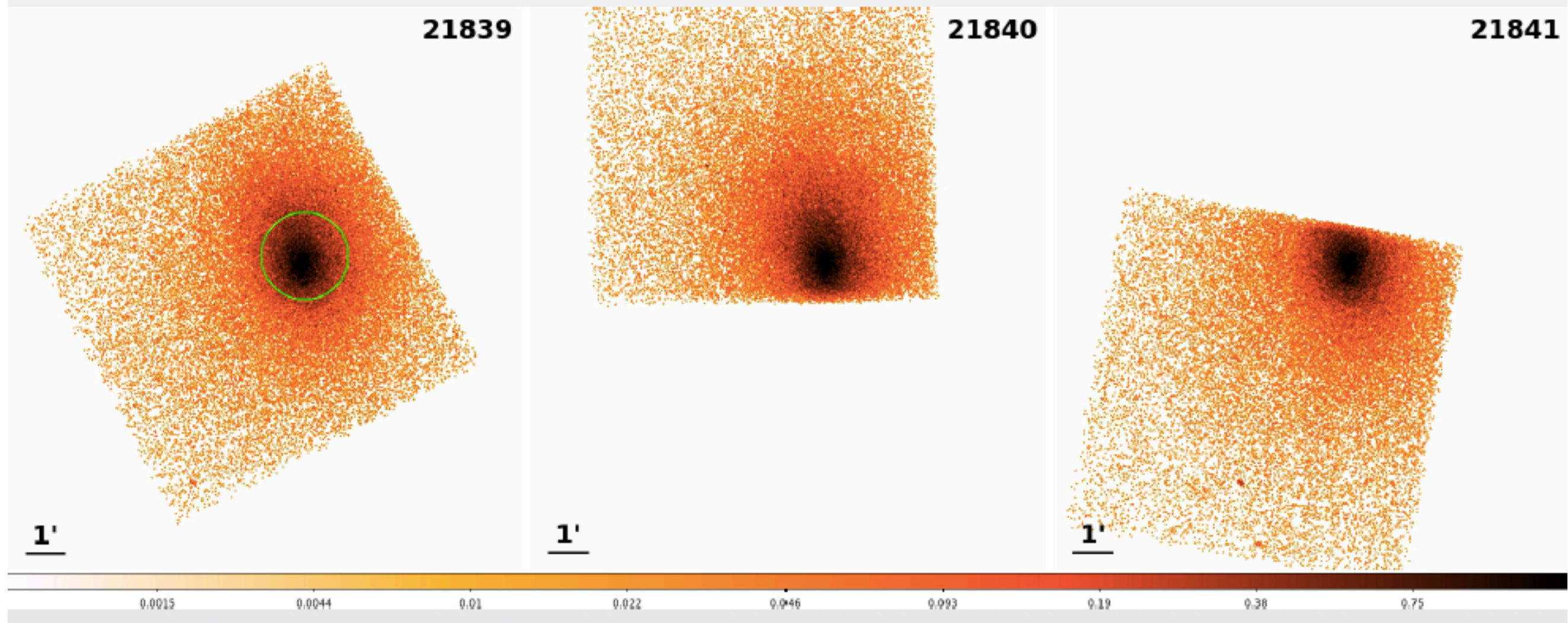


Chandra Contaminant Update: *No end in sight...*

Herman L. Marshall (HLM)
and
Akos Bogdan (SAO)

