

The XARM mission status and Science operations plan

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M. Nobukawa (XARM Xtend calibration), K. Mori (XARM Xtend PE),
and XARM pre-project team

XARM is the recovery mission of Hitomi

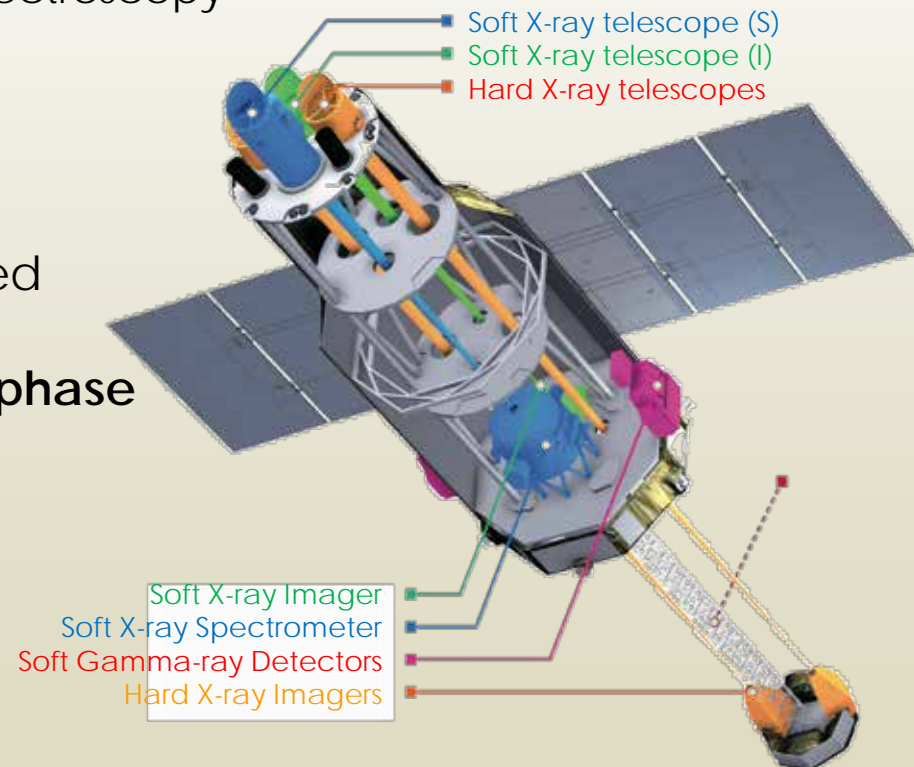
ASTRO-H/Hitomi Mission

- X-ray observation in 0.3 ~ 600 keV
 - ü High resolution spectroscopy
 - ü Wide FOV Imaging
 - ü Hard X-ray Imaging spectroscopy
 - ü Super sensitive gamma-ray spectroscopy
- 2003 NeXT project
- 2005~ ASTRO-H mission
- 2016.2.17 Launch
- 2016.3.26 lost communication
- 2016.4.28 Operation terminated

Objects observed during check-out phase

- Perseus Cluster of galaxies
- N132D
- IGR J16318-4848
- RX J1856.5-3754
- G21.5-0.9
- Crab

(so call, IACHEC objects!!)



