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# Evolution of temperature-dependent CTI correction for ACIS

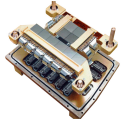
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R. Nick Durham (CXC)



# Summary

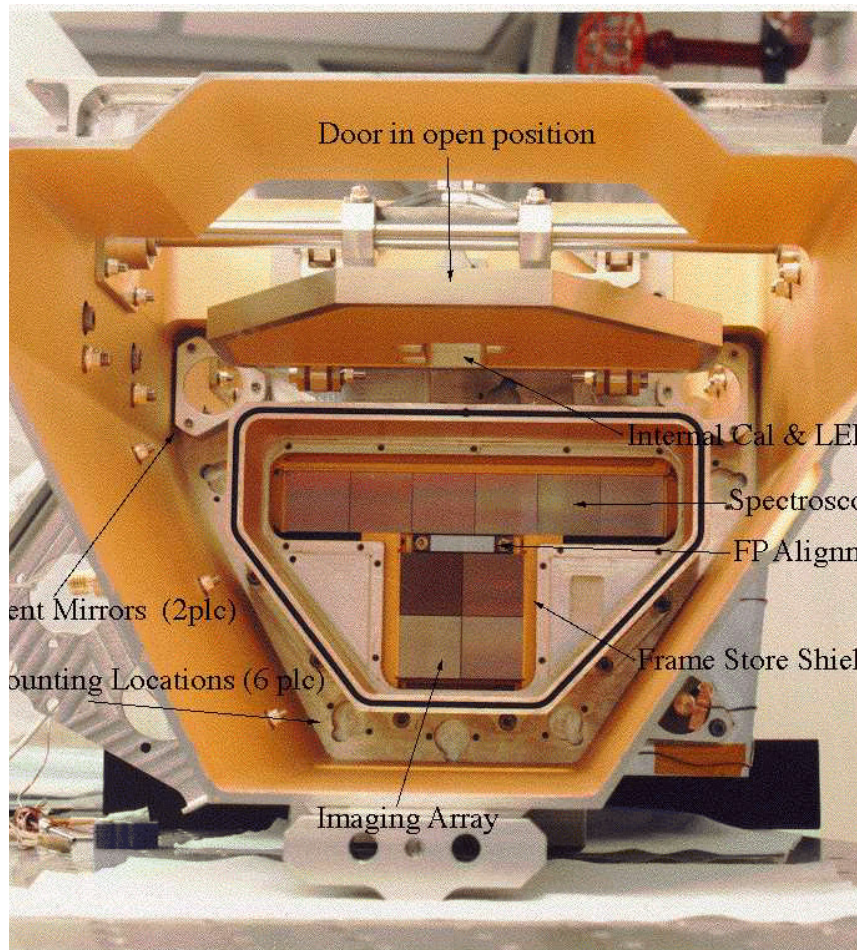
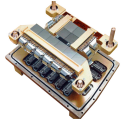
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- ACIS CTI correction last calibrated in 2005
  - Pulseheight corrected with time-dependent gain
  - Width changes are smaller, not currently corrected
- Temperature-dependence added in 2010
- Continuing radiation damage:
  - $dCTI \sim 2 \times 10^{-6} / \text{yr}$  (FI);  $1 \times 10^{-6} / \text{yr}$  (BI)
- Changing thermal environment
- How's the current calibration doing?



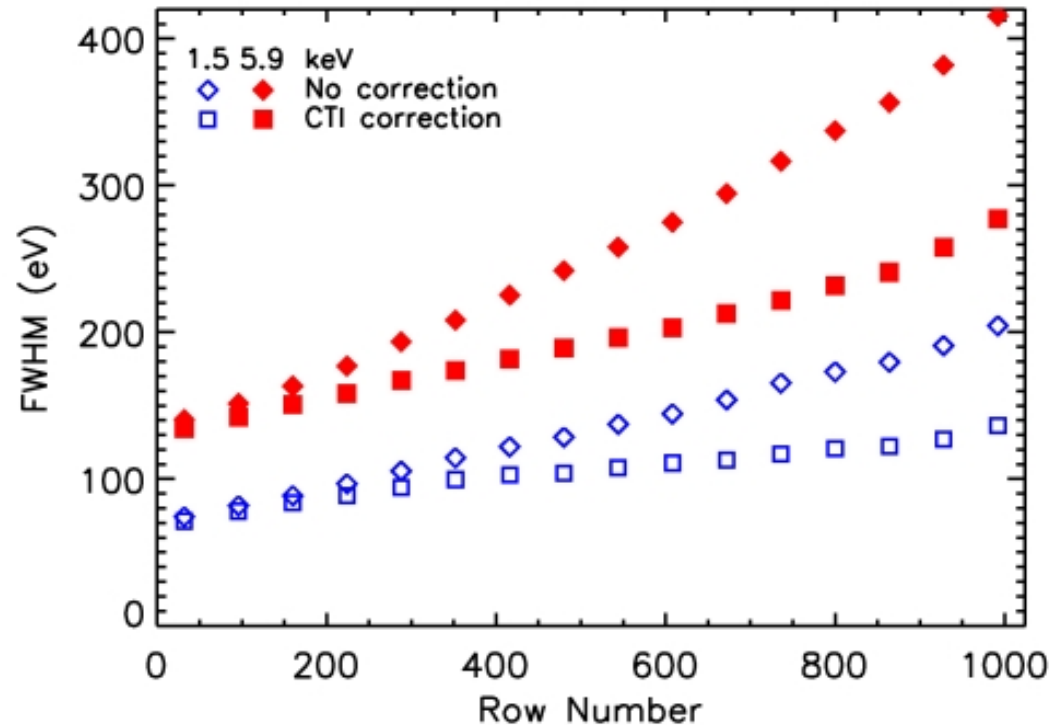
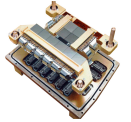
# Advanced CCD Imaging Spectrometer



