



JEREMY J. DRAKE AND THE CXC CALIBRATION GROUP

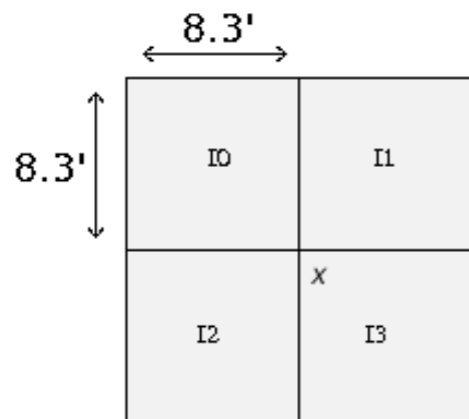
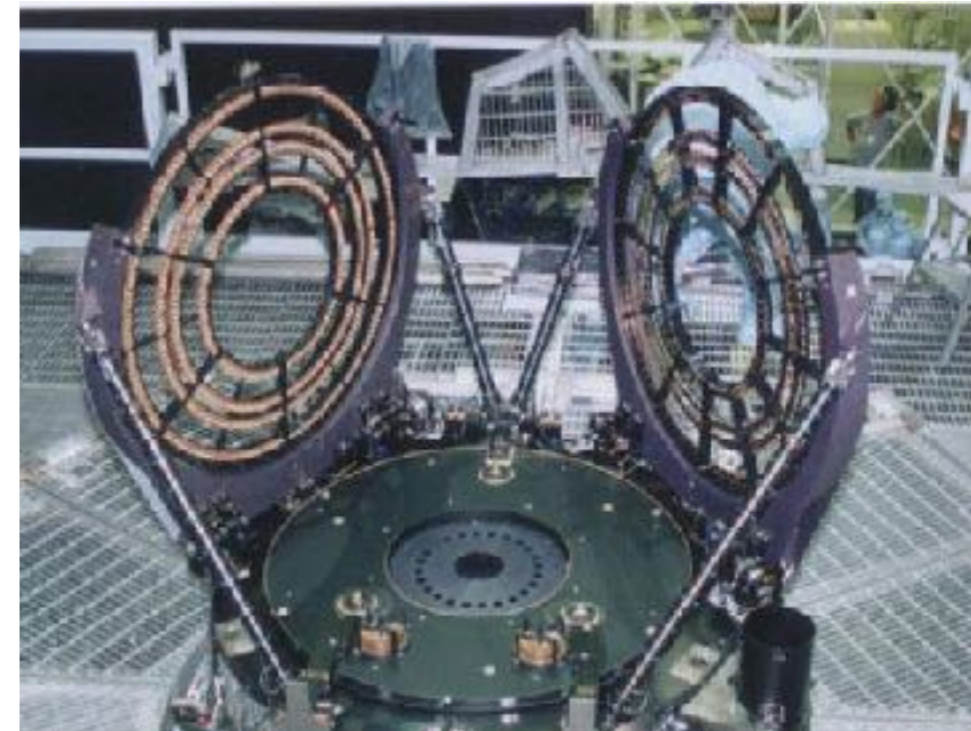
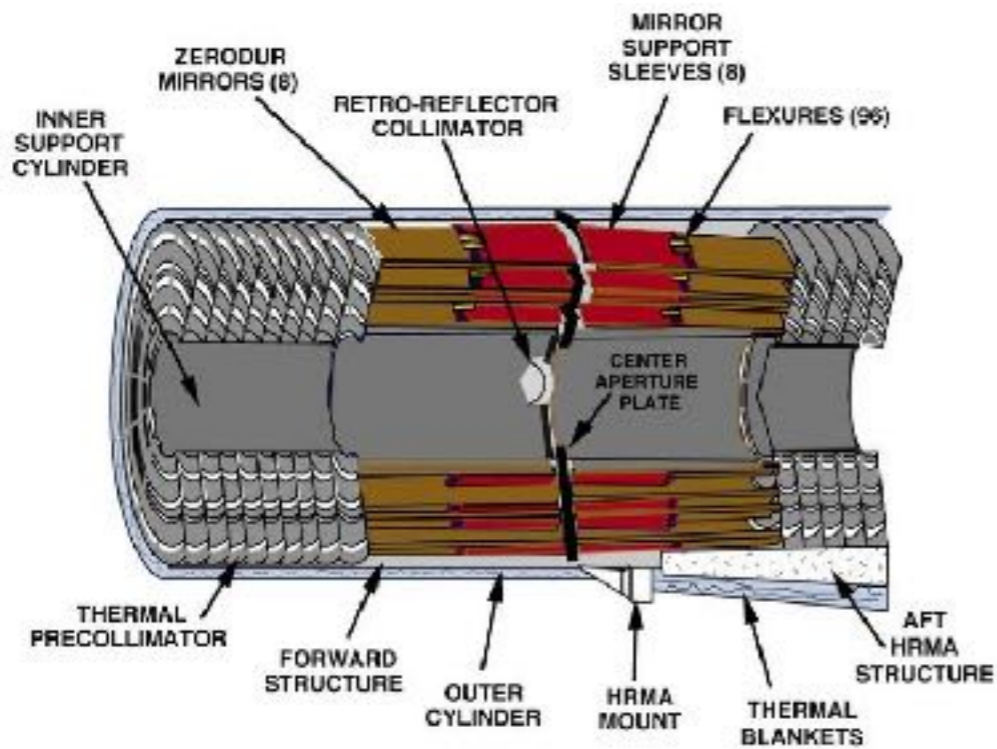
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# CHANDRA CALIBRATION STATUS

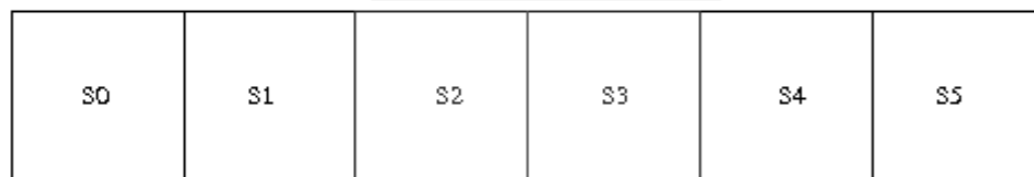


iachee 2018

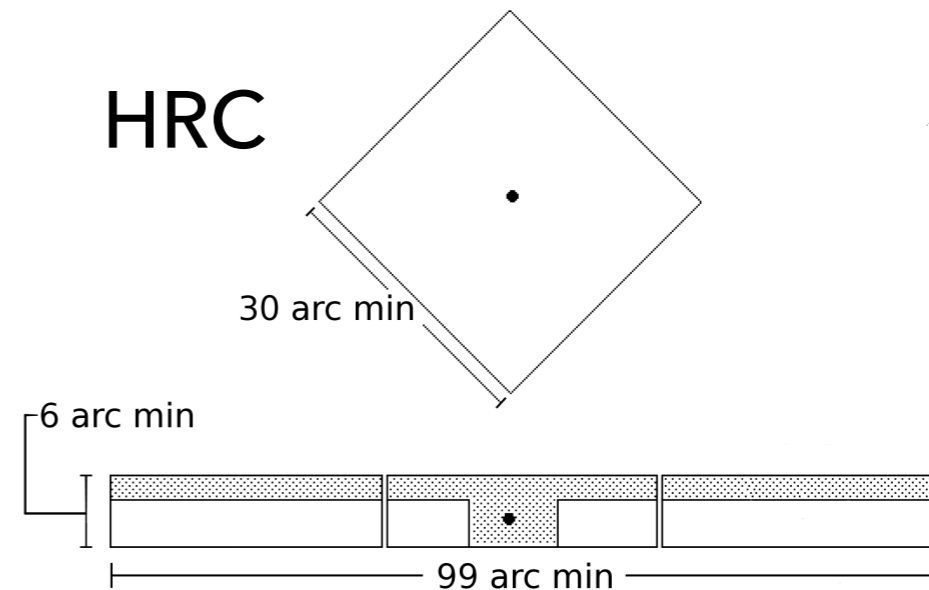
# CHANDRA HARDWARE COMPONENTS



ACIS



HRC



# OUTLINE

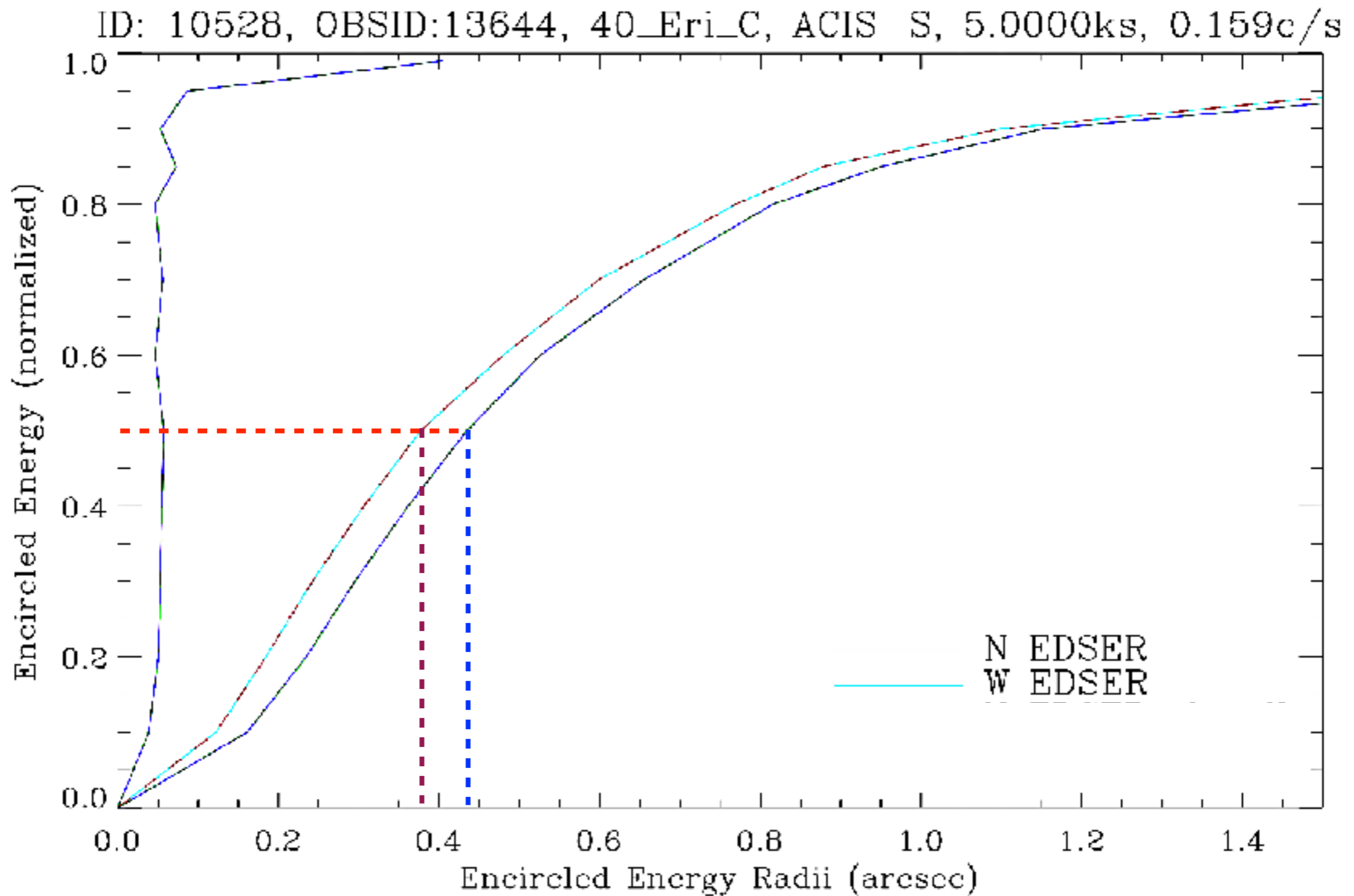
- ▶ Point Spread Function
  - ▶ calibrating EDSER; empirical PSFs
- ▶ ACIS
  - ▶ mid-chip gain droop; QEU; contamination
- ▶ HRC-S,I
  - ▶ QE decline; gain decline

# POINT SPREAD FUNCTION

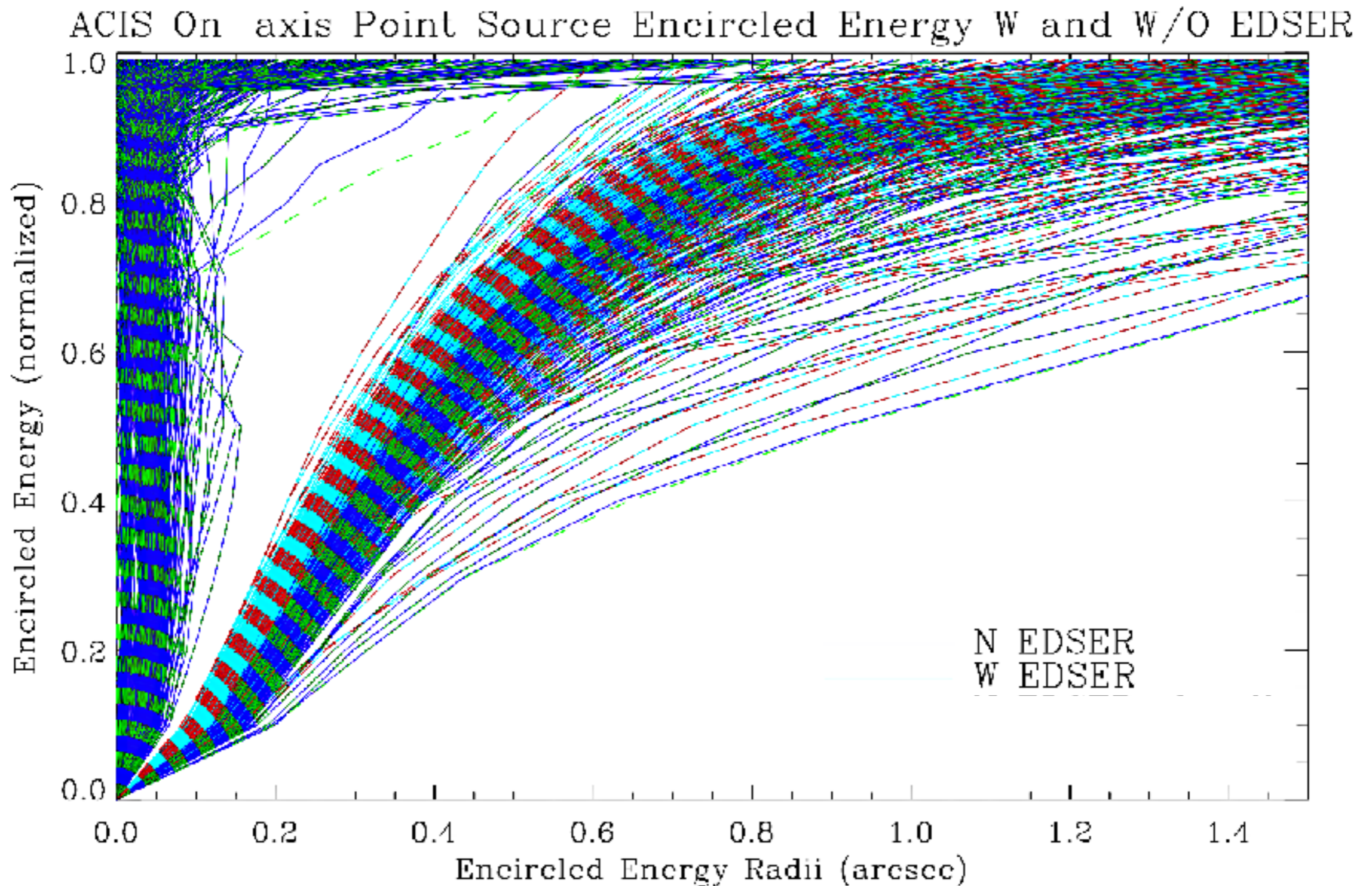
### ACIS PSF WITH EDSER (V. KASHYAP, P. ZHAO, D. JERIUS)

