

Measurement of X-ray, UV, optical transmission of Contamination Blocking Filter on board Hitomi SXI

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14th Workshop of International Astronomical Consortium for High Energy Calibration

20- 23, May 2019, Shonan Village Center

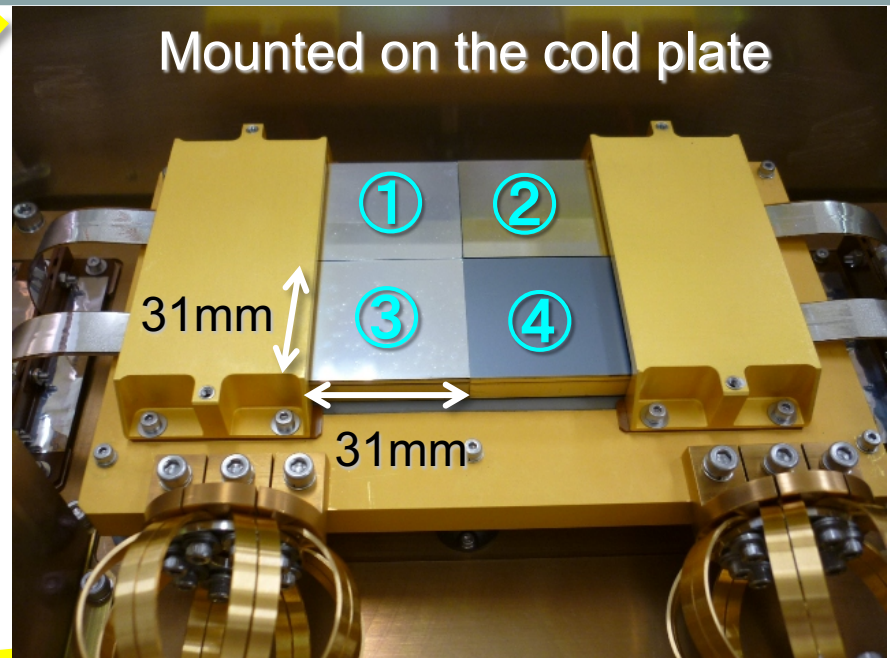
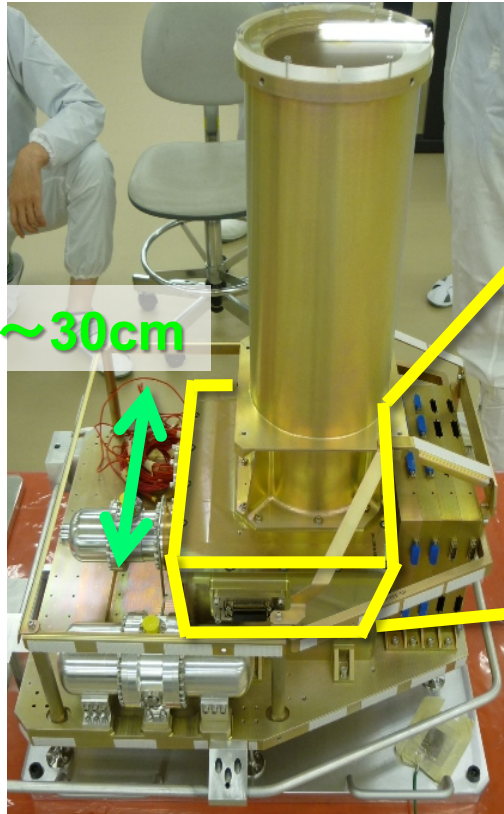


Outline of this talk

- 1.Hitomi SXI
- 2.Contamination Blocking Filter (CBF)
- 3.Optical transmission of CBF
- 4.X-ray and UV transmission of CBF
- 5.Discussion
- 6.Conclusion

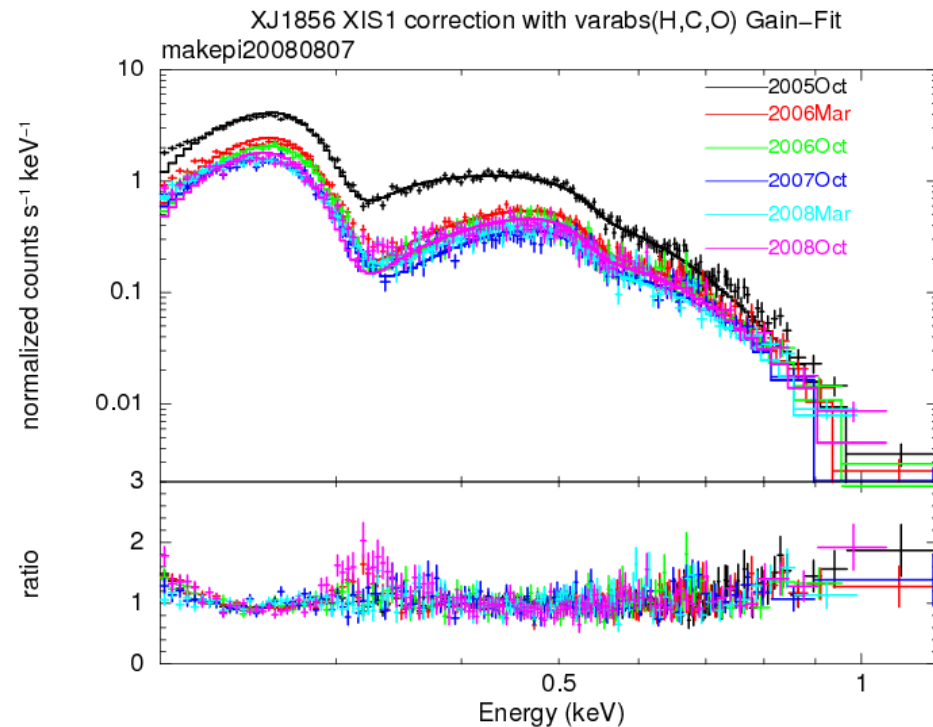
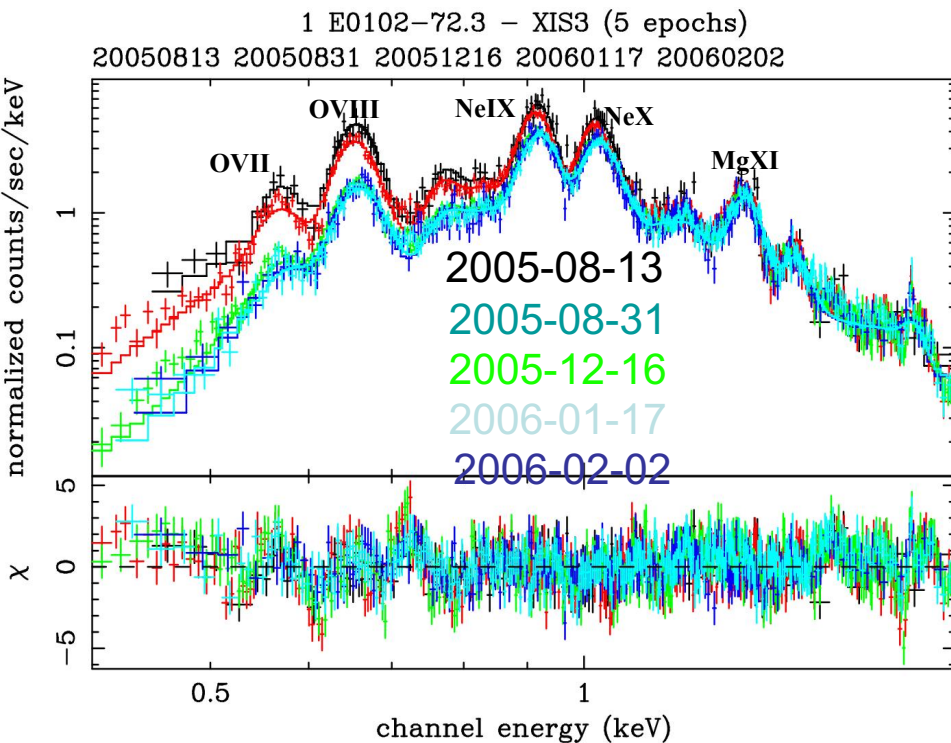