- Frame transfer design
  - 1024 x 1024 pixels / CCD
  - 24  $\mu\text{m}$  / pixel
- 8 front-illuminated CCDs (FI)
  - Depletion depth: 64–76  $\mu\text{m}$
- 2 back-illuminated CCDs (BI)
  - Depletion depth: 30–40  $\mu\text{m}$
- Temperature set point  $-120^{\circ}\text{C}$
- Initial CTI
  - CTI (FI)  $\ll 10^{-6}$ ; CTI (BI)  $\sim 10^{-5}$
- Radiation belt passages in 1999
  - dCTI (FI)  $\sim 10^{-4}$ ; dCTI (BI)  $\sim 0$
- Continuing radiation damage
  - dCTI (FI)  $\sim 3 \times 10^{-6}$  / yr
  - dCTI (BI)  $\sim 1 \times 10^{-6}$  / yr



# ACIS CTI correction

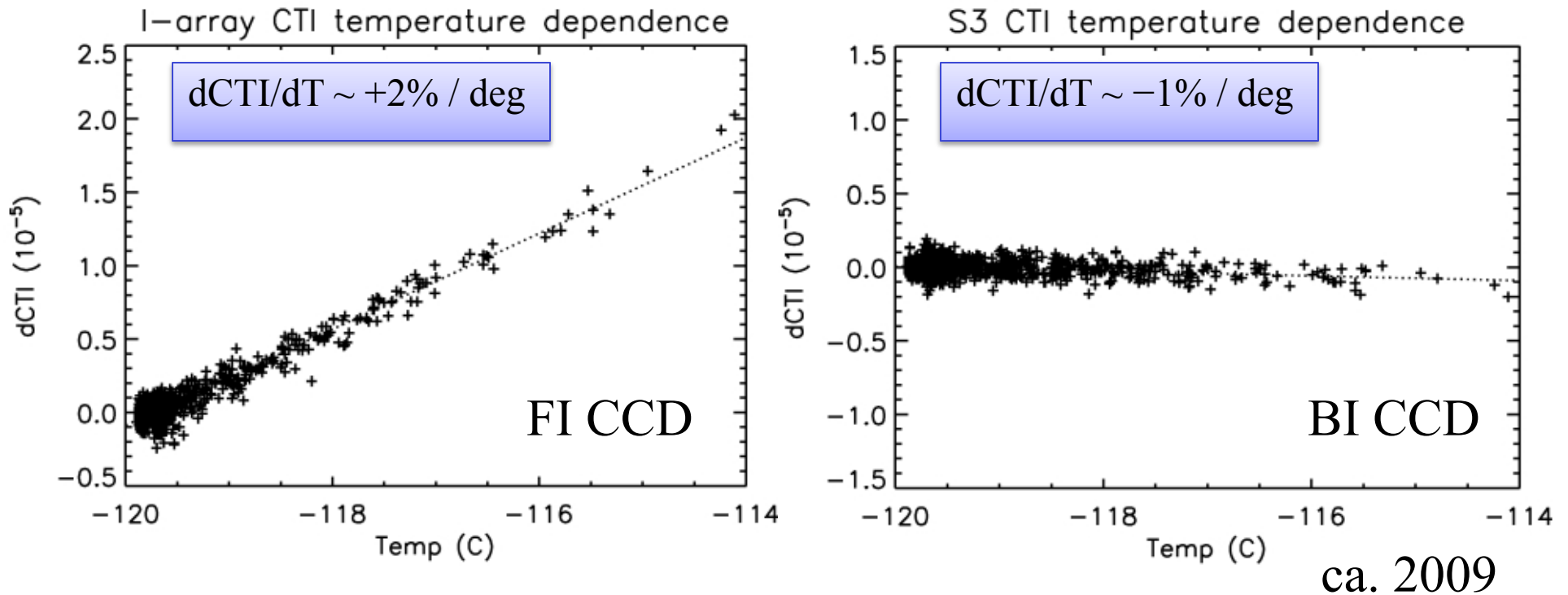
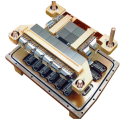


- Incorporated into *CIAO* tool `acis_process_events`
- Reconstruction of original X-ray event island in the absence of CTI
- Removes position dependence of pulseheight
- Significantly improves spectral resolution and detector uniformity
- Charge loss is stochastic – cannot recover all of lost performance

Grant+ 2004, Proc. SPIE 5501



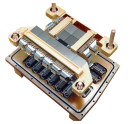
# CTI dependent on temperature



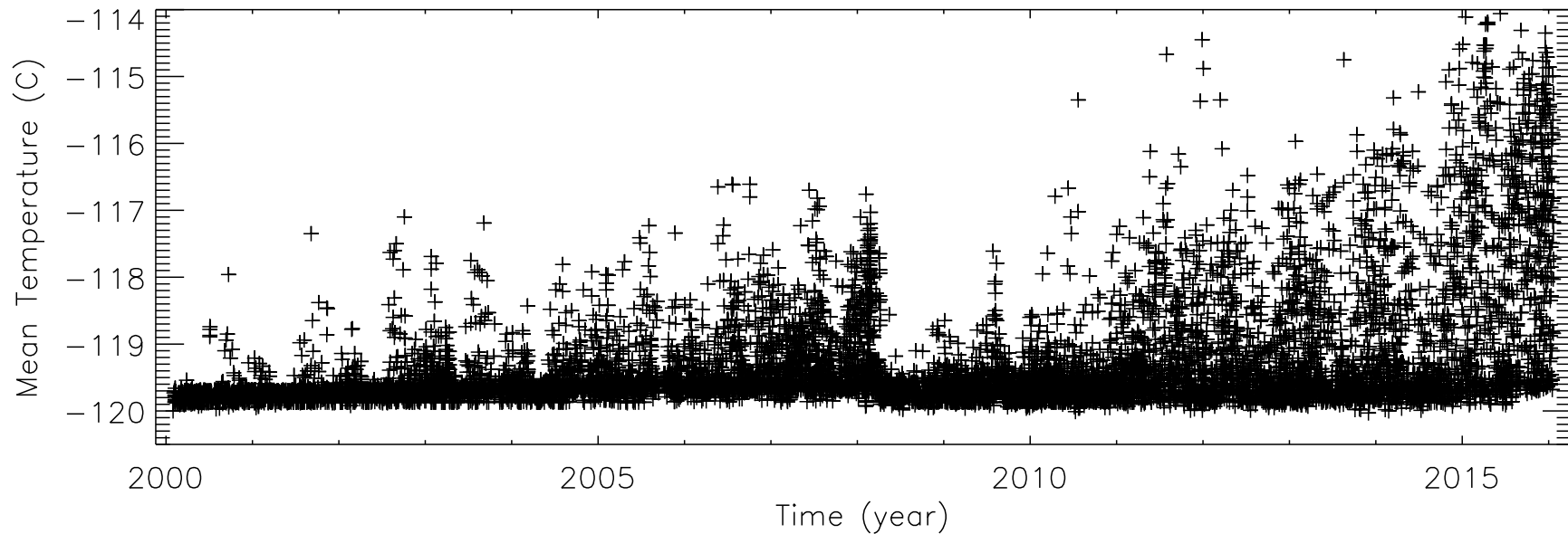
- Charge traps have temperature dependent re-emission time constants
- Roughly linear for small temperature deviations
- Causes temperature-dependent performance
- More important for FI than BI CCDs