- ▶ Energy Dependent Subpixel Event Repositioning (EDSER)
  - ACIS images can be sharpened significantly at sub-pixel resolutions
- ▶ Applies corrections to event locations based on photon energy and grade (Li et al. 2004, ApJ 610, 1204)
- ▶ BUT: EDSER'd PSFs have not yet been calibrated

# QUANTIFYING EDSER (V. KASHYAP, P. ZHAO, D. JERIUS)

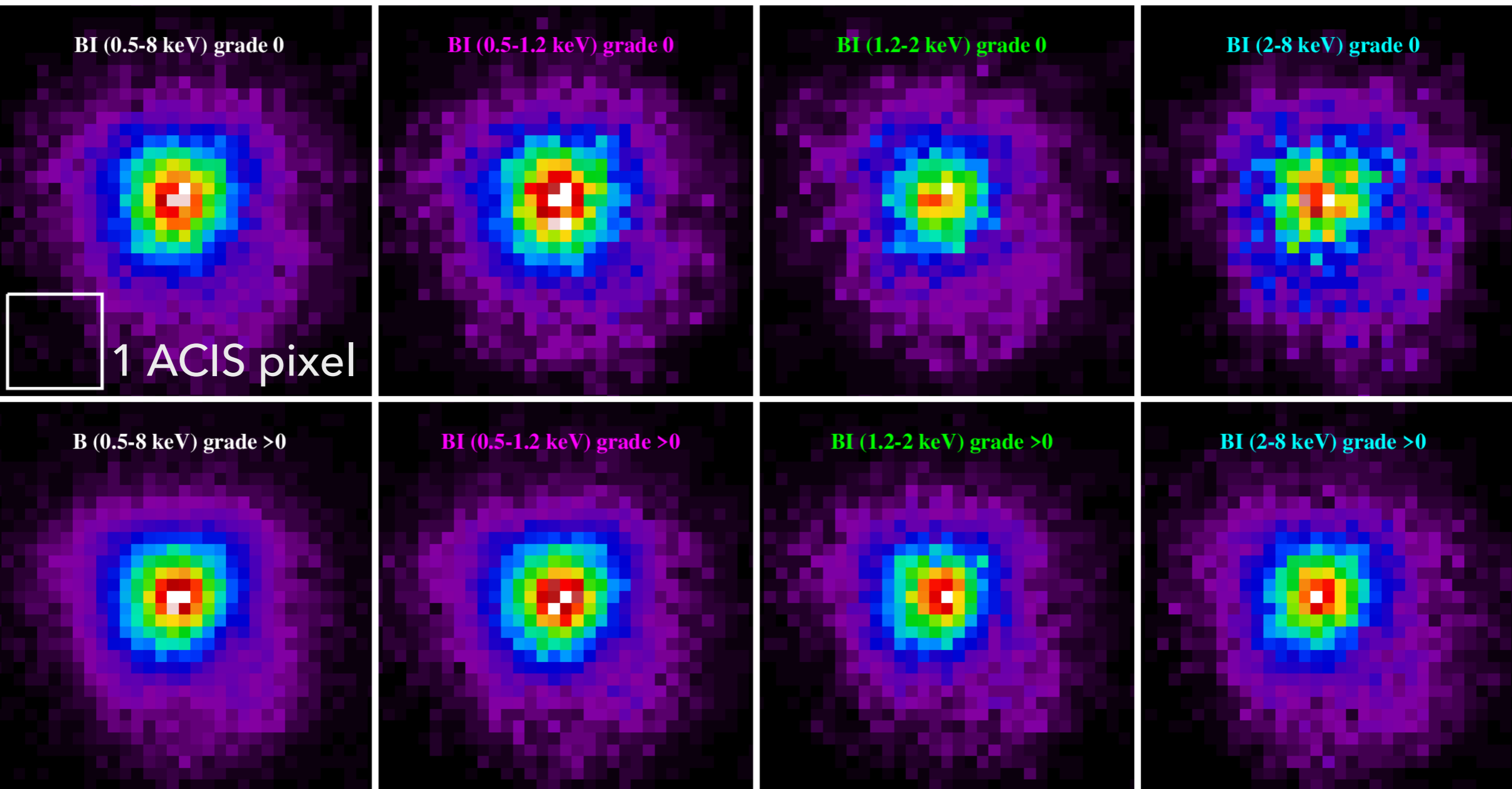


# QUANTIFYING EDSER (V. KASHYAP, P. ZHAO, D. JERIUS)

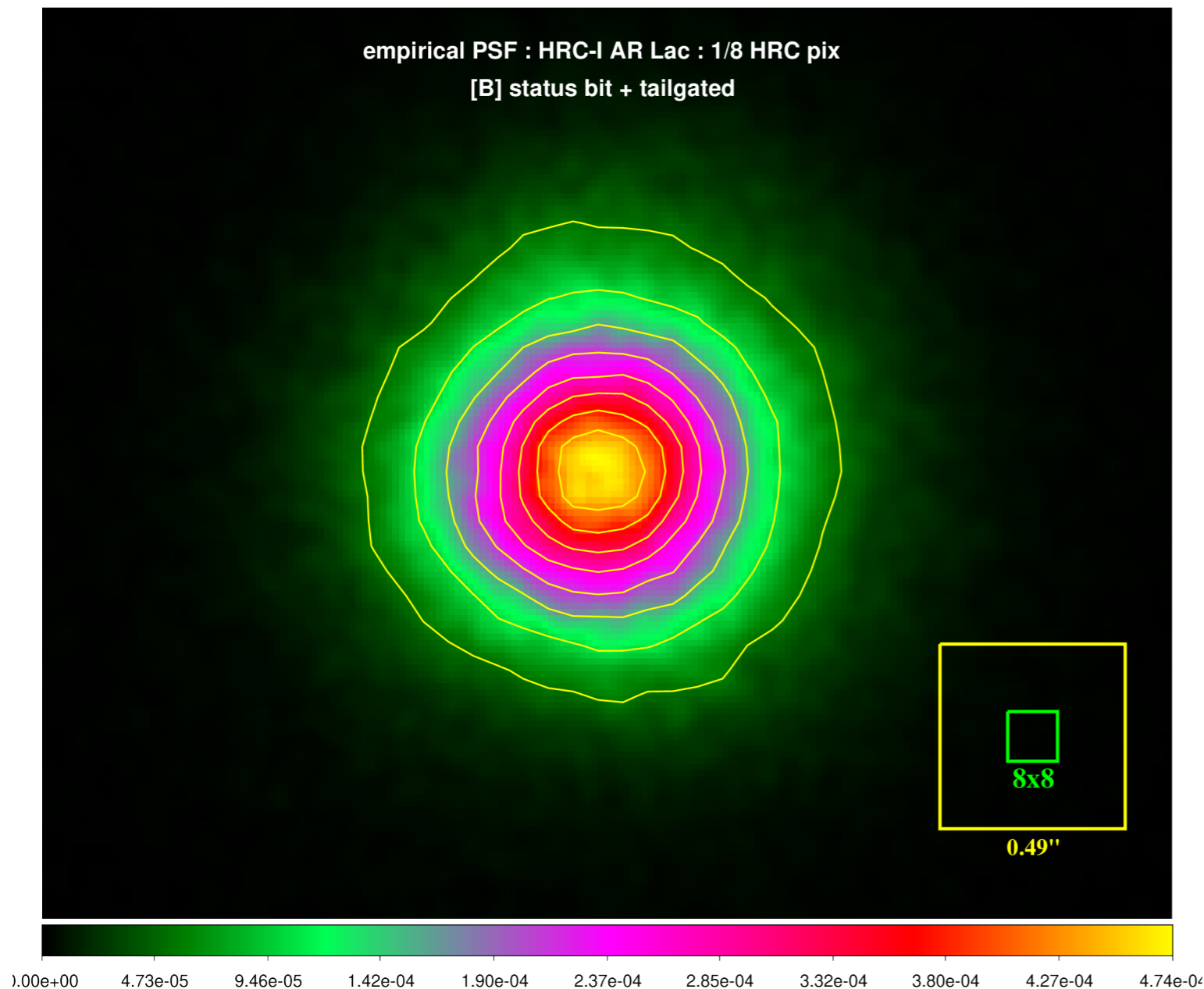




# EMPIRICAL PSF WITH EDSER (V. KASHYAP, P. ZHAO, D. JERIUS)



## HRC-I EMPIRICAL PSF (V. KASHYAP)

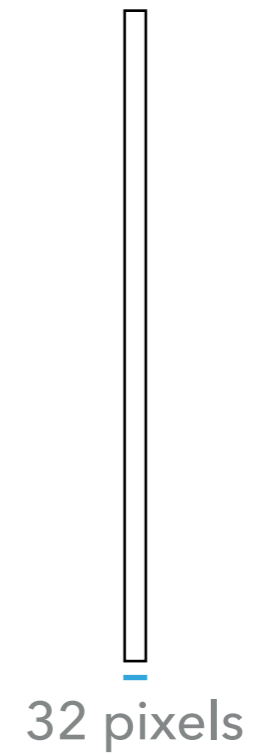
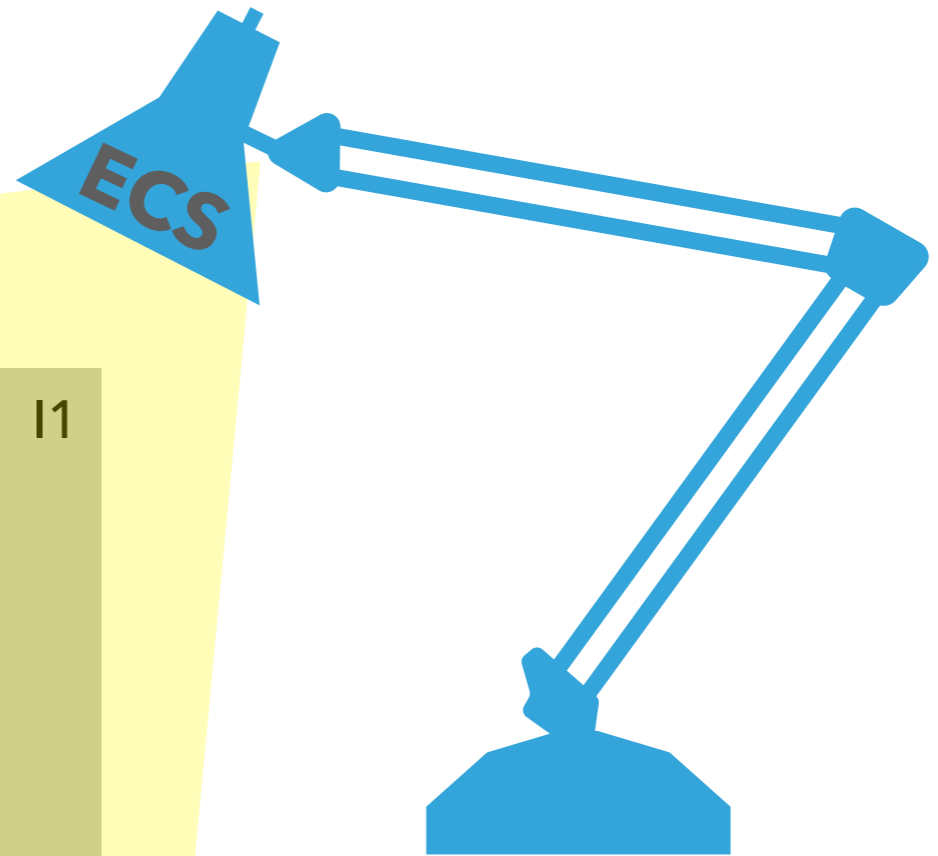
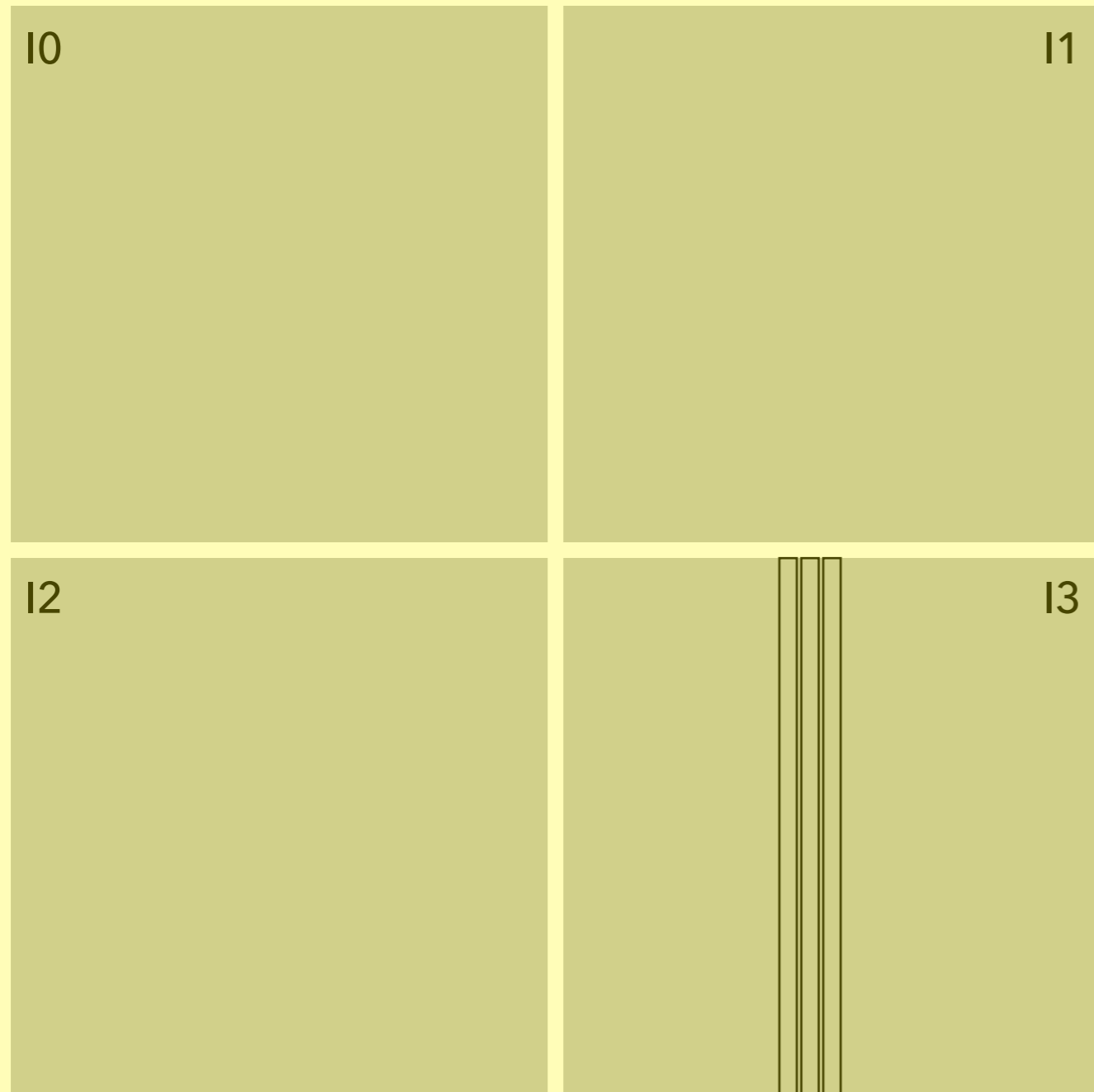