1. Hitomi SXI



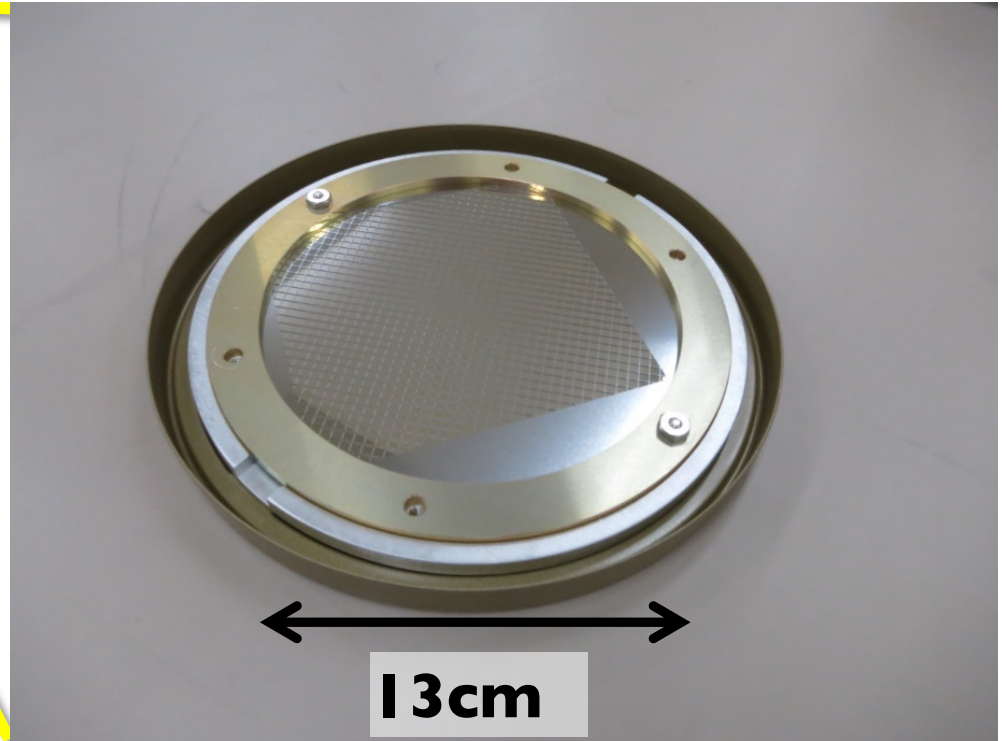
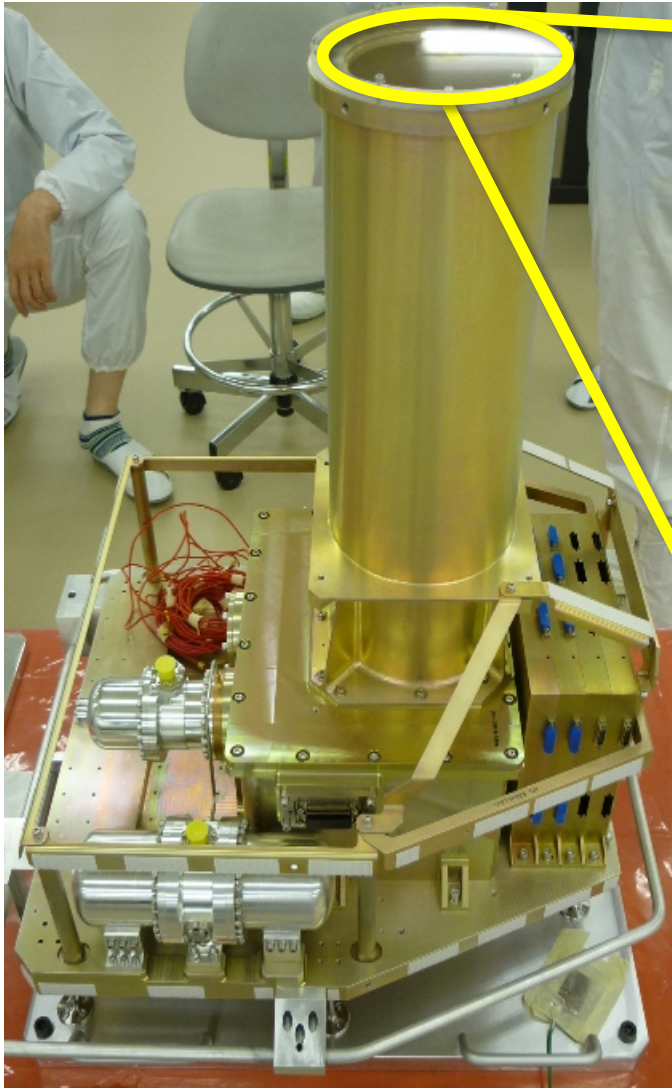
- P-channel CCD manufactured by Hamamatsu Photonics
- BI with depletion layer of $200\mu\text{m}$
- $24\mu\text{m}$ square pixel
- $\sim 62\text{mm} \times 62\text{mm}$

2. Contamination Blocking Filter



- From our Suzaku XIS experiment in space, we noticed the degradation of QE due to the contamination of material in the satellite.

2. Contamination Blocking Filter

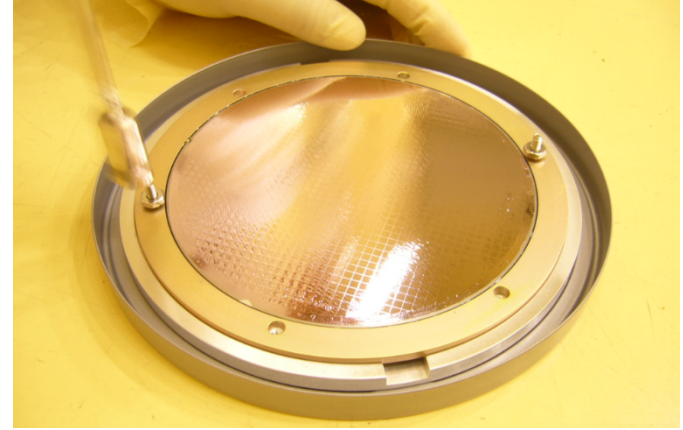


CBF (**C**ontamination **B**locking **F**ilter)
was made by Luxel CLDT
Al ; 80nm and 40nm
Polyimide ; 200nm

2. Contamination Blocking Filter

【 EUV shielding 】

- **CBF** : **C**ontamination **B**locking **F**ilter
- Pre-FM (before 2015)
200nm thick **Polyimide** was covered with 30nm thick **Al** layers.



【 Optical Blocking 】

- **OBL** : **O**ptical **B**locking **L**ayer
- 100nm thick **Al** is directly coated over the BI-CCD wafer manufactured by Hamamatsu.

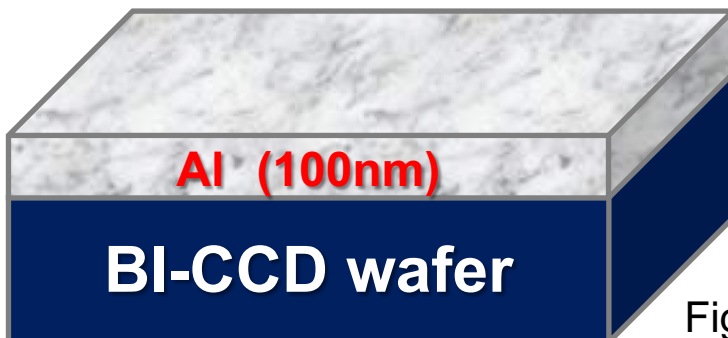
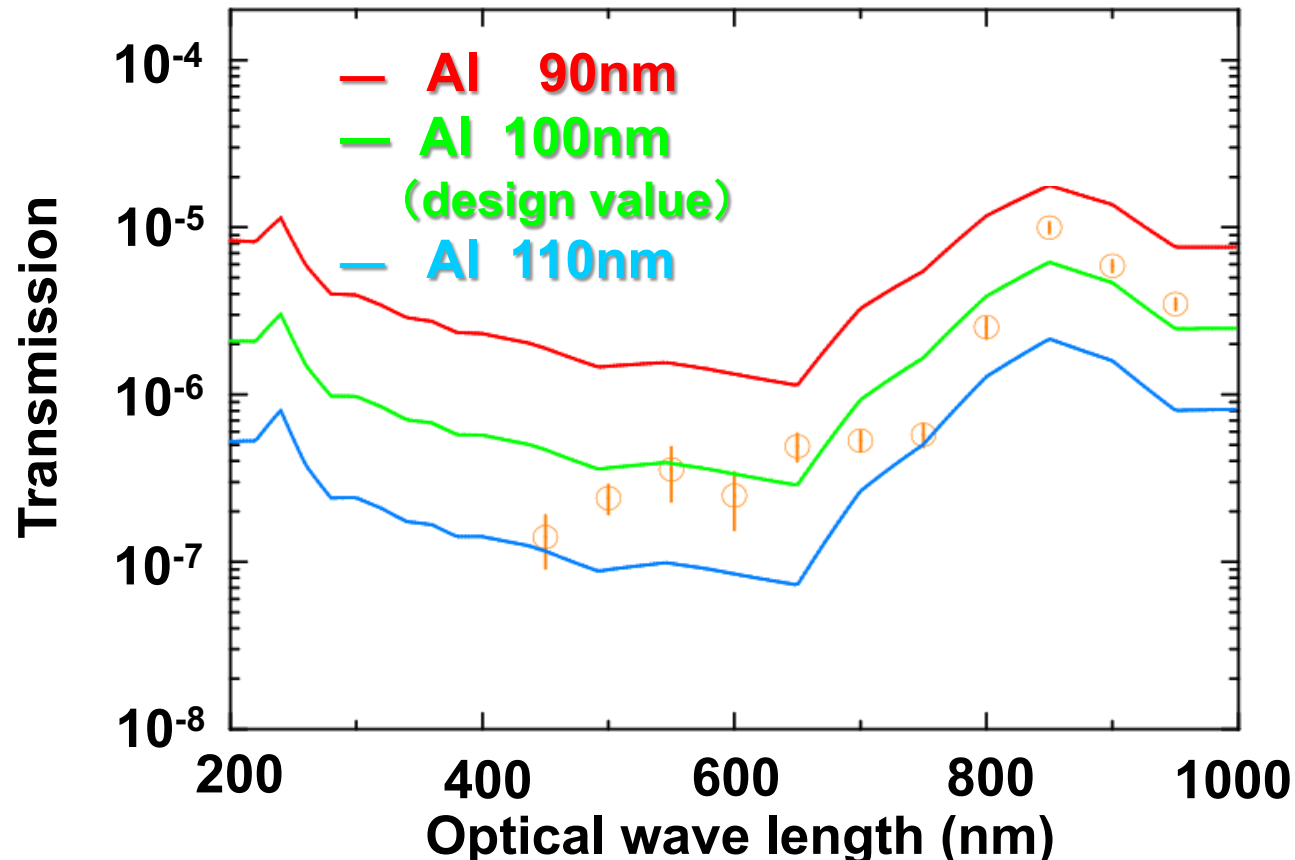