Scientific Outputs from Hitomi mission

11 papers

| Perseus Cluster | | | |
|--|---------------|--------|------------|
| The Quiescent Intracluster Medium in the Core of the Perseus Cluster | A.Fabian | Nature | 2016 July |
| Hitomi constraints on the 3.5 keV line in the Perseus galaxy cluster | M. Markevitch | ApJL | 2016 July |
| Solar abundance ratios of the iron-peak elements in the Perseus cluster | H.Yamaguchi | Nature | 2017 Nov |
| Measurements of resonant scattering in the Perseus cluster core with Hitomi SXS | K.Sato | PASJ | 1710.04648 |
| Atmospheric gas dynamics in the Perseus cluster observed with Hitomi | Y.Ichinohe | PASJ | 1711.00240 |
| NGC1275 | | | |
| Hitomi Observation of Radio Galaxy NGC 1275: The First X-ray Microcalorimeter Spectroscopy of Fe-K(α) Line Emission from an Active Galactic Nucleus | H.Noda | PASJ | 1711.06289 |
| N132D | | | |
| Hitomi Observations of the LMC SNR N132D: Highly Redshifted X-ray Emission from Iron Ejecta | E.Miller | PASJ | 1712.02365 |
| RXJ1856-3754 | | | |
| (calibration paper only) | -- | -- | -- |
| IGR J16318-4848 | | | |
| Glimpse of the highly obscured HMXB IGR J16318–4848 with Hitomi | H.Nakajima | PASJ | 1711.07727 |
| G21.5-0.9 | | | |
| Hitomi X-ray Observation of the Pulsar Wind Nebula G21.5 ϕ -0.9 | H.Uchida | PASJ | 1802.05068 |
| Crab | | | |
| Search for Thermal X-ray Features from the Crab nebula with Hitomi Soft X-ray Spectrometer | M.Tsujimoto | PASJ | 1707.00054 |
| Hitomi X-ray studies of Giant Radio Pulses from the Crab pulsar | Y.Terada | PASJ | 1707.08801 |



**PASJ Hitomi Special
Coming soon.**

Instruments papers



| | Title | Author | Special Issue |
|--|--|-----------------|---------------|
| AH | The Hitomi (ASTRO-H) x-ray astronomy satellite | T. Takahashi | JATIS |
| SXS | Thermal analyses for initial operations of the soft x-ray spectrometer onboard the Hitomi satellite | H. Noda | JATIS |
| | Porous plug phase separator and superfluid film flow suppression system for the soft x-ray spectrometer onboard Hitomi | Y. Ezo | JATIS |
| | Calibration sources and filters of the soft x-ray spectrometer instrument on the Hitomi spacecraft | Cor P. de Vries | JATIS |
| | In-orbit operation of the soft x-ray spectrometer onboard the Hitomi satellite | M. Tsujimoto | JATIS |
| | Performance of the helium dewar and the cryocoolers of the Hitomi soft x-ray spectrometer | R. Fujimoto | JATIS |
| | Design, implementation, and performance of the Astro-H SXS calorimeter array and anticoincidence detector | C. Kilbourne | JATIS |
| | Design, implementation, and performance of the Astro-H soft x-ray spectrometer aperture assembly and blocking filters | C. Kilbourne | JATIS |
| | Vibration isolation system for cryocoolers of soft x-ray spectrometer on-board ASTRO-H (Hitomi) | Y. Takei | JATIS |
| | In-flight performance of pulse-processing system of the ASTRO-H/Hitomi soft x-ray spectrometer | Y. Ishisaki | JATIS |
| | ★ In-flight performance of the soft x-ray spectrometer detector system on Astro-H | F. S. Porter | JATIS |
| ★ In-flight calibration of Hitomi Soft X-ray Spectrometer. (1) Background | C. A. Kilbourne | PASJ | |
| ★ In-flight calibration of the Hitomi Soft X-ray Spectrometer. (2) Point spread function | Y. Maeda | PASJ | |
| ★ In-flight calibration of Hitomi Soft X-ray Spectrometer. (3) Effective area | M. Tsujimoto | PASJ | |
| SXI | Soft X-ray Imager aboard Hitomi (ASTRO-H) | T. Tanaka | JATIS |
| ★ | In-orbit performance of the soft X-ray imaging system aboard Hitomi (ASTRO-H) | H. Nakajima | PASJ |
| SXT | ★ Ground-based x-ray calibration of the Astro-H/Hitomi soft x-ray telescopes | R. Iizuka | JATIS |