On-axis HRC-I AR Lac data, filtered on status-bits and tailgated events, derolled, and stacked (contours at 10% intervals from peak)

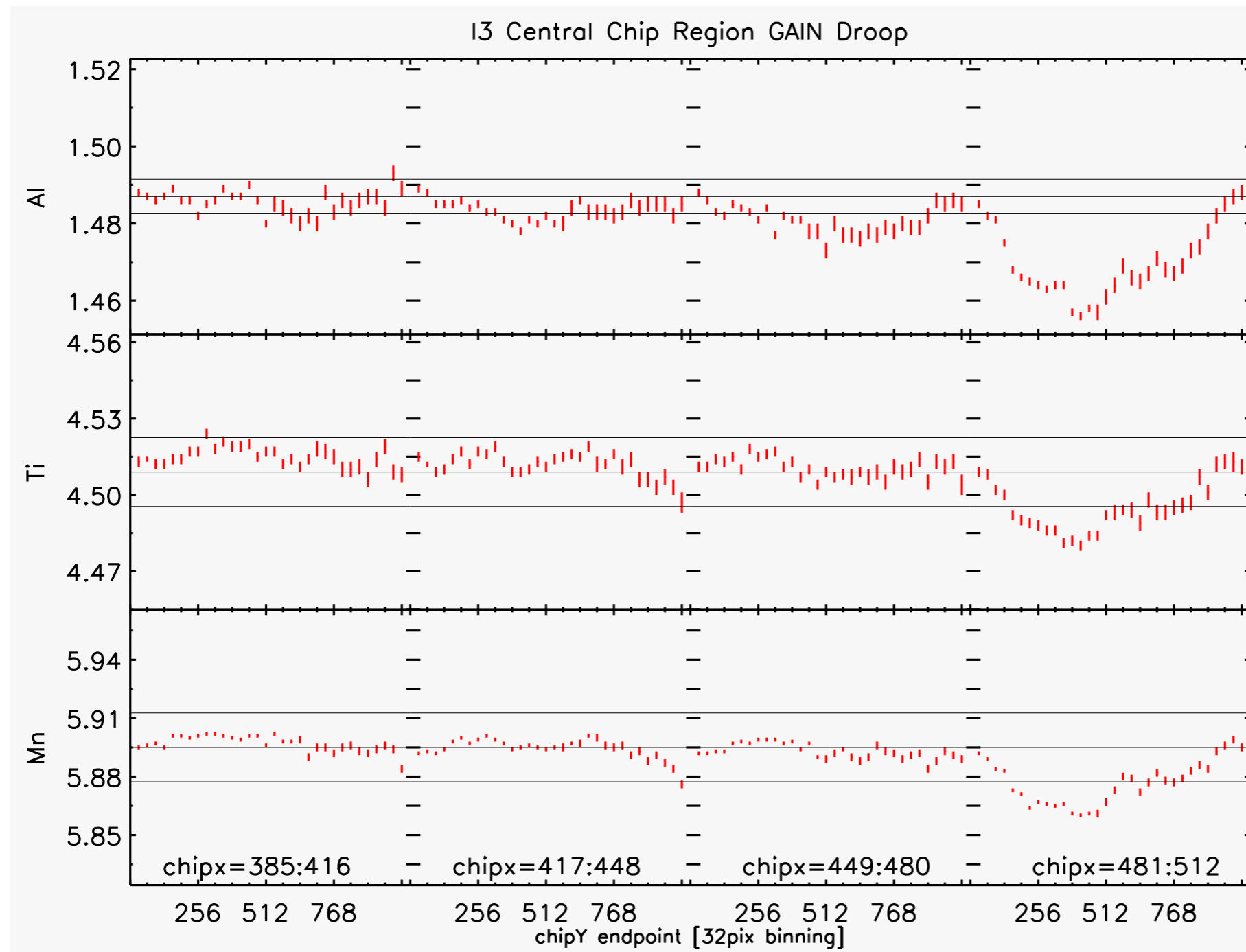
- ▶ HRC-I is the best for high spatial resolution analysis
- ▶ *pros*: no pixels, no pileup, low(ish) background
- ▶ *cons*: colorblind, tailgated events broaden PSF

# ADVANCED CCD IMAGING SPECTROMETER (ACIS)

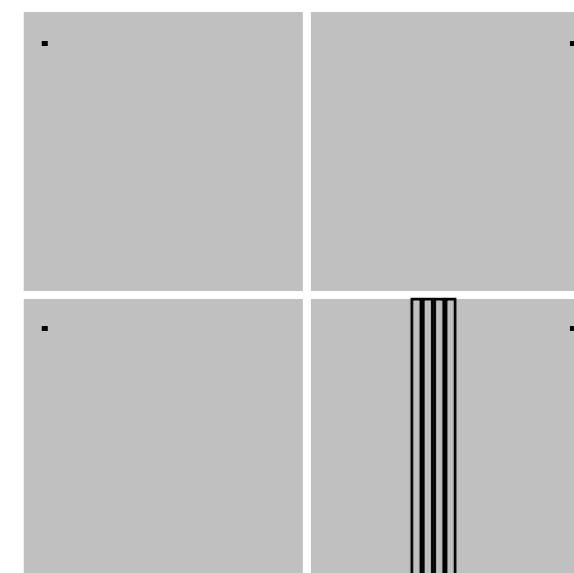
# MID-CHIP GAIN DROOP (T. GAETZ)



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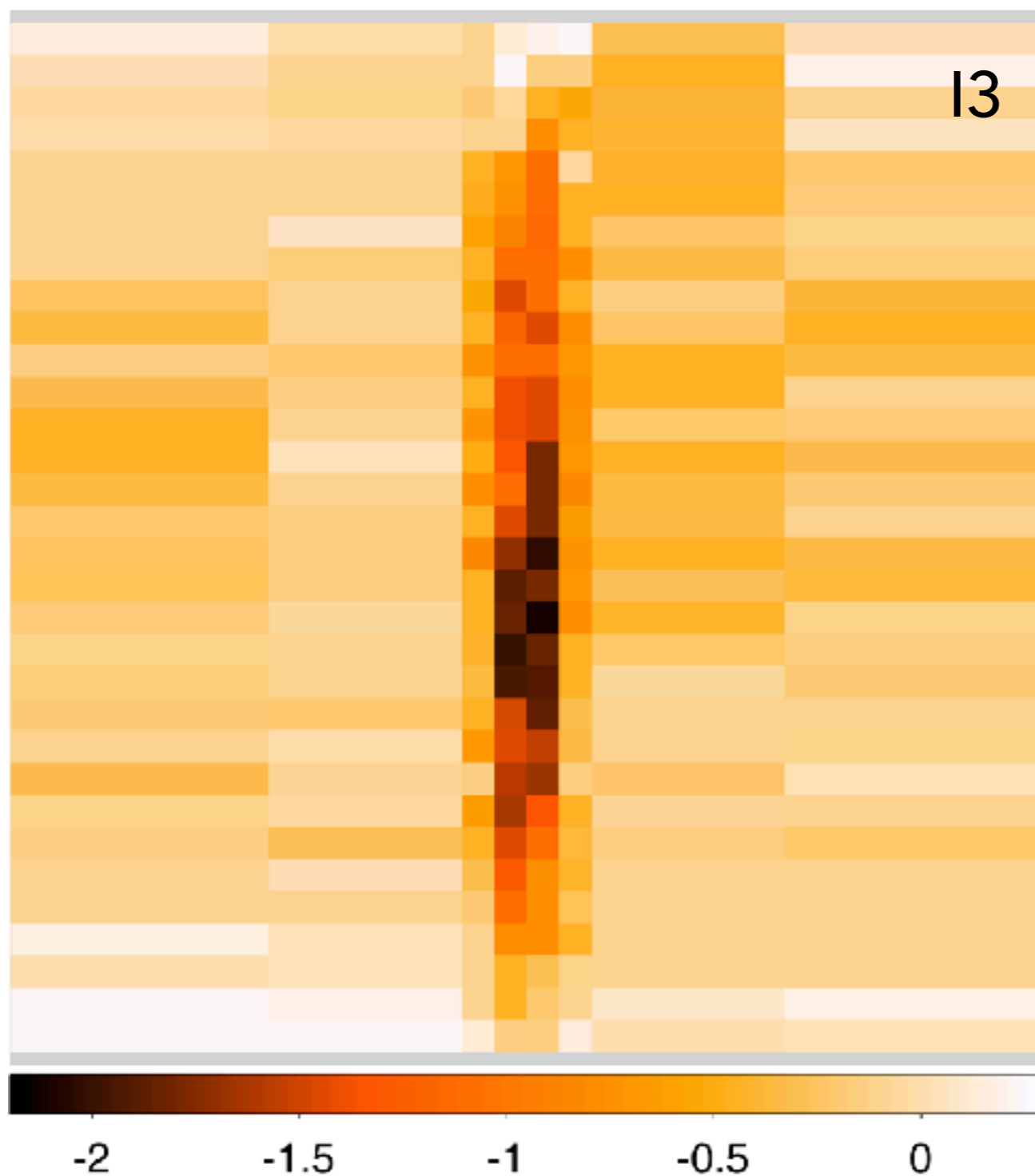


Epoch 1, -120.19  
to -119.19 C

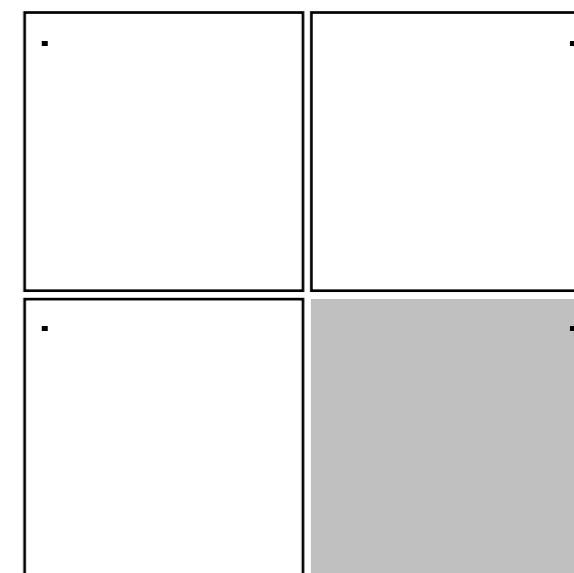


# MID-CHIP GAIN DROOP (T. GAETZ)

Al K



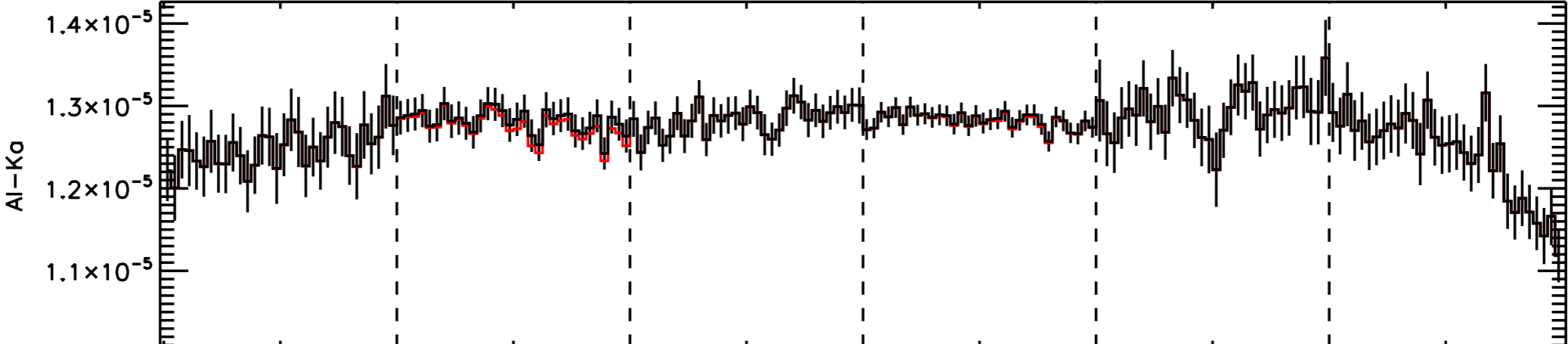
Epoch 1, -120.19  
to -119.19 C



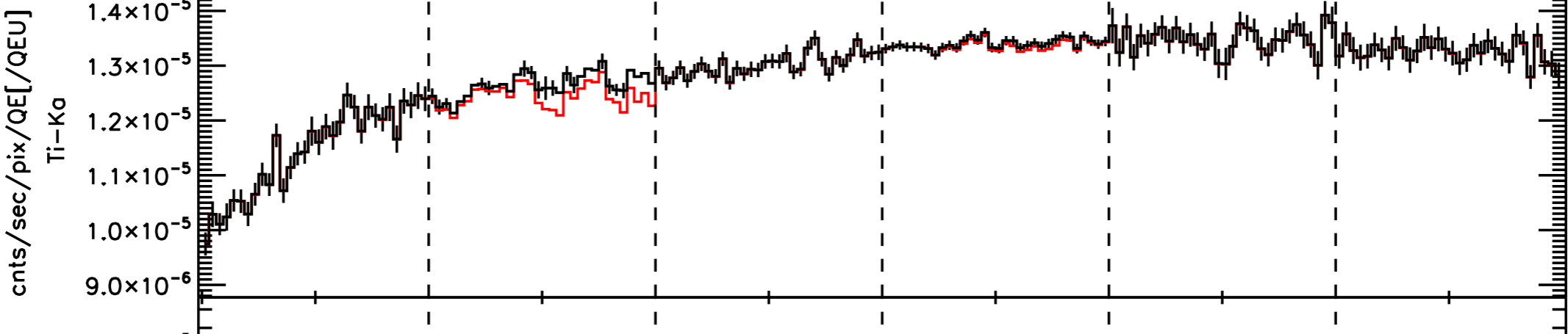
# QUANTUM EFFICIENCY UNIFORMITY (R. DURHAM, P. PLUCINSKY)

