# **ASTROSAT CALIBRATION**

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## ASTROSAT

A Satellite Mission for Multi-wavelength Astronomy Indian Space Research Organisation http://astrosat.iucaa.in/

India's first dedicated astronomical observatory in space

- Five science payloads covering Opt/UV to hard X-ray bands for simultaneous multi-wavelength timing and spectroscopy
- All science payloads delivered
- Final integration ongoing
- Being readied for Oct 2015 launch















Orbit: 650 km 6° - 8°









ASTROSAT will be a proposal-driven observatory-class mission (AO will be opened I year after launch; I st 6m PV, followed by 6m of GT)

Major interest areas of X-ray instruments will include

- Strongly Magnetic Neutron Stars: Cyclotron Spectra, HE continuum (structure and evolution of neutron star magnetic fields, radiation processes)
- Wideband Spectral Variability Monitoring (accretion disk geometry, emission mechanism, QPO origin, disk-jet connection)
- Transients



#### **ASTROSAT** Simulation

Target : Vela X-1







### Ground Calibration of ASTROSAT X-ray Payloads

- Spectral channel to energy relation
- Spectral resolution
- Effective area
- Timing
- Imaging and FOV

**Radioactive Sources** 

(simulations; need PV for final calibration)

Radioactive Sources, X-ray gun, optical

Supplemented by Geant4 simulations

**Products generated** 

Response of detector elements,

Individual and Collective

Dependence on instrument settings, temperature

**HEASARC CALDB** format

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LAXPC: proportional counters (3 units)



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0

0

 $10^{4}$ 

Collimator Profile from source scan across the detector

#### Timing properties: power spectrum

Frequency (Hz)

3×104

2×104

 $4 \times 10^{4}$ 

5×104

#### CZTI: (Cadmium Zinc Telluride pixellated detector array behind Coded Mask)



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#### **SXT:** (CCD at the focus of foil mirror optics)







Optical PSF

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-20

Energy (keV)

-20

Positions-X, mm

-400

-200

Positions-Y, mm

Energy (keV)

200

400

IACHEC 2015 Beijing

10

20

Position along the wire (mm)

0

-10

## **In-flight Calibration**

- Initial calibration during Performance Verification (PV) phase: 6 months
- Periodic calibration later in flight (2% of mission time reserved)

Calibration objectives during PV

SXT	LAXPC	CZTI	SSM
<ul> <li>Alignment</li> <li>PSF</li> <li>Effective Area</li> <li>Spectral Response</li> <li>Background</li> <li>Timing</li> <li>Contamination</li> <li>CTI</li> </ul>	<ul> <li>Alignment</li> <li>FOV</li> <li>Effective Area</li> <li>Spectral Response</li> <li>Background</li> <li>Timing</li> </ul>	<ul> <li>Alignment</li> <li>FOV</li> <li>Effective Area</li> <li>Spectral Response</li> <li>Background</li> <li>Timing</li> <li>CAM Response</li> </ul>	<ul> <li>Alignment</li> <li>FOV</li> <li>Effective Area</li> <li>Spectral Response</li> <li>Background</li> <li>Timing</li> <li>CAM and Wire</li> <li>Platform Rotation</li> </ul>

- Target classes: Stars, Isolated WD & NS, CVs, XRBs, SNRs, AGNs, Clusters
- Several bright, hard sources will be used during these calibration runs
- Many are variable essential to establish source characteristics via simultaneous observations with other missions, e.g. SWIFT, NuSTAR

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## **ASTROSAT PV** phase schedule

#### Time after launch (months)



- Request short (~Iks), simultaneous SWIFT observations of selected bright targets from ASTROSAT PV phase list
- Will plan ASTROSAT observations concurrent with scheduled NuSTAR targets request collaboration for calibration
- Post-PV, will participate in periodic multi-mission calibration campaigns