# Focal plane temperature excursions



ACIS Science Observations

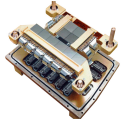


- Thermal control has become more difficult with time
  - ACIS cooling is less efficient for some spacecraft orientations
  - Constraints of other Chandra components not always favorable for ACIS
- Observations with “cold” temperatures ( $T < -119.2^{\circ}\text{C}$ )
  - 99% in 2000; 68% in 2007; 33% in 2015



# Implication for Calibration

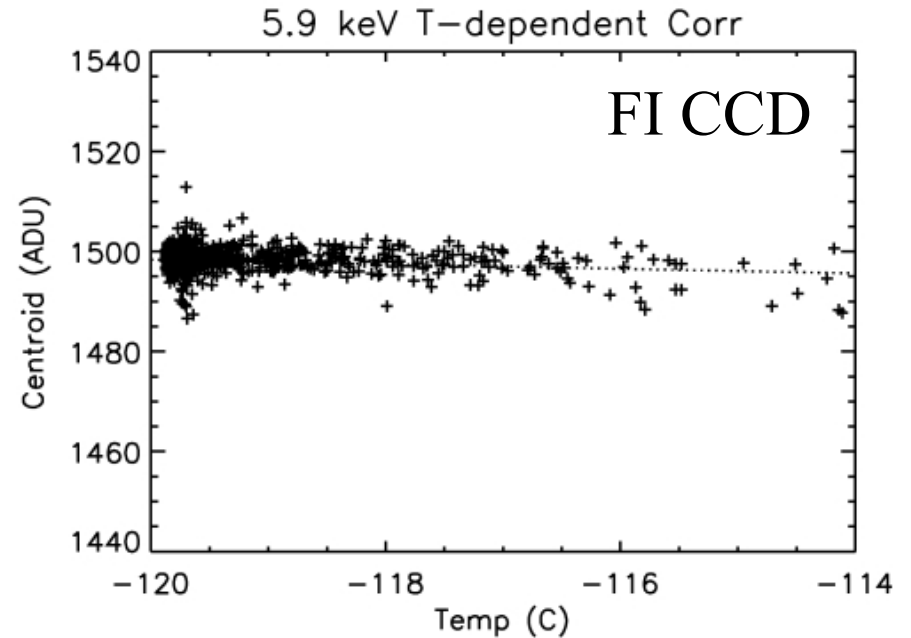
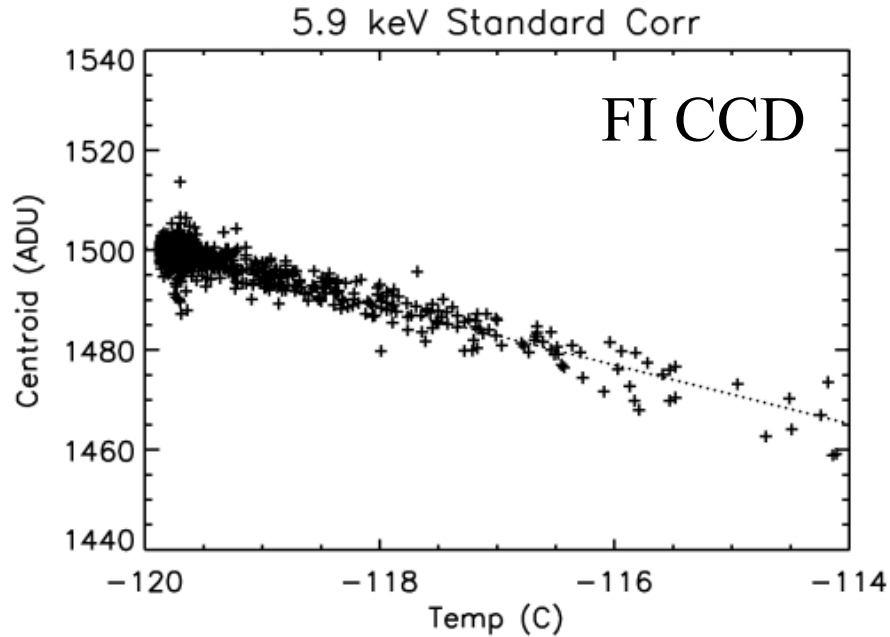
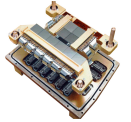
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- Temperature variations can produce significant calibration changes for some locations on some CCDs
- Warmer temperatures are uncontrolled
  - Variation within a single observation as high as 4-5°C
- Scientific impact varies:
  - High: line-rich spectrum, ACIS-I (FI CCD), high S/N
  - Low: continuum spectrum, ACIS-S3 (BI CCD), low S/N
- Mitigation strategies:
  - ✓ T-dependent CTI correction added to standard pipeline in 2010
  - ✓ Mission Planning has constraints on orientation to keep  $T < -114^{\circ}\text{C}$ 
    - Increase nominal FP temperature (degrades performance; requires lengthy recalibration)