★ Calibration, performance

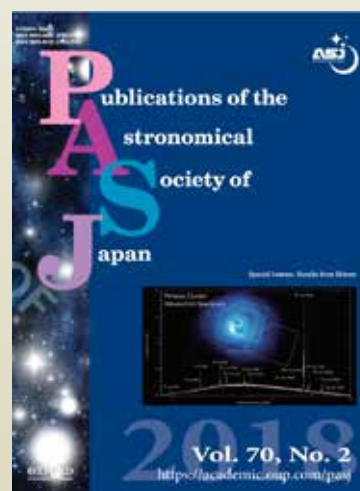
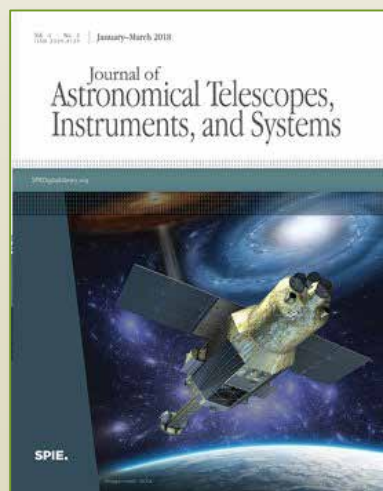
Continue.

Instruments papers (cont.)

| | Title | Author | Special Issue |
|------|--|-----------------|---------------|
| HXI | The hard x-ray imager onboard Hitomi (ASTRO-H) | K. Nakazawa | JATIS |
| ★ | In-orbit performance and calibration of the hard x-ray imager onboard Hitomi (ASTRO-H) | K. Hagino | JATIS |
| HXT | Supermirror design for Hard X-Ray Telescopes on-board Hitomi (ASTRO-H) | K. Tamura | JATIS |
| ★ | On-ground calibration of the Hitomi Hard X-ray Telescopes | H. Mori | JATIS |
| ★ | In orbit performance of the Hard X-ray Telescope (HXT) on board the Hitomi (ASTRO-H) satellite | H. Matsumoto | JATIS |
| CAMS | In-flight performance of the Canadian Astro-H Metrology System | L. Gallo et al. | JATIS |
| ★ | Design and performance of Soft Gamma-ray Detector onboard the Hitomi (ASTRO-H) satellite | H. Tajima | JATIS |
| ★ | Time assignment system and its performance aboard the Hitomi satellite | Y. Terada | JATIS |
| ★ | Astro-H/Hitomi data analysis, processing, and archive | L. Angelini | JATIS |

