Chandra Calibration Status



IACHEC Meeting May 12, 2014

Contamination on the ACIS Filters

Abell 1795



Blazars

Mkn 421 PKS 2111-304

E0102-72









Contamination on the ACIS filters



Plots are based on data acquired through 2013.



From LETG/ACIS-S data





Contamination on the ACIS filters

Spectral fits with ACIS contamination model V7



Contamination on the ACIS filters

April 2014 observations fit with ACIS contamination model V7



Contamination on the ACIS Filters

Mkn421 LETG/ACIS-S big dither observations (March 2014)



ACIS Temperature-Dependent CTI Correction



ACIS Temperature-Dependent CTI Correction





12: compare Obsids: 13994, 13995

There are short term gain variations on I0 and I2 that require self-calibration.

HRMA Thermal Gradients



Thermal gradients across the HRMA in the top-to-bottom direction will affect the PSF in imaging data and the LSF in gratings data.



ACIS Low energy Gain

Plots are based on a set of LETG/ACIS-S observations taken at different off-sets along the dispersion axis.

Ratio of photon energy derived from detector gain to that derived from the dispersion relation



Chandra Internal Cross-Calibration with G21.5-09



Mkn 421 Interleaved 0.85-1.5 keV



Mkn 421 Interleaved 1.5-4 keV



Present Calibration Activities

- Update ACIS contamination model, including: new time-dependence, elemental ratios, spatial gradients for each element and possibly a N component.
- Release revised gain tables for the BI chips (S1 and S3) to reduce residuals at low energies.
- Release a new OSIP file consistent with the new gain table.
- > Adjust 0th order HETG efficiency.
- Adjust the HRC-S QE and LETG 1st order efficiency.

LETG/HRC-S Calibration

Uncorrected

Corrected



Time-dependent tilt in the dispersion axis



LETG/HRC-S Calibration



After correcting for timedependent tilts and 0th order position. With these corrections the extraction region and background can be reduced by 25%



ew Version

HETG 0th Order Calibration



SN 1987a 0.5-5keV band



Triangles – fit to 0th order Empty squares – model independent flux Filled squares – simultaneous fit to 0th and 1st order.

Chandra Internal Cross-Calibration with Mkn 421