Figure. Schematic view of OBL

2. Contamination Blocking Filter

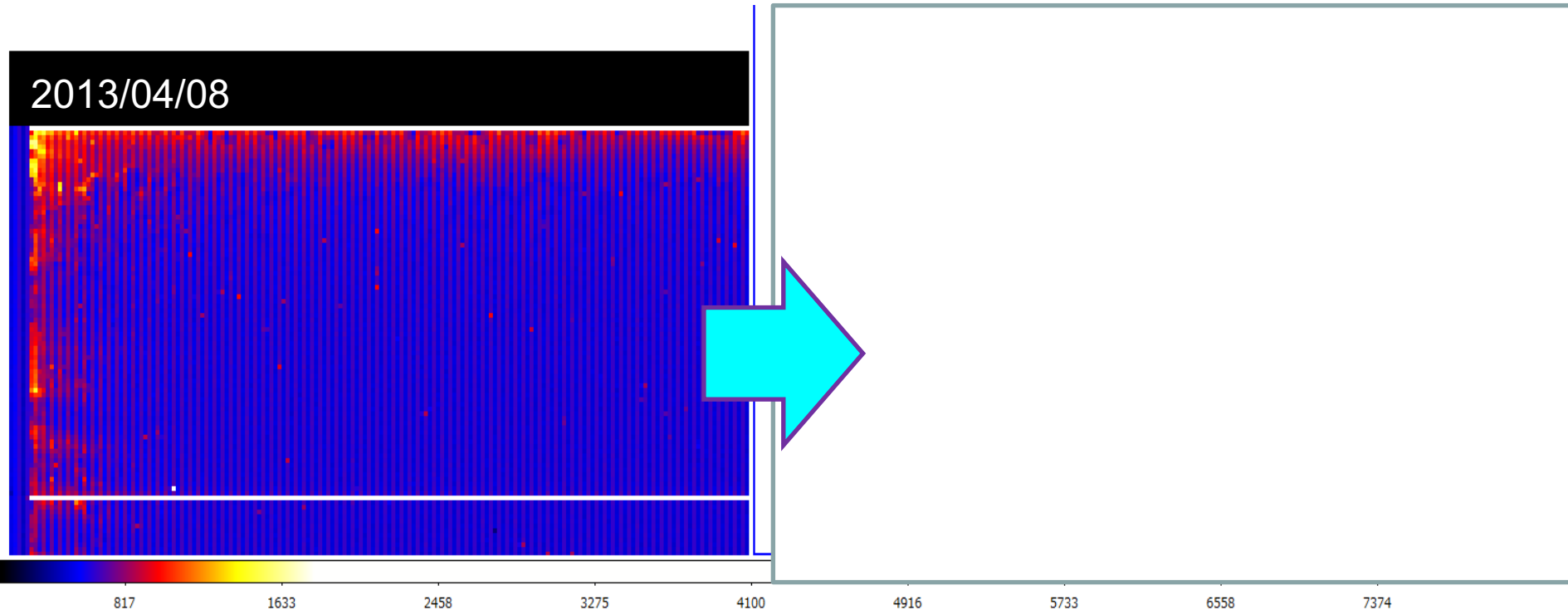
Optical transmission of OBL



- We found the optical transmission was $< \sim 10^{-5}$ which was equivalent value that we had expected from the design thickness of OBL.

2. Contamination Blocking Filter

Pinholes in Al Coating of **OBL** for Hitomi Satellite

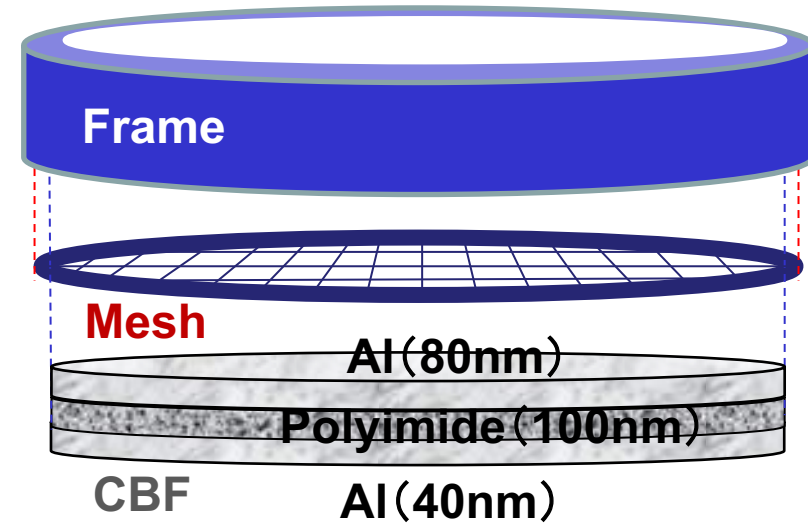


- We had checked the pin holes on OBL by irradiating the optical light.
- But, the number of pin holes was increasing during five months.
- We decided to make the thickness of CBF thicker to block the optical light.

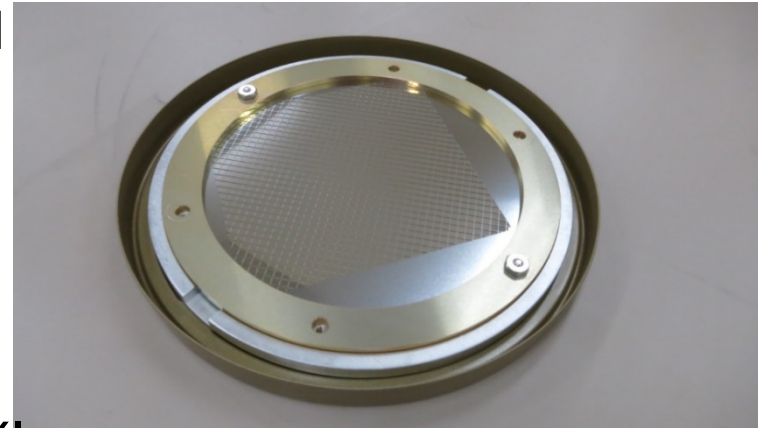
2. Contamination Blocking Filter

FM-CBF (after 2015)

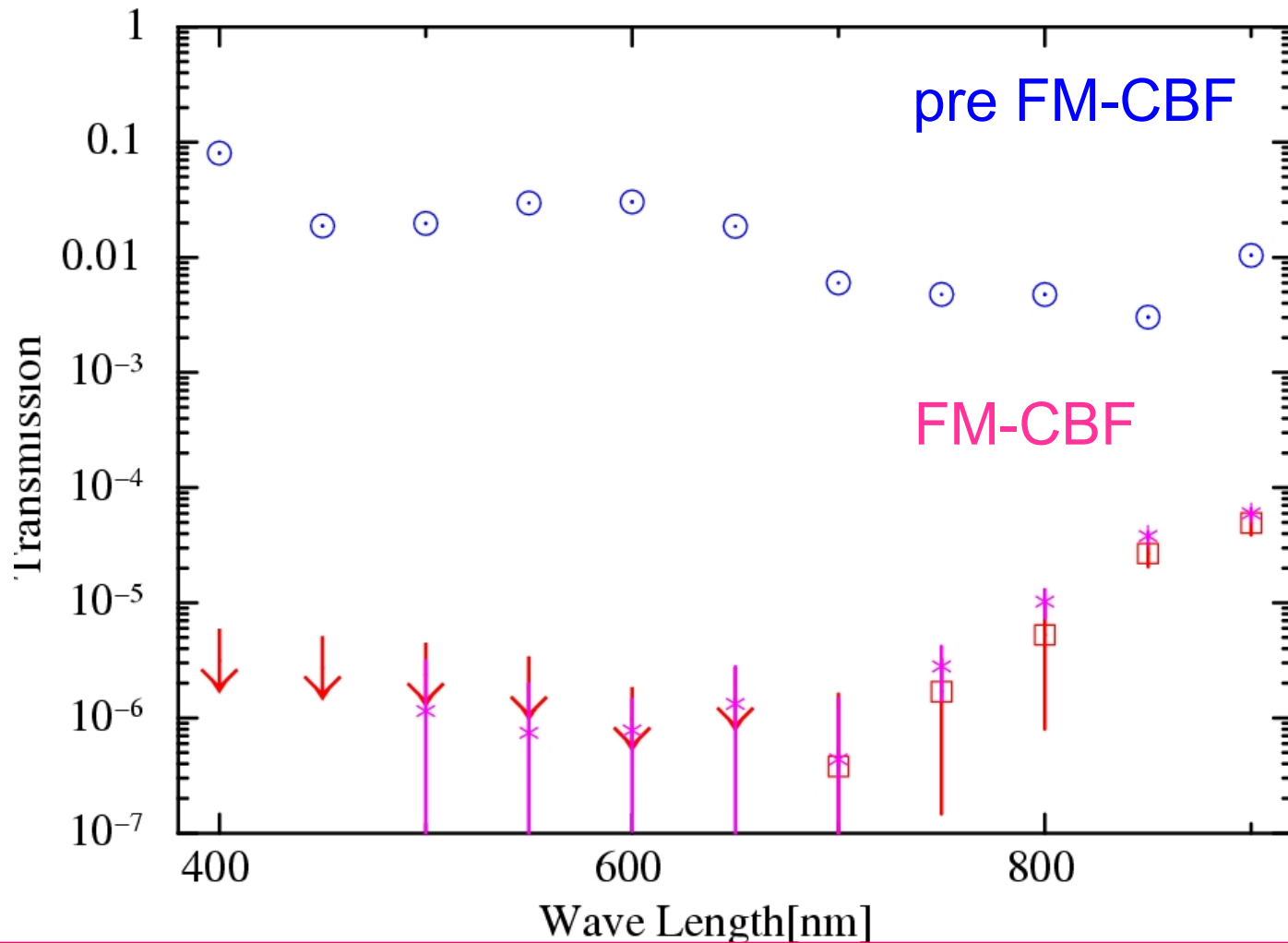
| | Pre-FM-CBF | FM-CBF[nm] |
|-----------|------------|------------|
| Al | 30 | 80 |
| Polyimide | 200 | 200 |
| Al | | 40 |



- CBF was supported by the mesh with guard to block solar X-ray scattered middle plate in the satellite.
- The thickness of mesh was 0.3mm, and its interval and width of mesh was 3mm and 0.15mm, respectively.
- This guard did not cut the field of view of SXI.