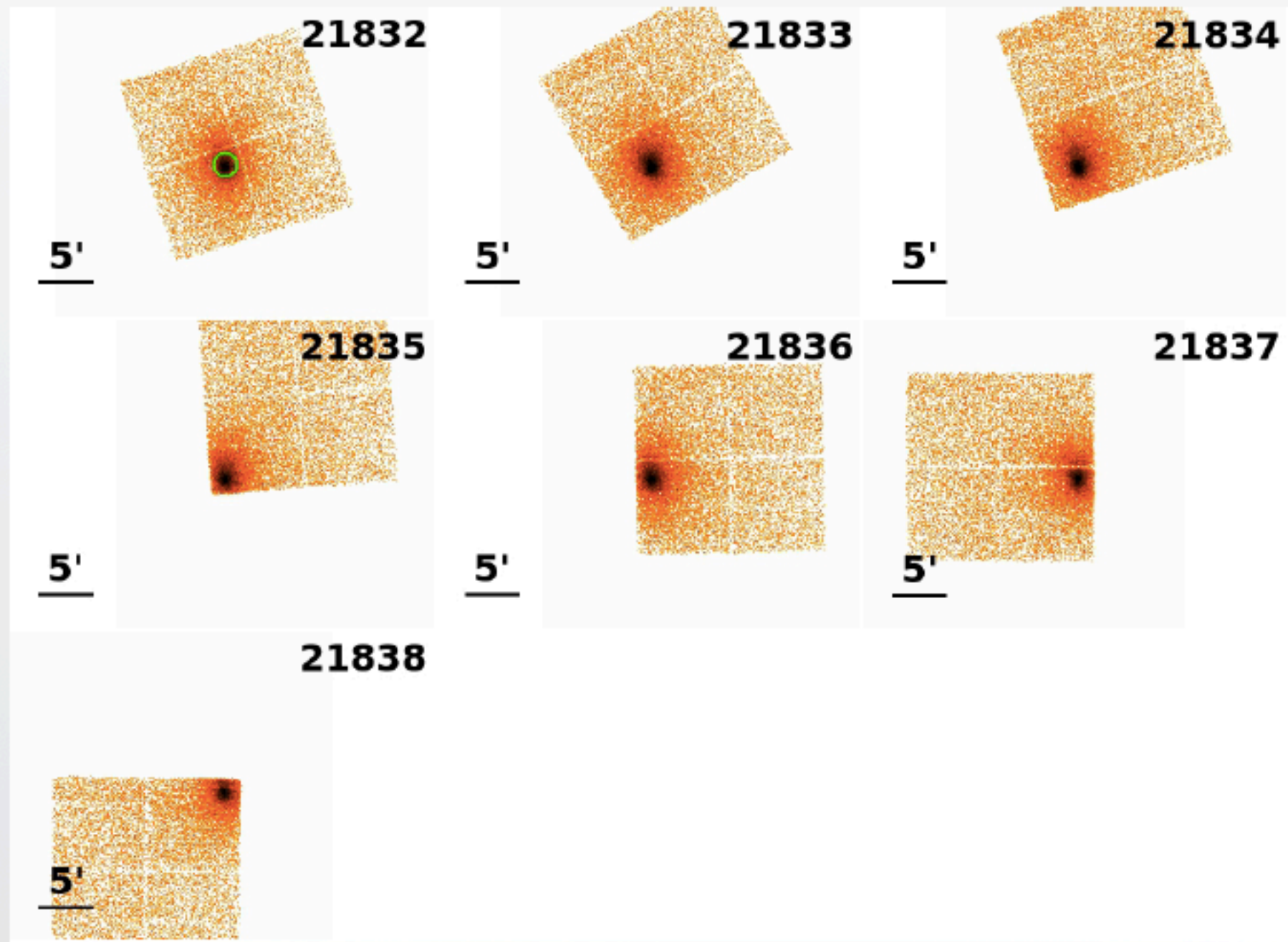
ACIS-S observations of A1795 in 2019

- April 2019: 3 pointings on S3 (center, top, bottom)
- November 2019: 1 pointing (center)
- Exposure time is 15-20 ks

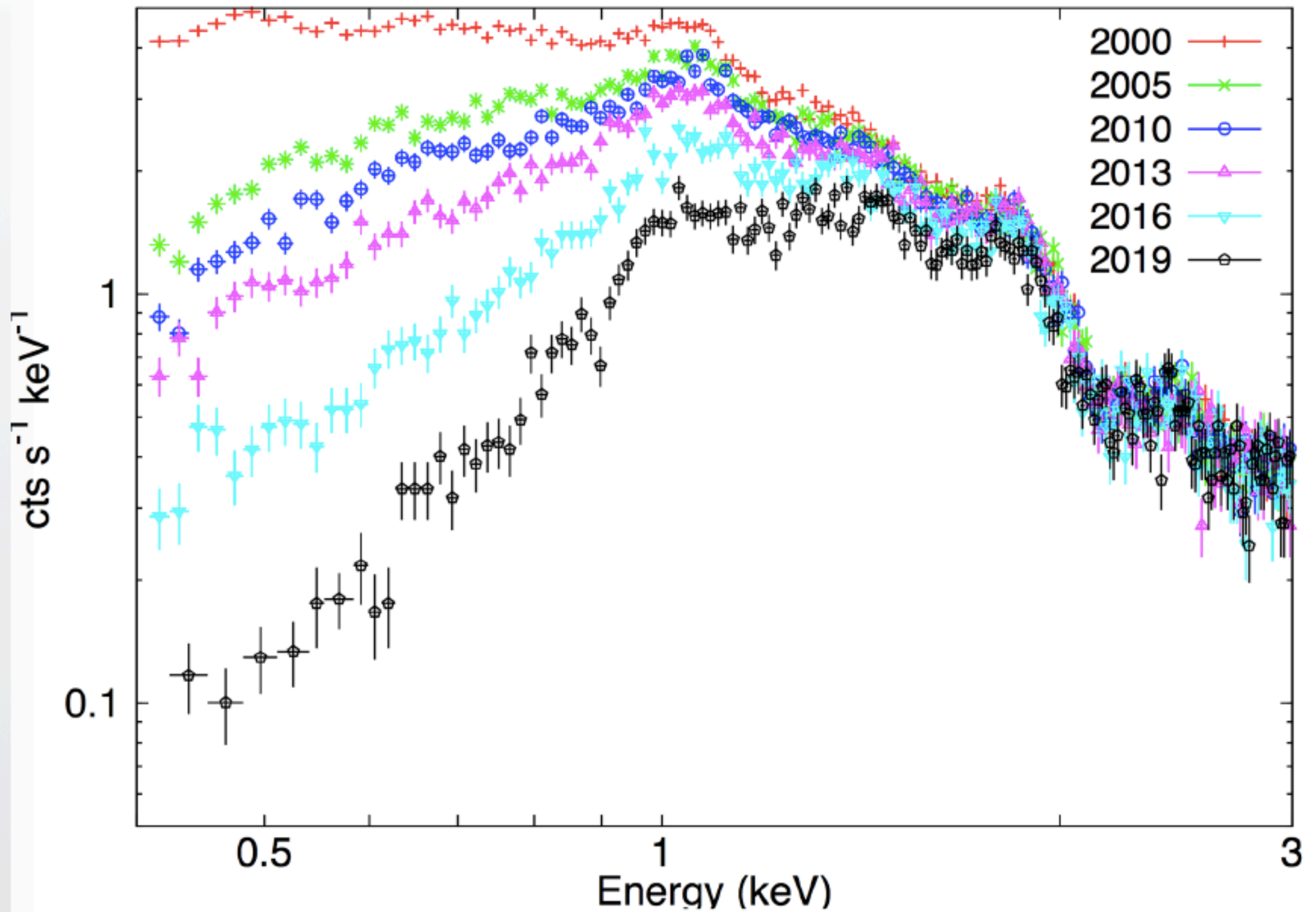


ACIS-I observations of A1795 in 2019

- April 2019: 7 pointings on ACIS-I array
- November 2019: 1 pointing (aim point)
- Exposure time is 20-25 ks



