13th IACHEC meeting closure

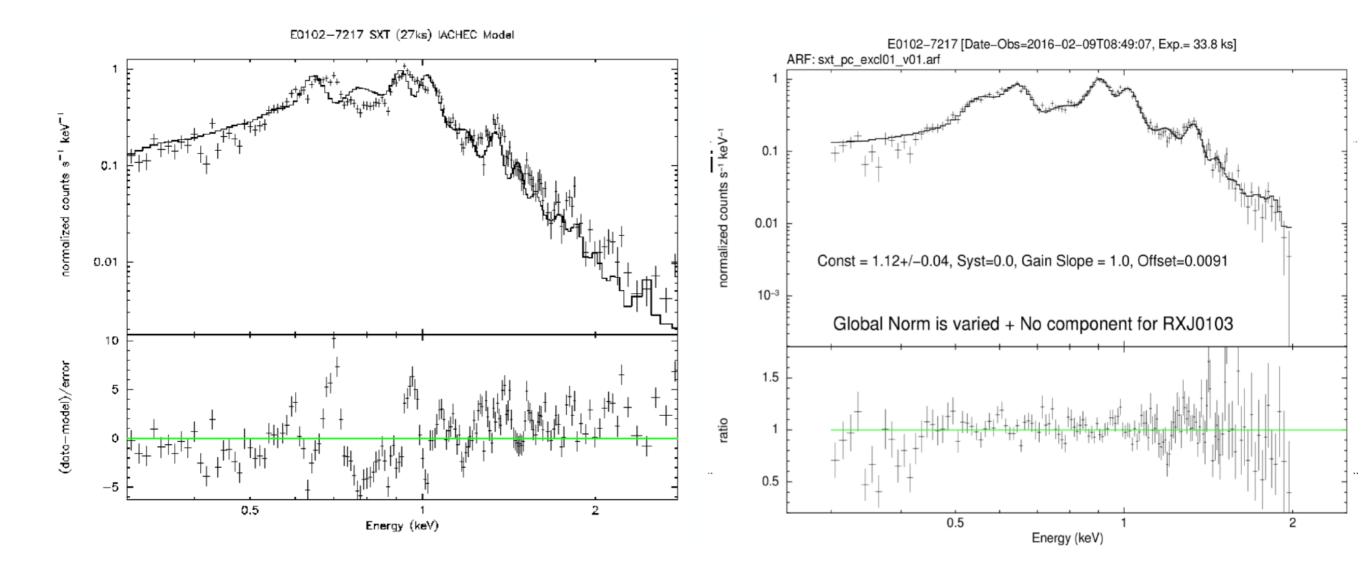
Matteo Guainazzi, Kristin Madsen



["... singing together in perfect harmony", P.Kretschmar]

Scope of the IACHEC

 Provide a forum where calibration teams meet, openly present issues and achievements, and discuss possible solutions



Evolution of attendance

IACHEC year	Number of participants	Number of talks	
2006 (Iceland)	36	26	
2007 (California)	35	30	
2008 (Germany)	36	26	
2009 (Japan)	35	34	
2010 (Massachusetts)	45	50	
2011 (Italy)	44	38	
2012 (California)	40	29	
2013 (UK)	36 (+6 seq.ed)	48 (20 plenary)	
2014 (Virginia)	51	54 (24 plenary)	
2015 (RPC)	57 + 37 students	45 (25 plenary)	
2018 (Italy)	52	50 (30 plenary)	

Scope of the IACHEC

1. Provide a forum where calibration teams meet, openly present issues and achievements, and discuss possible solutions

2. Define calibration standard and procedures

The IACHEC aims to provide standards for high energy calibration and supervise cross calibration between different missions. This goal is reached through working groups, where IACHEC members cooperate to define calibration standards and procedures. The scope of these groups is primarily a practical one: a set of data and results (eventually published on refereed journals) will be the outcome of a coordinated and standardized analysis of references sources ("high-energy standard candles"). Past, present and future high-energy mission can use these results as a calibration reference.