# Performance of adjusted corrector



Temperature-dependent pulseheight change (% / deg)

ca. 2009

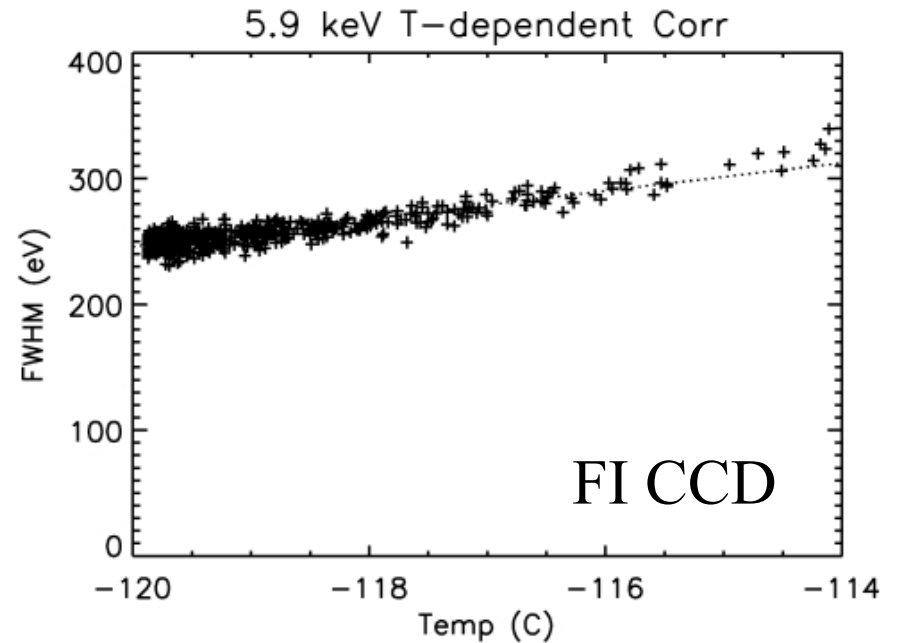
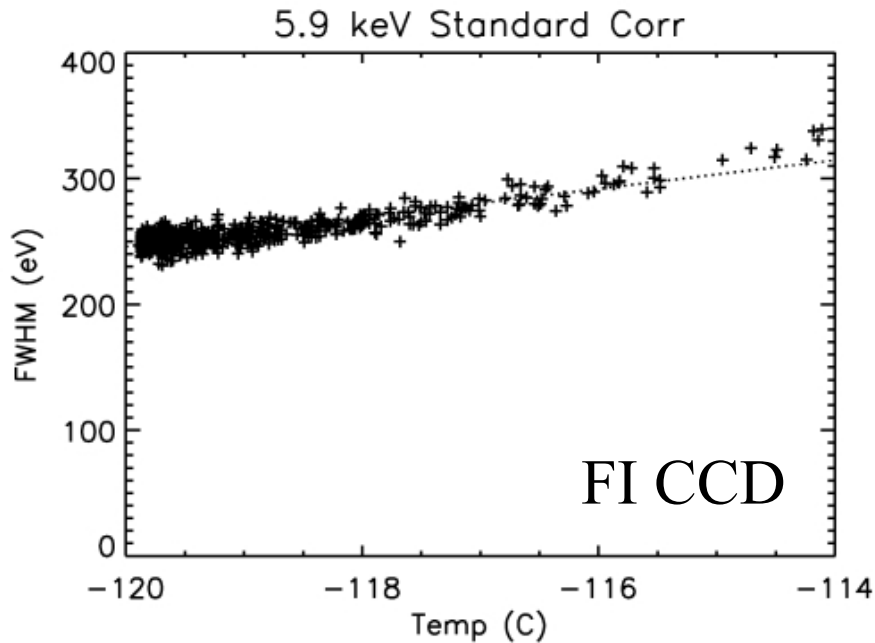
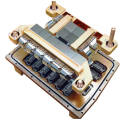
|             | 1.5 keV | 6 keV  |
|-------------|---------|--------|
| Standard    | -0.7%   | -0.4%  |
| T-dependent | +0.03%  | -0.07% |

- Top 64 rows of CCD (worst case)
- Smaller effect at lower rows
- Calibration accuracy goal is 0.3%
- Reduces temp-dependence of pulseheight





# Performance of adjusted corrector



Temperature-dependent FWHM change (eV / deg)