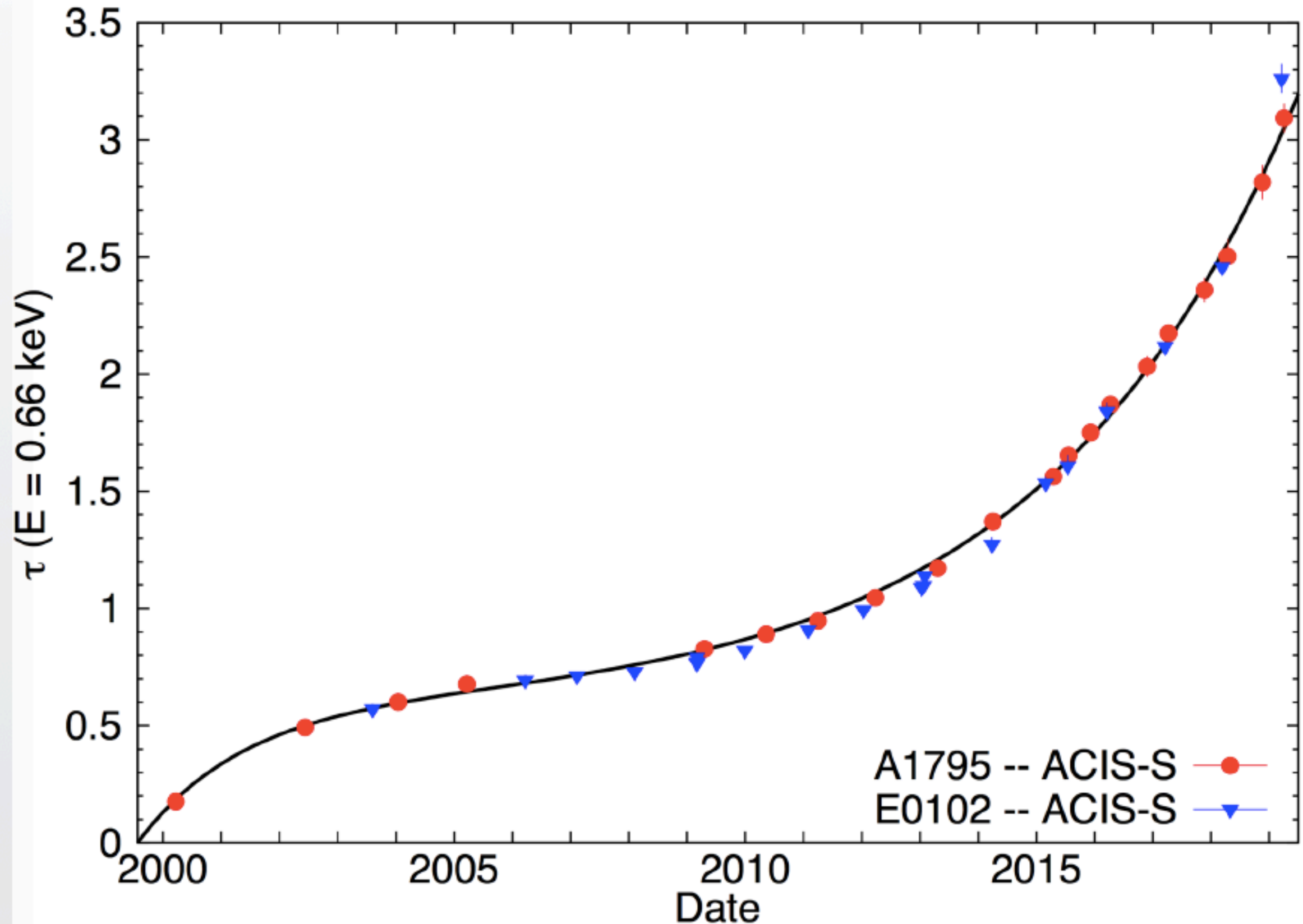
ACIS-S spectra of A1795



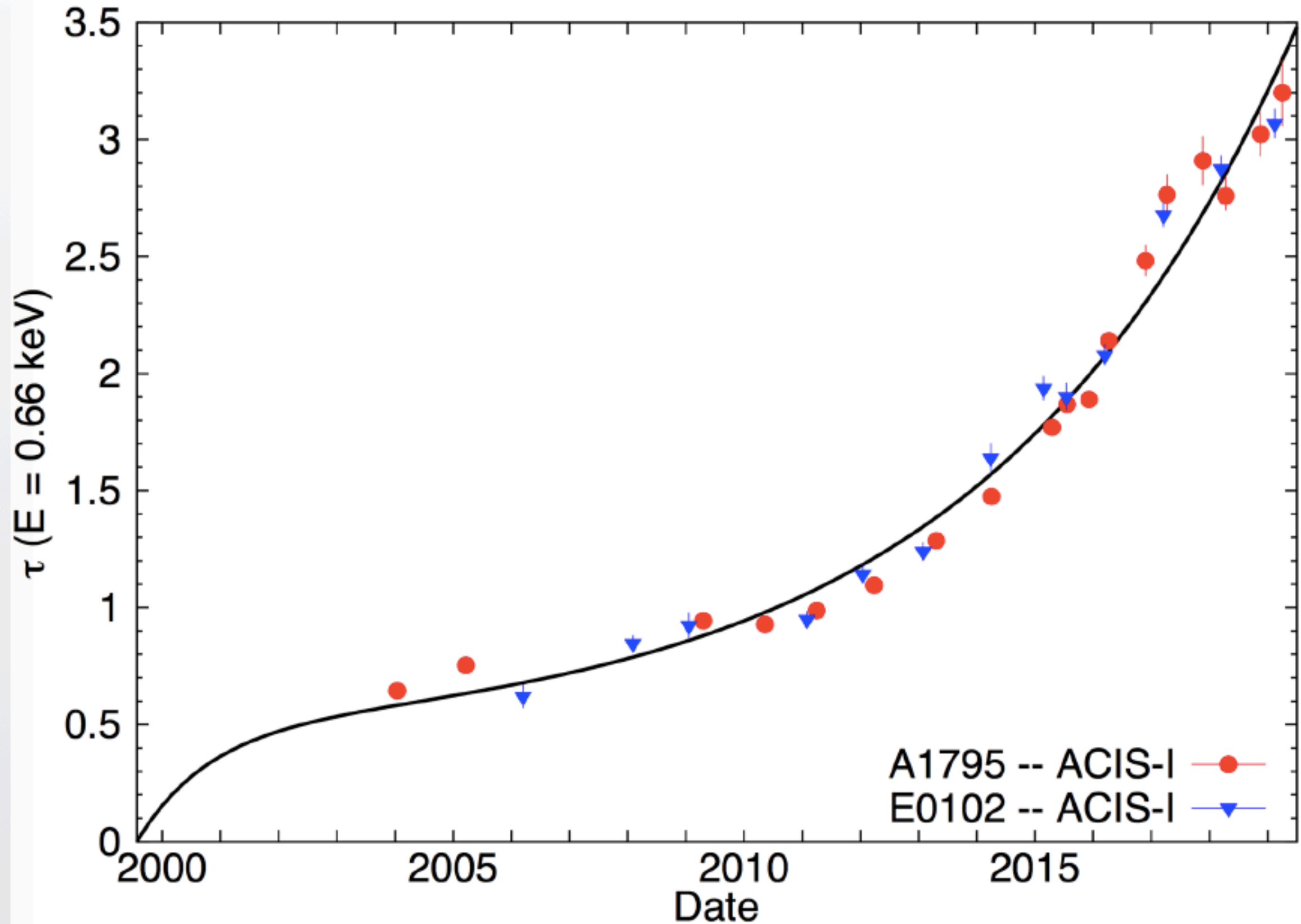
ACIS-S observations of A1795

- Uniform data analysis procedure
- Computing the time dependence:
 - point sources excluded
 - spectral characteristics of Obs ID 494 (December 1999) used as reference by extracting circular region with 65" radius centered on A1795