[IACHEC web page]

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- To-be IACHEC Standard Candles: Capella, N132D, RXJ1856-3754

IACHEC paper status

Paper	Status last year	Status this year	
C/X blazar sample	wrapping up	wrapping up	
Crab	advanced draft	in preparation	
3C273 (INTEGRAL)	_	to be started	
3C273 (NICER)	_	to be started	
3C273 (NuSTAR)	_	to be started	
G21.5-0.9	_	to be started	
N132D	not discussed	in preparation	
Concordance Project	_	submitted/in preparation	
MMS (galaxy clusters)	to be started	in preparation	
RXJ1856-3754	_	under consideration	
Timing (Crab)	_	under consideration	
Crab phase-resolved	_	under consideration	

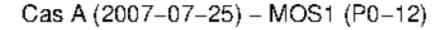


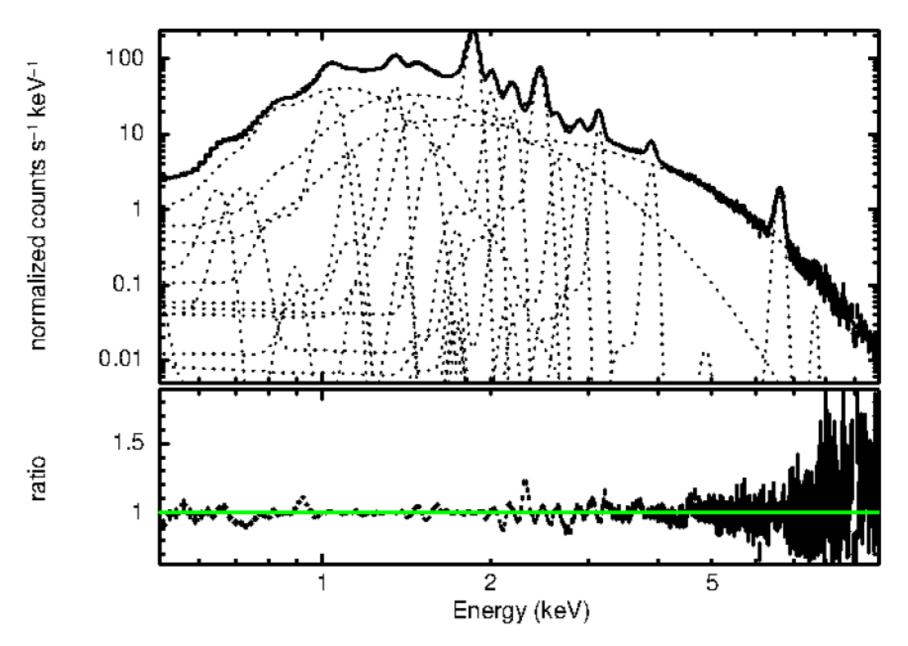
Cas A Remnant Averaged Model



A.Beardmore's presentation

- In the spirit of IACHEC
 - MOS1 small window spectrum
 - model tbabs * (3 brems + 15 gaus)







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Y.Terada's presentation

XARM is the recovery mission of Hitomi

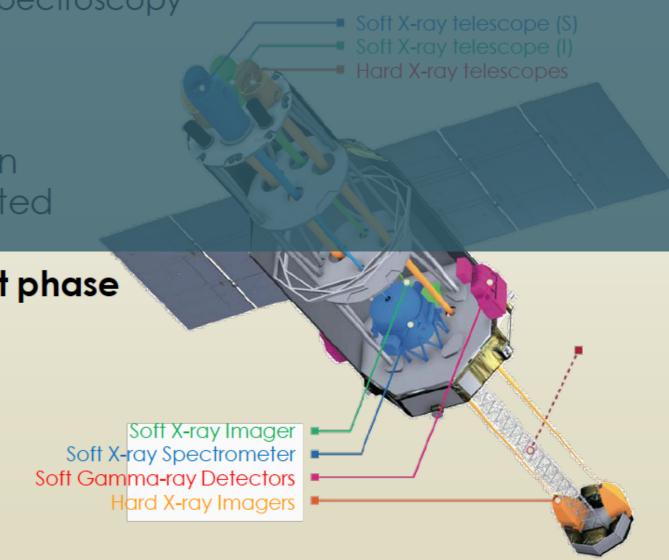
ASTRO-H/Hitomi Mission

- X-ray observation in 0.3 ~ 600 keV
 - ✓ High resolution spectroscopy
 - ✓ Wide FOV Imaging
 - ✓ Hard X-ray Imaging spectroscopy
 - ✓ Super sentive gamma-ray spectroscopy
- 2003 NeXT project
- 2005~ ASTRO-H mission
- 20/16.2.17 Launch
- 2016.3.26 lost communication
- /2016.4.28 Operation terminated