3. Optical Transmission of CBF



- We confirmed the optical transmission was $< \sim 5 \times 10^{-5}$ which was expected from design thickness of new CBF.

4. X-ray and UV Transmission of CBF

- We measured Soft X-ray and UV transmission of FM-CBF at the synchrotron facility in Japan (KEK Photon Factory).

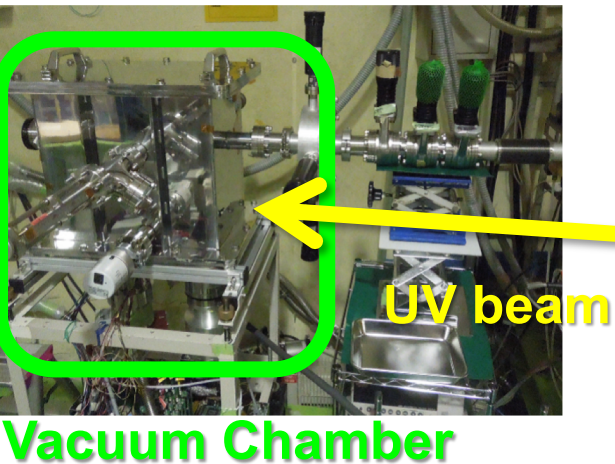


Fig: BL-20A (20-60eV)

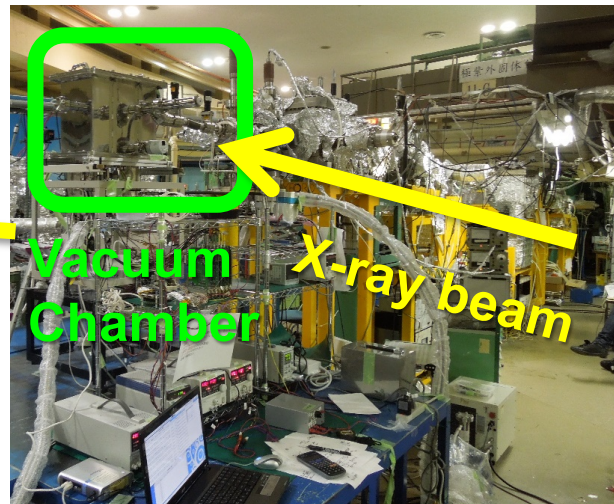


Fig: BL-11A (150-1800eV)

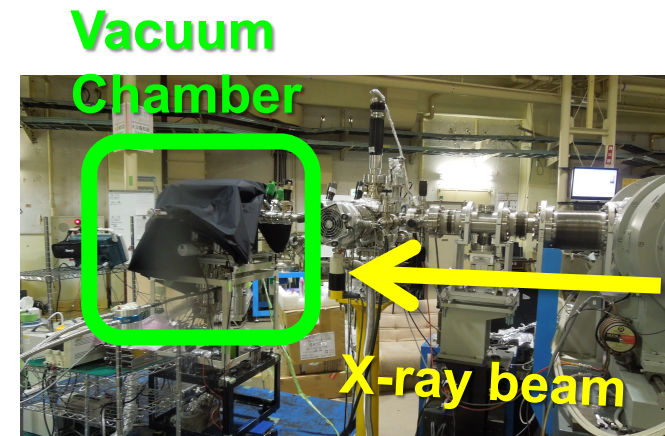
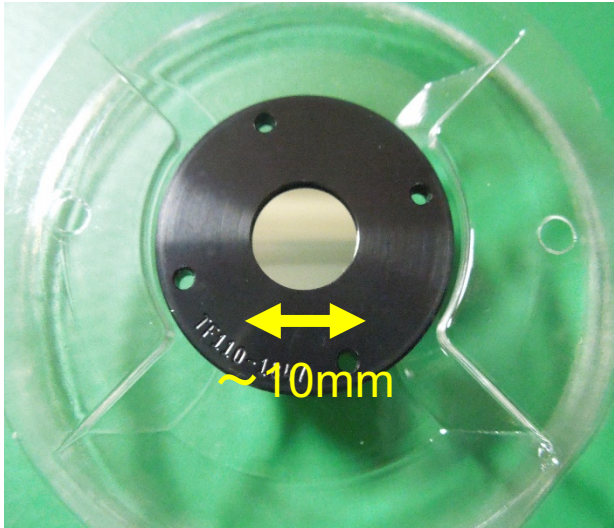


Fig: BL-11B (1730-5000eV)

We measured the X-ray (UV) flux using photo diode with/without CBF, and calculated the X-ray (UV) transmission of CBF derived from the flux ratio between two them.

4. X-ray and UV Transmission of CBF

Witness Filter



Experimental Set up

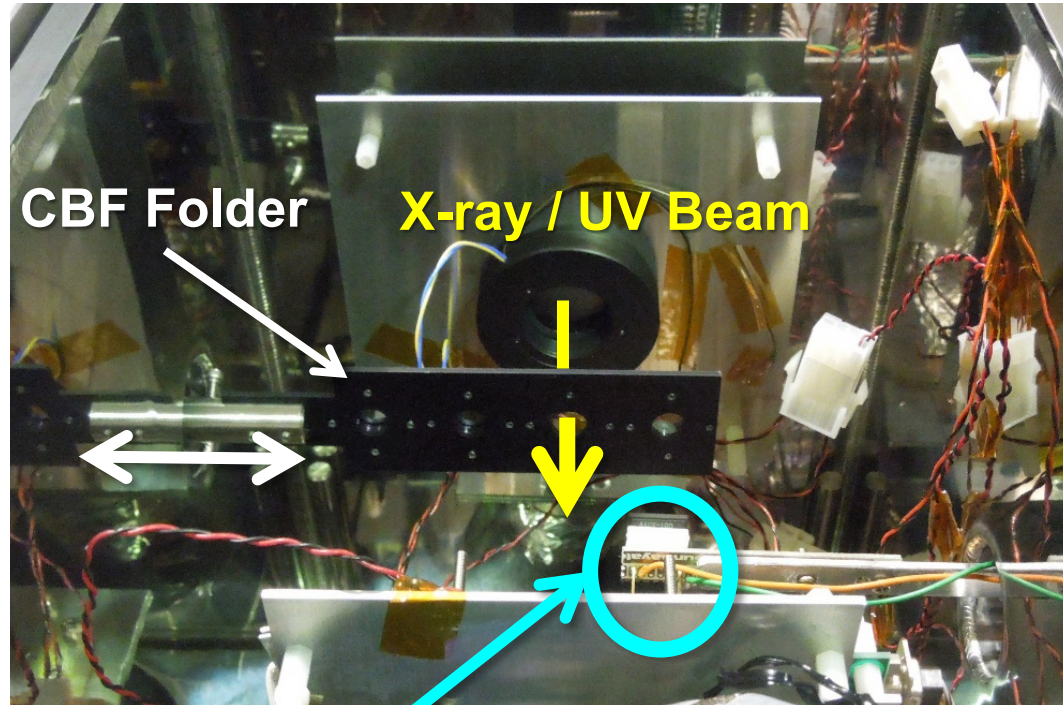
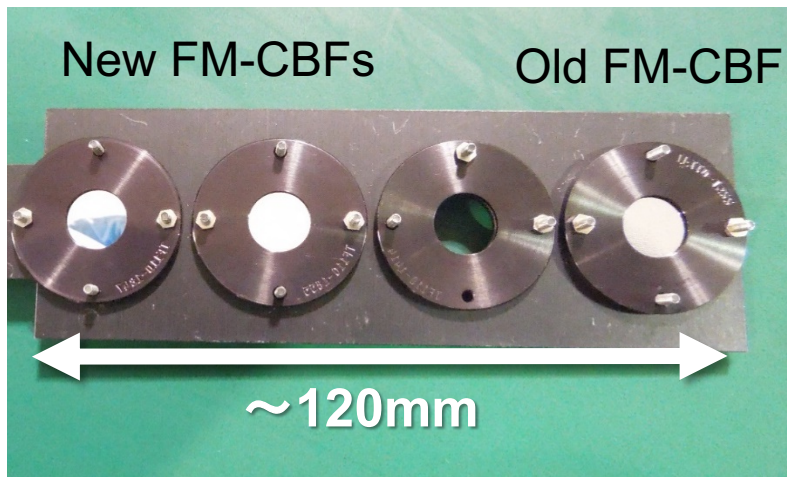