 - spectrum described with Galactic column density, APEC models, and ACIS contamination with fixed O/C and F/C ratios
 - For subsequent observations the spectra of the same 65" circular region is extracted with the contamination correction turned off
 - The follow-up spectra are fit with best fit spectrum obtained from Obs ID 494 and additional ACIS contamination

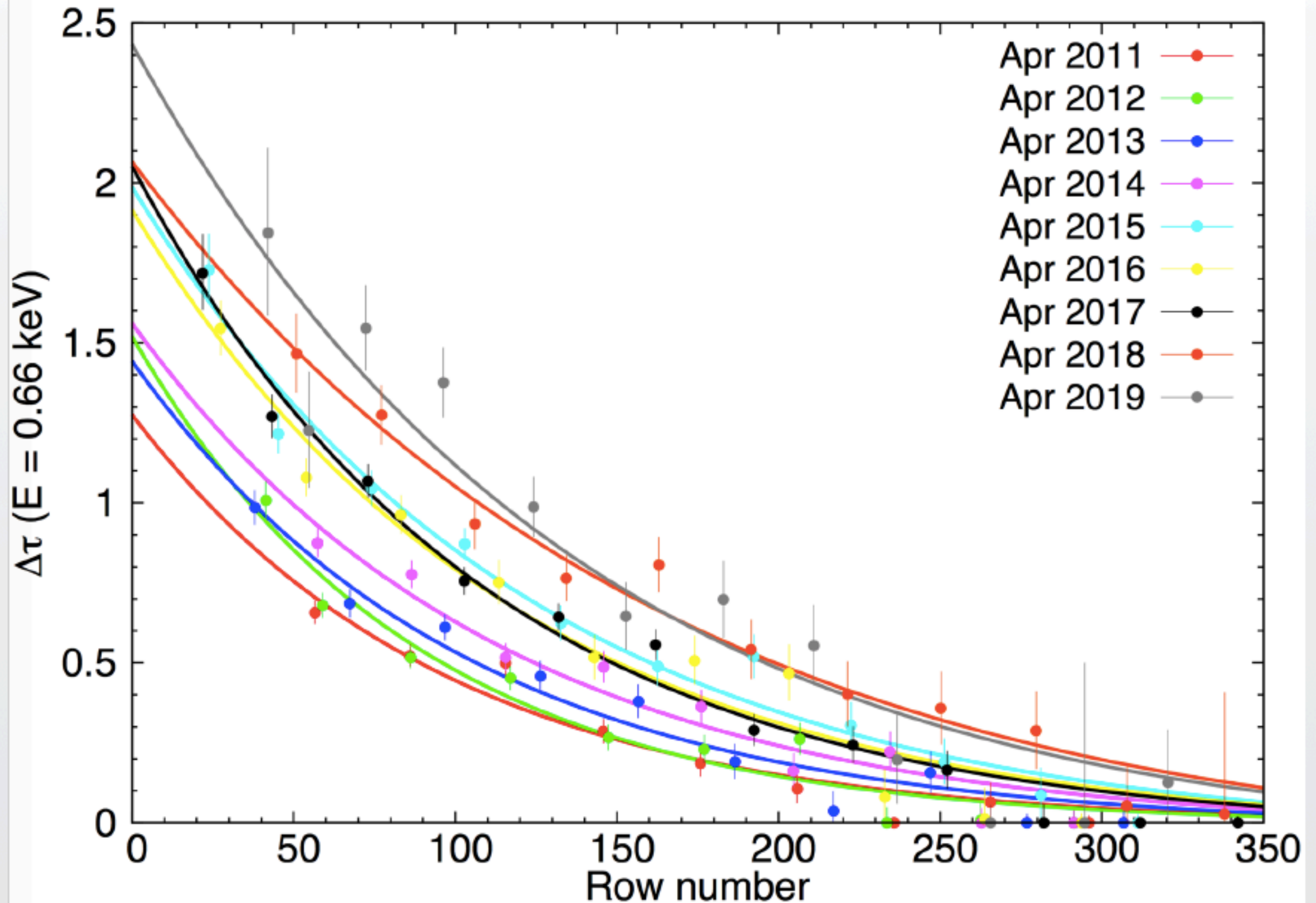
Time dependence of contaminant in the center of ACIS-S using A1795 data