★ Calibration, performance

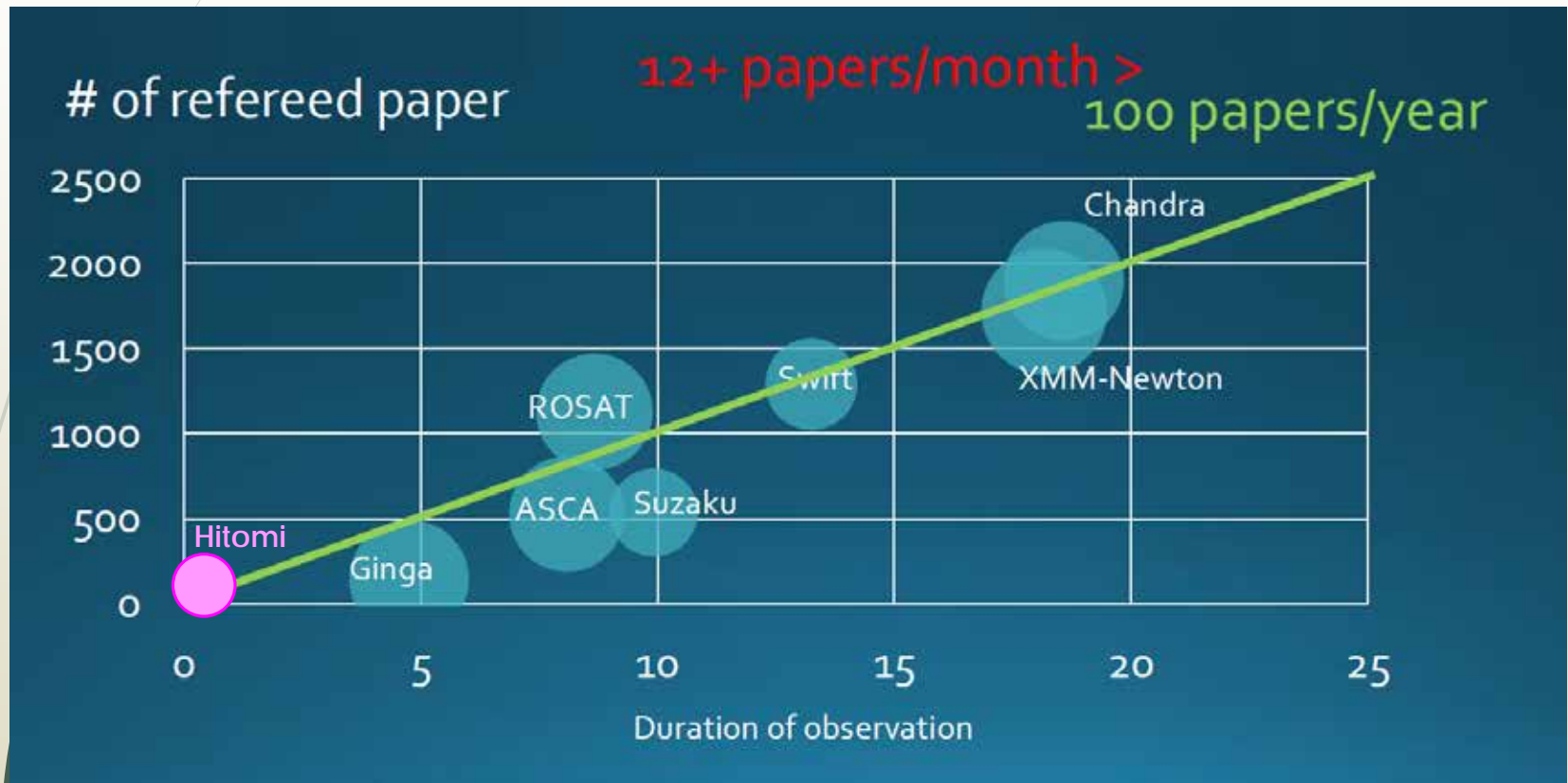
26 papers



JATIS Hitomi Special On-line!

Paper Productivity

© M.Tashiro, (from HEAPA meeting 2018)

