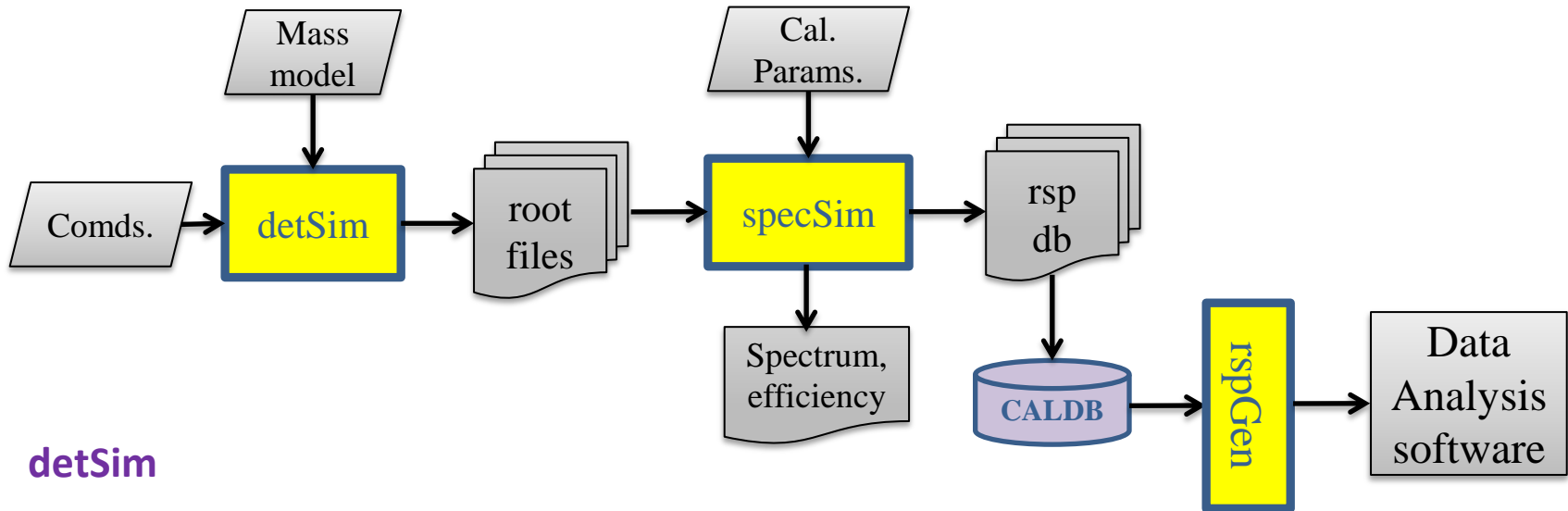
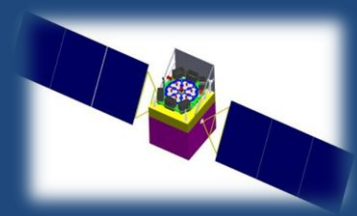
1. HXMT



- **Scientific Objectives:**
 - X-ray all sky survey
 - pointing obs. of X-ray source
- **Orbit: circle, 550km**
 - orbital period: ~90mins
 - orbital inclination: 43 degree
- **Instruments:**
 - HE: NaI(Tl)/CsI(Na), 5000cm², 20-250keV;
 - ME: Si-Pin, 952cm², 5-30keV;
 - LE: CCD, 384cm², 1-15keV



2. Flow of calibration data files



- **detSim**

- **Inputs:**

- Mass model of instruments and environment
 - Some commands to configure the simulation

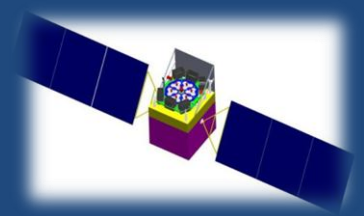
- **Outputs:**

- Raw event files(ROOT files)

- **External dependencies:**

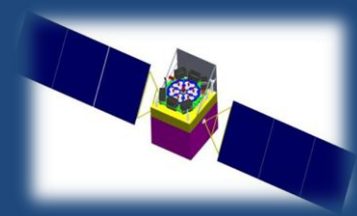
- GEANT4—toolkit for MC of particle interactions with matter from CERN
 - ROOT– Data analysis package from CERN.

Continue:



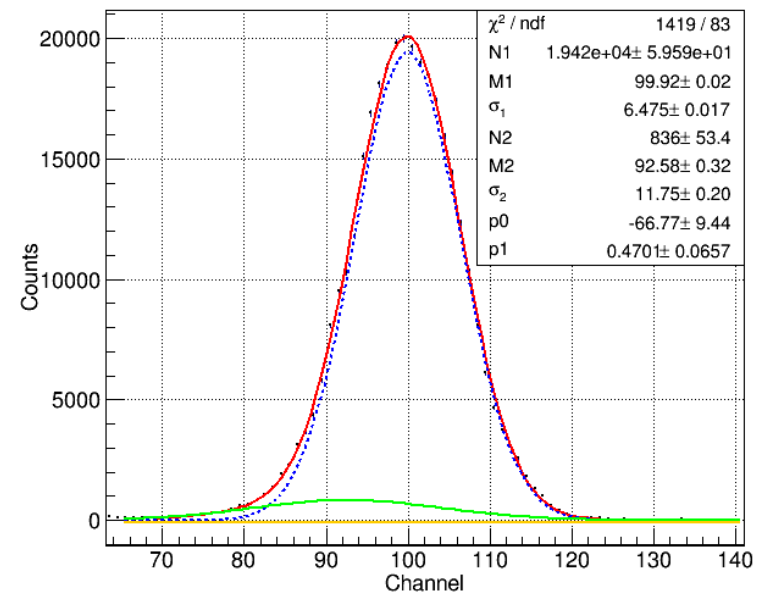
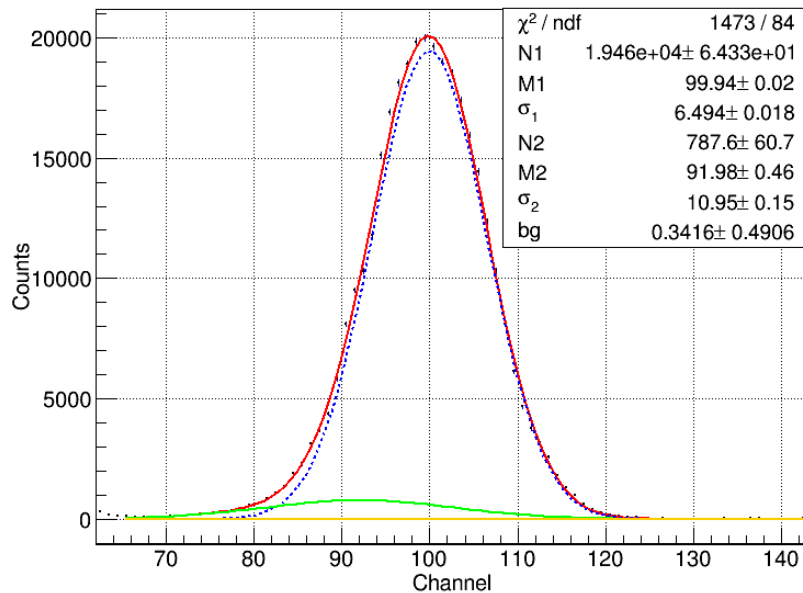
- **specSim**
 - **Inputs:**
 - Raw event root files
 - Calibration parameters from experiment data
 - **Outputs:**
 - Energy response curve and efficiency
 - FITS file, like rmf, arf.
 - **External dependencies:**
 - ROOT—Data analysis package from CERN.
 - CCFITS—FITS data file I/O
- **rspGen**
 - **Inputs:**
 - FITS file from CALDB
 - Some specific FTOOLS developed by HXMT group to generate the RMF of observe ID.
 - **Outputs:**
 - Applicated RMF files
 - **External dependencies:**
 - CCFITS—FITS data file I/O
 - FTOOLS of CALDB related.