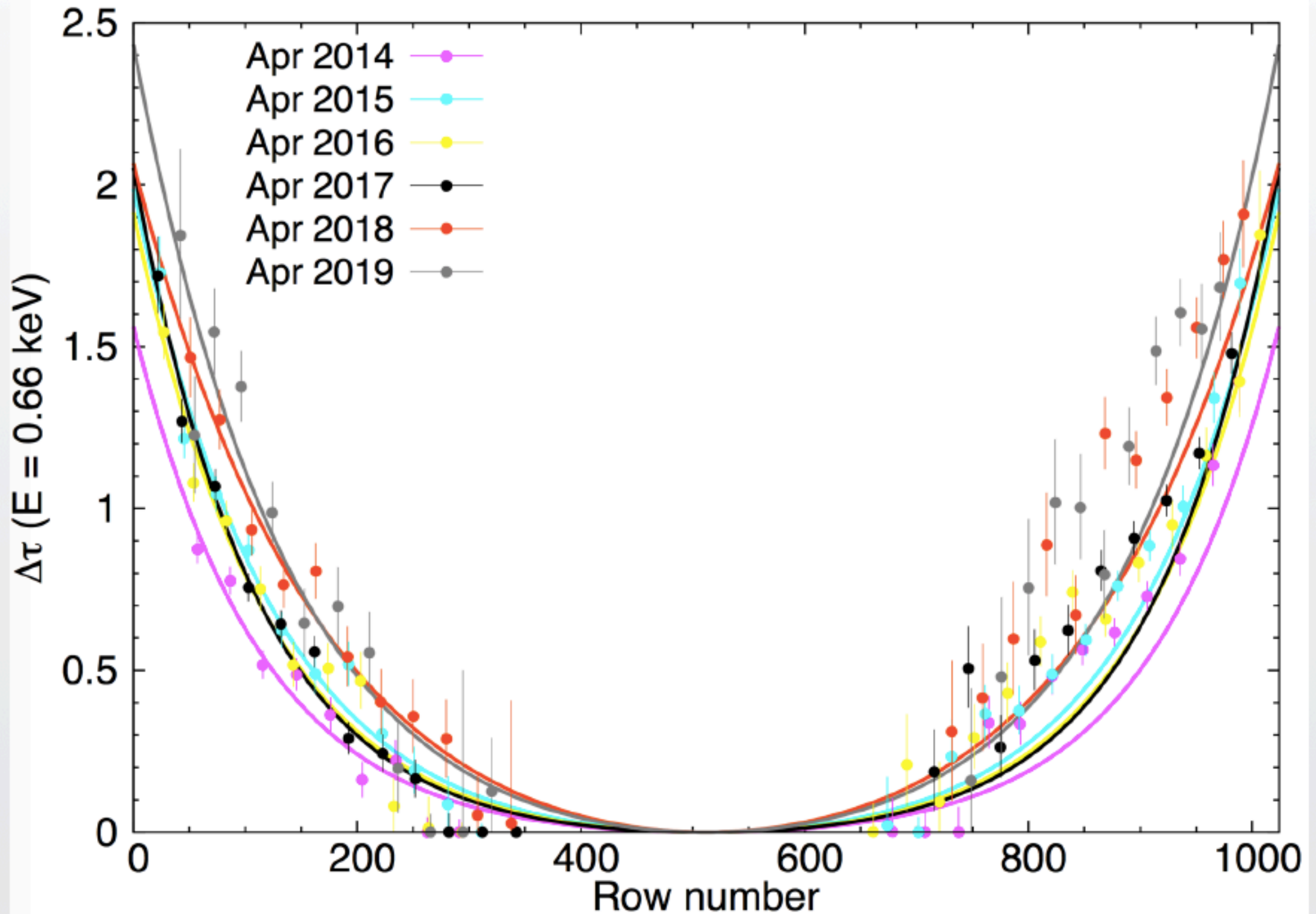
Time dependence of contaminant in the center of ACIS-I using A1795 data