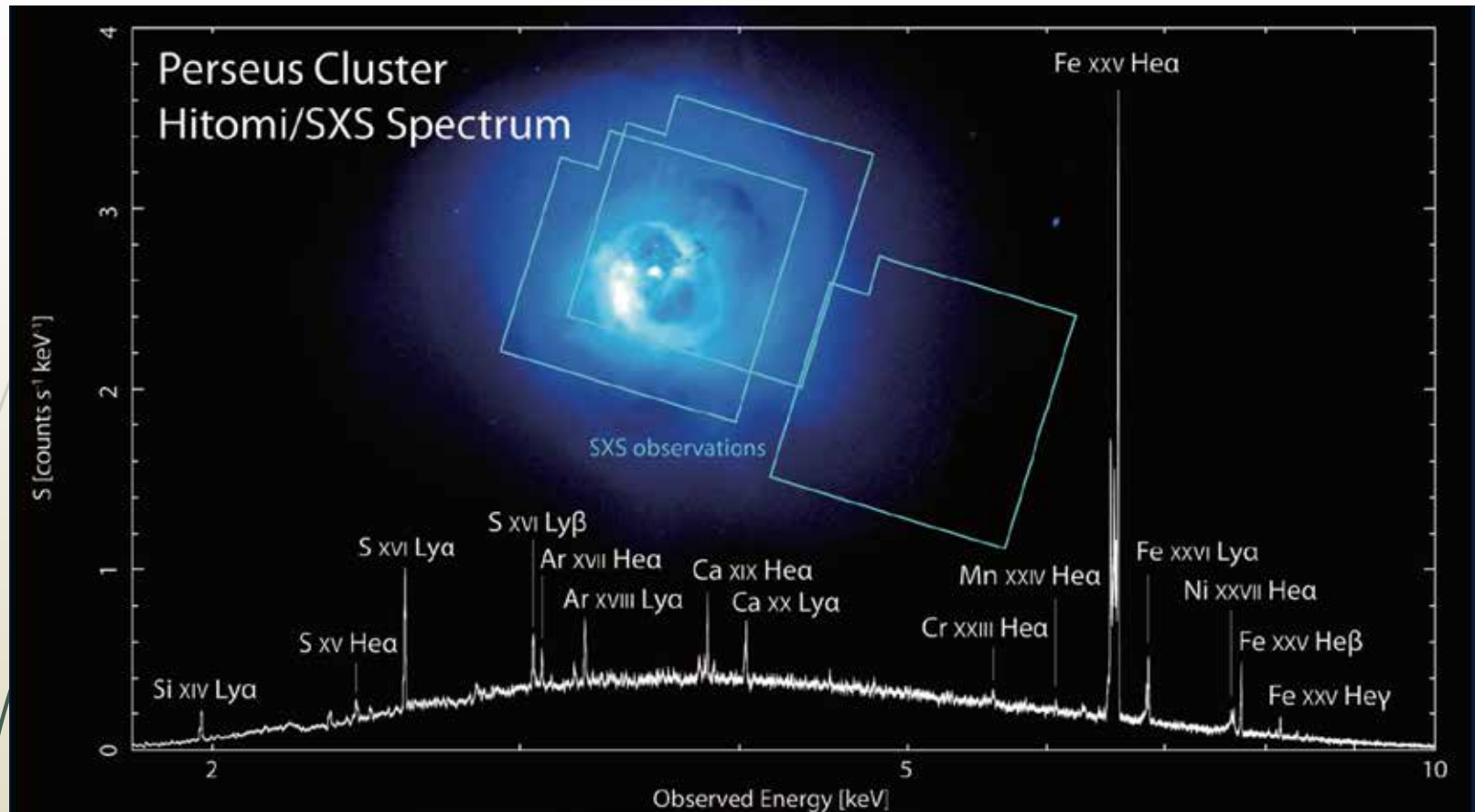


Hitomi after launch:

11 science papers / 1 month (operation) ~ 130 paper/year

Keeping this rate is grate!

Just we want to see ...



Challenge again,

X-ray Astronomy Recovery Mission (XARM)

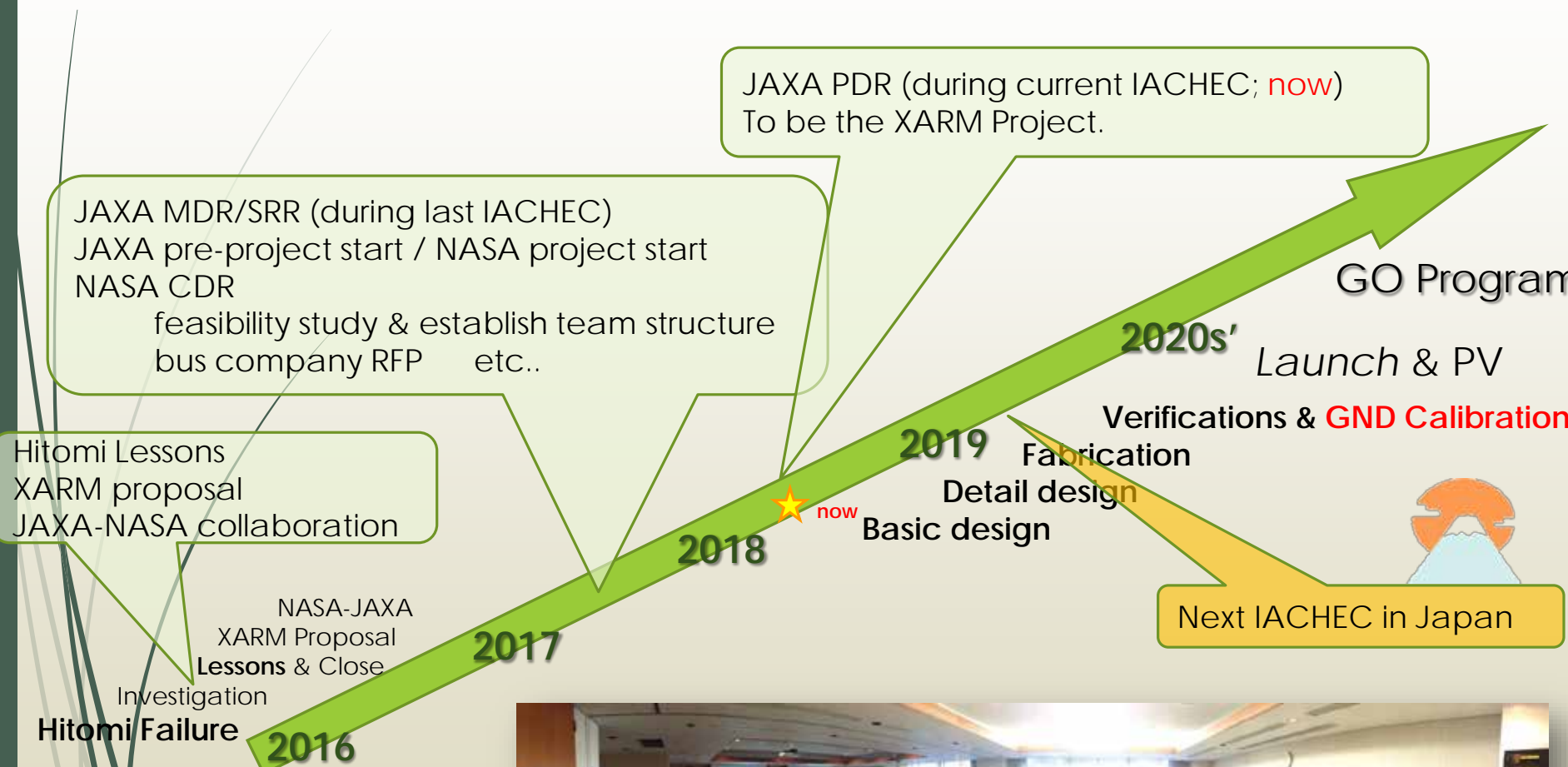
(& to be a Pilot mission for **Athena**)