ECS yr2000 QEU Correction chipY=33:64

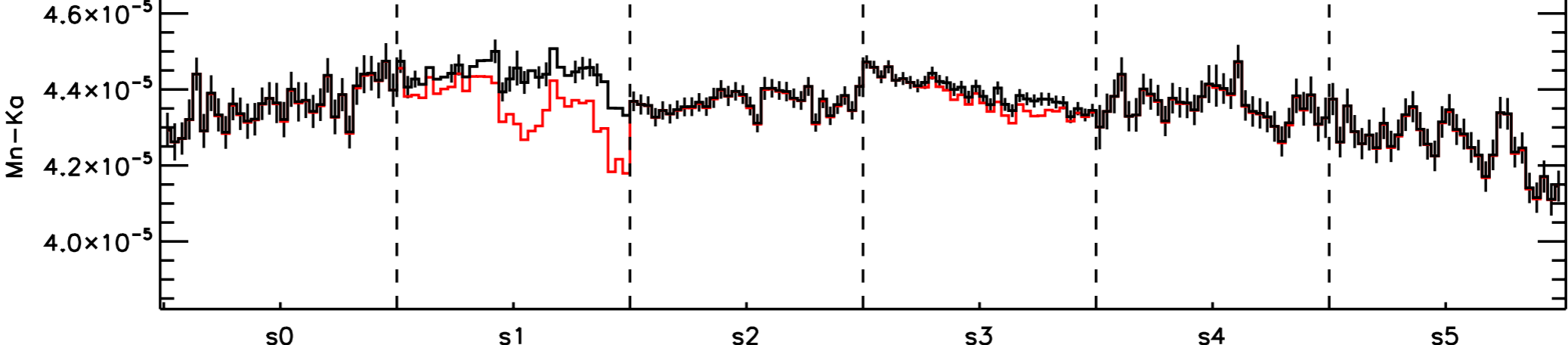
1.5 keV



4.5 keV

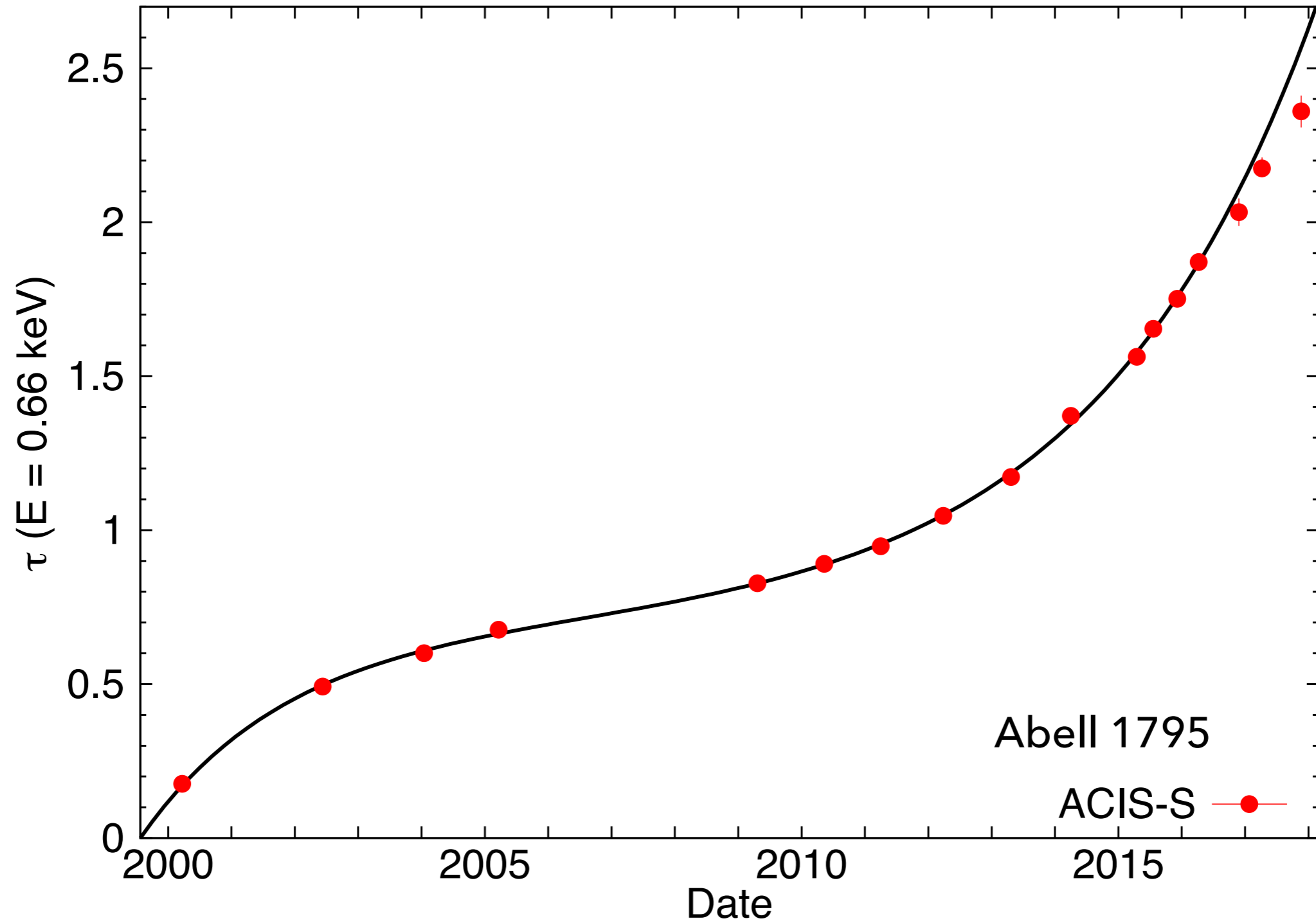


5.9 keV



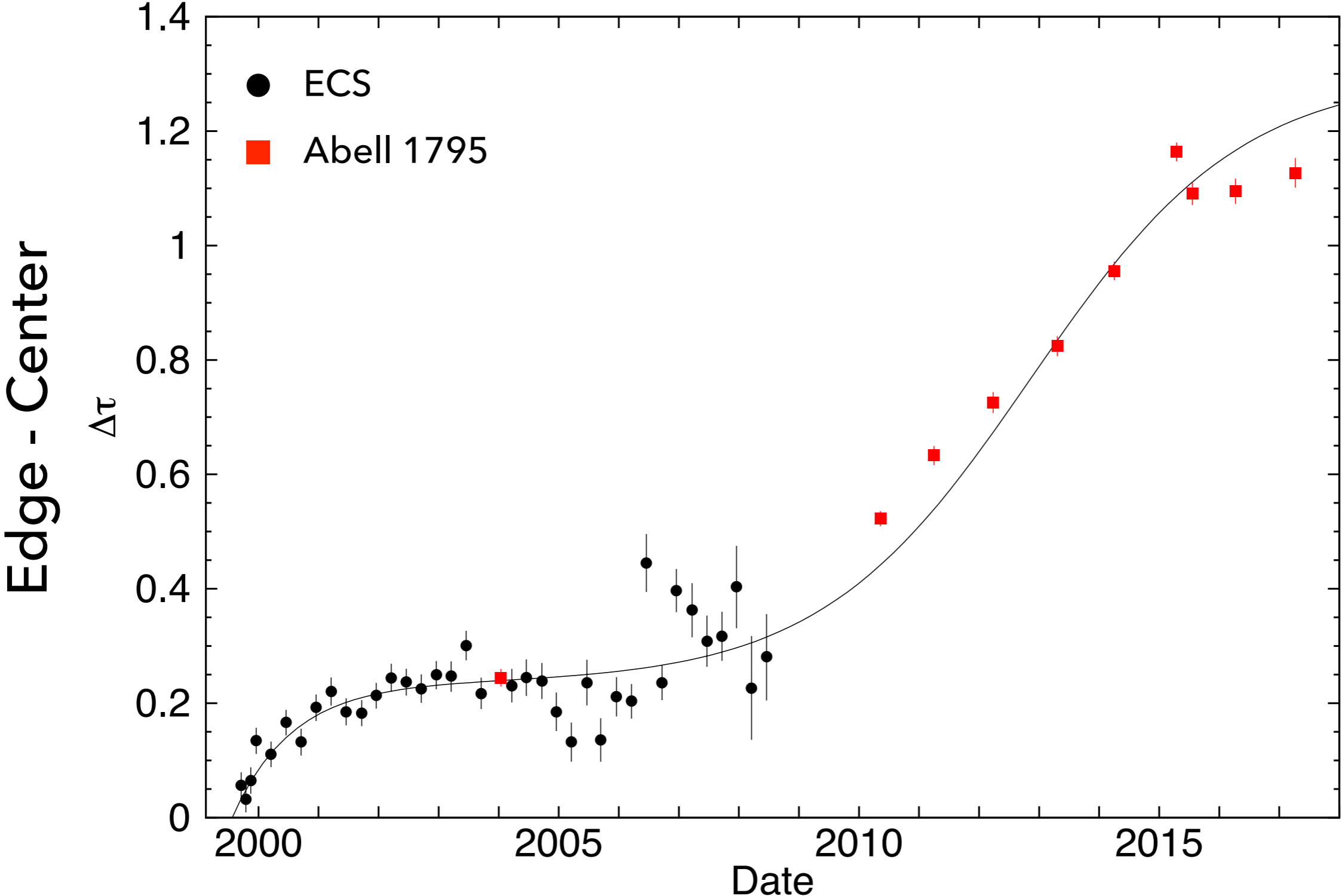
detX

# FILTER CONTAMINATION LAYER (A. BOGDAN, H. MARSHALL, P. PLUCINSKY ET AL)





# FILTER CONTAMINATION LAYER (A, BOGDAN, H. MARSHALL, P. PLUCINSKY ET AL)

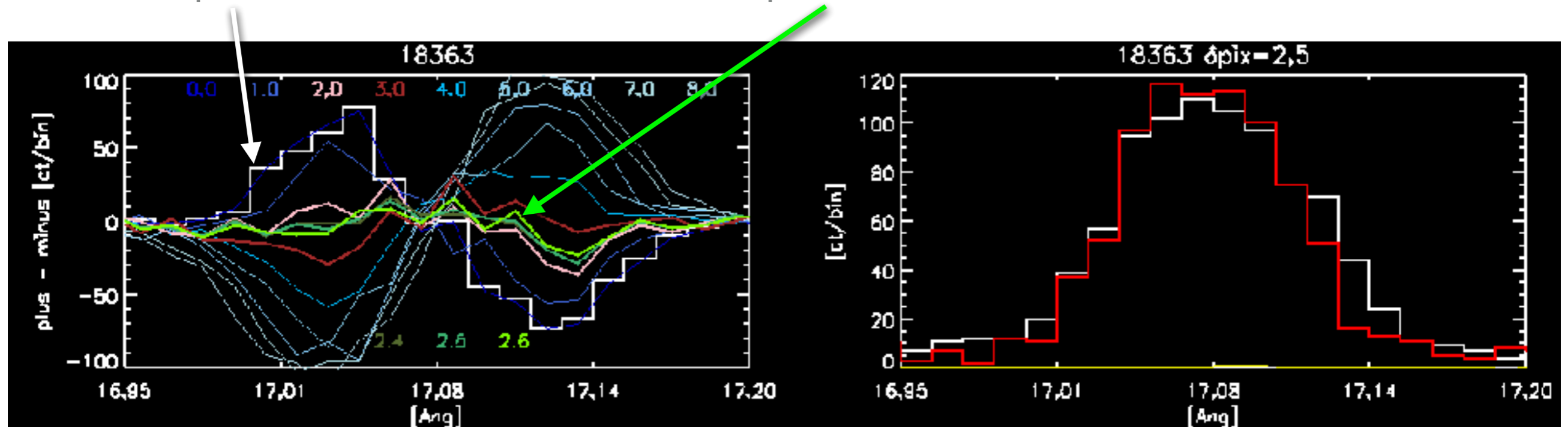


**HIGH RESOLUTION  
CAMERA**

# HRC-S DEGAP UPDATE (V. KASHYAP)

Old line profile difference

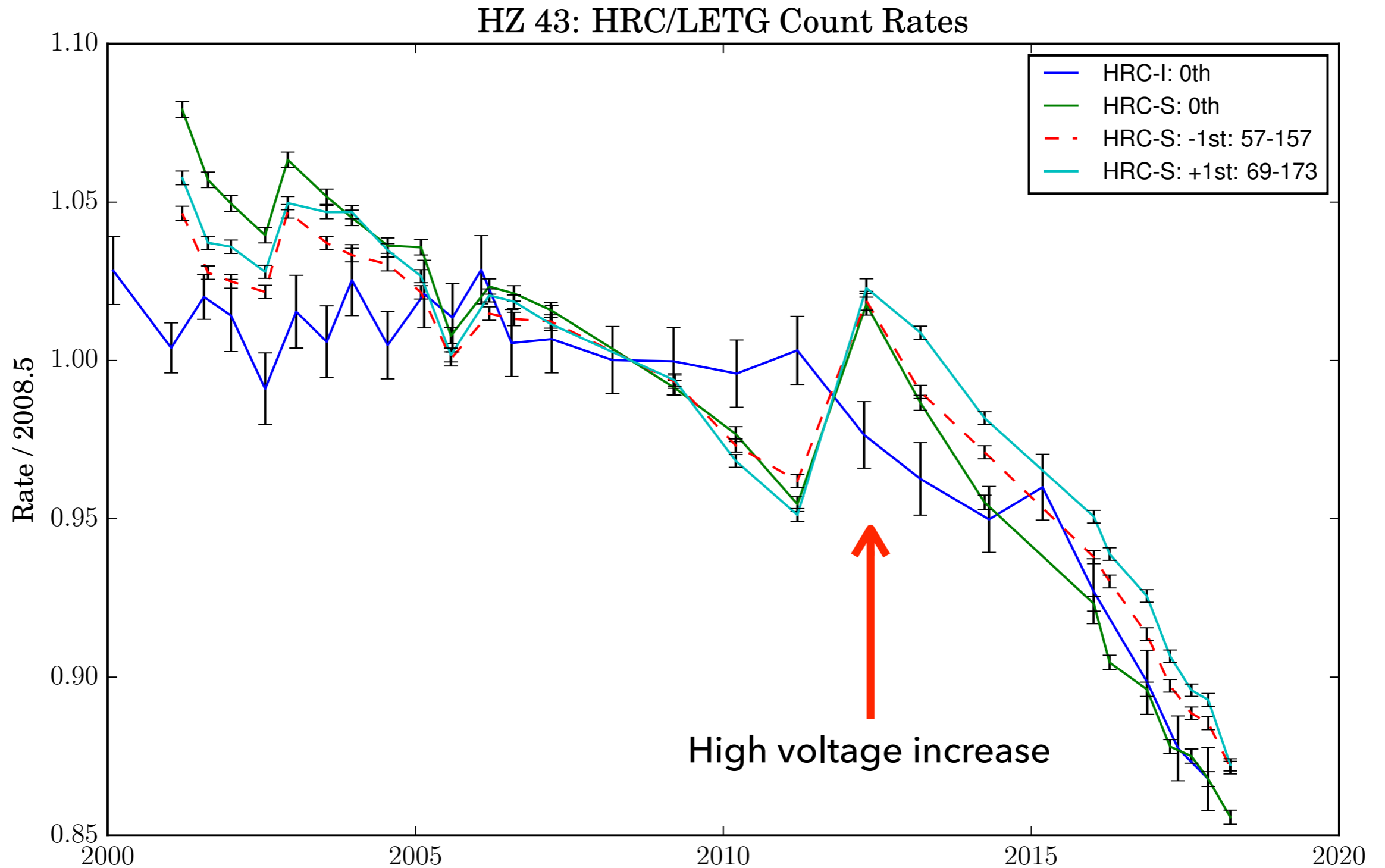
New line profile difference



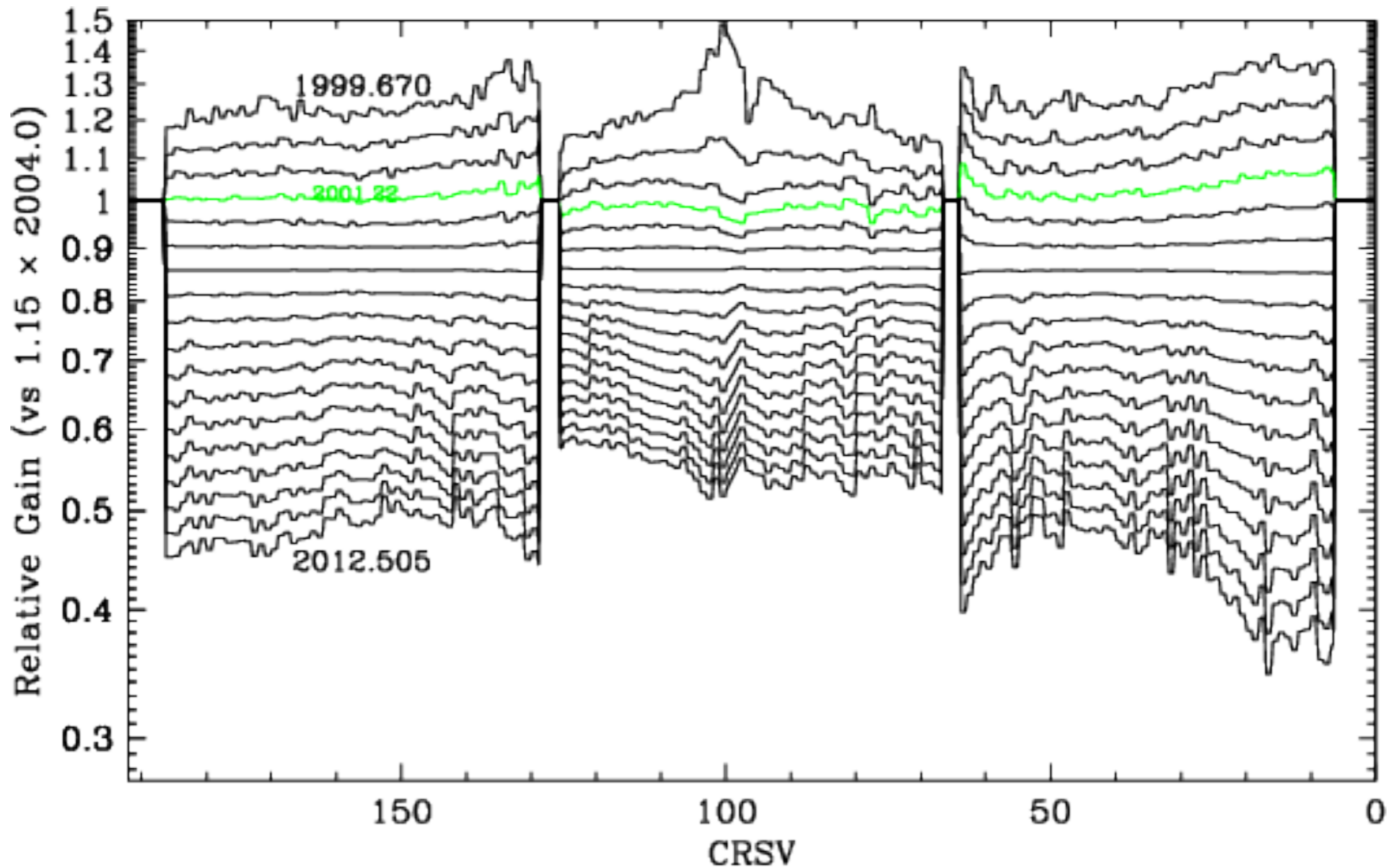
- ▶ HRC-S aim point degap solution improved in 2012 ==> sharper images