3. Calibration parameters

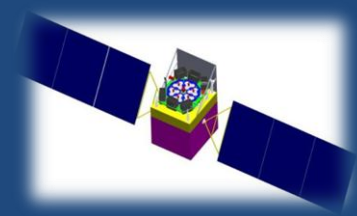


- Energy scale and resolution can be gained through radioactive sources and X beam experiments on ground.

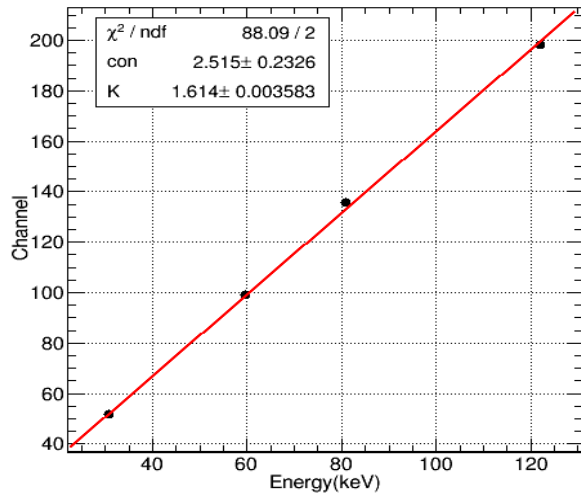
^{241}Am @ HE



- 59.5 keV full energy peak fit:
- The mean and sigma of full energy peak are influenced by the fit range and fit function which we used to describe the background.
- We change the different fit conditions ten more times and get the value and error of full energy peak and its sigma.



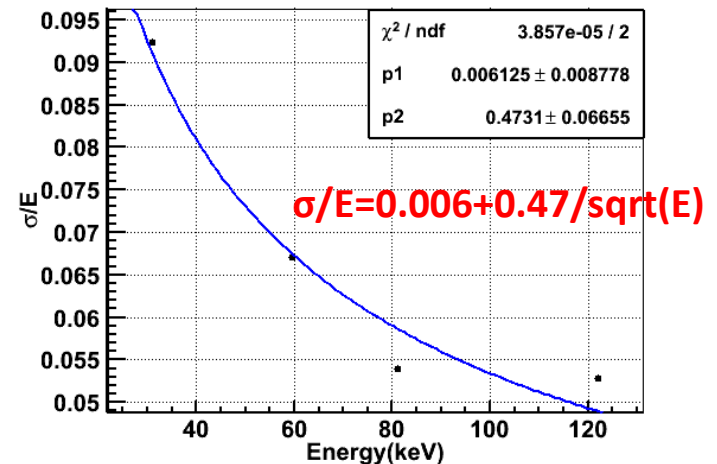
NIM A675(2012)78-83



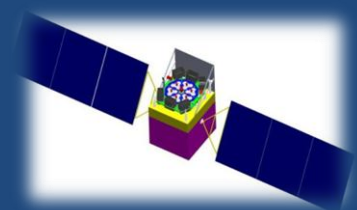
$C=2.515 \cdot E+1.614$

- Fit function for the energy-channel relation:
 - $C=K \cdot E+con$, nonlinear etc.
- The error of energy scale and resolution can be derived through the error propagation .

Fitted FWHM function	Literature		\bar{R}^2	
	Detector	References	NaI	LaBr ₃
$FWHM(E) = a + bE$	Ge	[3,21]	0.98011	0.99301
$FWHM(E) = a + b\sqrt{E}$	Ge	[3]	0.99534	0.99904
$FWHM(E) = a \cdot E^b$	NaI	[15]	0.99310	0.99884
$FWHM(E) = a + b\sqrt{E + cE^2}$	NaI, LaCl ₃	[16-18]	0.99581	0.99900
$FWHM(E) = a\sqrt{E} + bE$	NaI, LaBr ₃	[3,19]	0.99138	0.99873
$FWHM(E) = a\sqrt{E}$	NaI	[20]	0.98729	0.99850
$FWHM(E) = \sqrt{a + bE}$	NaI, Ge	[3,21]	0.97922	0.99923
$FWHM(E) = \sqrt{a + bE + cE^2}$	Ge	[3,21]	0.91769	0.99917
$FWHM(E) = a + bE + cE^2$	NaI, LaBr ₃	[22]	0.99713	0.99979
$FWHM(E) = a + b\sqrt{E} + c$	NaI, LaBr ₃	This study	0.99492	0.99922
$FWHM(E) = a + bE + cE^2 + d\sqrt{E}$	NaI, LaBr ₃	This study	0.99699	0.99980
$FWHM(E) = a + bE + c\sqrt{E}$	NaI, LaBr ₃	This study	0.99550	0.99904
$FWHM(E) = aE + bE^2$	NaI, LaBr ₃	This study	0.99695	0.98146