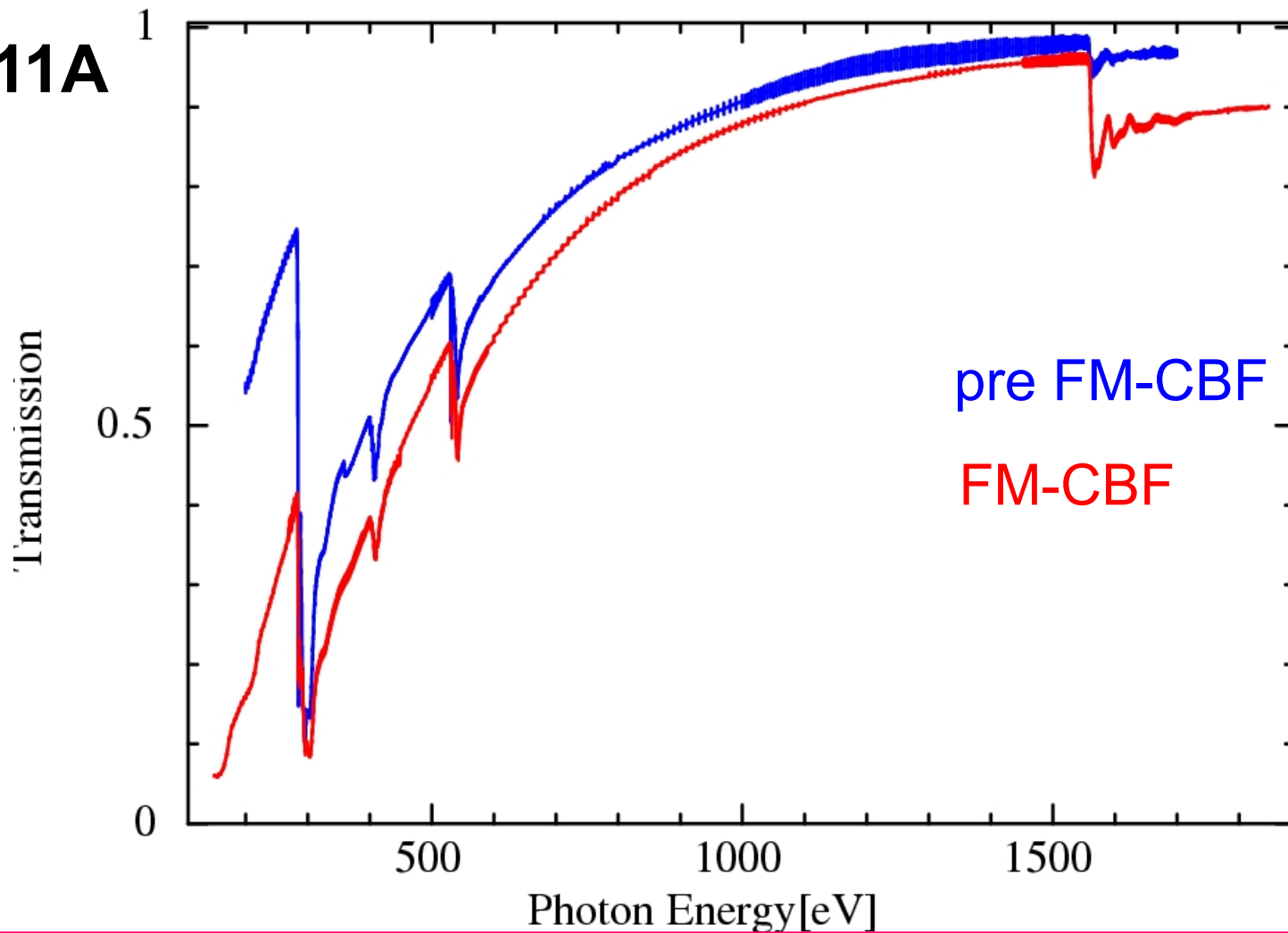
Photo Diode

CBF folder



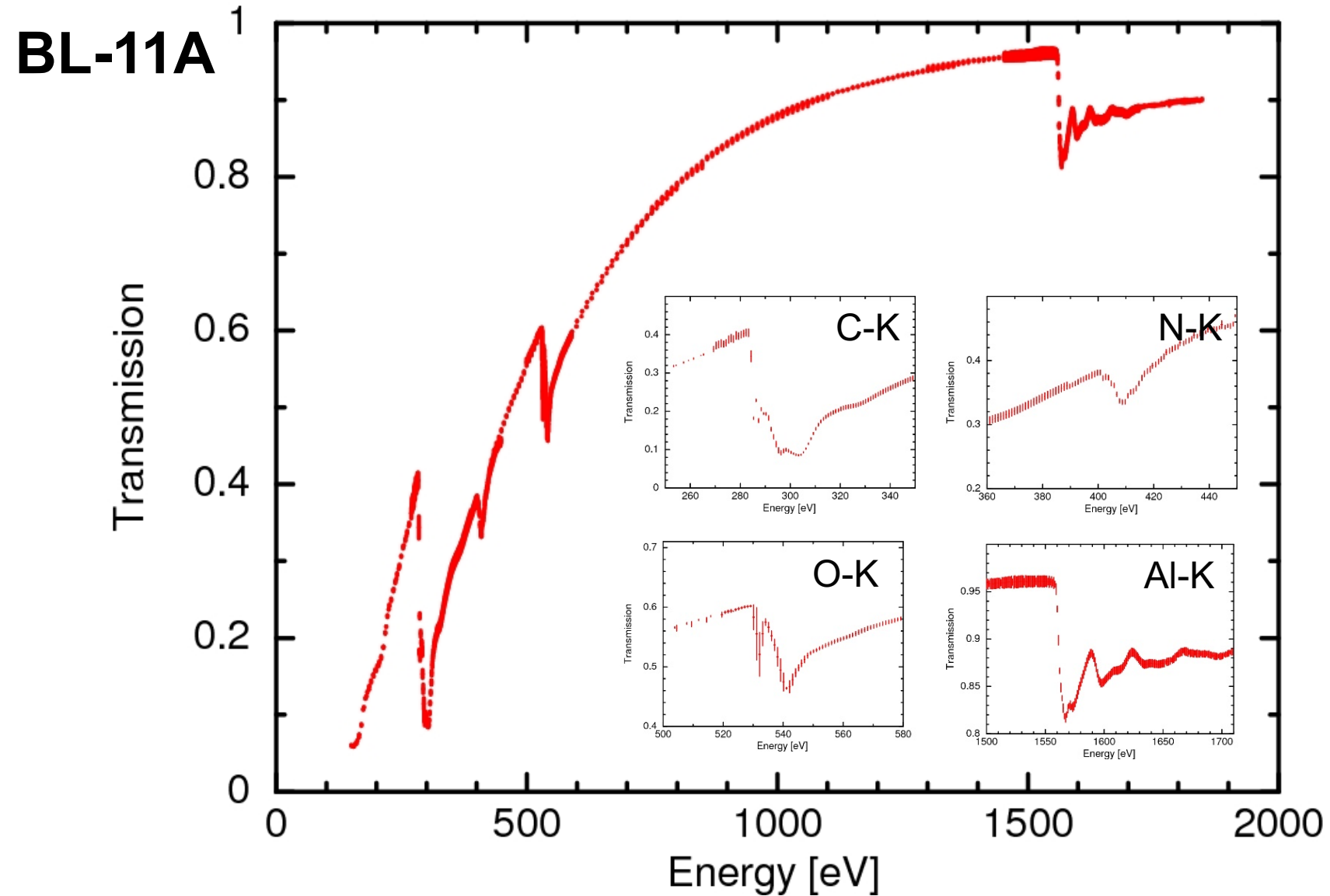
4. X-ray and UV Transmission of CBF

BL-11A



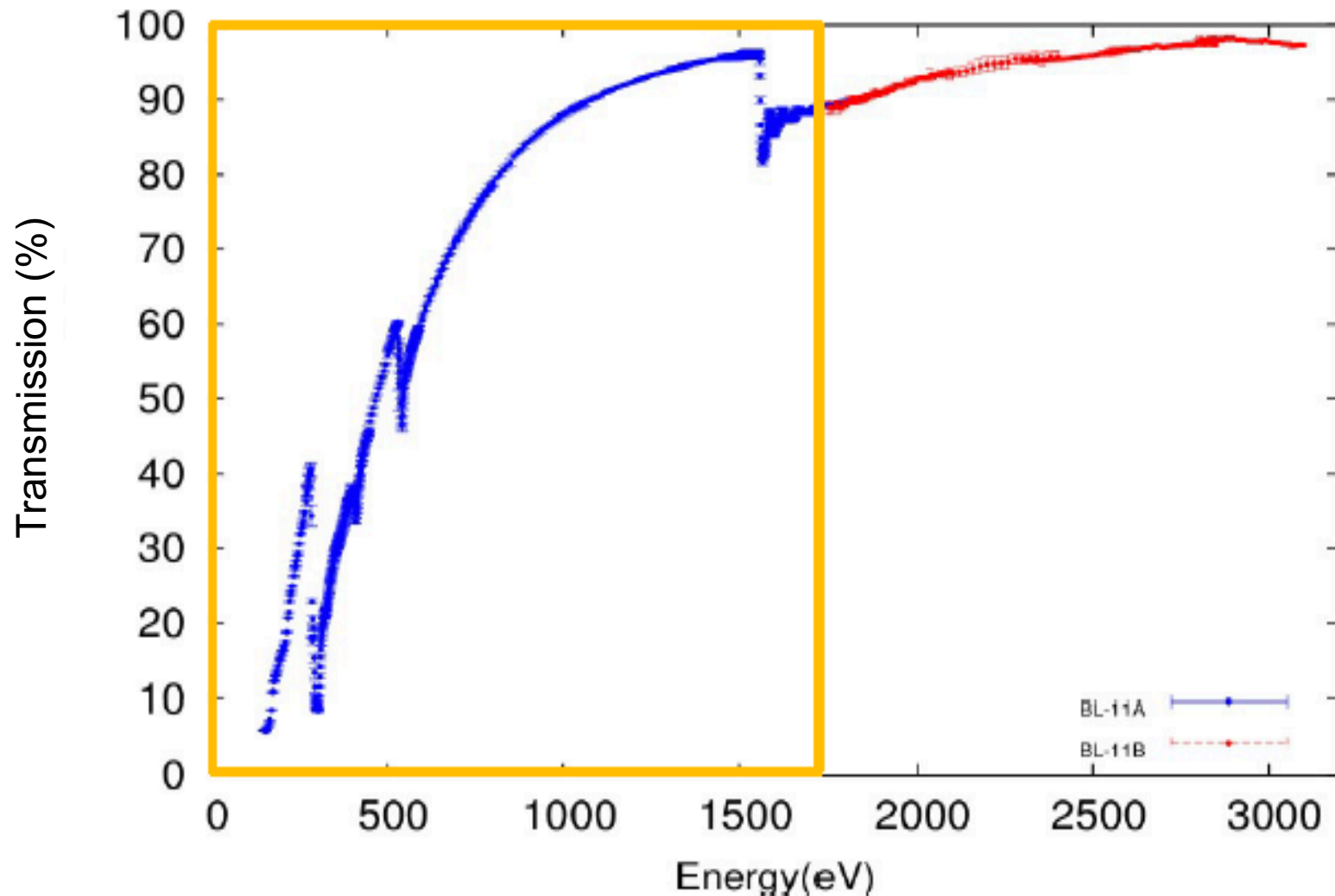
- The X-ray transmission of new FM CBF was lower than old CBF by $\sim 10\%$ at O-K due to the thicker Al.
- The X-ray energy was calibrated using absorption edges of old FM-CBF as a reference

