Bringing the High Energy Universe into Focus

NuSTAR Calibration Status

Karl Forster and the NuSTAR team

IACHEC-13, La Tenuta dei Ciclamini, April 9th, 2018

NuSTAR Observatory

- NASA small explorer astrophysics mission
- PI Fiona Harrison (Caltech)
- Partners: ASI, SSDC, DTK, HEASARC
- Launched on June 2012, 620 km, 6° orbit
- Orbital-ATK LeoStar-2 spacecraft bus





10.14m focal length Extendable Mast Conical Wolter-I approximation 133 shells (43 W/Si, 90 Pt/C) HPD = 1 arcminute FOV = 12' x 12'

No consumables

>10 year lifetime

• Single string



CdZnTe detectors 4x(32x32 pixels)

Resolution:

400 eV @ 6 keV 900 eV @ 60 keV 3 ms time resolution

NuSTAR Observatory

NuSTAR



NuSTAR Observatory

NuSTAR



Calibration & Status

Observatory status is green

2017 Calibration news

- Detector monitoring indicates gain has remained ~steady since 2016
 - Long term gain calibration of NuSTAR detectors Brian Grefenstette this afternoon
- Monthly Crab calibration observations focussed and straylight
 - Monitor detector absorption parameters, absolute normalization and vignetting
 - Update on Crab cross-calibration Kristin Madsen Wednesday afternoon
- Timing calibration improvements
 - Goal is to provide relative timing accuracy of a few μs

2018 Operations news

- New observing constraints implemented in 2018
 - Sun avoidance update Saa > 43° (was 39°, improvement in knowledge of astrometry star tracker behavior)
 - Mast leaning constraints (-X 90 ° < Saa < 110 °, +X 70 ° < Saa < 80 °)
 - Metrology calibration limit (104 ° < Saa < 106 °)
 - Optics thermal avoidance (145 ° < Saa < 160 °)
 - Mast adjustment planned for 2018 May
- Increase in multi-observatory coordinations for calibration (and science!)
 - Improvement in communication between operations teams, particularly for ToO's

Long term gain monitoring – Highlights

Brian Grefenstette

- Large dynamic range (2-200 keV) of NuSTAR CZT detectors requires two methods to determine any variations of the transfer function with time.
 PI = PI₀ * SLOPE + OFFSET
- Method 1: Use background lines at 105, 121, and 144 keV to monitor the SLOPE variations with time.
- Method 2: Use routine observations of Kepler to monitor any OFFSET variations with time
- Gain has plateau'd since 2016.
- Plot shows the measured line centroid per year for Det0 on FPMA and FPMB with linear 0.2% / year SLOPE change.
- Investigating last data point and overall trend (non-linear?)



Long term gain monitoring – Highlights

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 PI = PI₀ * SLOPE + OFFSET
- Method 1: Use background lines at 105, 121, and 144 keV to monitor the SLOPE variations with time.
- Method 2: Use routine observations of Kepler to monitor any OFFSET variations with time
 - Offset shows no timedependence.
 - Will continue to monitor Kepler at ~6 month intervals to confirm stability of offset.
 - Multi-epoch fits to Kepler show OFFSET variations over a few years of ~20+/-10 eV.

Full discussion by Brian G. in the Detectors and Background WG this afternoon



Crab calibration campaign

Kristin Madsen

Re-calibration of detector absorption, absolute normalization, and the vignetting files Scope: The new Crab campaign expands upon our focused Crab data base (50 observations) and complements it with StrayLight observations.

- Absolute normalization: StrayLight observations ٠
- Detector absorption: StrayLight observations ٠
- Vignetting files: focused Crab observations from the entire mission lifetime



Full discussion by Kristin M. in the Coordinated Obs and Non-Thermal SNR WG's

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ning calibration status

Matteo Bachetti, Craig Markwardt, Eric Gotthelf

- Current Method:
 - Based on point measurements of clock drift during ground-station passes.
 - Results in absolute and relative timing uncertainties on the order of a few milliseconds via barycorr.
 - Residual orbit-timescale variations in photon arrival times observed for millisecond pulsars.
- New method
 - Inclusion of a new engineering data point (the clock oscillator temperature) and trying a "first principles" clock correction.
 - Will result in some barycorr-like FTOOL for *NuSTAR*.
 - Work in progress; aim to have results by this summer's SPIE meeting.



Pulse-phase drift for PSR B1821-24 (~3 ms period)

Time shift (μs)

New observing constraints

- Avoid observing at certain ranges of Solar aspect angle (Saa)
 - Translates into scheduling constraints



New observing constraints

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 - Translates into scheduling constraints



Mast adjustment - 2018

- We plan to perform another adjustment of the optics / mast interface
 - Periodic adjustments using the mast adjustment mechanism were expected
 - MAM adjustment performed twice before, during IOC and in 2013 September
 - Shift by a few mm in two directions
- Should bring Metrology system back into calibrated range for extremes of mast bending



Optics thermal constraints

- Avoid observing at certain ranges of Solar aspect angle (Saa)
 - Translates into scheduling constraints



Optics thermal behavior

- Change in behavior occurred around 2017-09-25
- Most obvious on outermost sensor on optic 0
- Change not seen in optic 1

Optics thermal constraints

- Avoid observing at certain ranges of Solar aspect angle (Saa)
 - Translates into scheduling constraints
- Apparent degradation has occurred near the thermal sensor on the rear exterior of Optic 0
- Likely a hole or tear in optics cover
- Optics bench qualified to 70 °C
- Current Red limit set at 55 °C
- Avoid observations in range

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145 ^{\circ} < Saa < 160 ^{\circ}
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 Waivers for short observations of time constrained targets may be granted



Multi-observatory coordination

- Steadily improving coordination between high energy missions
 - Regular group notification of acceptance and timing unanticipated ToO observations
- The **Crab calibration** observations on 2018-03-13/14 Is an excellent example
 - Scheduled as monthly NuSTAR calibration observation
 - Request from XMM-Newton for simultaneous observations -> schedule adjustment
- Including straylight exposures during XMM-Newton exposures
 - To determine if adjustments in EPIC-pn calibration should be made
- Notification to other teams resulted in major cross calibration event
 - Astrosat, Insight (HXMT), INTEGRAL, Swift-BAT, Fermi
- Crab exhibited gamma-ray flare episode just before start of observations
 - Simultaneous radio telescope observations (Japan)
- Major transient events are now regularly scheduled with *unanticipated* contemporaneous observations
 - E.g. Swift J1658.2-4242, MAXI J1820+070
- Coordination will be improved by development of VO protocols for observatory operations
 - Led by ops team members at ESA (INTEGRAL + XMM) and CXC

Summary

- Calibration update planned for later this year
 - Absolute normalization, vignetting, detector absorption parameters
 - New clock correction FTOOL relative timing accuracy improvement
- Yearly IACHEC cross calibration campaign on 3C 273 July 2018
- Mast adjustment should solve metrology calibration limits for Saa ~ 105°
 - Mast bending will continue to be monitored
 - Saa constraints remain in place
- Optics thermal behavior also monitored
 - Saa avoidance remain in place
- The additional scheduling constraints will have minor impact on science (or calibration) observations
- Recommendation of NASA senior review in 2019 will determine if mission is extended another 3 years
- Observatory status is green

