

Chandra ACIS-I3 Response Width

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Introduction

- degradation of thermal blankets \Rightarrow more and more difficult to maintain cold focal plane
- last few years, more observations with some or all of the exposure at elevated temperatures
- study how ACIS response depends on focal plane temperature using ECS data
 - combine ECS data for Epochs 60, 61, 62, 63, 64
 - slice into temperature intervals (currently, 1° C wide):
 - initial study:
 - 120.19 to -119.19
 - 118.19 to -117.19
 - 116.19 to -115.19
 - where sufficient data is available, fit the ECS lines (width parameters scaled and tied so that of Al-K α width)
 - fit Al-K α width vs. $\ln(T_p)$ (per node)

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Chandra Approach to Modeling ACIS Response

ACIS response factored into:

- **p2_resp**
 - ideal (undamaged) CCD response matrix (basic CCD physics: escape peak, shoulders, etc.)
 - “scatter matrix” (response broadening due to CTI)
- **det_gain**
 - $\text{p2_resp} \times \text{det_gain} \rightarrow \text{Energy}$
 - $\text{p2_resp} \times \text{Energy} \rightarrow \text{p2_resp}$
- **p2_resp and det_gain tightly coupled:**
 - $\text{p2_resp} + \text{det_gain} \rightarrow \text{Energy}$
 - $\text{p2_resp} \times \text{Energy} \rightarrow \text{det_gain}$
- **p2_resp and det_gain calibrated using (Epoch 1) External Calibration Source (ECS) spectra**

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- **det_gain**
 - pha of response peak \Rightarrow Energy
 - pha \Rightarrow Energy \Rightarrow pi
- p2_resp and det_gain tightly coupled:
 - $p2_resp + det_gain = \text{constant}$
 - $pha \text{ peak vs. } det_gain \Rightarrow \text{energy} = det_gain$
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 - p2_resp + det_gain \Rightarrow rmf
 - rmf **peak vs** chipx, chipy, energy \Rightarrow det_gain
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FI CCD CTI Scatter matrix:

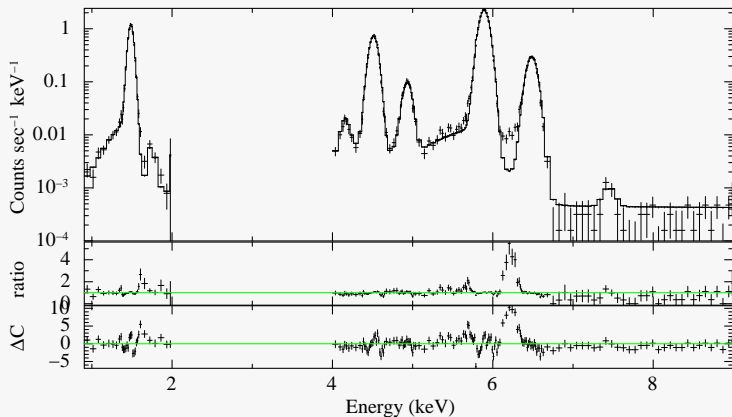
- 256×32 tiles (chip coords):
- PHACHAN(pha channel [ADU])
 - 40, 60, 80, 100, 128, 180, 230, 300, 380, 460, 512, 670, 1025, 1535, 2050, 3100
- L1_WIDTH [ADU] for each (PHACHAN)

Introduction

External Calibration Source (ECS) spectra

- $K\alpha$, $K\beta$ lines for Al, Ti, and Mn, covering $\sim 1.5 - 6$ keV.

e001: 119-120: c3n3_00



Introduction

- FI chip scatter matrix: broadening by an asymmetric Lorentz function

$$R_{scat}(E) = 1 / \left[1 + [(E - E_0) / \Delta(E_0)]^2 \right]^\alpha$$

where

$$\alpha = \begin{cases} 3.70 & \text{for } E < E_0 \\ 1.90 & \text{for } E > E_0 \end{cases}$$

ECS spectra using

- asymmetric Lorentz function broadening profile (Xspec local model)
- RMF **without** CTI broadening

Fitting ECS data

Fit main ECS lines (plus $\Gamma = 0.2$ powerlaw for background):