XARM Instruments

| Instrument | FOV/pix | E (FWHM @6 keV) | Energy band |
|--|--------------------------|--------------------------------------|--------------|
| Resolve (XMA + X-ray micro calorimeter) | 2.9' / 6 x 6 pix | 7 eV (goal 5 eV) | 0.3 – 12 keV |
| Xtend (XMA + X-ray CCD) | 38' / 1280 x 1280 pix | < 250 eV at EOL (< 200 eV at BOL) | 0.4 – 13 keV |

Basically the same instruments as Hitomi,
but **no hard X-ray and soft gamma-ray instruments.**

XARM Mission status



XARM team Japan kick off meeting on Dec. 26, 2017

Team structure & Science Outputs

Team Structure

- Project management by PM, PI, PE
- Sub-systems:
 - Resolve team, Xtend team (**Instrument teams**)
 - Bus system, GND system, and Mission Operation team (MOT)
 - **Science Operations team (SOT)** + Science Management Office

Next slide

Notes on Calibration activity

- Responsibilities of calibration activities are on instrument teams.
- SOT supports the calibration activity by instrument team(s);
 - Review of calibration requirement and recommendation
 - Support actual ground measurements
 - Management of in-orbit calibration plan & execution.

Science Management Office

- Mission PI/co-PI + Instrument PI + science category leads
- “XARM Participating Scientists” selected by agencies (NASA/JAXA/ESA)
- *PV-phase Collaborating-Scientists-Program is planed.*

Science Operations Team

Goal

- Enhance science outputs, supporting science operations by PI.
- Non volunteer base, well defined tasks until the end of mission.

Scopes in 4 categories

- Data & software:
 - Daily operations of planning of scientific objects, telemetry monitoring, pipeline processing and archive
- Instruments & performance:
 - Support **calibration activities** by instrument teams, management of in-orbit calibration plan & execution
 - **Performance enhancement studies**,
 - Systematic studies of behavior of instrument to have better analyses tools/methods required for scientific analyses
 - Simulation studies & behavior studies of instruments etc.
 - (instrumental studies what you in this room are doing now.)
- GO : User support
 - Proposal support/Operation support/Analyses support
 - Education of young astronomers, etc