Shape of the spatial structure of the contaminant on ACIS-S

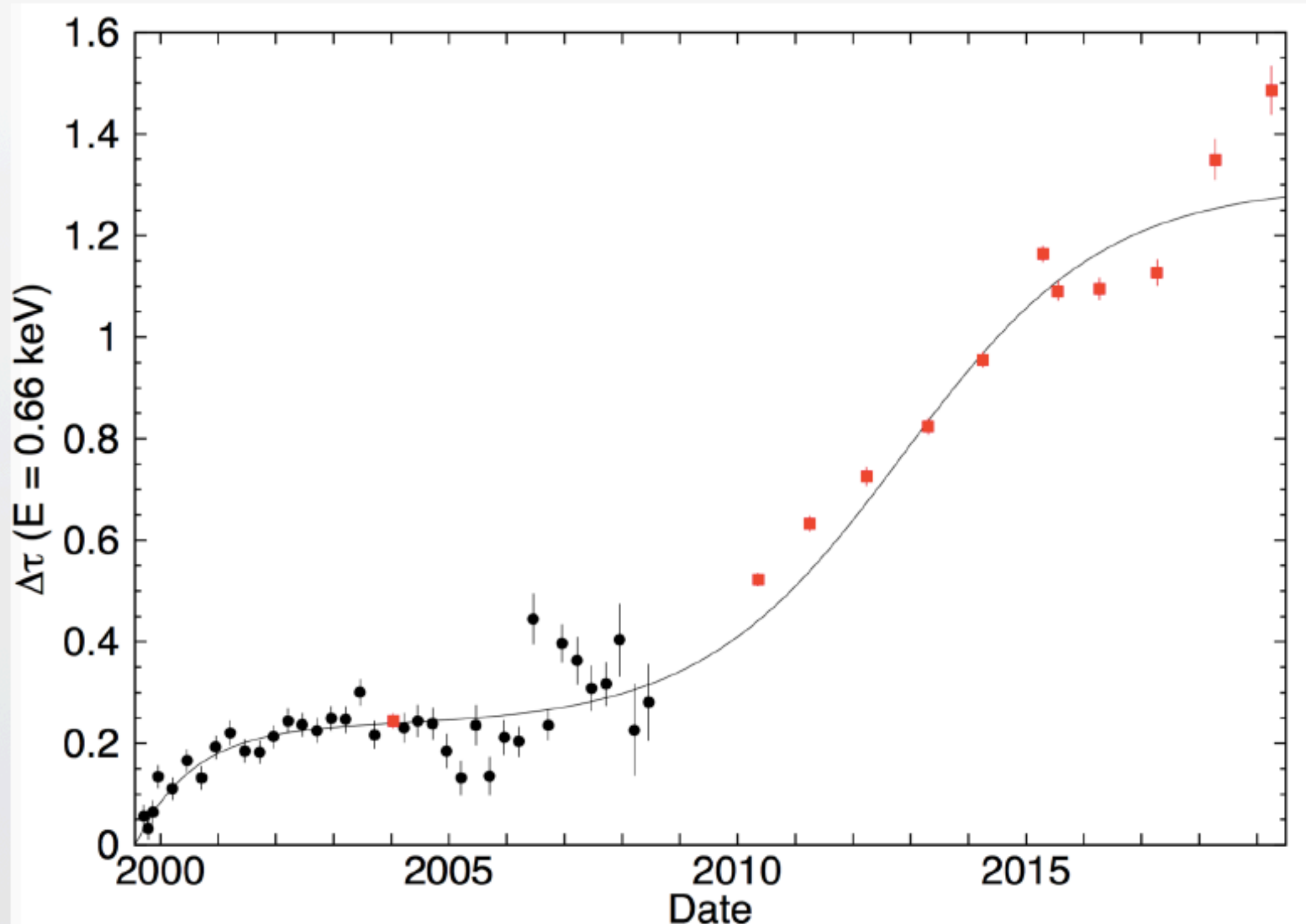


Shape of the spatial structure of the contaminant on ACIS-S



Edge-to-center difference at E=0.66 keV

- Plot shows optical depths relative to the center at E=0.66 keV
- ECS data at E=0.66 keV (black circles) and A1795 data (red boxes)



Big Dither Process

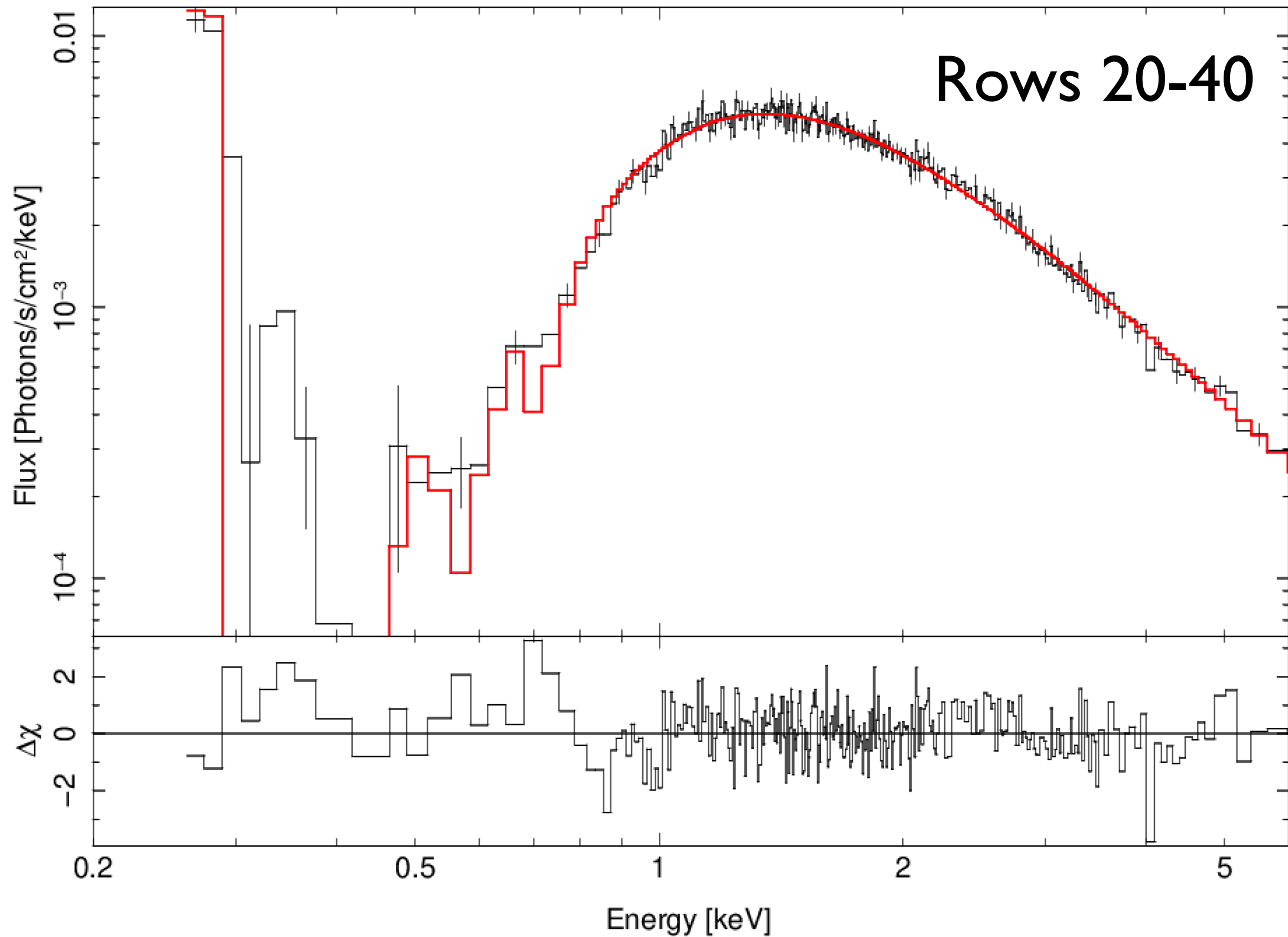
- New Big Dither Data: 6-7/2018, 1-2/2019
- Extract 15 spectra of 20-80 rows each
- Fit COF optical depths in isis
 - fixed slope, curvature per group
- Fit $N * \cosh[(y-512)/120] + C$; each element