- ▶ But...! Caused a shift in derived wavelengths relative to 0<sup>th</sup> order location...

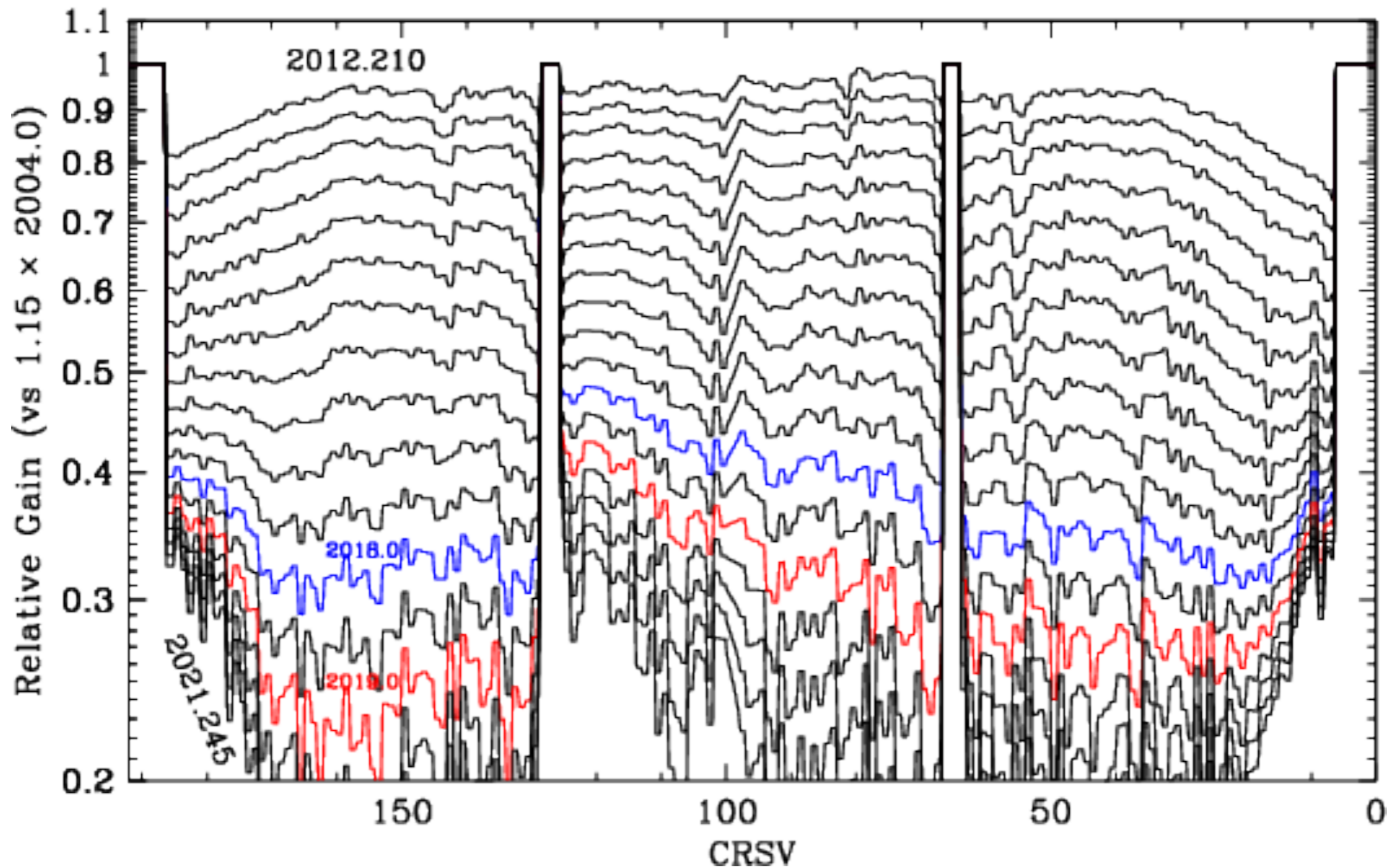
# QUANTUM EFFICIENCY DECLINE (B. WARGELIN, P. RATZLAFF, V. KASHYAP, J. DRAKE)



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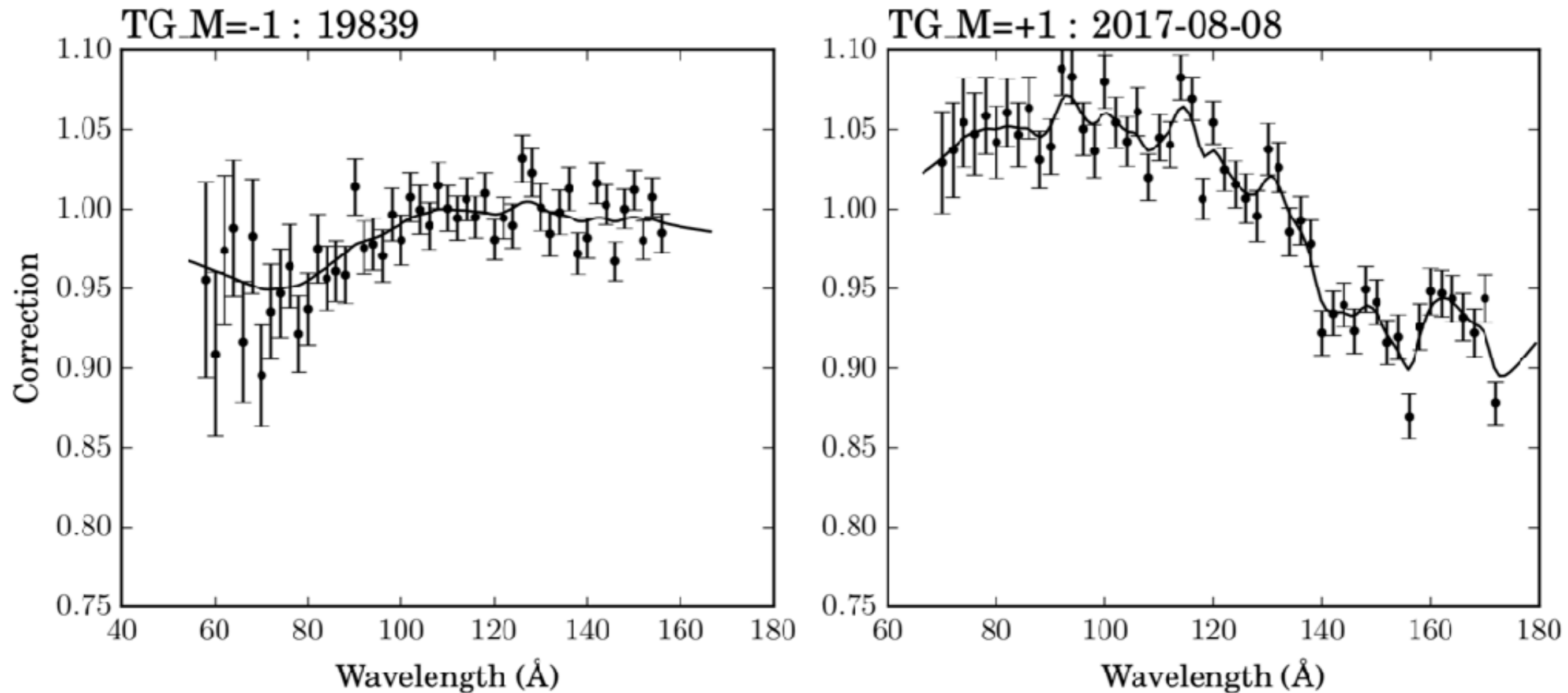


# QUANTUM EFFICIENCY DECLINE (B. WARGELIN, P. RATZLAFF, V. KASHYAP, J. DRAKE)

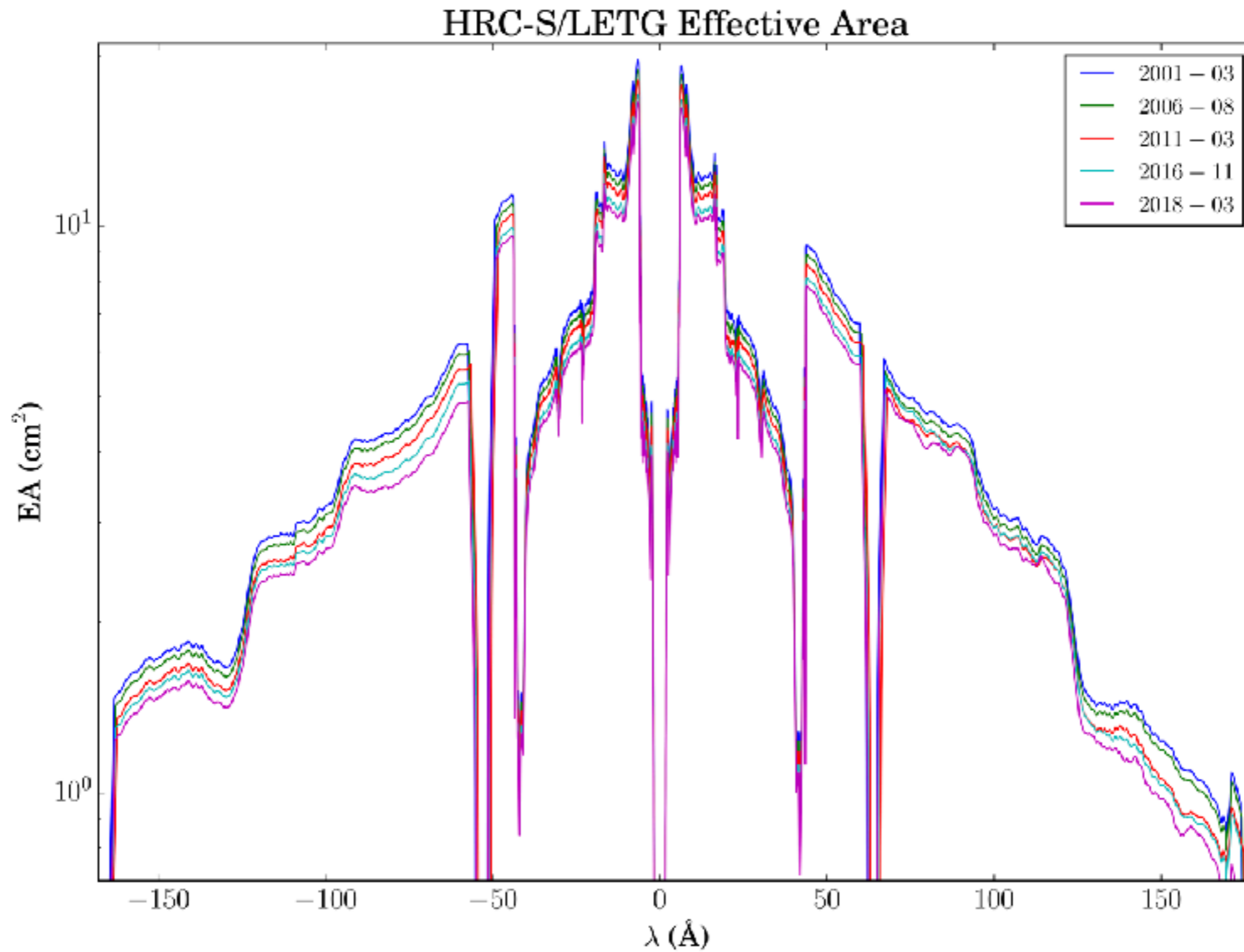


# QUANTUM EFFICIENCY DECLINE (B. WARGELIN, P. RATZLAFF, V. KASHYAP, J. DRAKE)

## HZ43 Empirical QEU Corrections



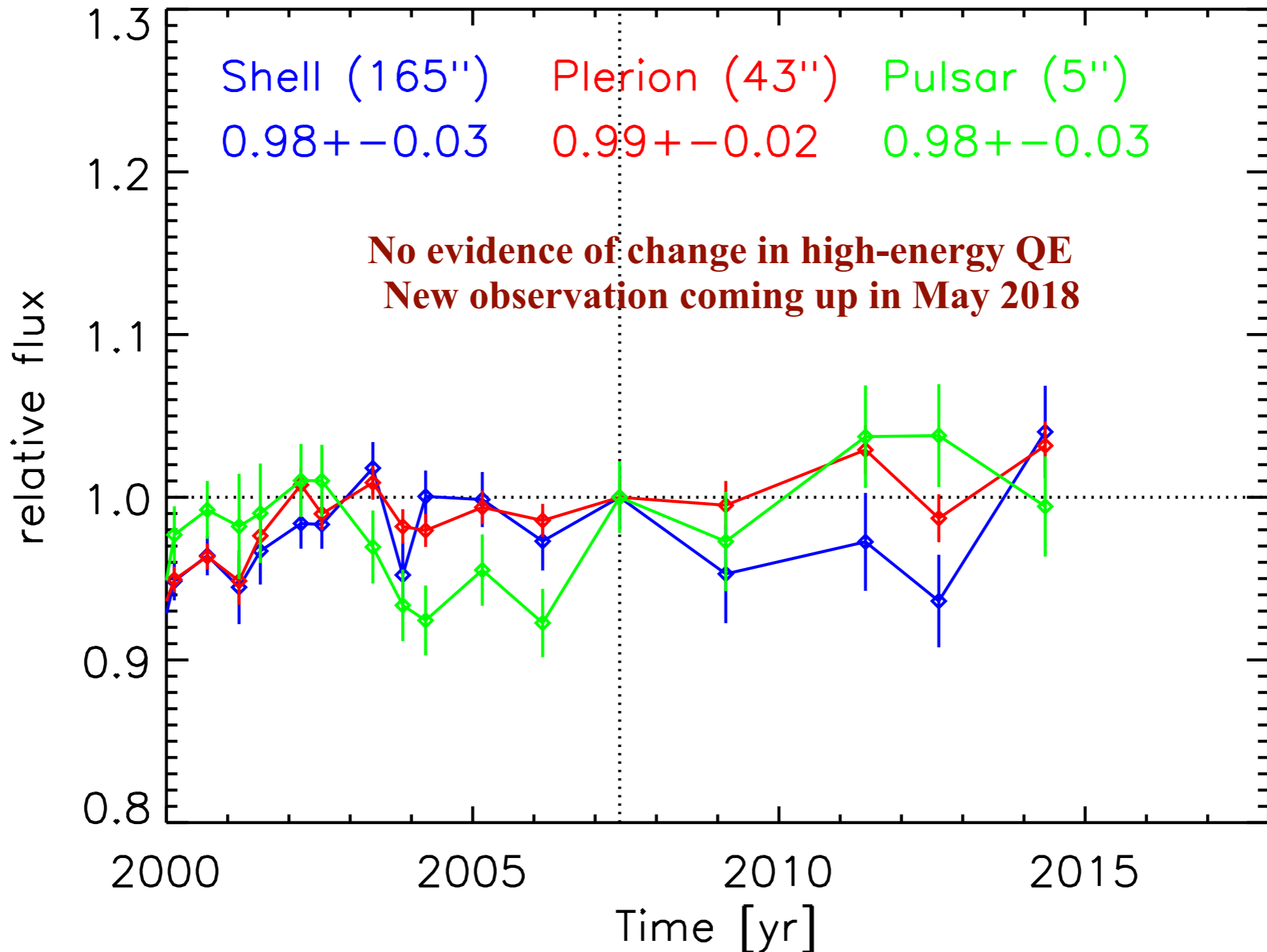
# QUANTUM EFFICIENCY DECLINE (B. WARGELIN, P. RATZLAFF, V. KASHYAP, J. DRAKE)