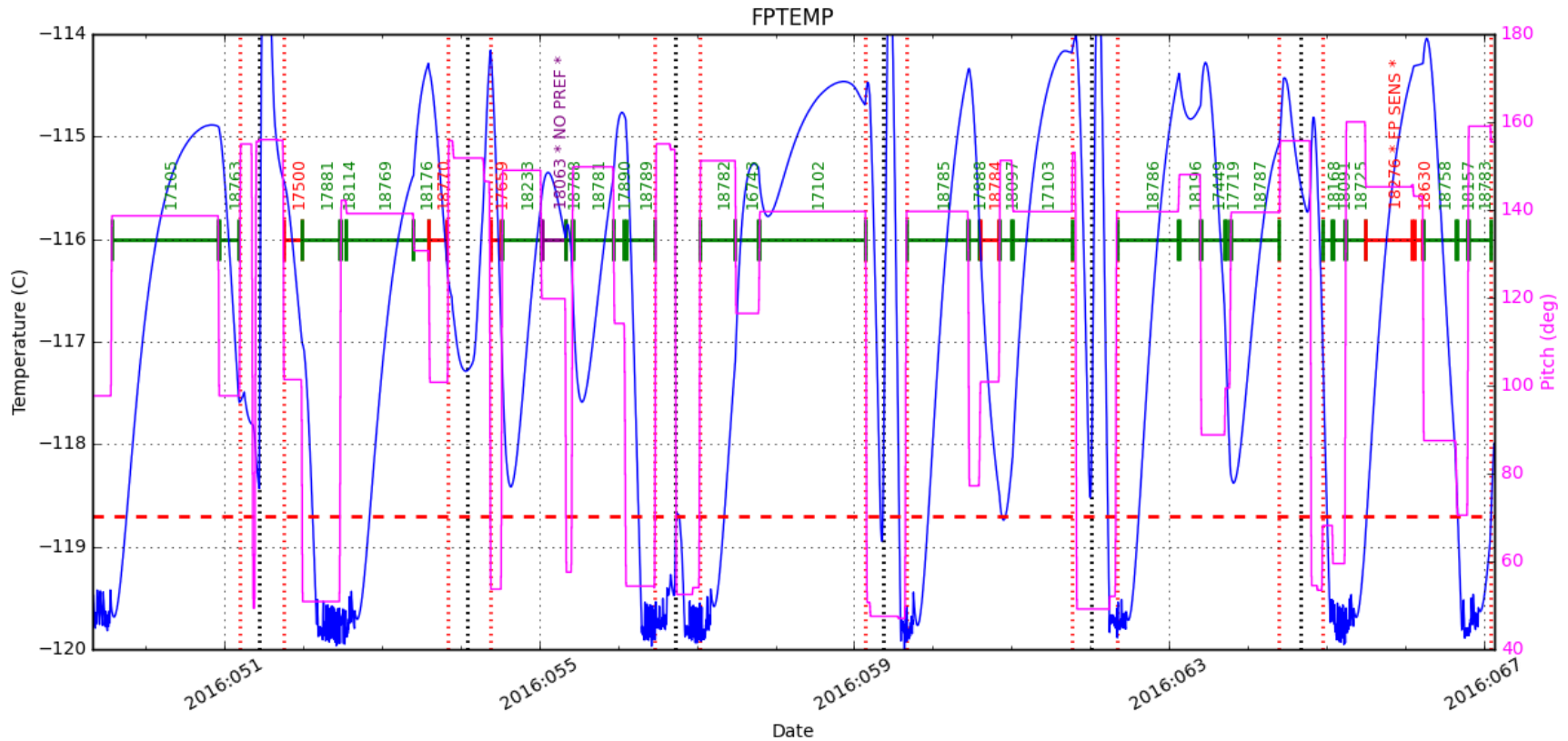
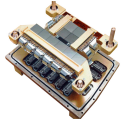
ca. 2009

|             | 1.5 keV | 6 keV    |
|-------------|---------|----------|
| Standard    | +3.8 eV | +11.2 eV |
| T-dependent | +3.2 eV | +10.6 eV |

- Very small reduction in temperature dependence of line width
- Stochastic charge loss – may not be possible to do much better
- FWHM change negligible for ACIS-S3 (BI CCD)



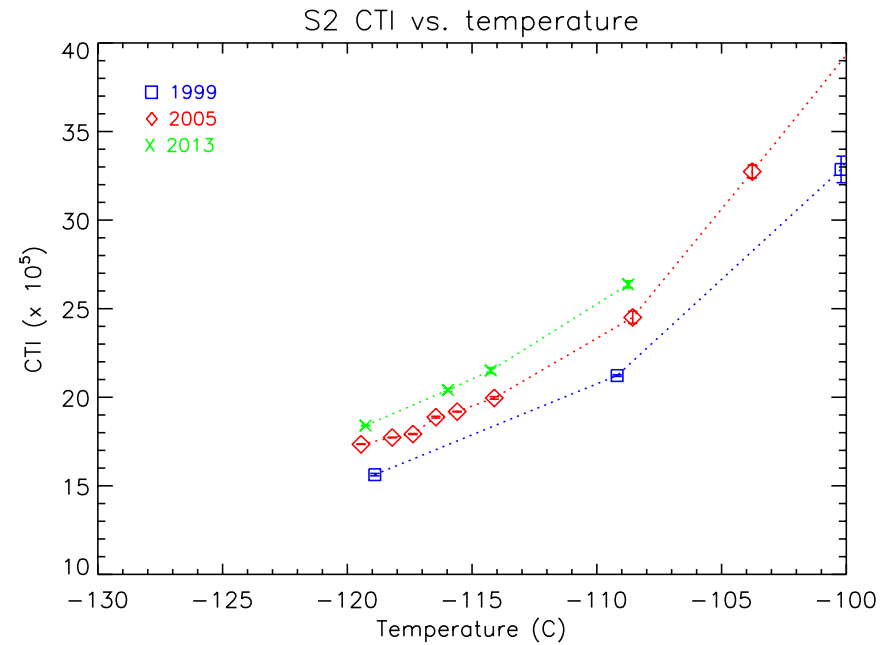
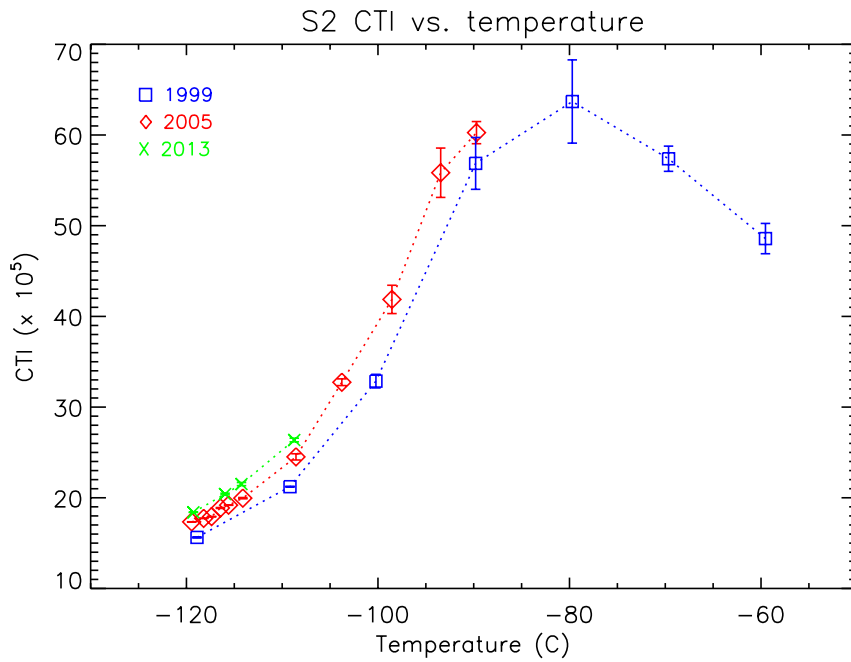
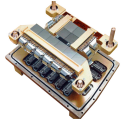
# Current thermal conditions