The Approach

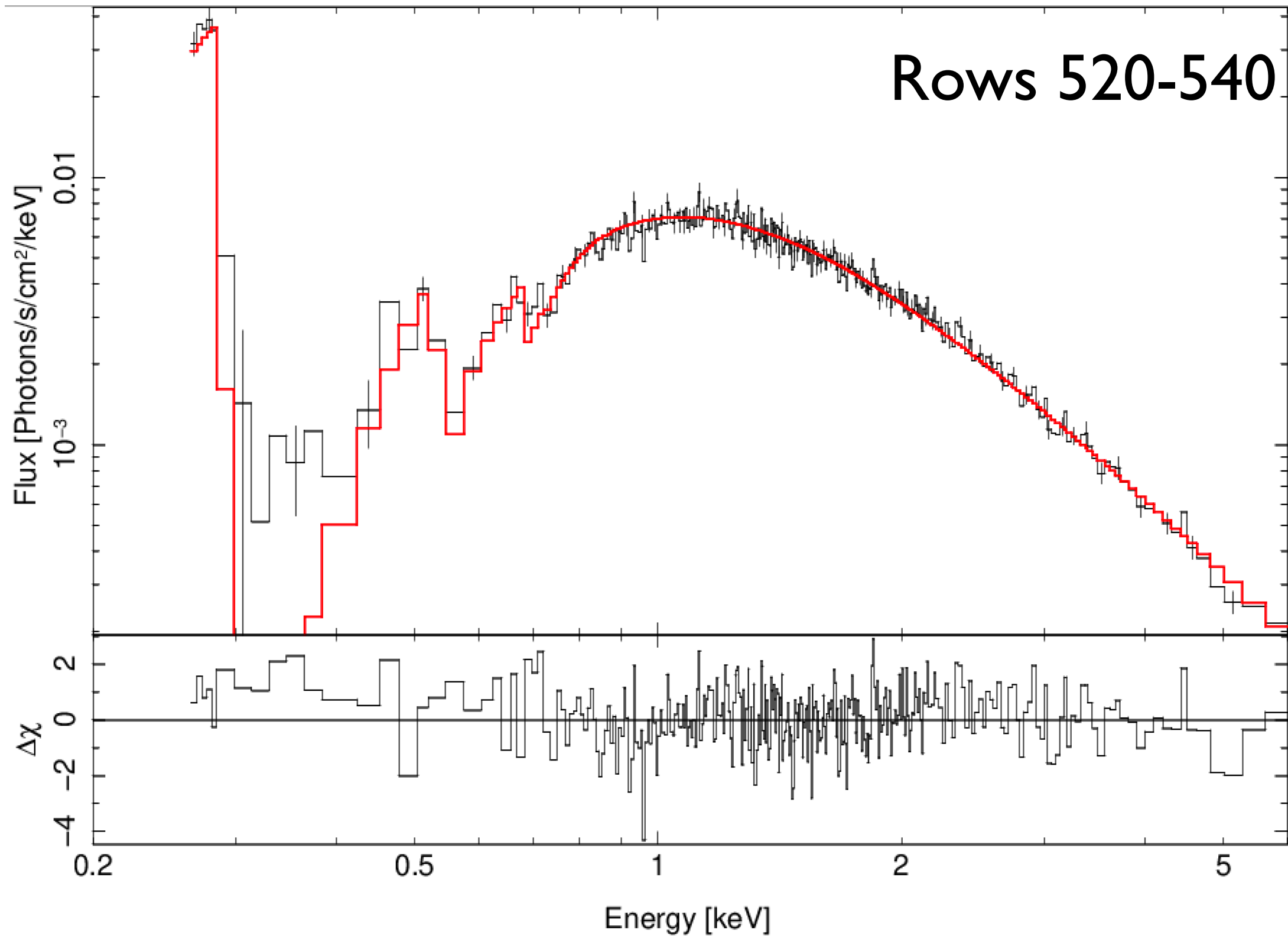
- Model: $\tau_C = \tau_{C0} + t_{C1}(t) f(x,y)$
- LETG/ACIS of blazars, 'Big Dither'
 - measure O-K as $h(t,y)$
 - measure F-K as $k(t,y)$
- Determine $\tau_{C, \text{Henke}}$ from cluster data, corrected for τ_{O-K}, τ_{F-K}
- Adjust C-K edge

$$f(x, y) = e^{-y/a_1} + e^{(y-1024)/a_2} - e^{-512/a_1} - e^{-512/a_2}, a_1 = 106.25, a_2 = 129.62$$