# HRC-I HIGH ENERGY QE (V. KASHYAP, P. RATZLAFF)

G21.5-0.9 [HRC-I]



# SUMMARY

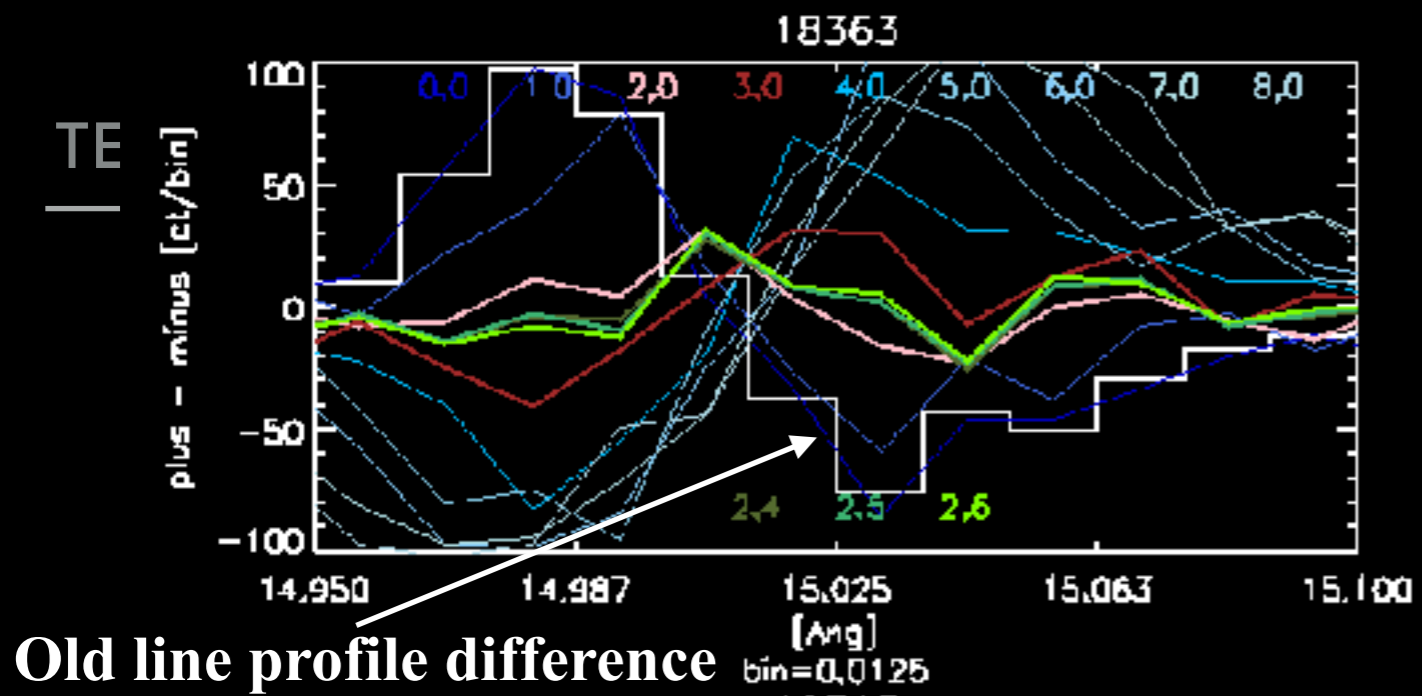
- ▶ EDSEER PSF calibration well underway; empirical PSFs soon
- ▶ ACIS mid-chip gain droop calibration ongoing
- ▶ ACIS QEU improvements underway
- ▶ ACIS contamination model will be updated: slower rate of increase; uniform rate of increase across detector
- ▶ Continuing HRC-S QE secular changes calibrated; HRC-I QE re-calibration underway
- ▶ HV increase on HRC-S.... Only a matter of time



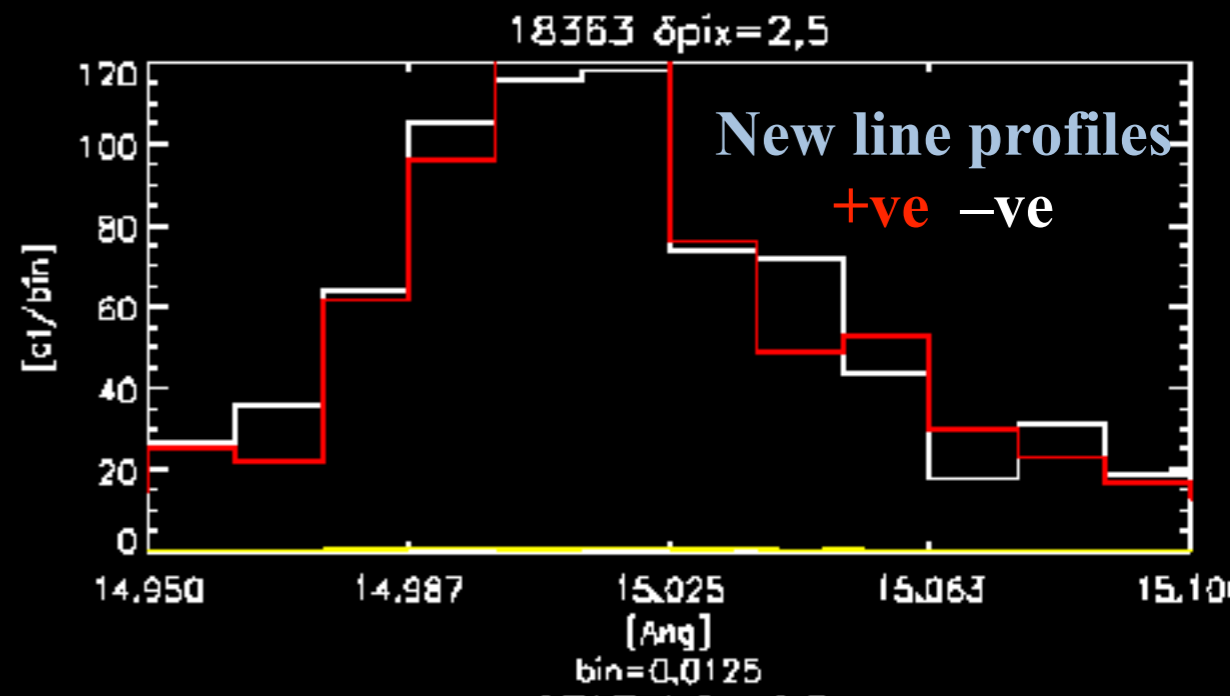
So, how is the concordance project coming along?

# SUPPLEMENTARY MATERIAL

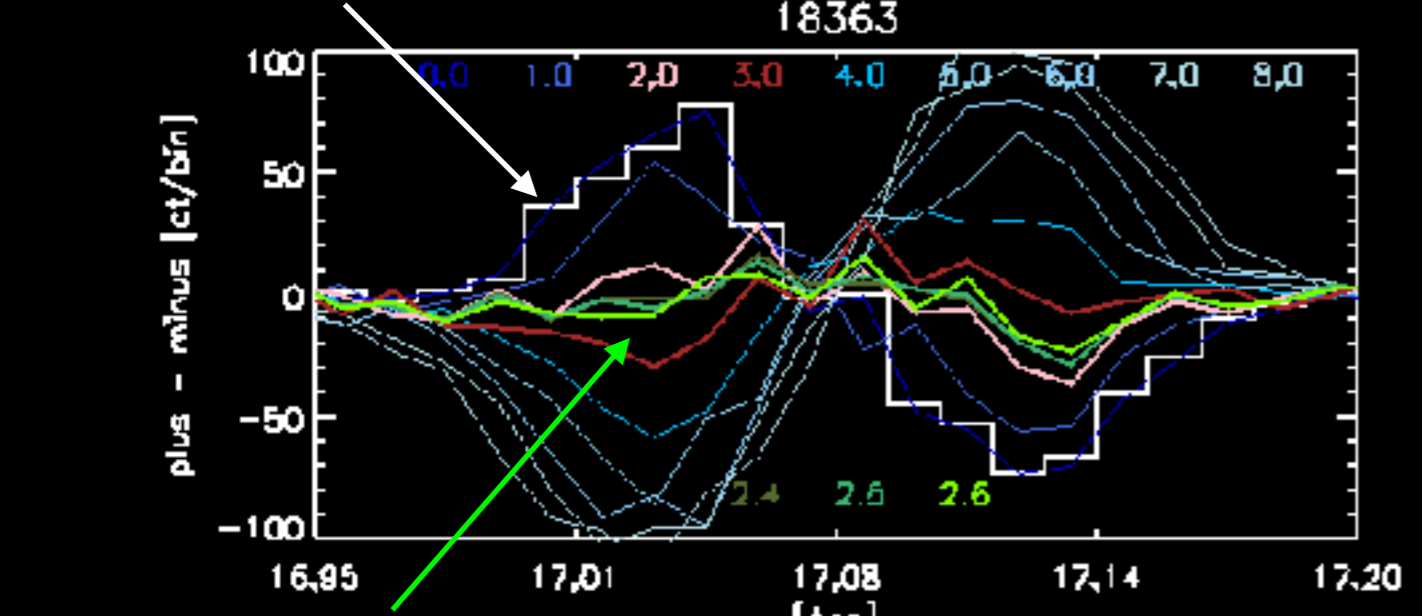
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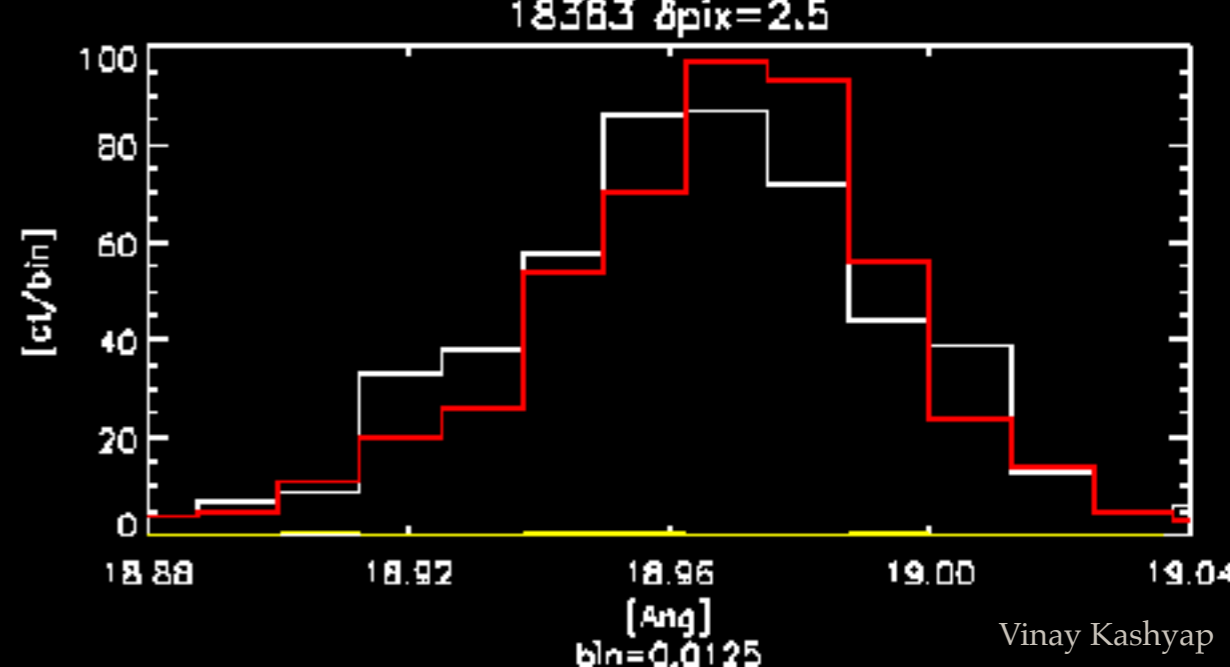
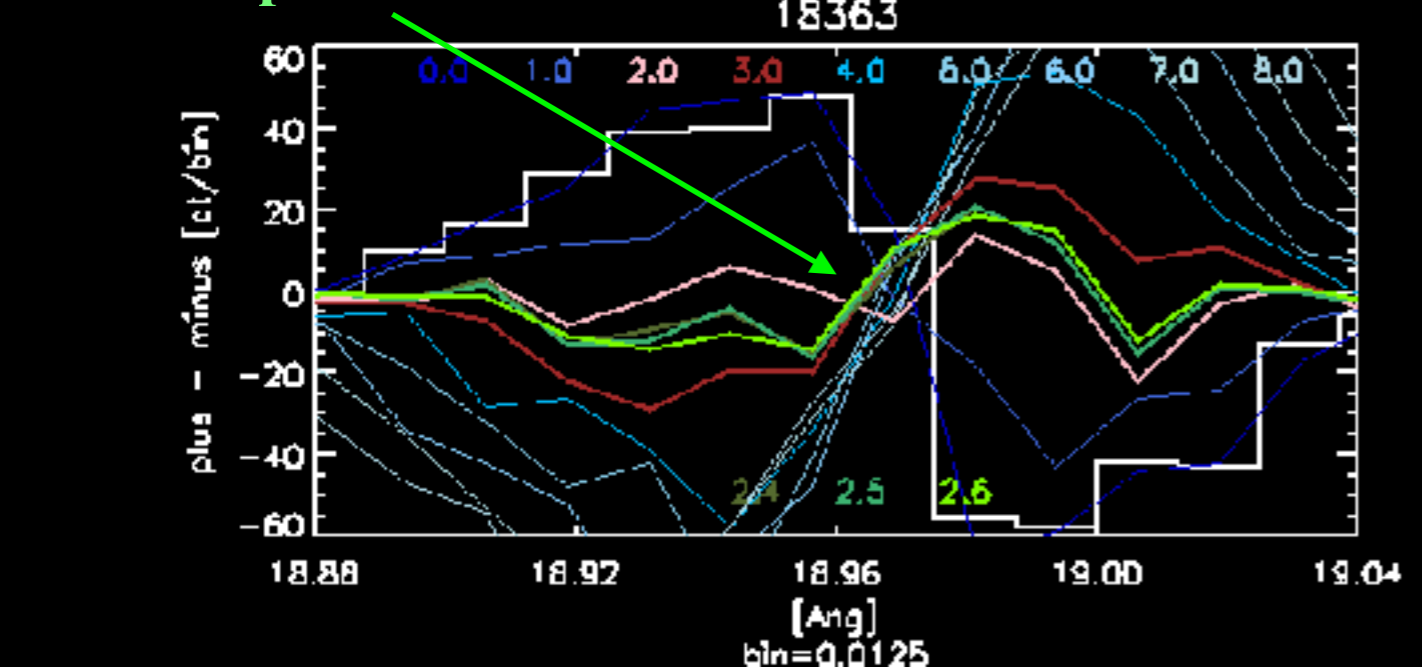
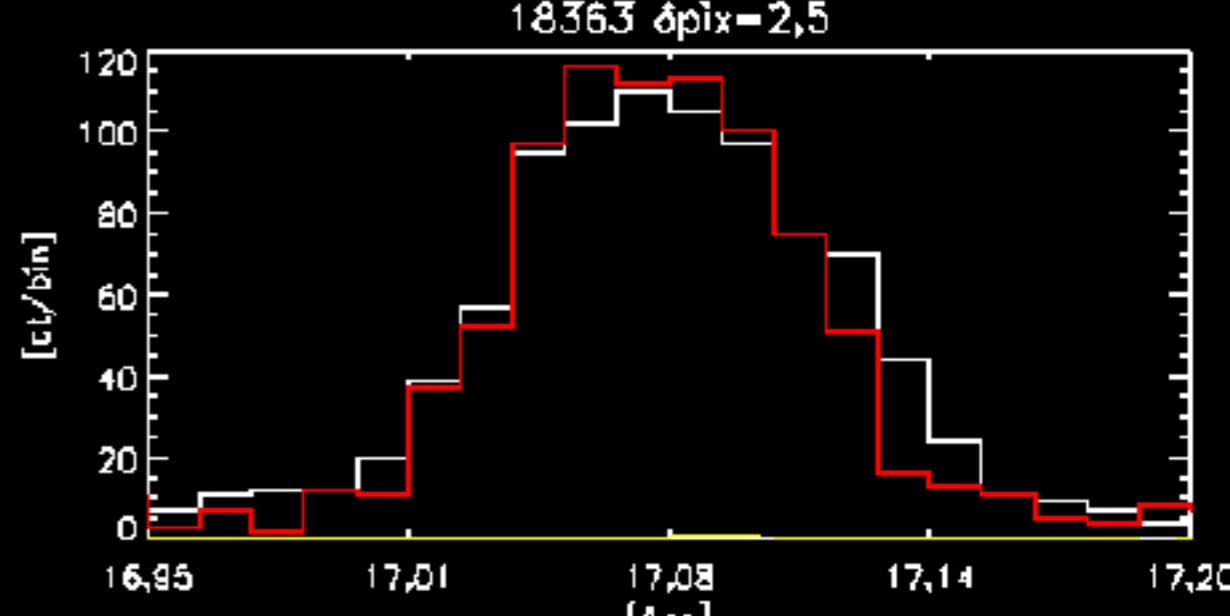
Old line profile difference