- Algorithm assumes deviations are small; temperature-dependence is linear
- Original calibration data primarily  $T < -116^{\circ}\text{C}$
- Neither is true at present
- If trap population is changing, T-dependence may be evolving



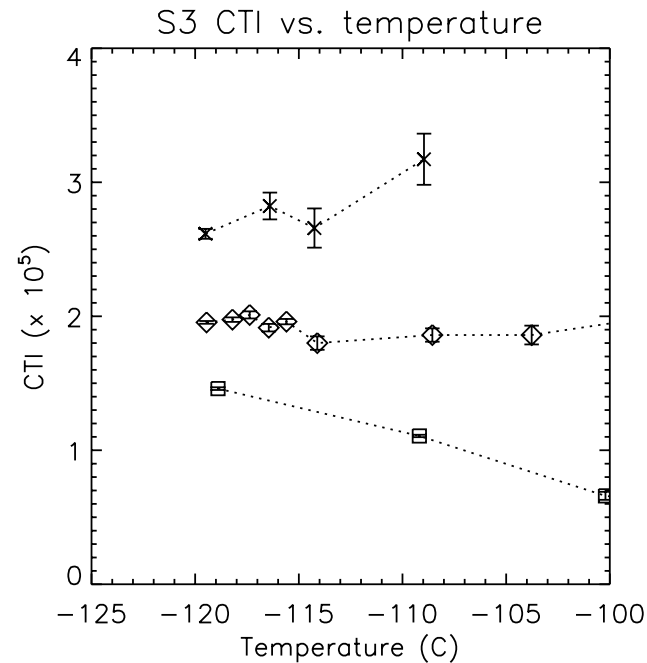
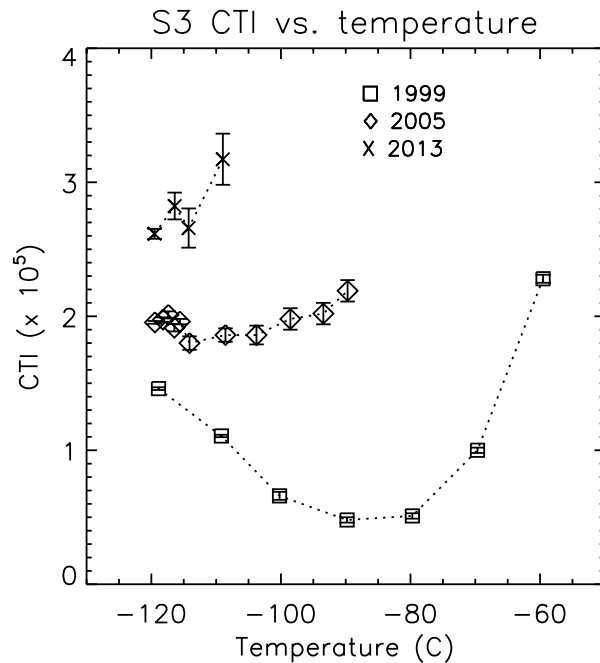
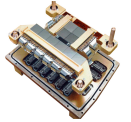
# Is temperature-dependence changing?



- For FI CCDs, change is minimal, but...



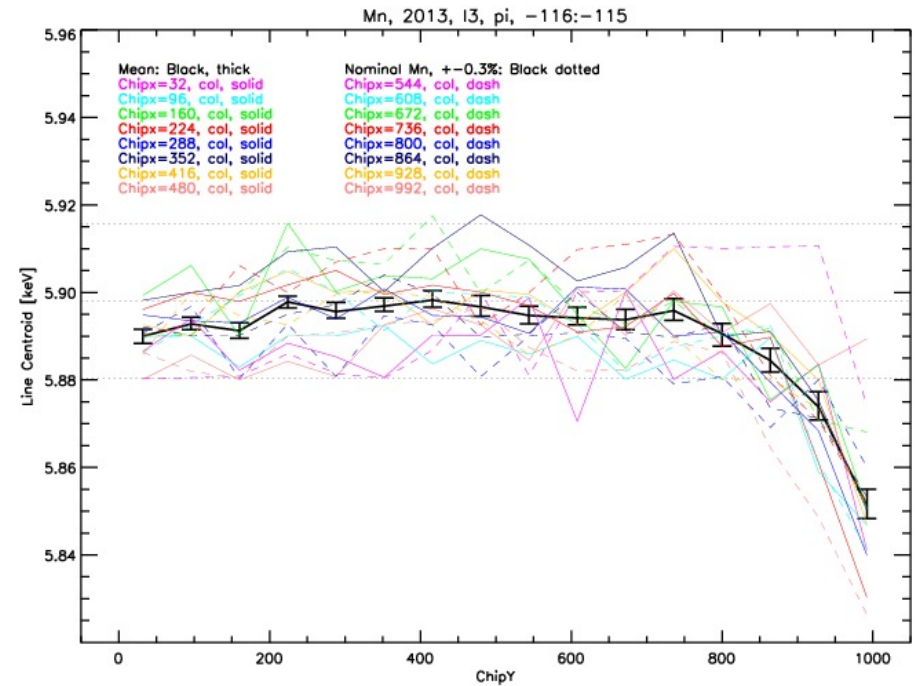
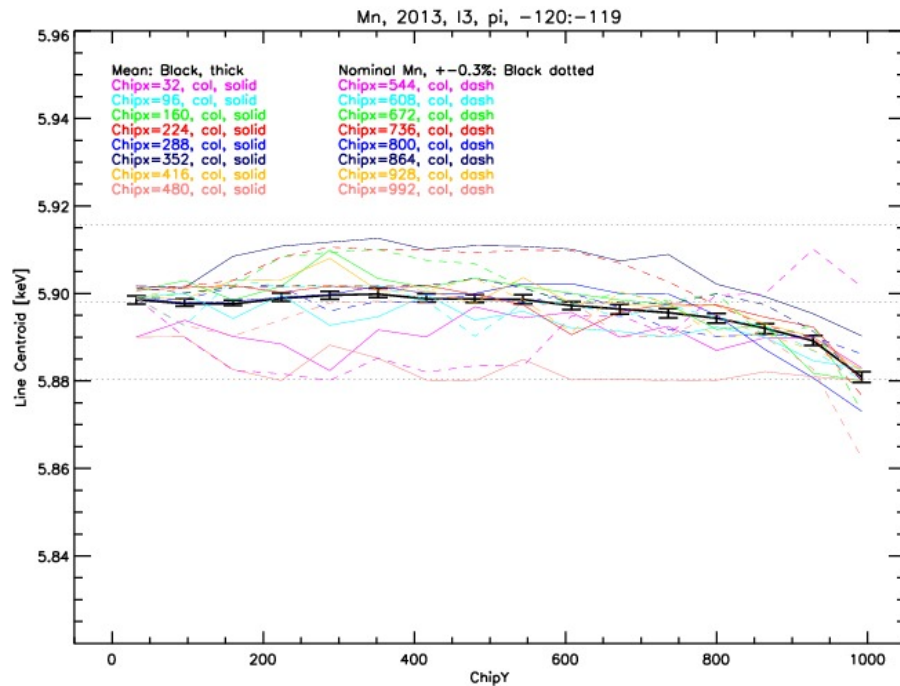
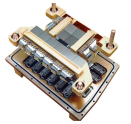
# Is temperature-dependence changing?