Performance/calibration requirement

XARM Scientific Objectives

1. Structure formation of the Universe and evolution of clusters
2. Circulation history of baryonic matters in the Universe
3. Transport and circulation of energy in the Universe
4. New science with unprecedented high resolution X-ray spectroscopy



Flow down

Mission Requirements + Success criteria



Flow down

Design of Spacecraft and Instruments

- ü Design parameters of Resolve/Xtend
- ü Calibration requirements

| High resolution X-ray spectrometer | | | | | | X-ray imaging spectrometer | | | | | | | | | |
|---|--|--|--|--|---|---|---|---|--|---|--|--|---|---|---------|
| Angular resolution HPD ≤ 1.7 arcmin | Effective area $\geq 210 \text{ cm}^2 @ 6 \text{ keV}$ $\geq 160 \text{ cm}^2 @ 1 \text{ keV}$ | Field of view $\geq 2.9 \text{ arcmin} \times 2.9 \text{ arcmin}$ | Energy range of spectroscopy 0.3 - 12 keV | Absolute energy scale $\leq 2 \text{ eV}$ | Energy resolution $\leq 7 \text{ eV FWHM @ 6 keV}$ | Non X-ray background $\leq 2 \times 10^{-3} \text{ c/s/keV/array}$ | Highest photon rate per array to handle $\geq 150 \text{ c/s/array}$ | Absolute time tagging accuracy $\leq 1 \text{ ns}$ | Angular resolution HPD ≤ 1.7 arcmin pixel size $\leq 100 \mu\text{m}$ | Effective area $\geq 300 \text{ cm}^2 @ 6 \text{ keV}$ | Energy range of spectroscopy 0.4 - 13 keV | Field of view $\geq 30 \text{ arcmin} \times 30 \text{ arcmin}$ | Energy resolution $\leq 250 \text{ eV FWHM @ 6 keV}$ | Non X-ray background $\leq 1 \times 10^{-4} \text{ c/s/keV/arcmin}^2/\text{cm}^2$ (continuum component at 5 - 10 keV) | ASSISTs |



Flow down

Development Plan, Operation plan, and Calibration plan etc...

Resolve Calibration Plan



Hitomi SXS Calibration results

1. “Ground calibration of the Astro-H (Hitomi) soft x-ray spectrometer”, Eckart+, 2018, JATIS, in press
2. “In-flight verification of the calibration and performance of the ASTRO-H (Hitomi) Soft X-ray Spectrometer”, Leutenegger+, 2018, JATIS, in press
3. “In-flight Calibration of Hitomi Soft X-ray Spectrometer (1) Background”, Kilbourne+, 2018, PASJ, in press
4. “In-flight Calibration of Hitomi Soft X-ray Spectrometer (2) PSF”, Maeda+, 2018, PASJ, in press
5. “In-flight Calibration of Hitomi Soft X-ray Spectrometer (3) Effective Area”, Tsujimoto+, 2018, PASJ, in press
6. “Atomic data and spectral modeling constraints from high-resolution X-ray observations of the Perseus cluster with Hitomi”, Hitomi collaboration, 2018, PASJ, in press
7. “The Time Assignment System and Its Performance aboard the Hitomi Satellite”, Terada+, 2018, PASJ, in press

Some more in prep. Very useful for future X-ray u-cal in-flight cal.

We appreciate IACHEC colleagues for timing cross-cal campaign and your calibration results (we cited many of your work!).

Resolve Calibration Plan



Hitomi SXS In-flight calibration Highlights

| Cal item | Gain | E | Abs timing |
|----------|--|---|---|
| | <p>Flux (arbitrary unit)</p> <p>Energy (keV)</p> <p>Fe XXIV 2p-1s,2p²</p> <p>Fe XXIV 2s-1s,2s,2p² 2p-1s,2p²</p> <p>Fe XXIII 1s-2p²</p> <p>Fe XXIV 3p-1s,2p,3p 3d-1s,2p,3d</p> <p>SPEX v3.03 AtomDB v3.0.8 CHIANTI v8.0</p> <p>z y x w</p> <p>(a)</p> <p>Relative residual</p> <p>(b)</p> | <p>FWHM = 4.94 eV</p> <p>Counts/0.5 eV bin</p> <p>Energy (eV)</p> | <p>Counts/0.5 eV bin</p> <p>Phase</p> <p>TYPE : SXS Peak phase: 0.944eC1 +/- 2.000e-04</p> <p>TYPE : AXI Peak phase: 0.812eC3 +/- 1.302e-03</p> <p>TYPE : RAT Peak phase: 0.917eC3 +/- 1.080e-03</p> <p>TYPE : SGR Peak phase: 0.982eC3 +/- 1.000e-04</p> |
| Data | Perseus | Onboard ⁵⁵ Fe | Crab |
| Req | < 2 eV | < 1 eV | < 10 ms |
| Achieved | < 2 eV | << 1 eV(TBC) | 100 us |
| Ref | [6] | [2] | [2,7] |