Efficiency calibration



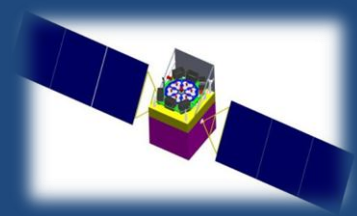
- One method is to use calibrated radioactive sources
 - The reference activities were provided in a calibration certificate by the supplier of the radioactive sources.
 - The activity at the day of measurements were calculated by taking into account the time elapsed since the calibration reference day.

$$\varepsilon(E) = \frac{N}{A * t * p(E) * T(E) * \exp - (\mu_{air} * d_{air}) * \Omega/4\pi}$$

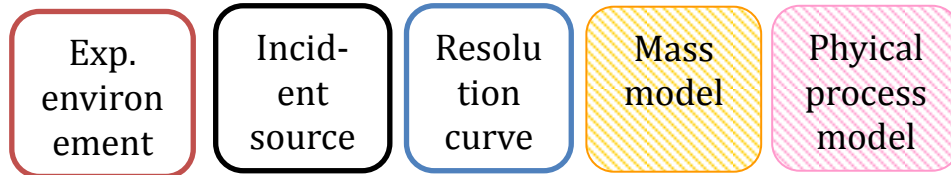
- N is the detected events, A is the activity of the source, T(E) is the transition probability of entrance window, p(E) is branch ratio.
- For HE, N must meet the energy cut(20—250 keV) and pulse width cut to get NaI-like events. For ME and LE, N should meet the energy cut.
- Error of efficiency can be estimated.

$$\delta\varepsilon_{\text{exp}} = \varepsilon_{\text{exp}} \sqrt{\left(\frac{\delta N}{N}\right)^2 + \left(\frac{\delta A}{A}\right)^2 + \left(\frac{\delta p}{p}\right)^2}$$

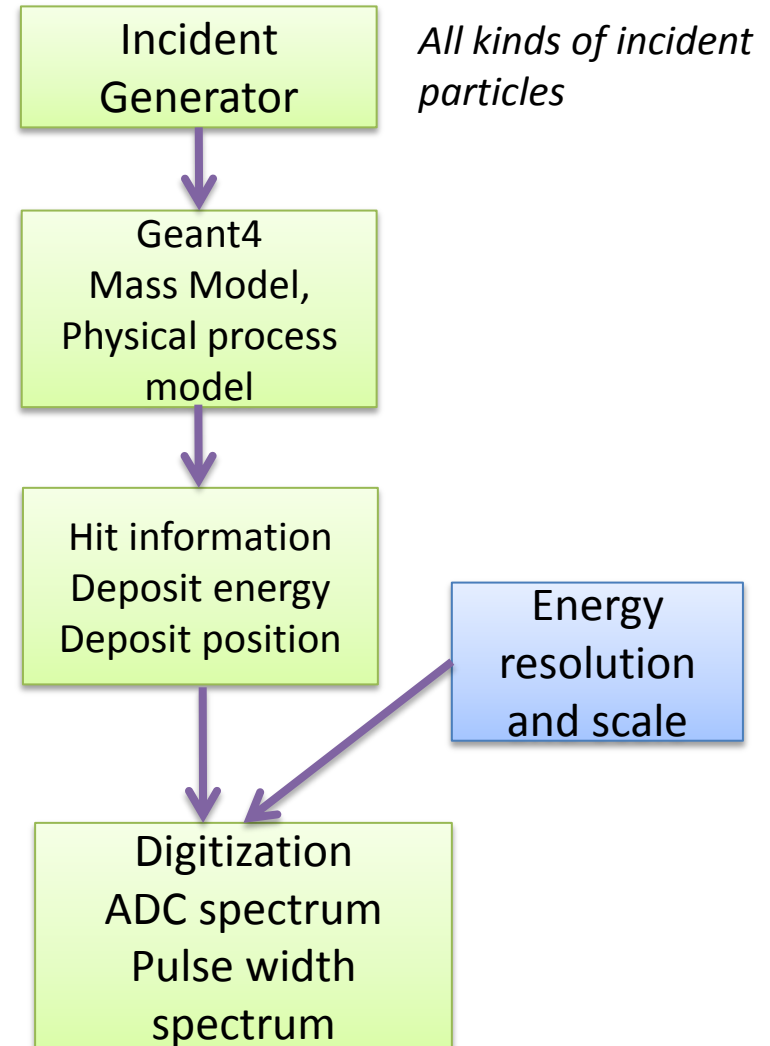
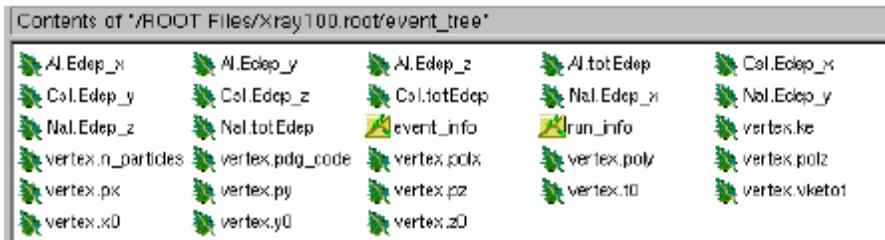
4. MC simulation based on G4



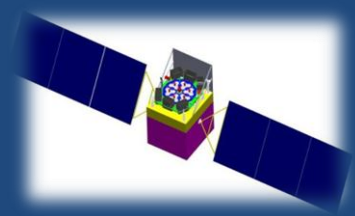
- Parameters for MC tuning:



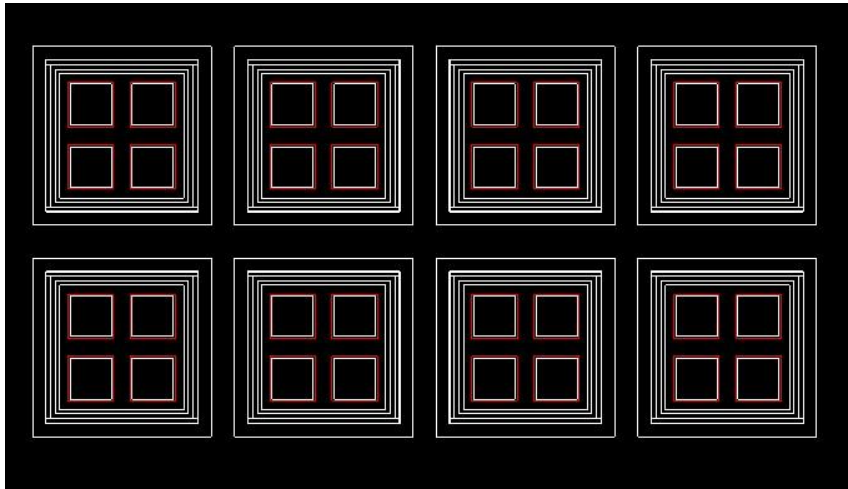
- Output hit information of MC:



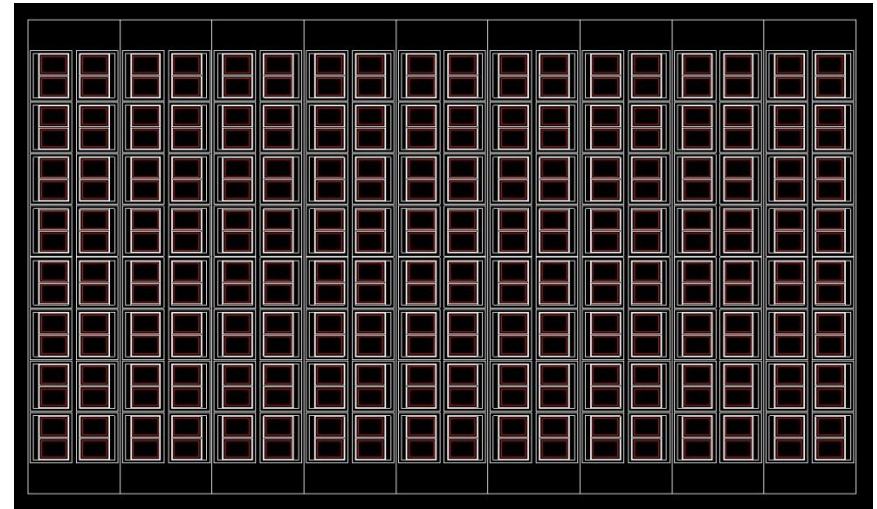
Visualization of HXMT detectors



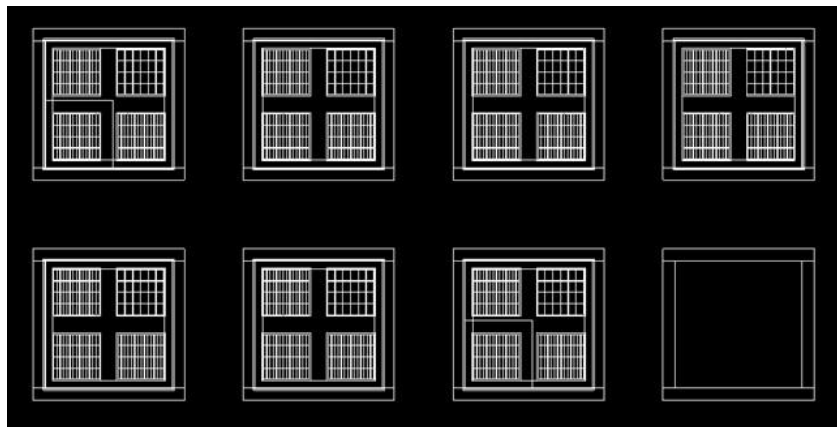
SCD detectors of LE in one box:



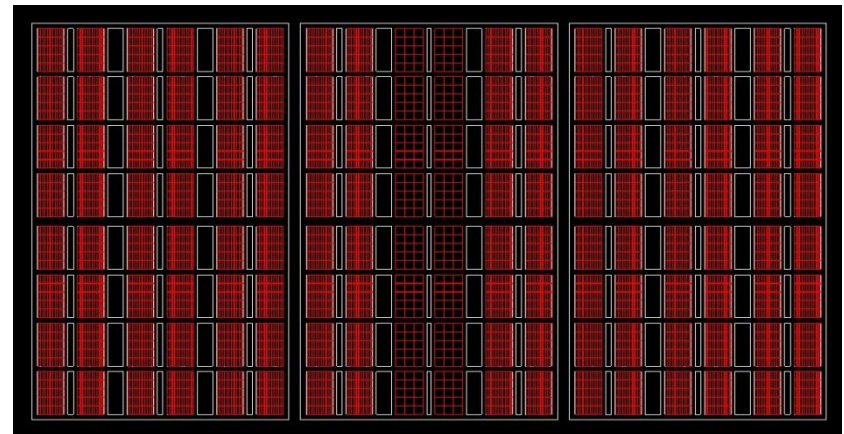
Si-Pin detectors of ME in one box:



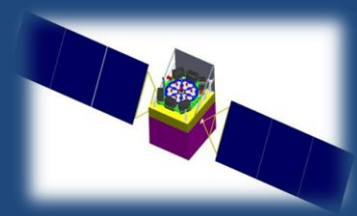
Collimators of LE:



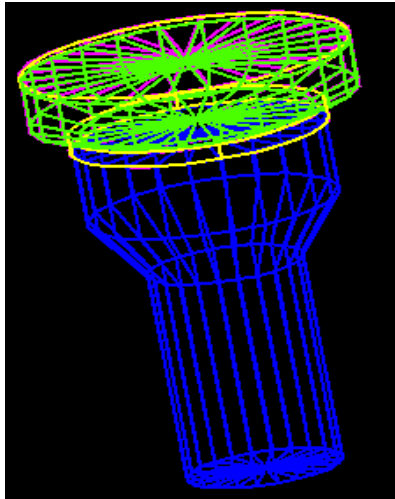
Collimators of ME:



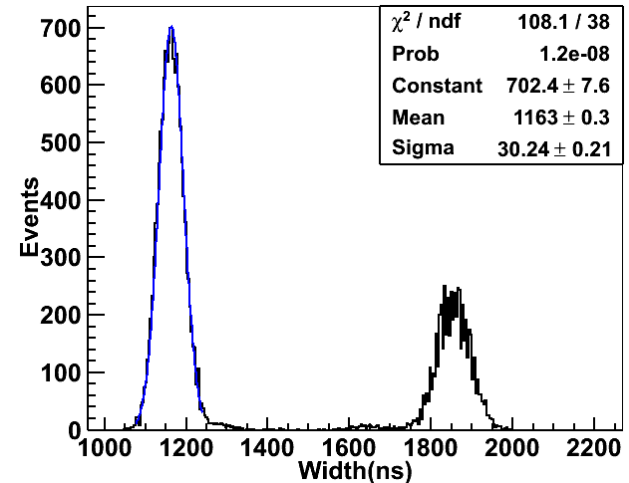
MC Spectrum of HE detector



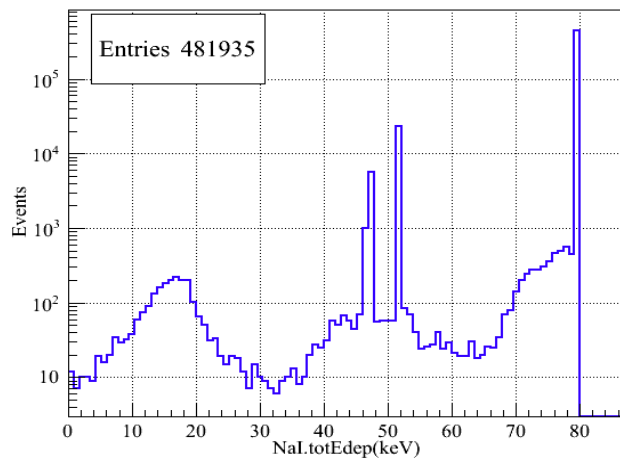
HE detector visualization:



Pulse width spectrum:

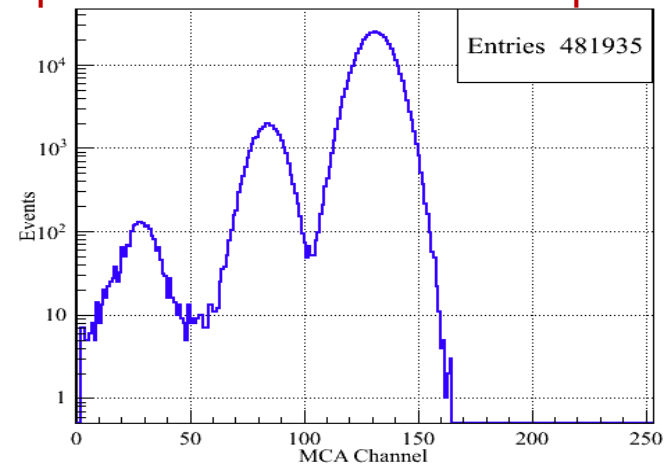


detSim—raw physical spectrum

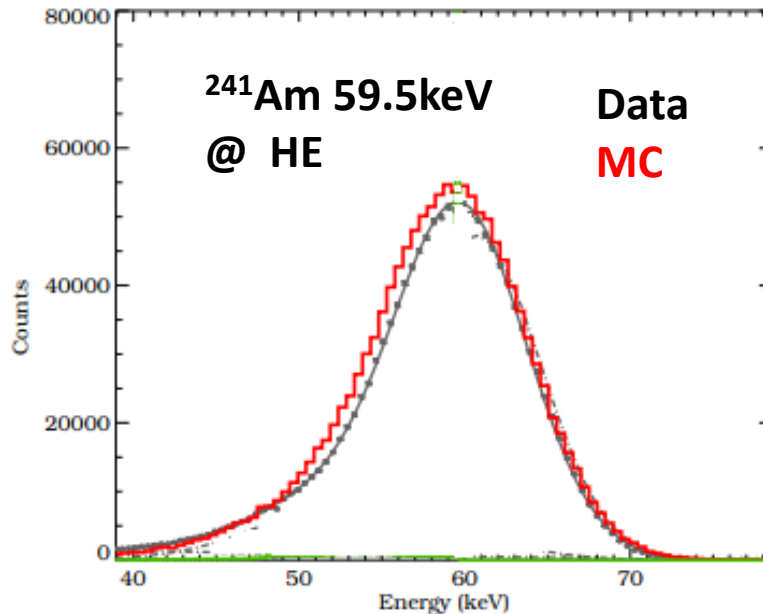
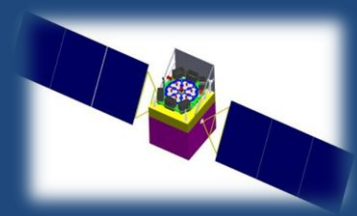


Energy scale
and
resolution

specSim—instrument-like spectrum

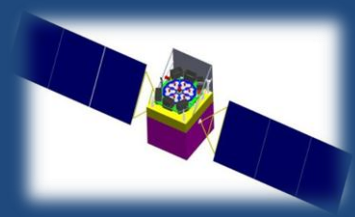


MC energy spectrum vs Data

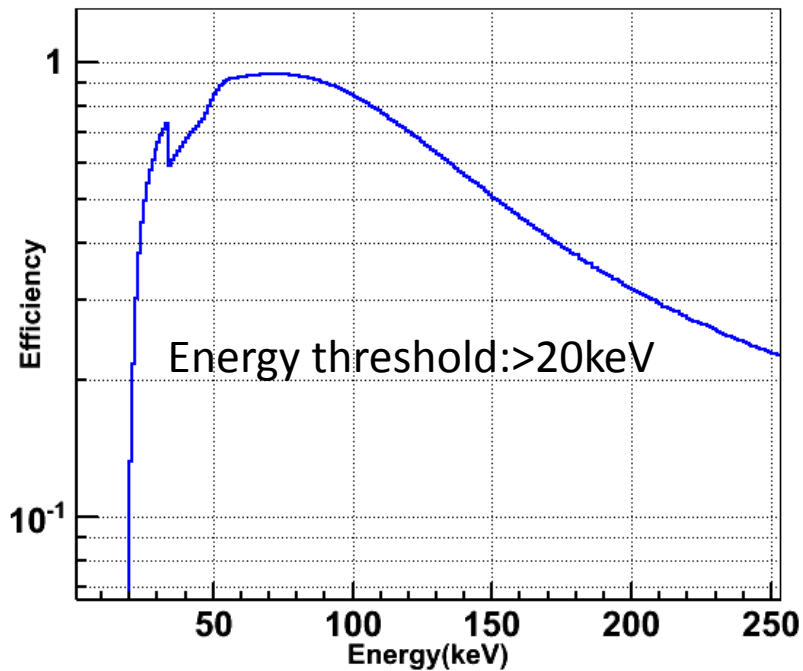


- The difference between MC spectrum and data decides the precision of the energy response.
- Experience to tune the MC parameters to reduce the difference .
- How big system error will be produced if we have several percent difference between the MC spectrum and data?