New line profiles  
+ve -ve



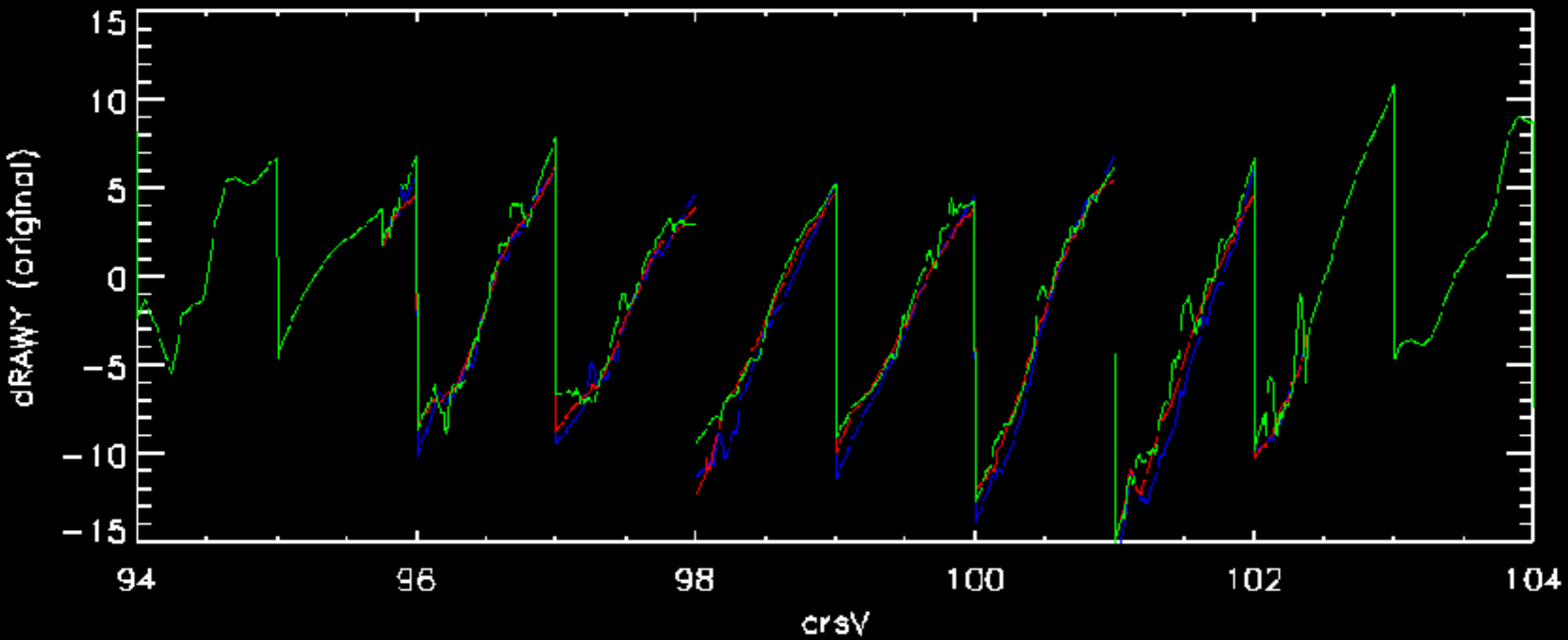
New line profile difference



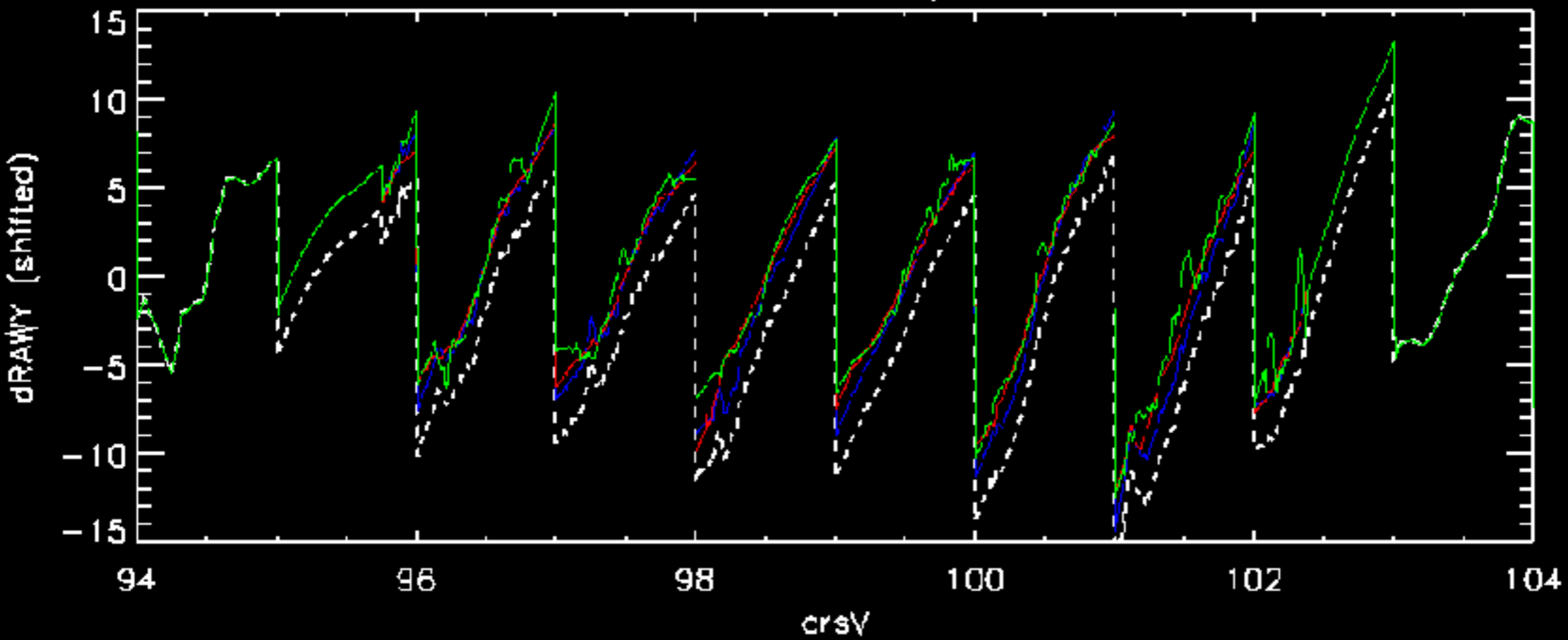
A simple fix: shift the degap for V taps near aim point by +2.5 pix

Top: old (blue / green / red are AMMP\_SF=1,2,3)

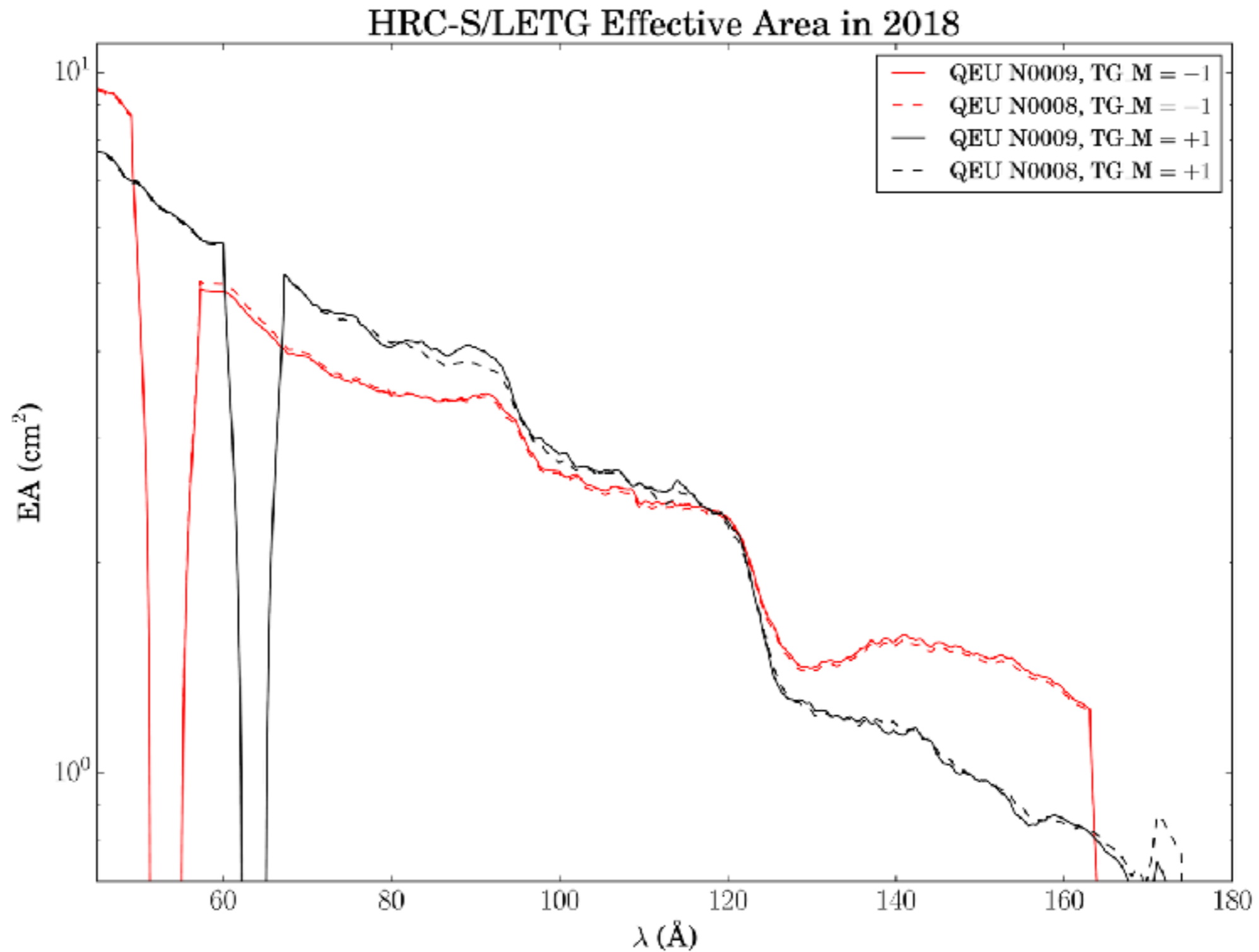
Bottom: new (dashed = old AMMP\_SF=1)