Resolve Calibration Plan



XARM Resolve Calibration Plan

Basically the same with SXS with some refinements based on SXS.

- Gate valve will be calibrated similarly to other flight components: In SXS, GV planned to be opened before PV, but all science obs made before that, leaving the largest uncertainty.
- Timing cal req'd and planned better. micro-calorimeter is also good at timing.
- Unified & Int'l team of detector, mirror, and MXS. They were separate in SXS. After all, cal can only be complete by evaluating end-to-end.
- Better coordination with ray-tracing, plasma code groups. We found that they are the ultimate limit for some calibration items.

Status

- Cal requirements ... Issued.
- Cal plan (on ground) ... issued & being executed. Some results ready by the next IACHEC meeting.
- Cal plan (in orbit) ... to be documented. Looking forward to discussing with IACHEC colleagues for cross-cal.

Xtend Calibration Plan

On-ground

XMA calibration plan is included in Resolve.

| Items | |
|--|--|
| Selection of FM CCDs | four CCD chips |
| Operation parameters | Microcode, bias voltage, ASIC gain, threshold, etc |
| Calibration parameters | Charge transfer efficiency, gain, energy resolution, etc |
| Performance Dependency on Radiation damage | Demonstration of radiation tolerance with proton radiation experiments |

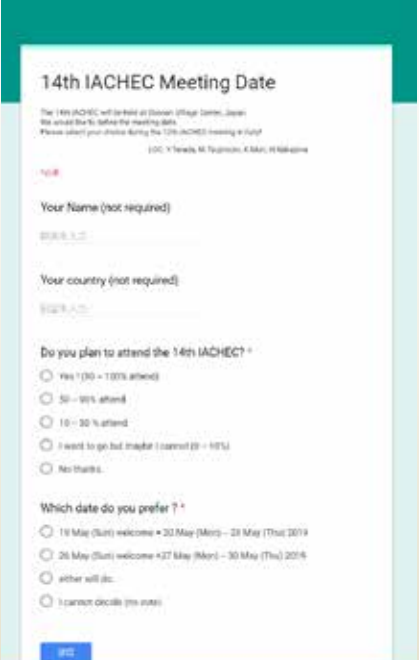
CALDB files for the initial phase in orbit will be prepared before launch.

In-orbit (steady operation)

- Calibration data will be taken from Fe-55 (5.9 keV line) and celestial objects (E0102-72, RX J1856, Perseus, TBD).
- Regularly update CALDB files.

Thank you

Please come to Japan in the 14th IACHEC in 2019!
Voting for the date is on going.



14th IACHEC Meeting Date

The 14th IACHEC will be held at Edoan Village Center, Japan.
We would like to gather the meeting date.
Please select your choice during the 13th IACHEC meeting in Italy!
JOC: Y. Tanaka, M. Tsujimoto, K. Mori, H. Nakamura

Your Name (not required)

Your country (not required)

Do you plan to attend the 14th IACHEC? *

- Yes (30 – 100% attend)
- 50 – 90% attend
- 10 – 50 % attend
- I want to go but maybe I cannot (0 – 90%)
- No thanks.

Which date do you prefer? *

- 18 May (Sat) welcome – 20 May (Mon) – 21 May (Tue) 2019
- 26 May (Sat) welcome + 27 May (Mon) – 30 May (Thu) 2019
- either will do.
- I cannot decide (no vote)

[\[Go\]](#)