Objects observed during check-out phase

- Perseus Cluster of galaxies
- N132D
- IGR J16318-4848
- RX J1856.5-3754
- G21.5-0.9
- Crab

(so call, IACHEC objects!!)





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Papers discussed/to discuss the comparison/consistency with IACHEC (Crab and G21.5-0.9) [Tsujimoto]

Crab

(so call, IACHEC objects!!)





Hard X-ray Imagers



Background: 1d

- During/after commissioning of all cameras: defined mixture of cameras with closed and (open) filter, high and low gain (TBC)
- as predefined set of commands (macro)
- similar to calibration source: one after another CLOSED filter during survey
- joint SRG background study for L2: eROSITA + ART-XC

Clusters of galaxies: 2d

- Are more subject of cross-calibration with other missions than actual eROSITA calibration (scientific like cluster T, not gain/CTI)
- eROSITA advantage: no chip gaps, large FOV!
- Preferred targets (1 low-T, 1 high-T): IACHEC recommendation A1795, A2029 (Coma (center to be defined), A1835, A2052, A2199)
- work to be expected, e.g.: derive T for eROSITA (7 temperatures should be the same within errors), XMM, Chandra determine differences and re-iterate (effective area, vignetting, EEF, RMF),... contribute to calibration parameters

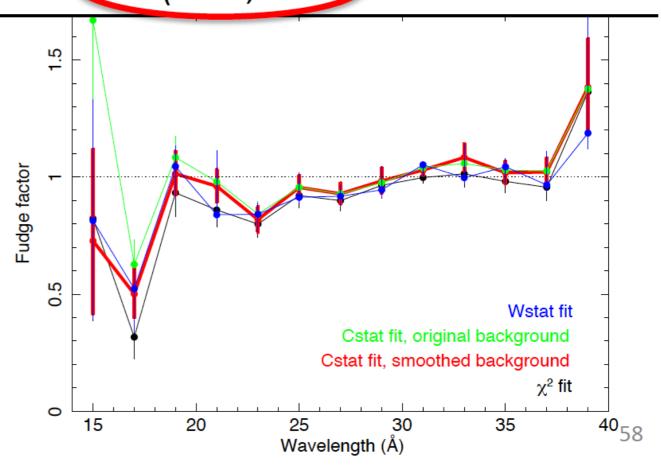
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- 1. Provide a forum where calibration team meet, openly present issues and achievements, and discuss possible solutions
- 2. Define calibration standard and procedures
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- 4. Represent an *autoritas* on calibration, inter-calibration, and cross-calibration, but also of technical and statistical knowledge in the analysis of data in general

Comparison methods

Method	χ²	Cstat (data bg)	Wstat	Cstat (Model bg)
Statistic	1188	1405	1154	1161
Dof/range	1102	1136±48	1102	1135±48
# evaluations	3084	2888	12873	3527
F (17 Å)	0.32±0.10	0.63±0.10	0.52(-0.18,+0.09)	0.50±0.10
F (27 Å)	0.90±0.05	0.93±0.04	0.92±0.04	0.93±0.04
F (37 Å)	0.96±0.06	1.03±0.06	0.97(-0.03,+0.14)	1.02±0.06

- χ^2 & Cstat with data bg:
 - → bias
- Wstat: unstable, asymmetric errors, 4 x more model evaluations; bias?



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- Several actions on pushing for funding laboratory measurements (HEAD, AHEAD, Athena Decadal WP ...)

Many thanks to:

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See you all in Shonan Village in May 2019!