shift = 2.5 pix



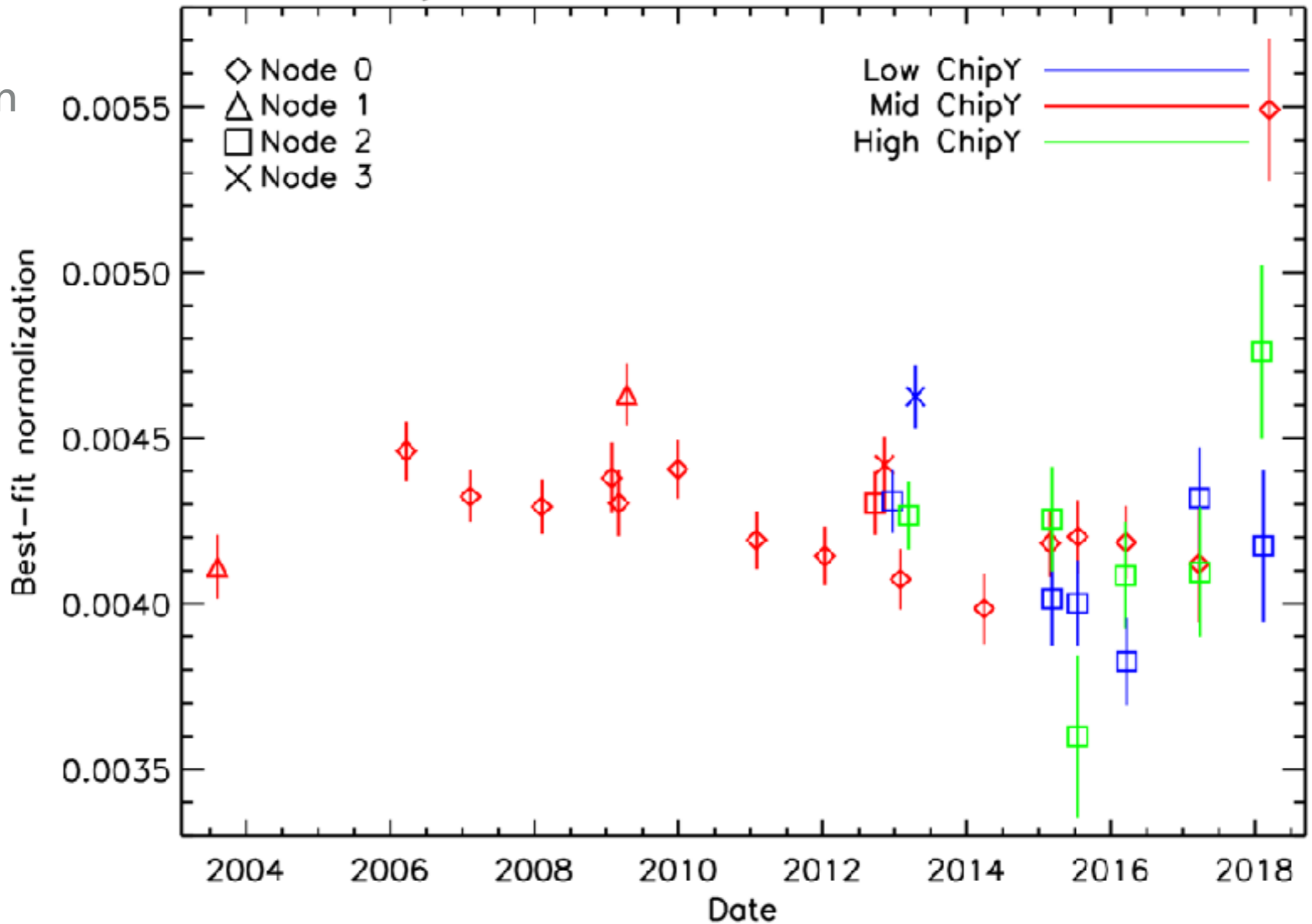
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# FILTER CONTAMINATION LAYER (A, BOGDAN, H. MARSHALL, P. PLUCINSKY ET AL)

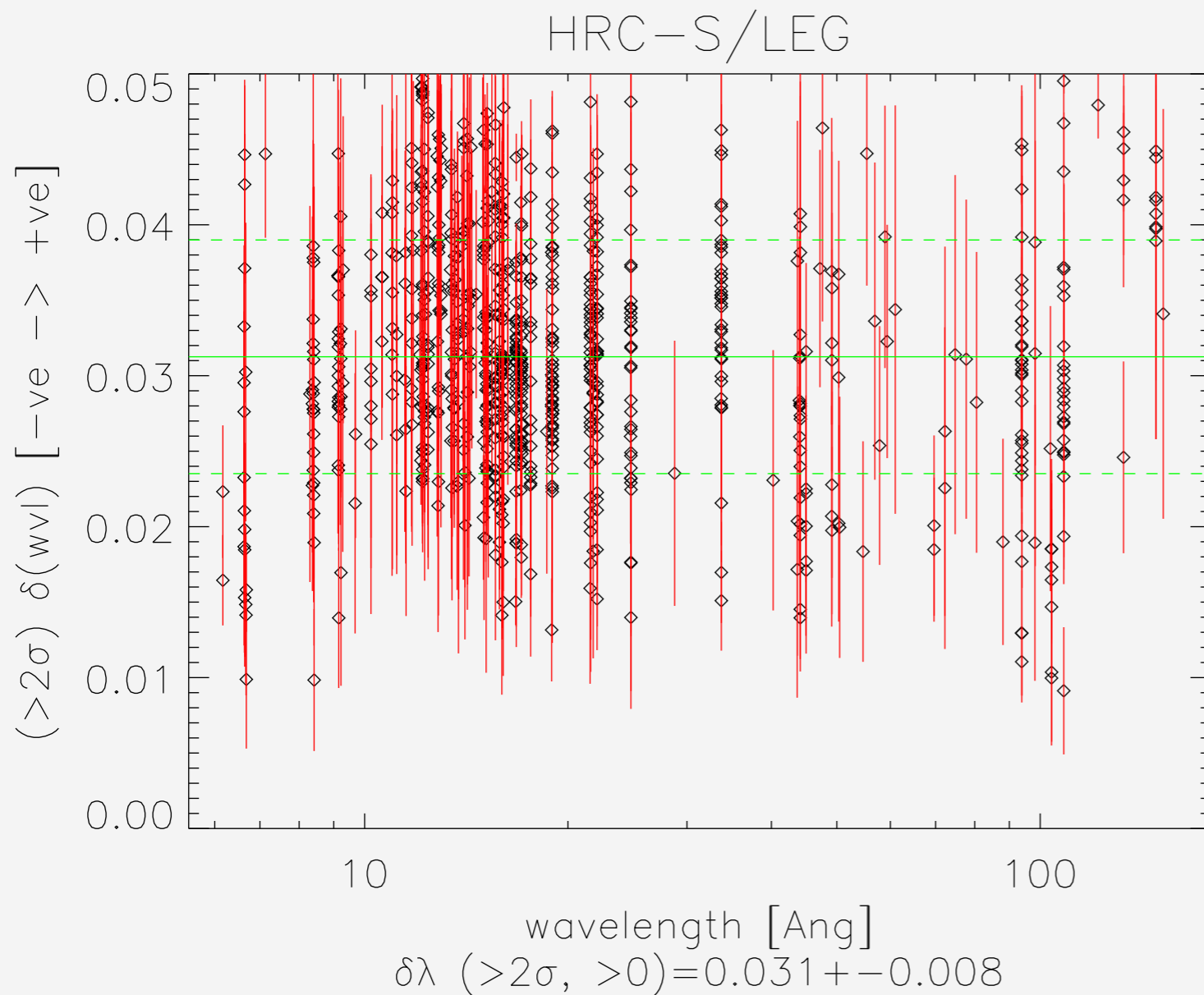
S3 subarray, N0010, CIAO 4.9, CALDB 4.7.8: 08 norm

E0102 O VIII  
normalisation





## HRC-S DEGAP UPDATE (V. KASHYAP)

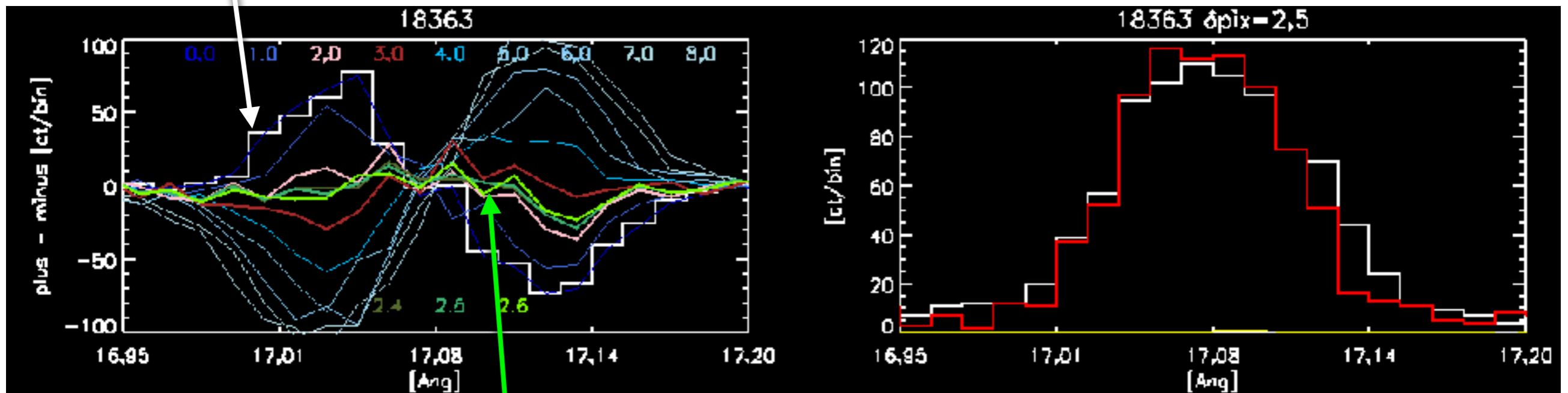


Measured shifts between strong lines seen in coronal sources observed with LETGS+HRC-S

- ▶ HRC-S aim point degap solution improved in 2012 ==> sharper images
- ▶ But introduced a 2-3 pixel offset in absolute astrometry!
- ▶ Caused a shift in derived wavelengths relative to 0<sup>th</sup> order location

# HRC-S DEGAP UPDATE (V. KASHYAP)

Old line profile difference

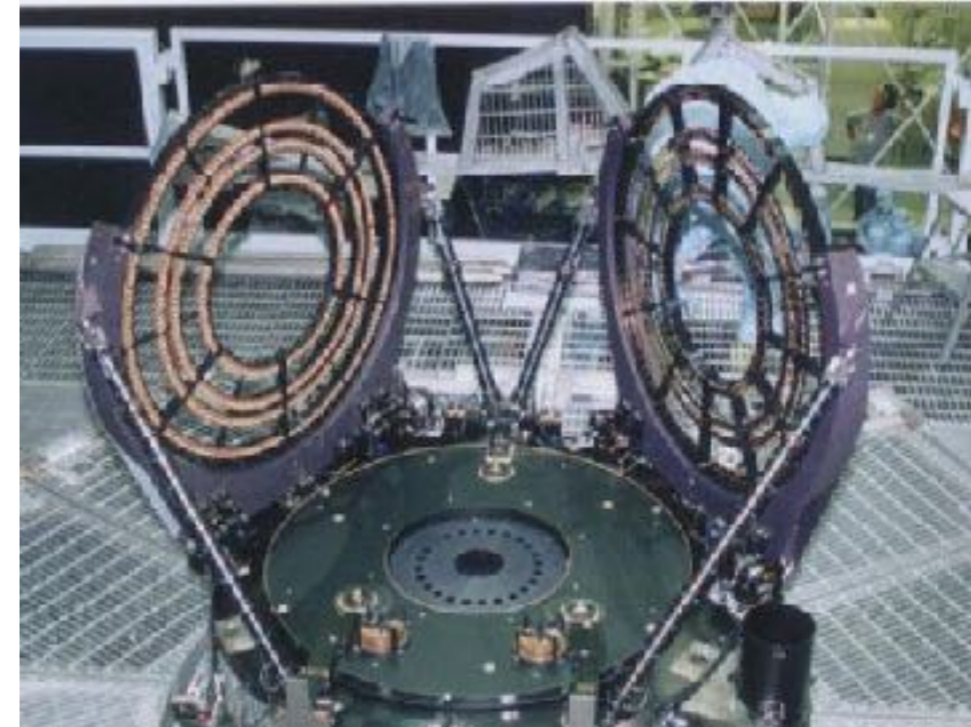
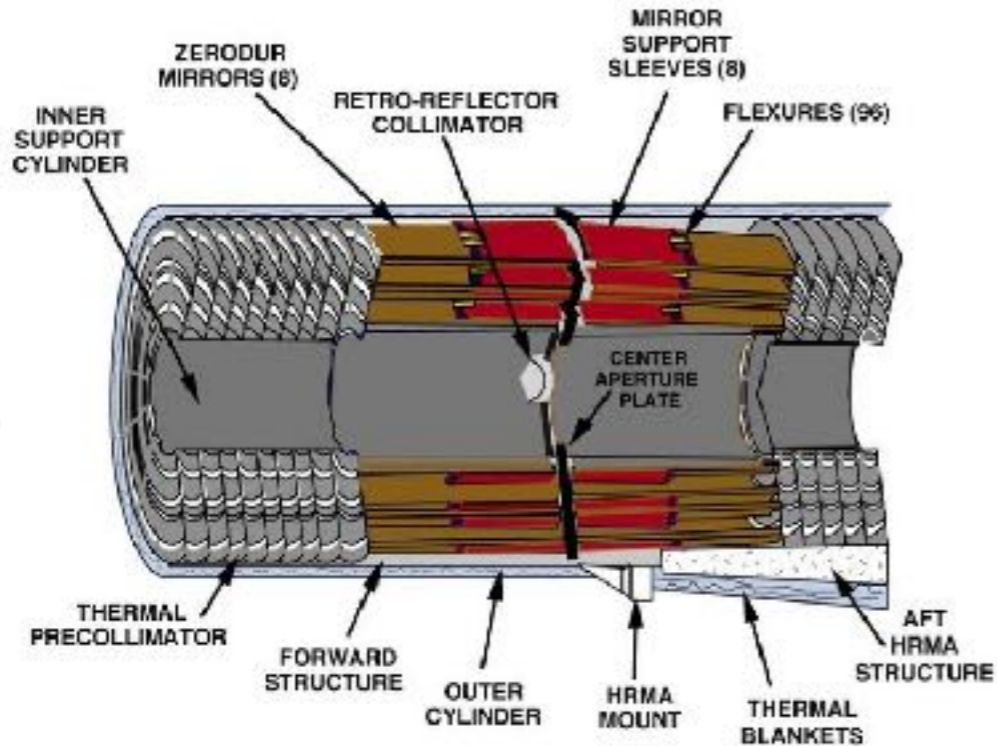


New line profile difference

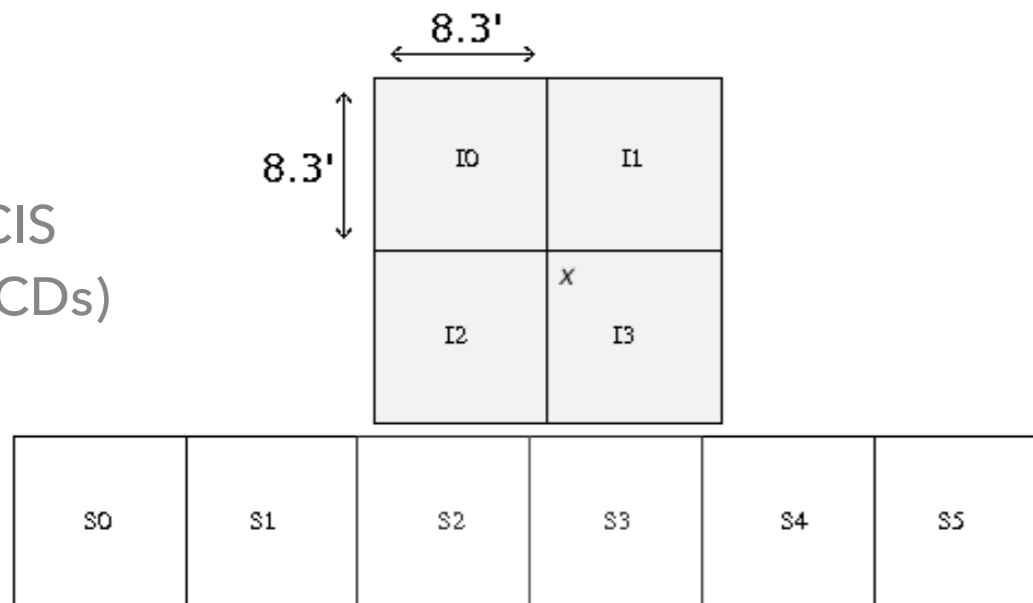
## CHANDRA HARDWARE COMPONENTS

LETG, HETG  
diffraction gratings

HRMA  
Mirrors



ACIS  
(CCDs)



HRC  
(Microchannel  
plates)