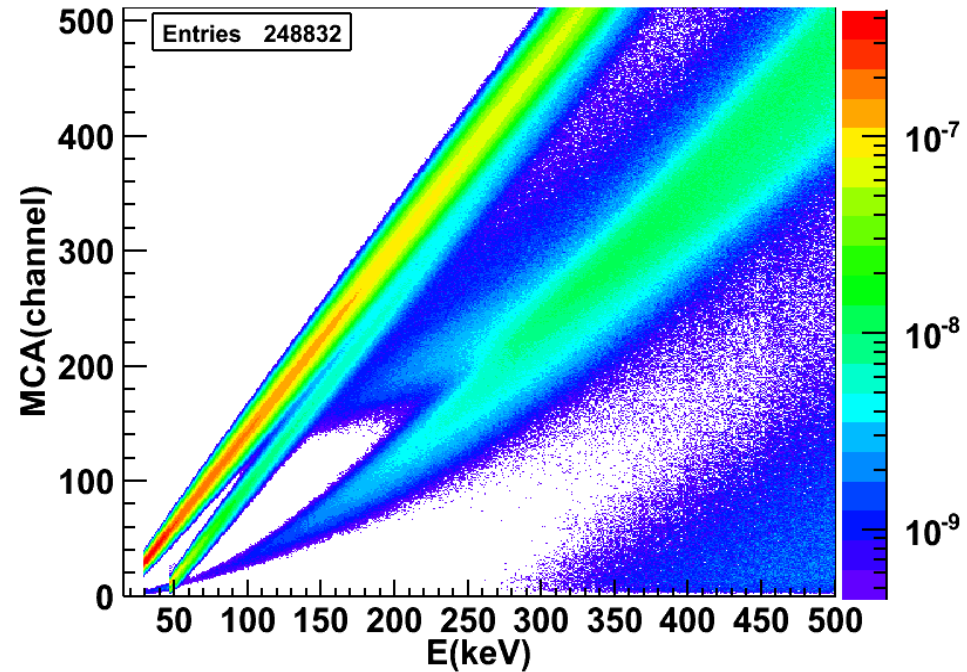
HE Detector response:



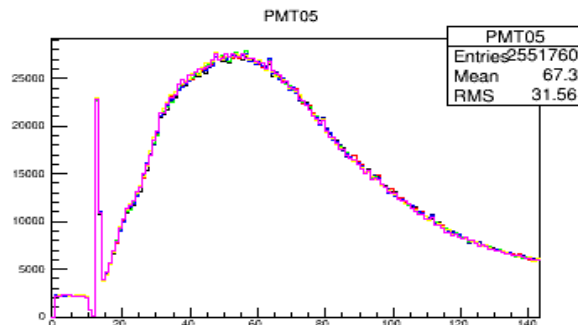
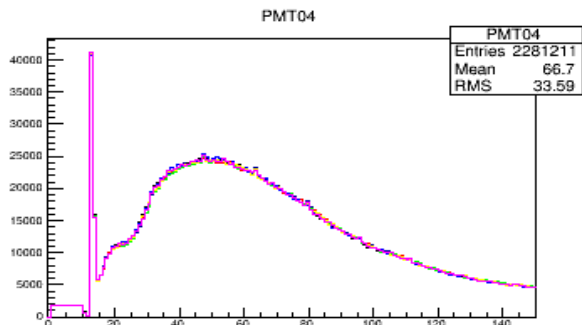
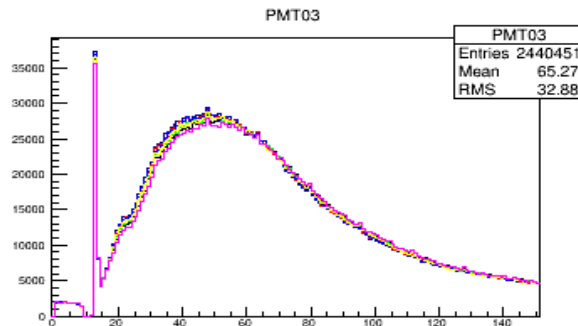
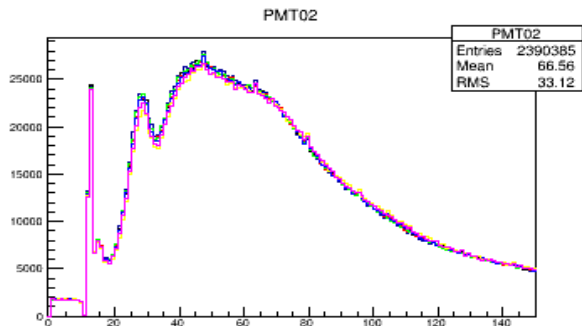
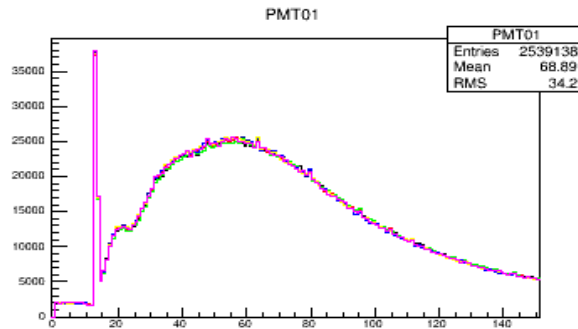
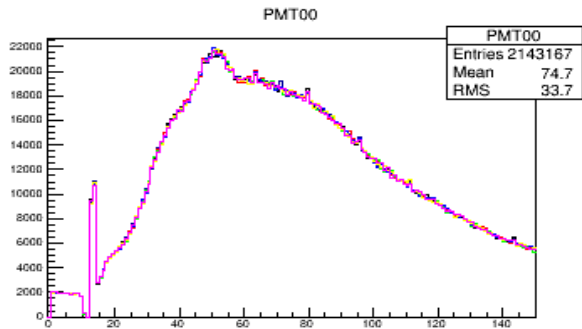
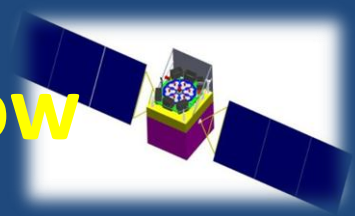
Efficiency:



RMF:



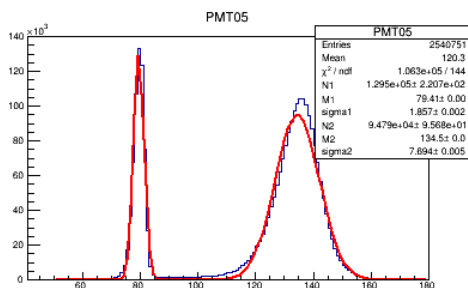
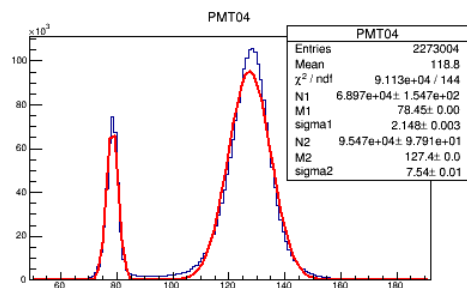
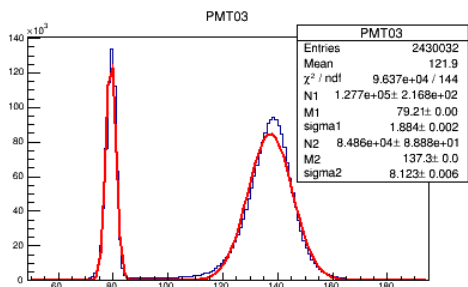
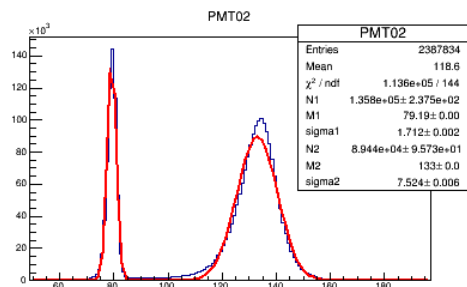
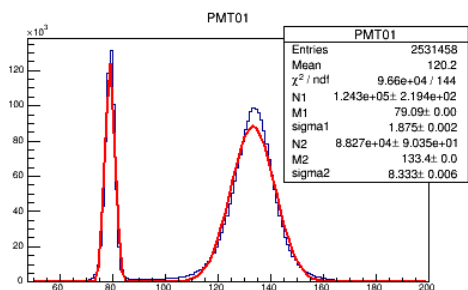
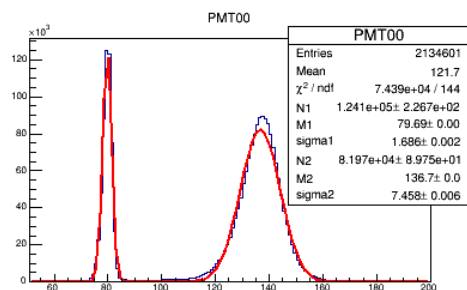
5. Performance of HXMT detectors now



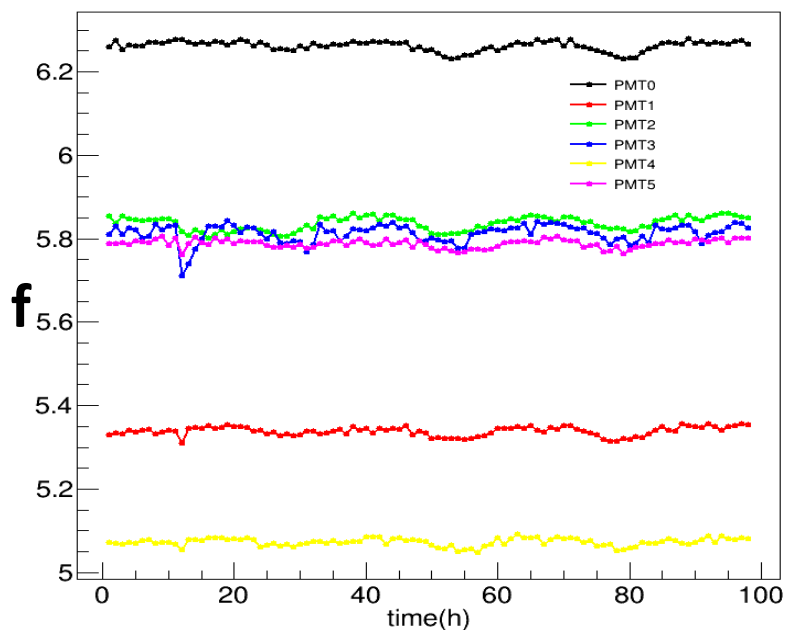
Energy spectrum of 6 detectors in one group.

One color represents one hour.

Pulse width spectrum

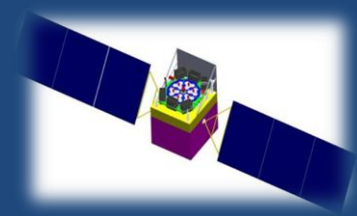


$$f = (M2 - M1) / (\sigma1 + \sigma2)$$

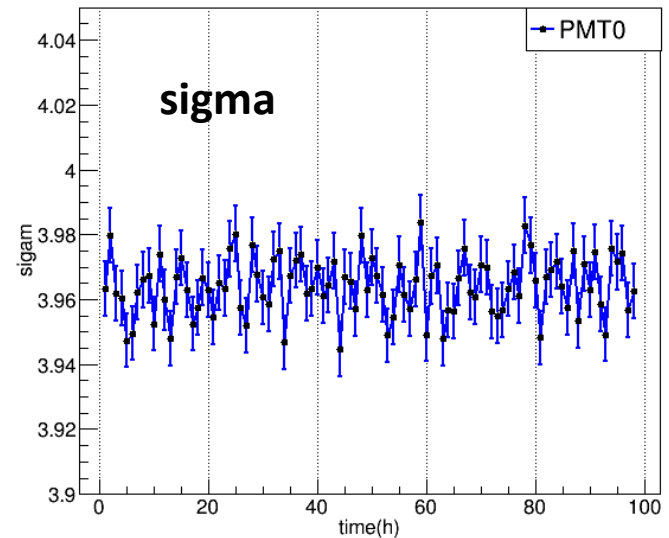
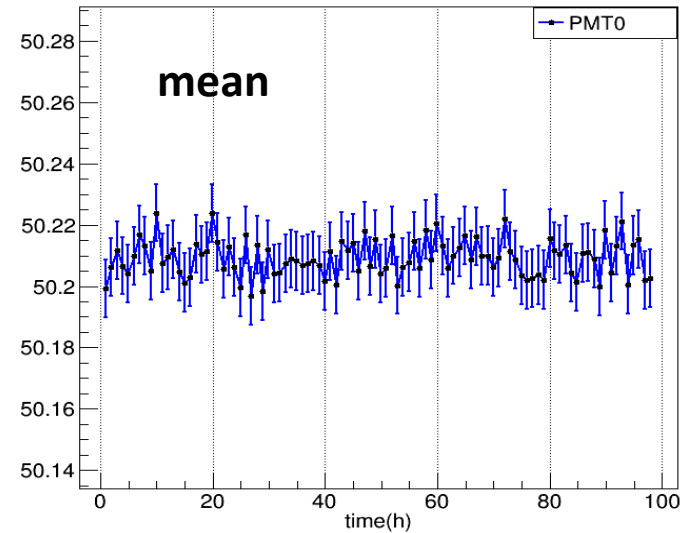
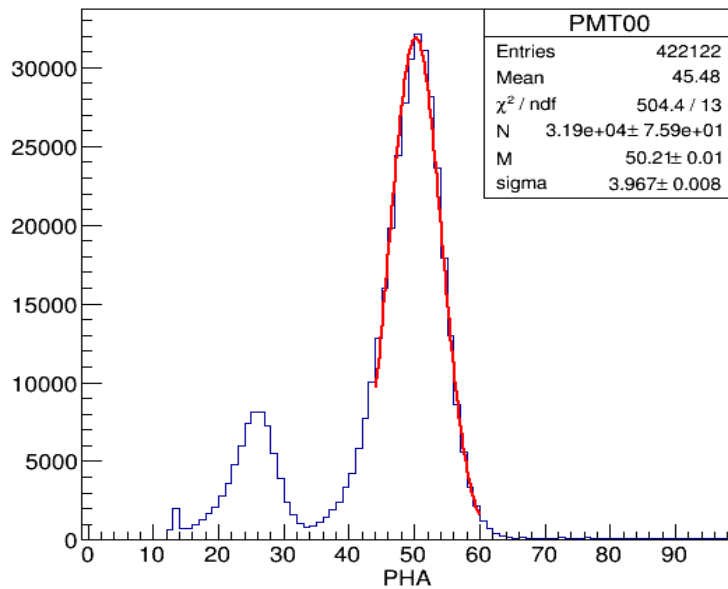