AlK $_{\alpha,\beta}$, TiK $_{\alpha,\beta}$, MnK $_{\alpha,\beta}$ (+ NiK $_{\alpha}$)

Width function:

$$\Delta \propto \left(E_0^{0.12} + 0.3 E_0 \right) \quad (1)$$

Allow $\Delta_{\text{Al-K}\alpha}$ to vary; tie the rest to $\Delta_{\text{Al-K}\alpha}$ using Eq. (1):

$$\begin{array}{cccccc} \Delta_{\text{Al-K}\alpha} & : & \Delta_{\text{Ti-K}\alpha} & : & \Delta_{\text{Ti-K}\beta} & : & \Delta_{\text{Mn-K}\alpha} & : & \Delta_{\text{Mn-K}\beta} \\ 1 & : & 1.7066 & : & 1.7995 & : & 2.0312 & : & 2.1397 \end{array}$$

K $_{\beta}$ line energies tied to K $_{\alpha}$ line energies (multiplicative factor); can probably be relaxed for high quality data.

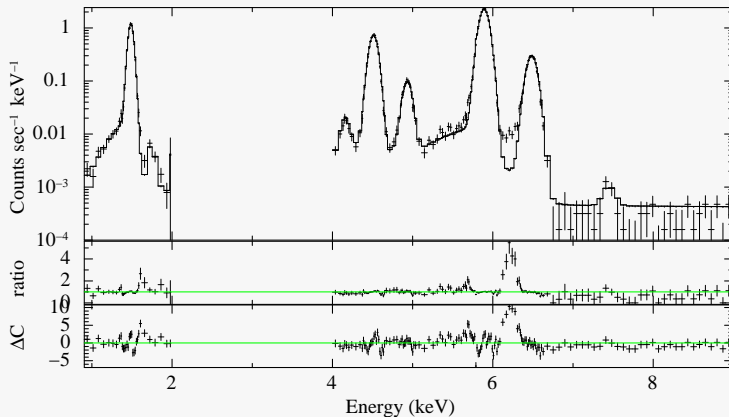
Line norm: Al-K $_{\beta}$ = 0.01 Al-K $_{\alpha}$