- For FI CCDs, change is minimal, but...
- BI CCDs have changed sign of  $dCTI/dT$
- Radiation-produced CTI now as important as initial CTI for BI CCDs
- FI CCDs had little initial CTI at launch ( $< 10^{-6}$ ), subsequent radiation damage similar enough to radiation damage from the belts



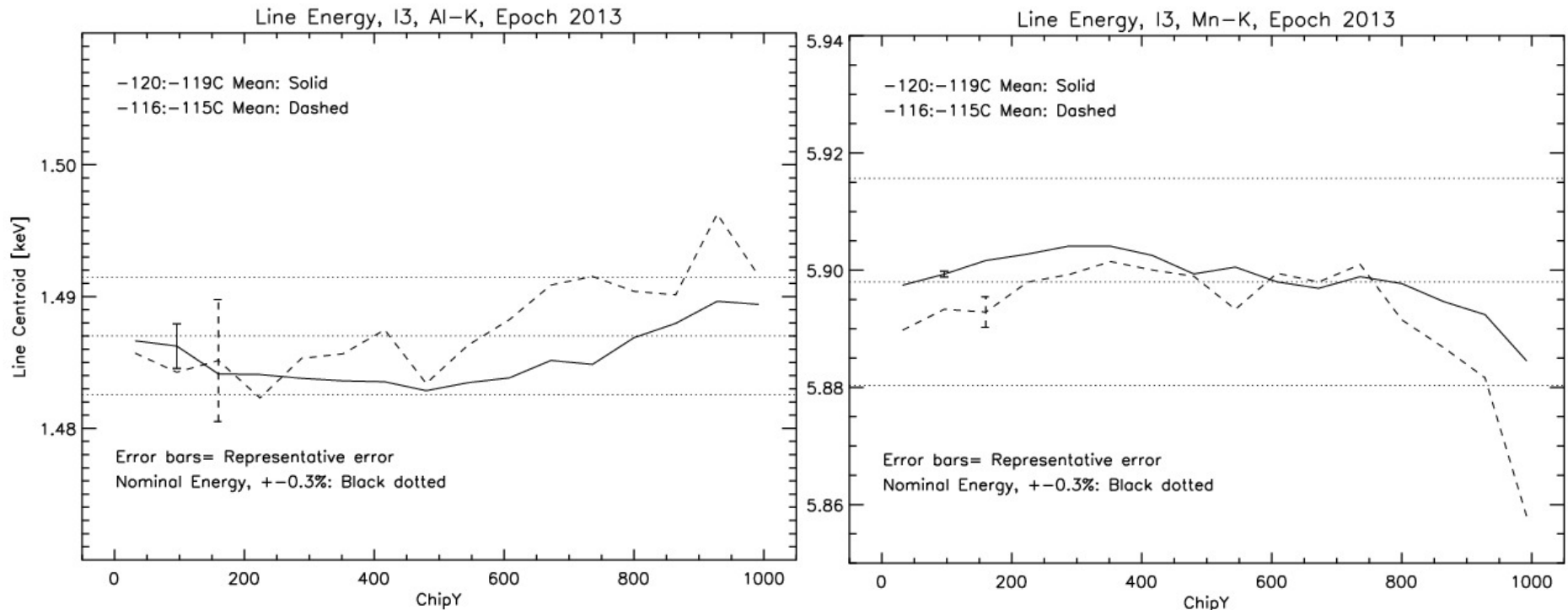
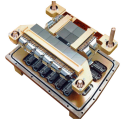
# Testing calibration accuracy



- Data goes through standard pipeline, same as user following CIAO threads
- 64 x 64 pixel regions, one-year time bins, one-degree temperature bins
- Fit Al-K and Mn-K $\alpha$  lines using standard response products
- Line energy should be 5.89 keV; width should be consistent with zero
- Calibration is OK for many regions even at high temp and late times!