100 hours of test for one group of HE.

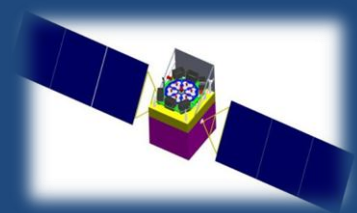
Automatic gain control



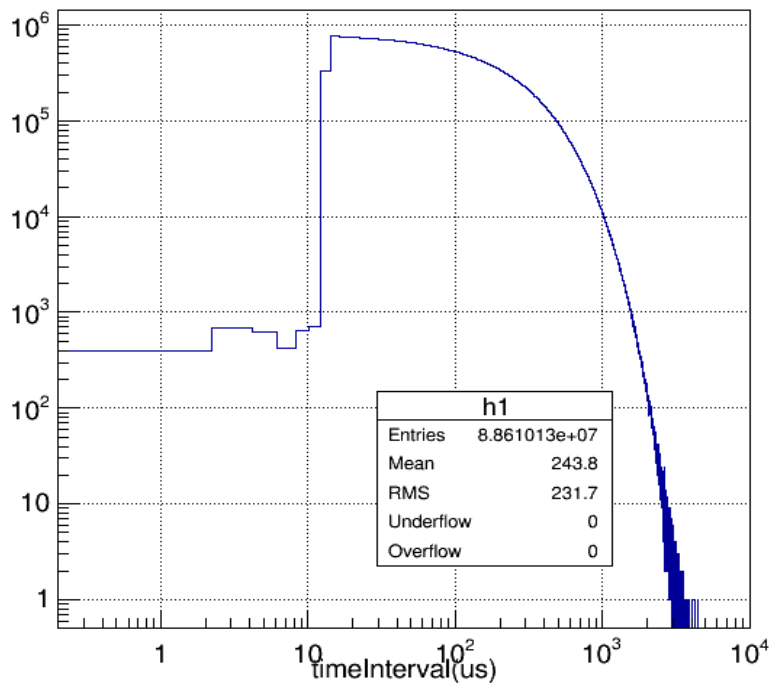
- ^{241}Am source of in-orbit calibration detector was used on one PMT in the 100 hours test.



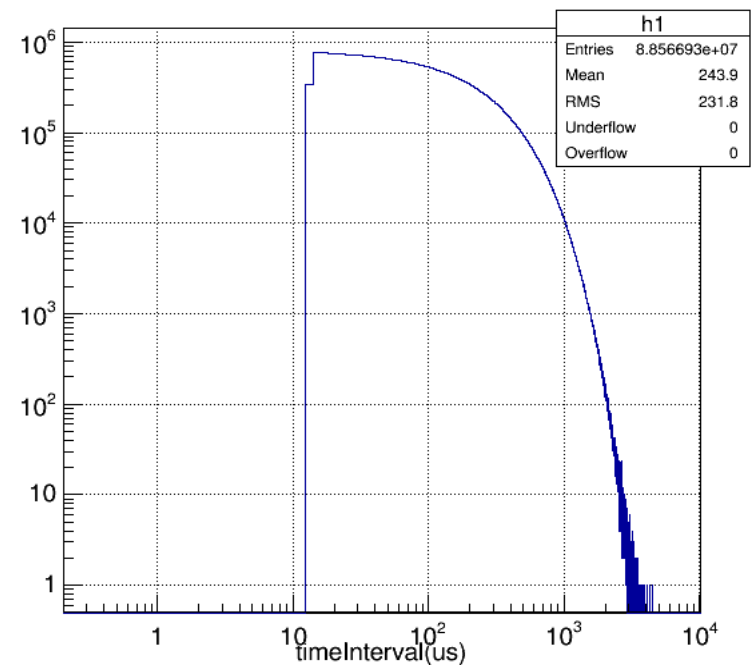
Time interval distribution



- Time Interval distribution of 6 HE detectors in one group.
- Time used to process one signal needs 12us.

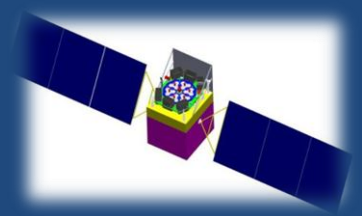


All Events, including GPS second events, physical events and radioactive events.



Physical events or radioactive events

6. Summary



- **Elaborate calibration experiments will be done in IHEP soon.**
- **Detector simulation:**
 - Simulation software ✓
 - Preliminary MC result vs experiment data ✓
 - Long-term work to tune MC parameters
- **Calibration parameters and their accuracy**
 - Energy scale, resolution and efficiency ✓
 - Timing related
- **CALDB**
 - Design the CALDB of HXMT ✓
 - Calibration software for users

Thank you for your attention!