Some Representative Fits

Epoch 1, -120.19 to -119.19

I3, node 3, 32 rows adjacent to framestore

e001: 119-120: c3n3_00

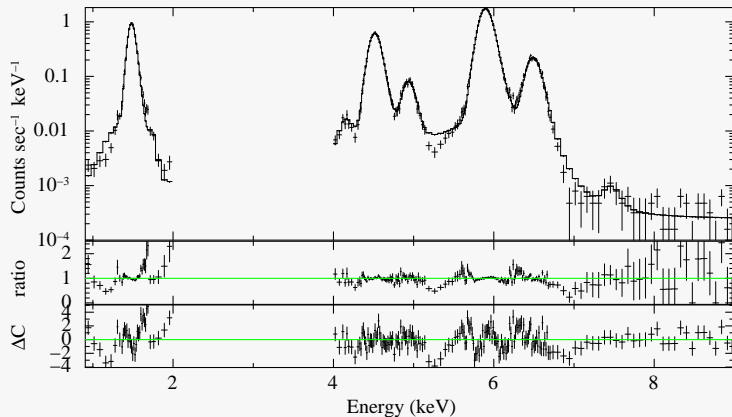


Some Representative Fits

Epoch 1, -120.19 to -119.19

I3, node 3, 32 rows, middle of the chip

e001: 119-120: c3n3_15

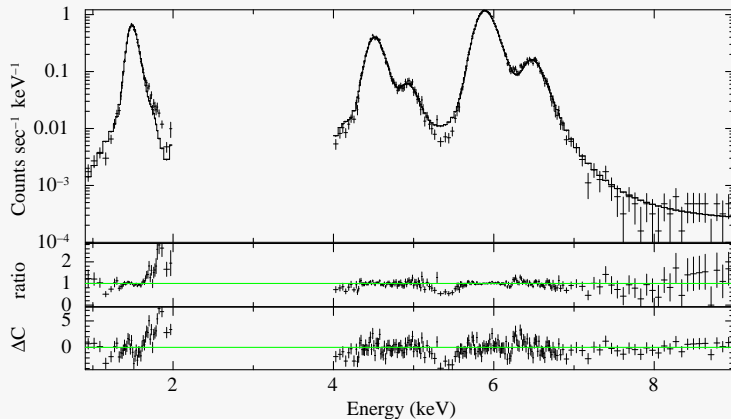


Some Representative Fits

Epoch 1, -120.19 to -119.19

I3, node 3, 32 rows far edge of the chip

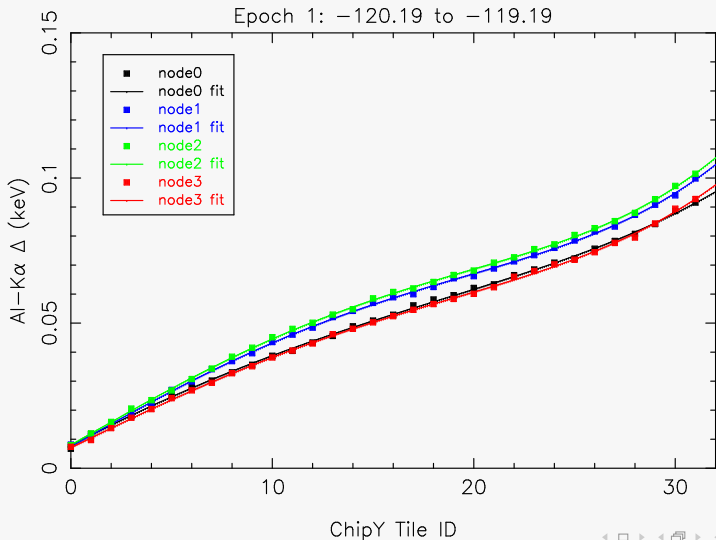
e001: 119-120: c3n3_31



Fit Width vs. `chipy` (4th degree Chebyshev)

Epoch 1, I3: -119.19C to -120.19C

Al-K α width vs. `chipy`



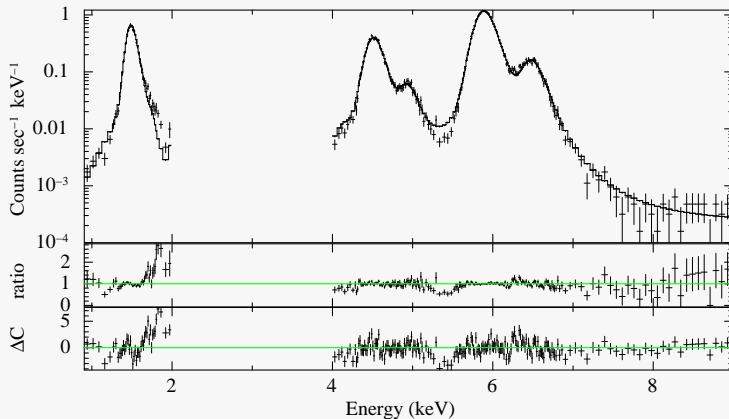
Variation With Focal Plane Temperature

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Variation With Focal Plane Temperature

I3 node 3: 32 rows adjacent to edge opposite to framestore
Epoch 1, -120.19 to -119.19

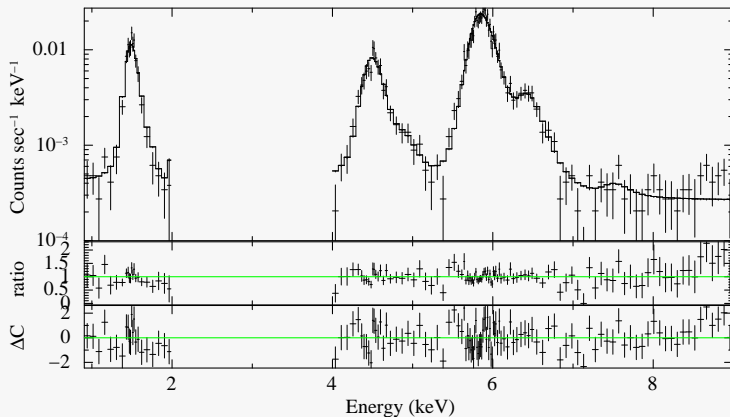
e001: 119-120: c3n3_31



Variation With Focal Plane Temperature

I3 node 3: 32 rows adjacent to edge opposite to framestore
Epoch 60 to 64, -118.19 to -117.19