4. X-ray and UV Transmission of CBF



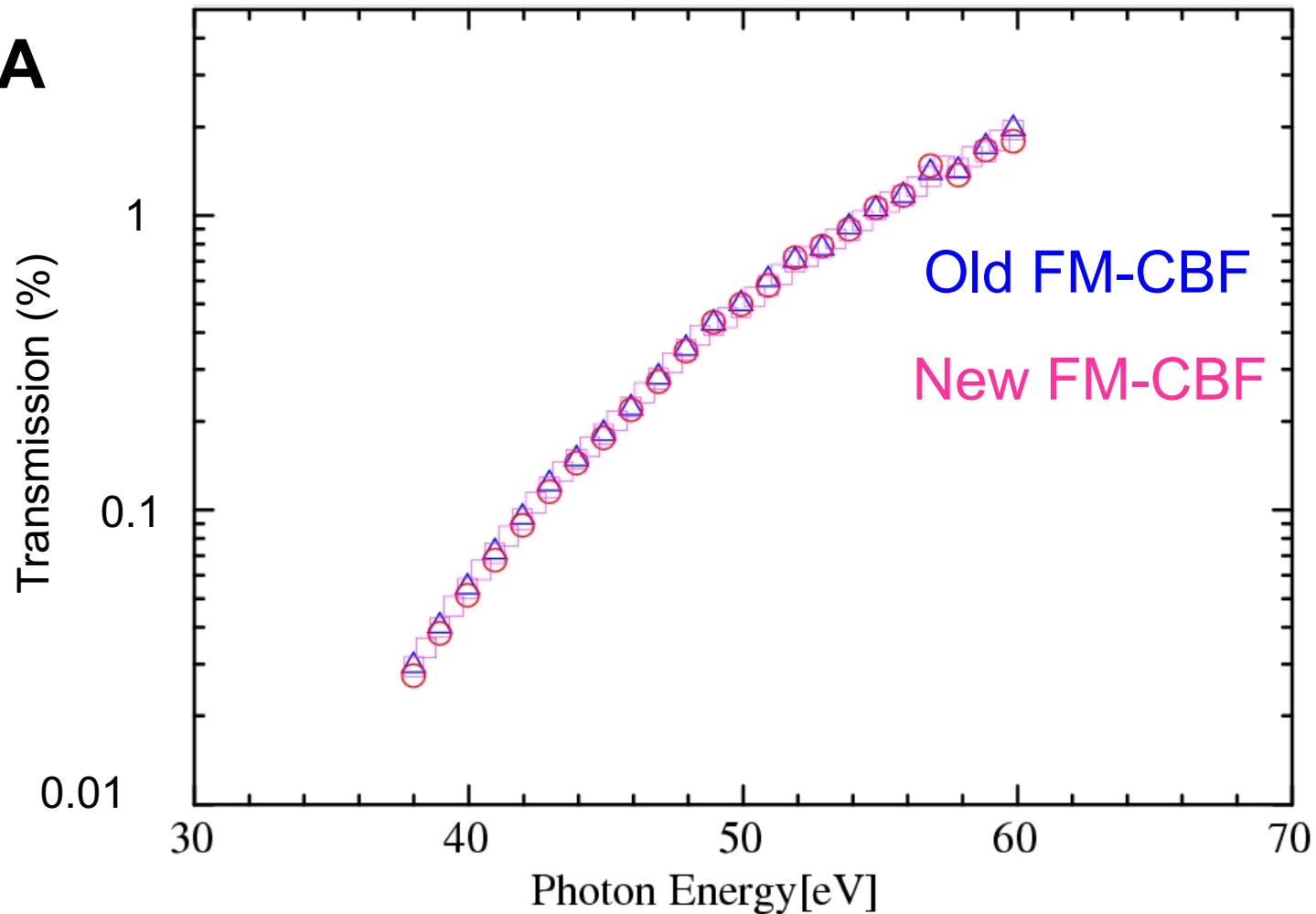
4. X-ray and UV Transmission of CBF

BL-11A+BL11B



■ ■ ■ 4. X-ray and UV Transmission of CBF

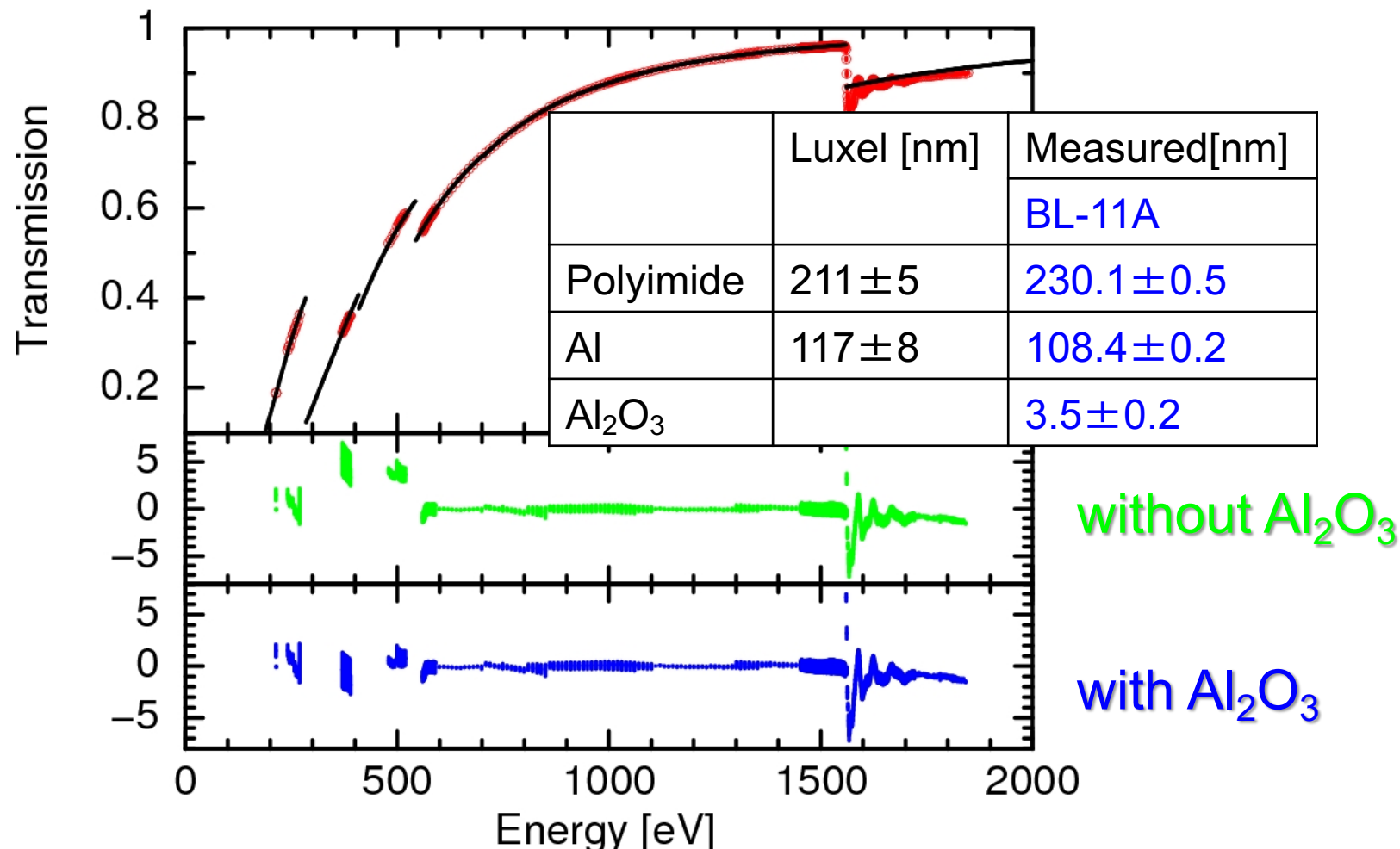
BL-20A



The UV transmission of new FM CBFs was same as its old ones because the thickness of polyimide was not changed.