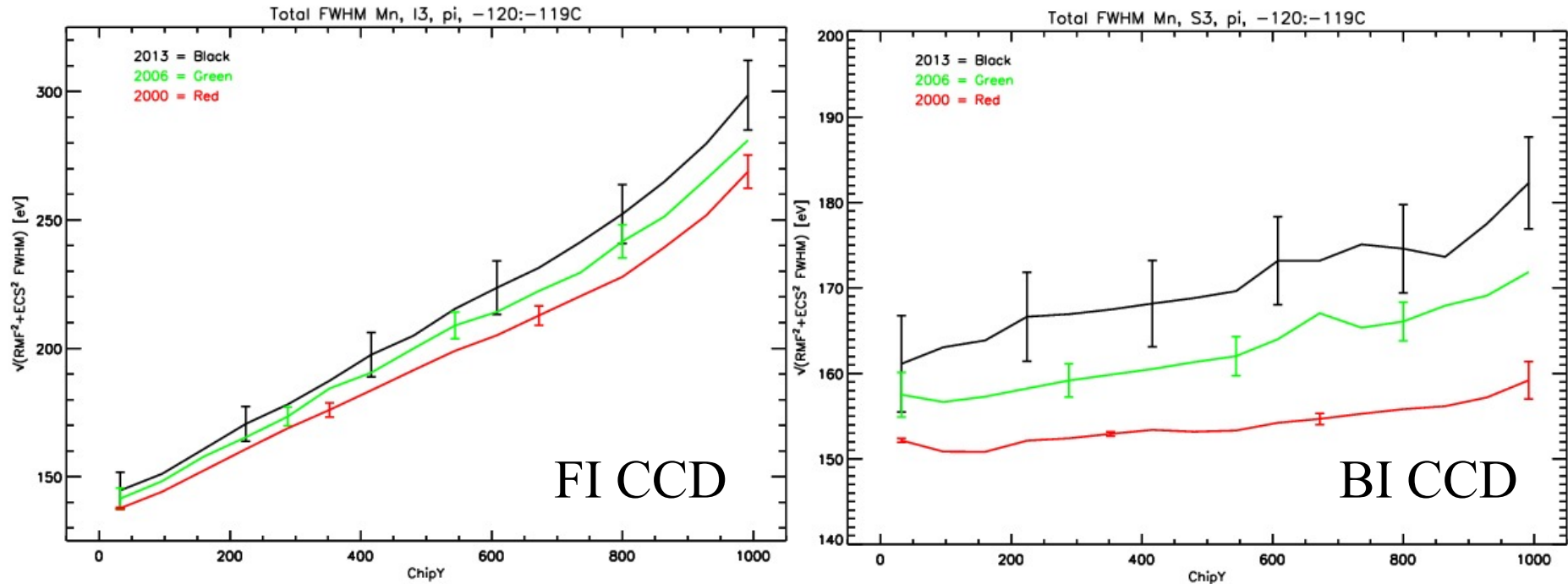
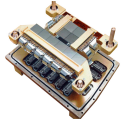
# Calibration Problem Areas: FI CCDs



- Calibration is OK for many regions even at high temp and late times!
- 5.9 keV undercorrects at high ChipY; 1.5 keV overcorrects at high ChipY
- Temperature worsens the problem
- Suggests CTI correction needs adjustment



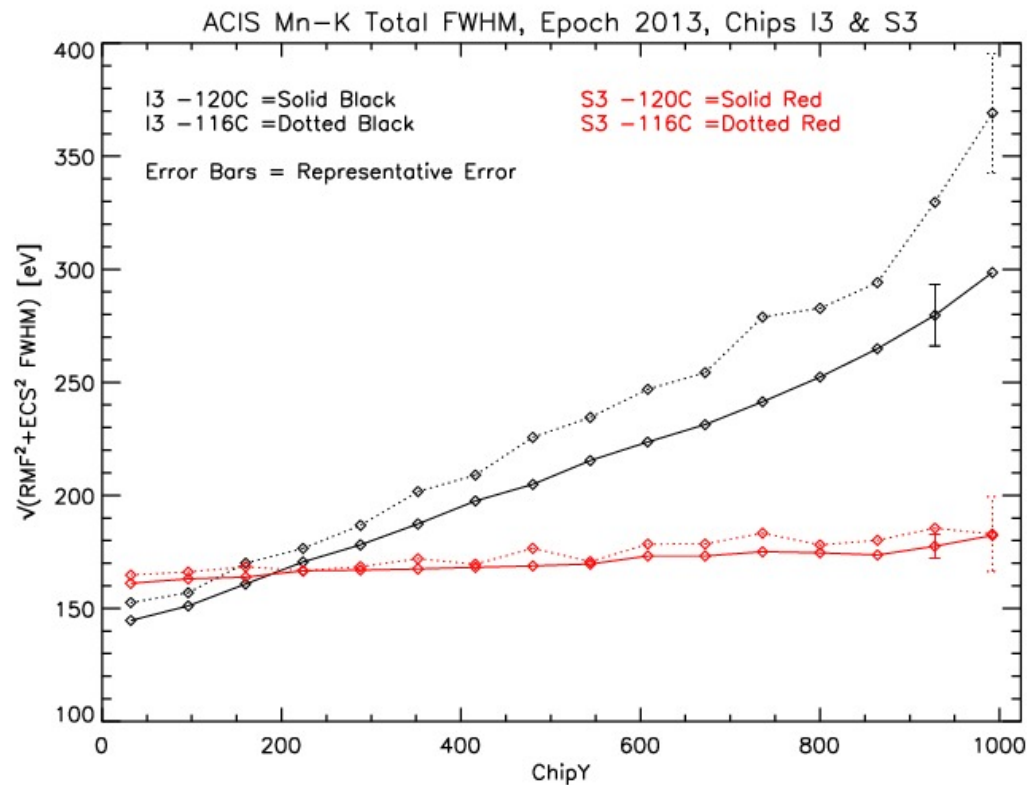
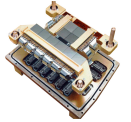
# Calibration Problem Areas: FWHM



- Line width increase with time for all regions/CCD type
- Increase larger at 5.9 keV than 1.5 keV
- Implies a recalibration of CTI correction may be necessary



# Calibration Problem Areas: FWHM



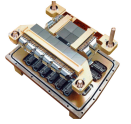
- As already shown, temperature-dependent CTI correction does not correct for additional line width
- Need separate response products for warm temperature data





# Summary

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- ACIS CTI correction last calibrated in 2005
- Temperature-dependence added in 2010
- Continuing radiation damage:
  - $dCTI \sim 2 \times 10^{-6} / \text{yr}$  (FI);  $1 \times 10^{-6} / \text{yr}$  (BI)
- Changing thermal environment
- How's the current calibration doing?
  - Energy calibration not bad except FI CCDs at high rows
  - Line width is harder to improve in the correction
  - Improved response products may be a better fix
- More to come at SPIE this summer