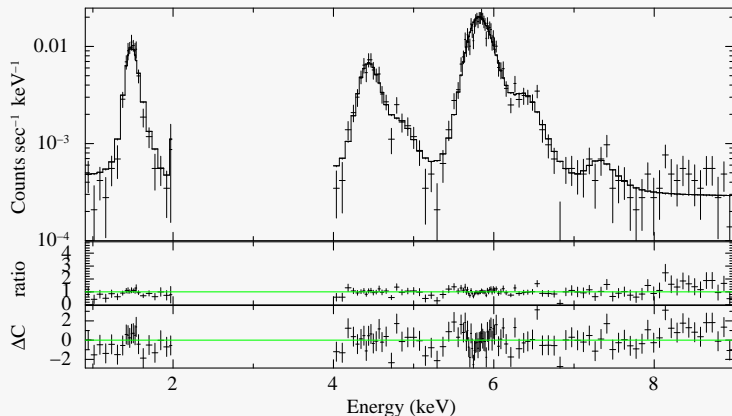
e060_061_062_063_064: 117-118: c3n3_31



Variation With Focal Plane Temperature

I3 node 3: 32 rows adjacent to edge opposite to framestore
Epoch 60 to 64, -116.19 to -115.19

e060_061_062_063_064: 115-116: c3n3_31

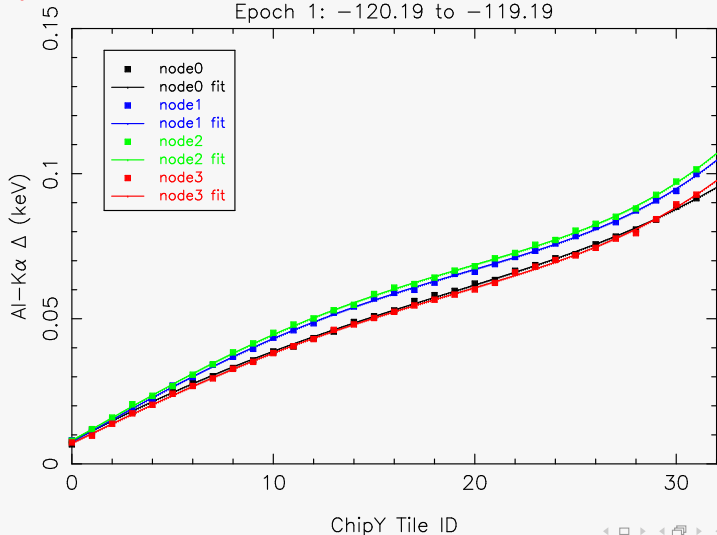


Al-K α Response Width

Variation vs. Focal Plane Temperature

Epoch 1, I3: -119.19C to -120.19C

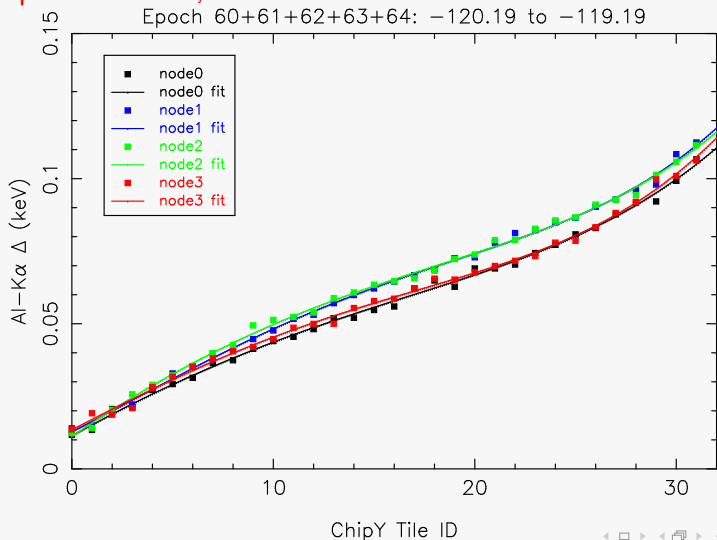
Epoch 1: -120.19 to -119.19



Al-K α Response Width

Variation vs. Focal Plane Temperature

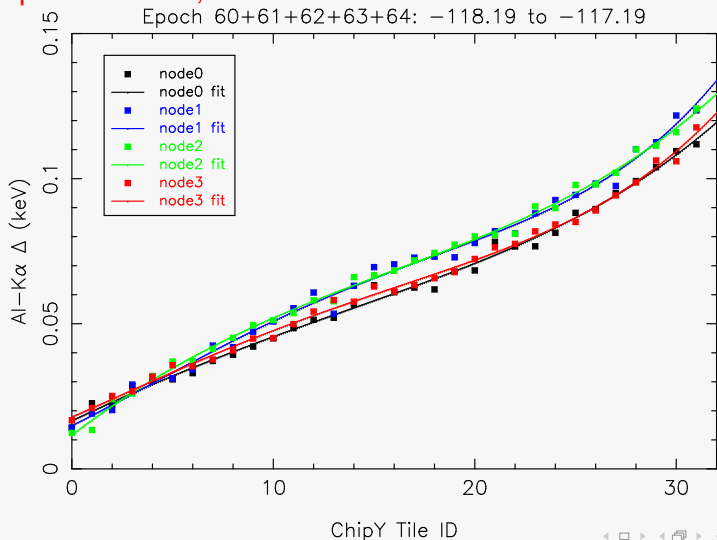
Epoch 60 to 64, I3: -120.19C to -119.19C



Al-K α Response Width

Variation vs. Focal Plane Temperature

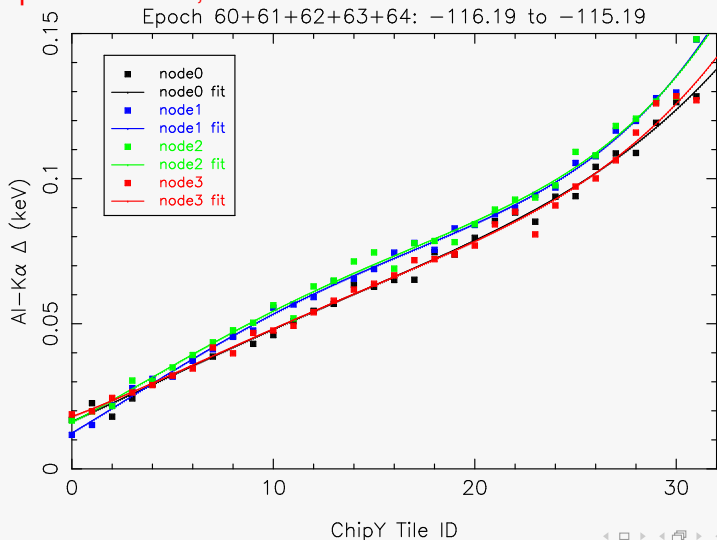
Epoch 60 to 64, I3: -118.19C to -117.19C



Al-K α Response Width

Variation vs. Focal Plane Temperature

Epoch 60 to 64, I3: -116.19C to -115.19C



Steps to CALDB p2_resp (FI)

- Scatter matrix tabulates width Δ in ADU vs. PHA
- ECS fits give $\Delta(E)$ at energy of Al-K α

So,

For each node and tile:

- $\Delta(\text{Al-K}\alpha) \Rightarrow \Delta(E)$ using width function, $\Delta \propto (E_0^{0.12} + 0.3 E_0)$
- $E_{\text{chan}} \Rightarrow \text{phachan}$ using `det_gain`
- map $\Delta(E_{\text{chan}}) \Rightarrow \Delta(\text{ADU})$ using `det_gain`

Inject new `L1_WIDTH` into appropriate `p2_resp` rows.

Summary

- Investigated response width vs:
 - node
 - `chipy`
 - focal plane temperature
- Preliminary results
 - Modest dependence on time?
 - Stronger dependence on focal plane temperature
- To be done:
 - examine intermediate times (say, every 5 years)
 - examine other temperatures
 - look at finer `chipx` gridding for `p2_resp` (with finer gridding in the matching `det_gain`)
 - construct modified `p2_resp` (and `det_gain`?)