4. Discussion

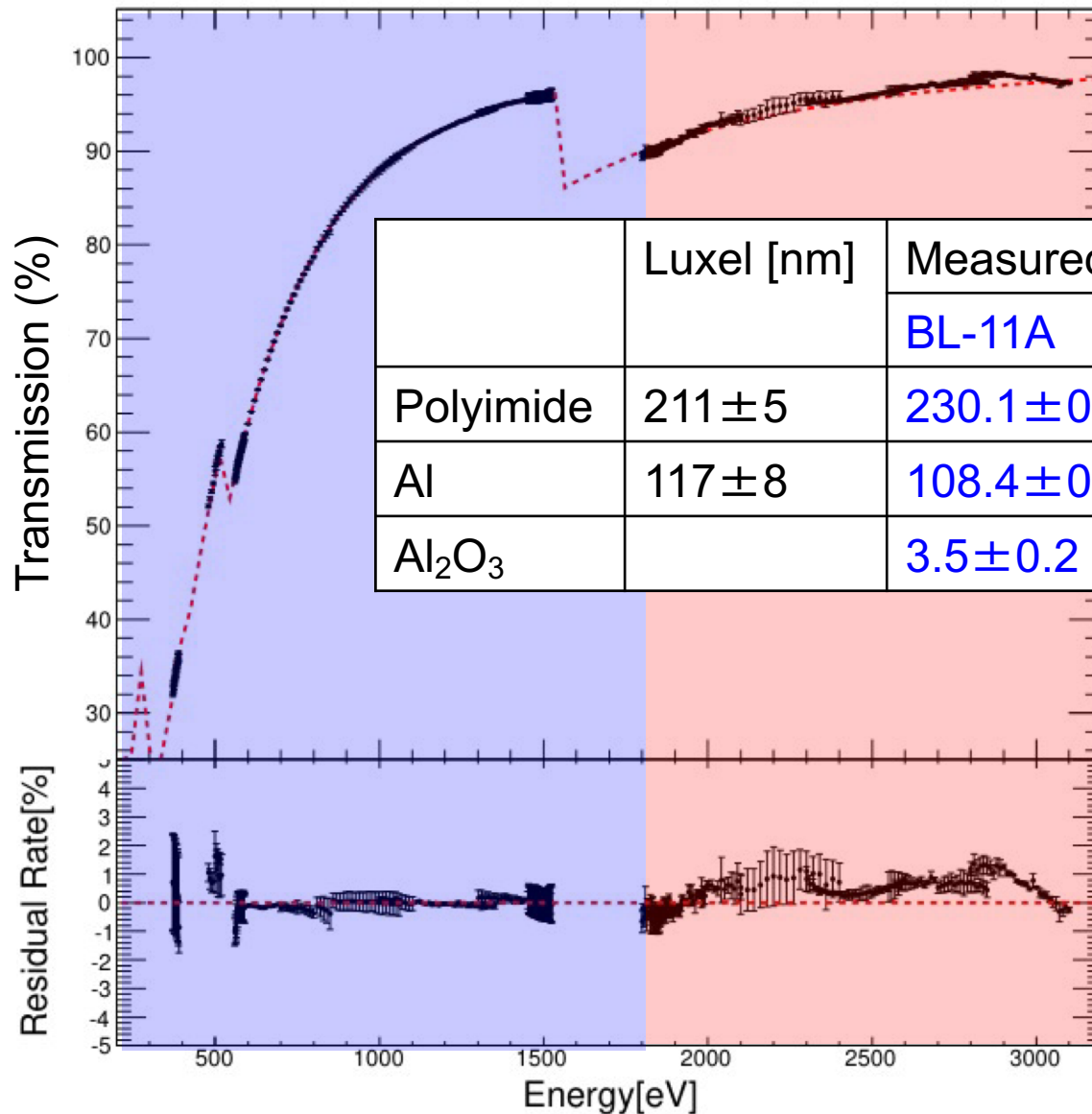
BL-11A



The thickness of polyimide and Al were slightly thicker and thinner value than their reported from Luxel Co LTD.

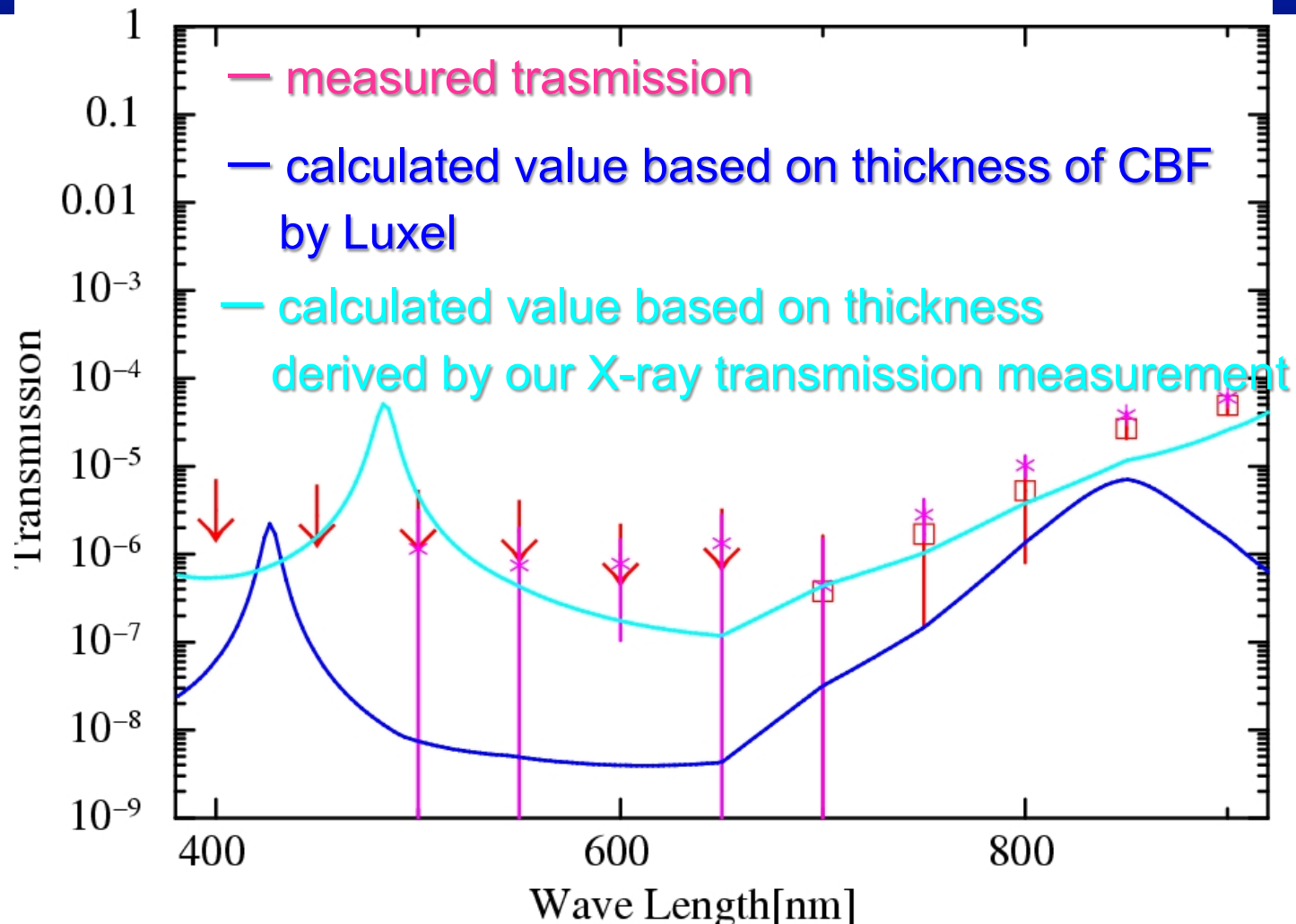
4. Discussion

BL-11A
+
BL-11B



| | Luxel [nm] | Measured[nm] | |
|--------------------------------|-------------|-----------------|-----------------|
| | | BL-11A | BL11A+BL11B |
| Polyimide | 211 ± 5 | 230.1 ± 0.5 | 231.7 ± 0.5 |
| Al | 117 ± 8 | 108.4 ± 0.2 | 107.6 ± 0.2 |
| Al ₂ O ₃ | | 3.5 ± 0.2 | 3.2 ± 0.2 |

4. Discussion



- The optical transmission was higher than what we expected from Al thickness reported from Luxel Co LTD.
- The thickness Al was thinner which we derived from both X-ray and optical transmission measurement.

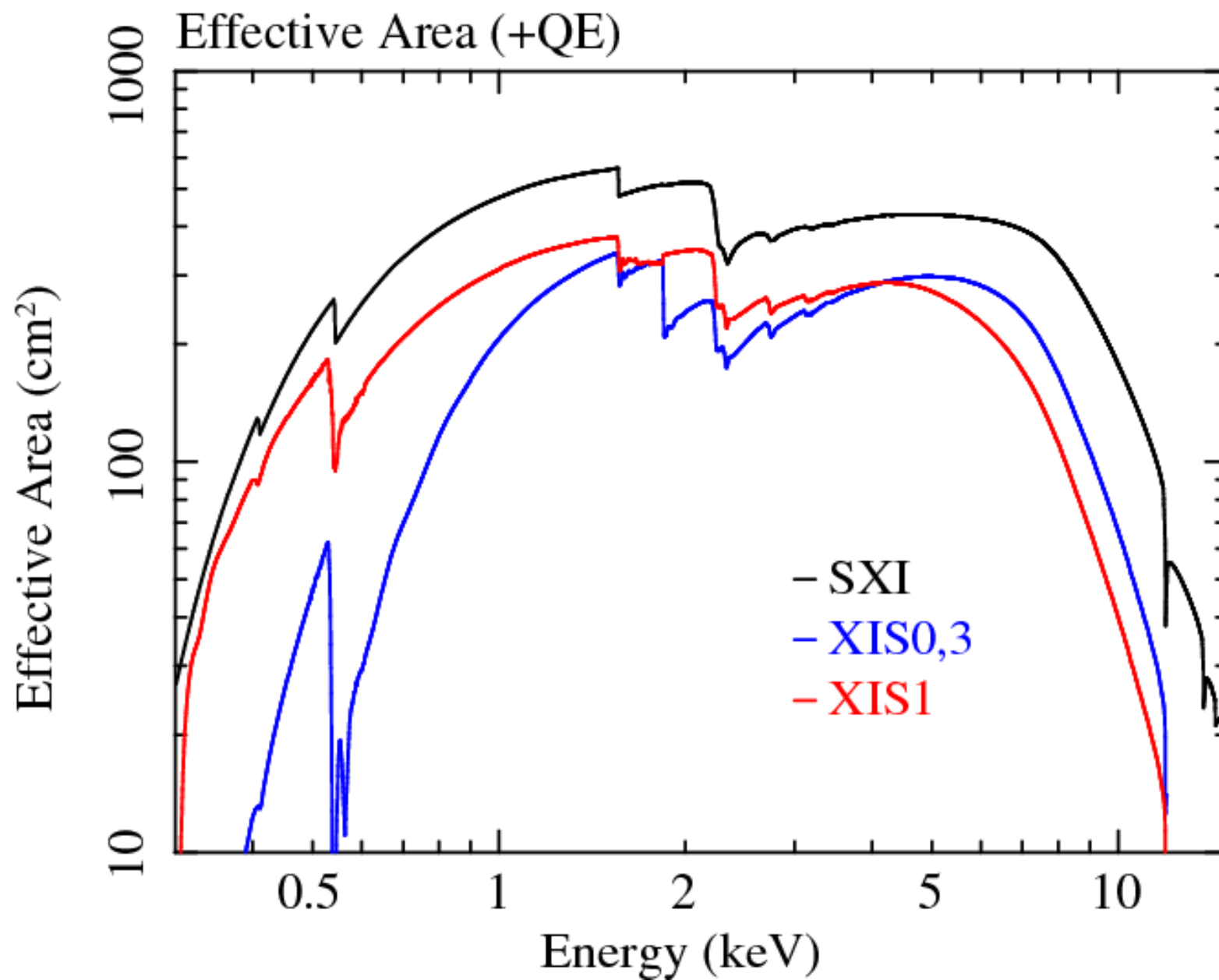


5. Conclusion

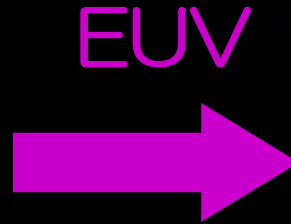
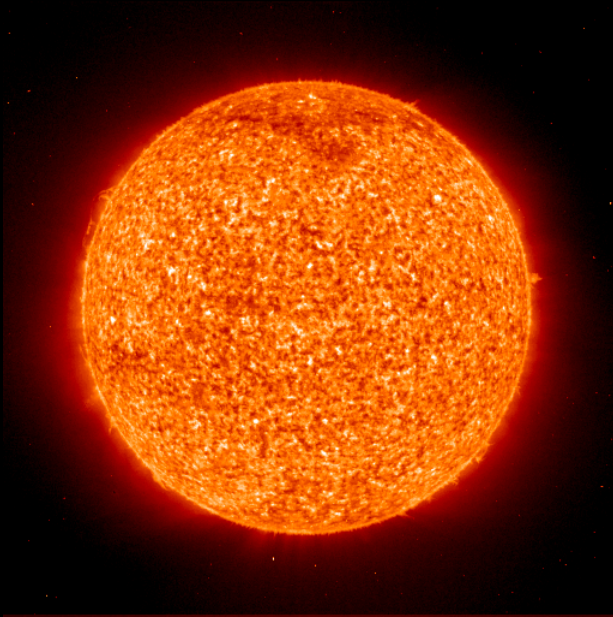
1. We remanufactured Contamination Blocking Filter of SXI due to the increase in the number of the pin holes on the OBL.
2. We measured the optical, soft X-ray, and UV transmission of FM-CBF for Hitomi SXI.
3. The optical transmission of FM-CBF were lower than 5×10^{-5} . This result meets the requirement of optical shield of SXI, but slightly higher than we expected from Luxel design thickness.
4. The X-ray transmission was also measured precisely including XAFS structure around C, N, O and Al absorption edges, and this results also suggests the thinner Al thickness which predicted the optical transmission.
5. We also measured the UV transmission and confirmed the low transmission in order to block the 304A UV light from sunlit atmosphere.
6. We will adopt CBF for XRISM SXI of the same design as FM-CBF Hitomi SXI.



Back up slide



SOHO EIT, He II line, 304 Å
October 11, 1996 at 20:24



Sun @ 304 Å observed
with SOHO satellite

The north pole image
of the earth @304 Å

**In order to observe X-ray in space,
we have to block this EUV light from
Sunlit atmosphere
as well as the optical light from objects.**

3. Optical Transmission of CBF

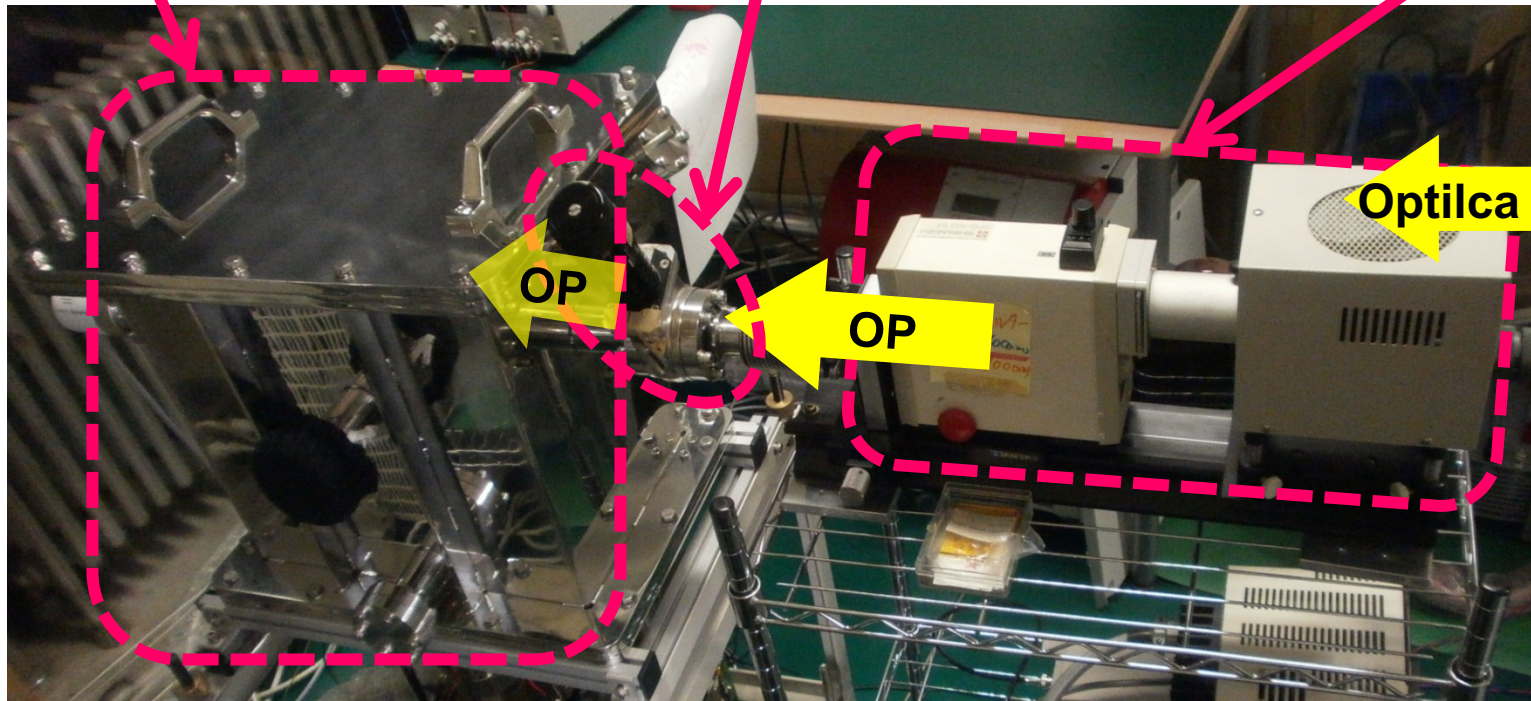
Vacuum Chamber

Al filter

Reducer of the optical light

Optical light source and
Grating spectrometer

450nm~950nm



Optilca Light (OP)

OP

OP

- We measured the optical flux using photo diode with/without CBF, and calculated the optical transmission of CBF derived from the flux ratio between two them.