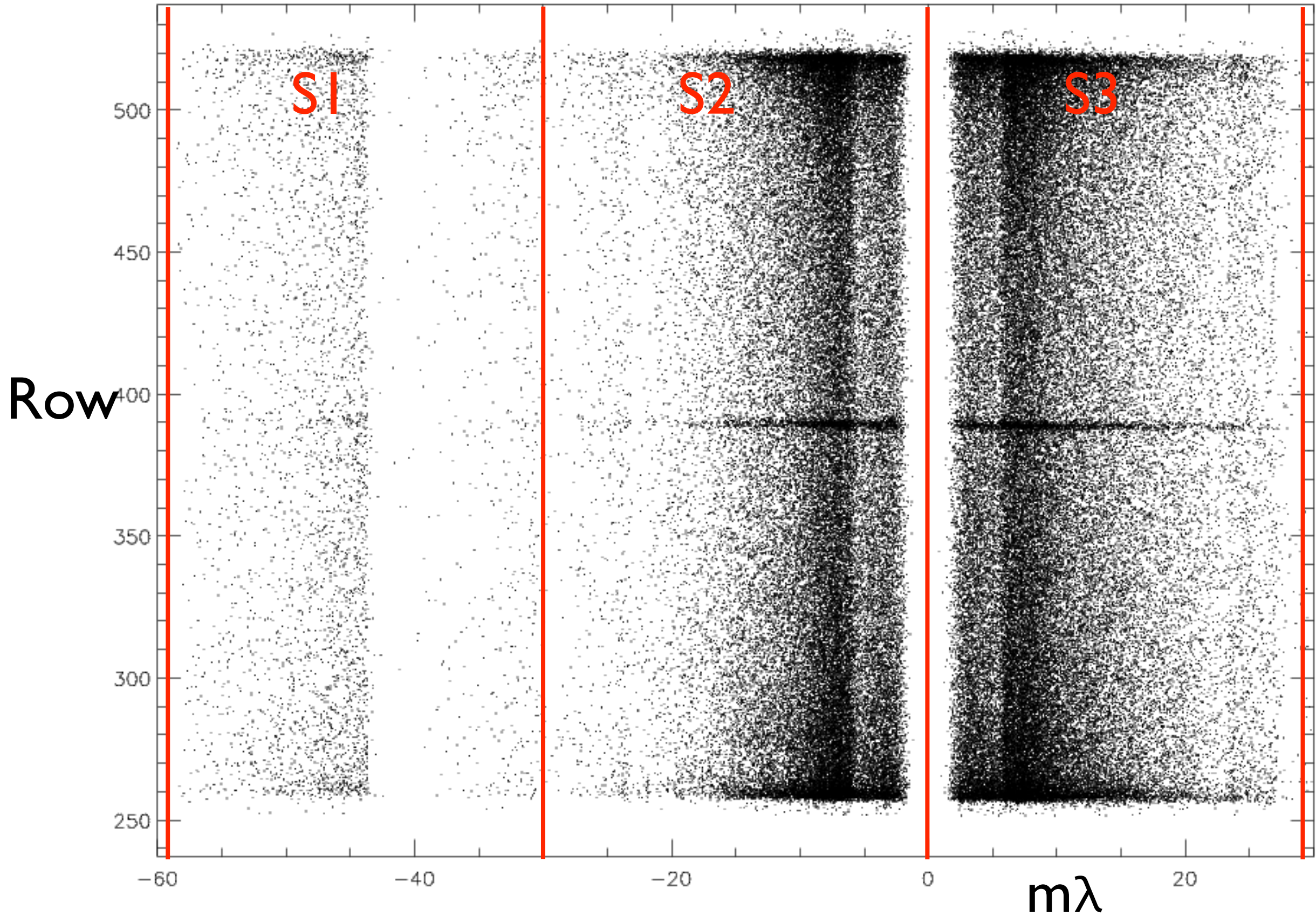
Example Spectral Fits



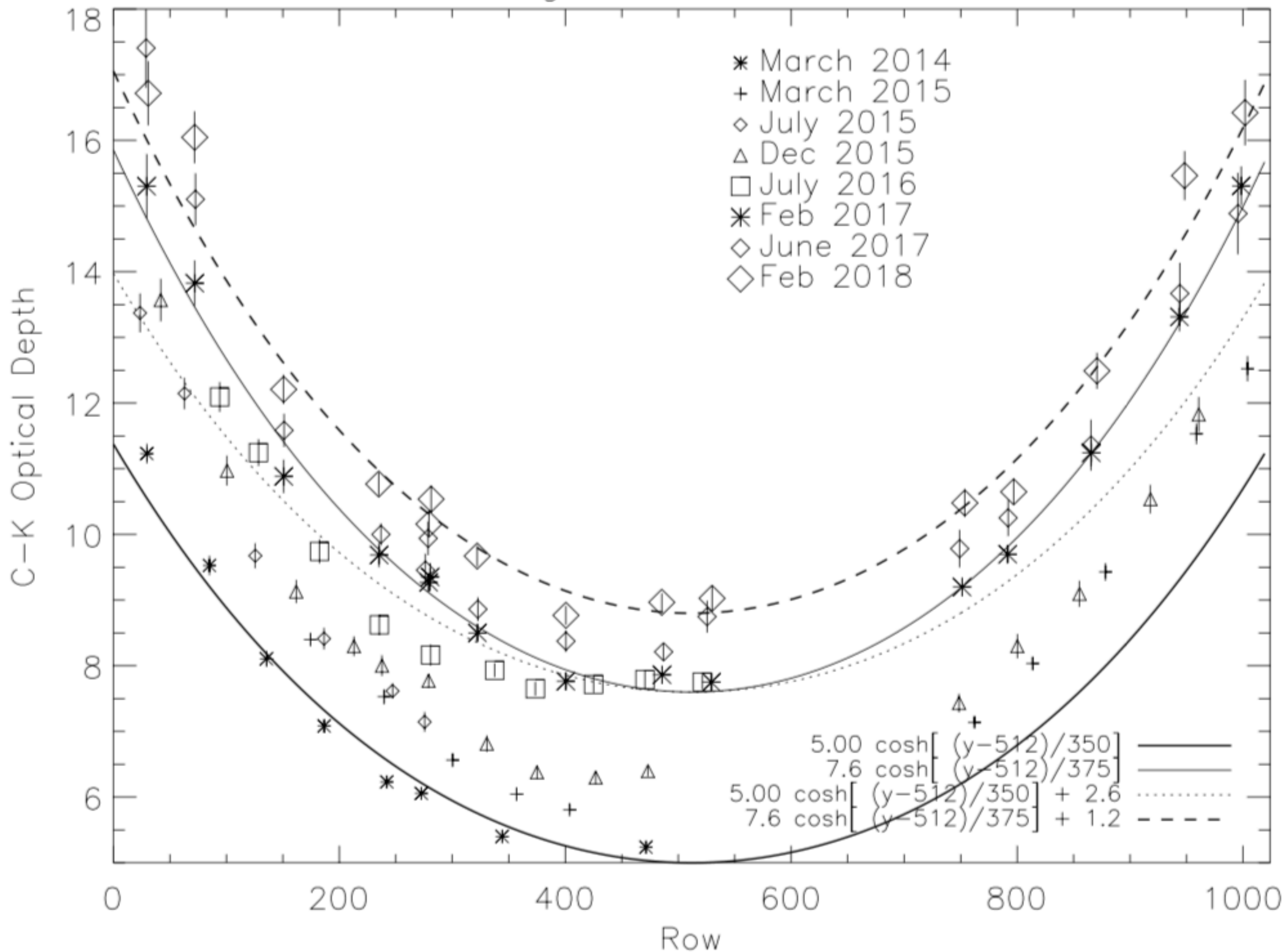
Example Spectral Fits



Mid-Row

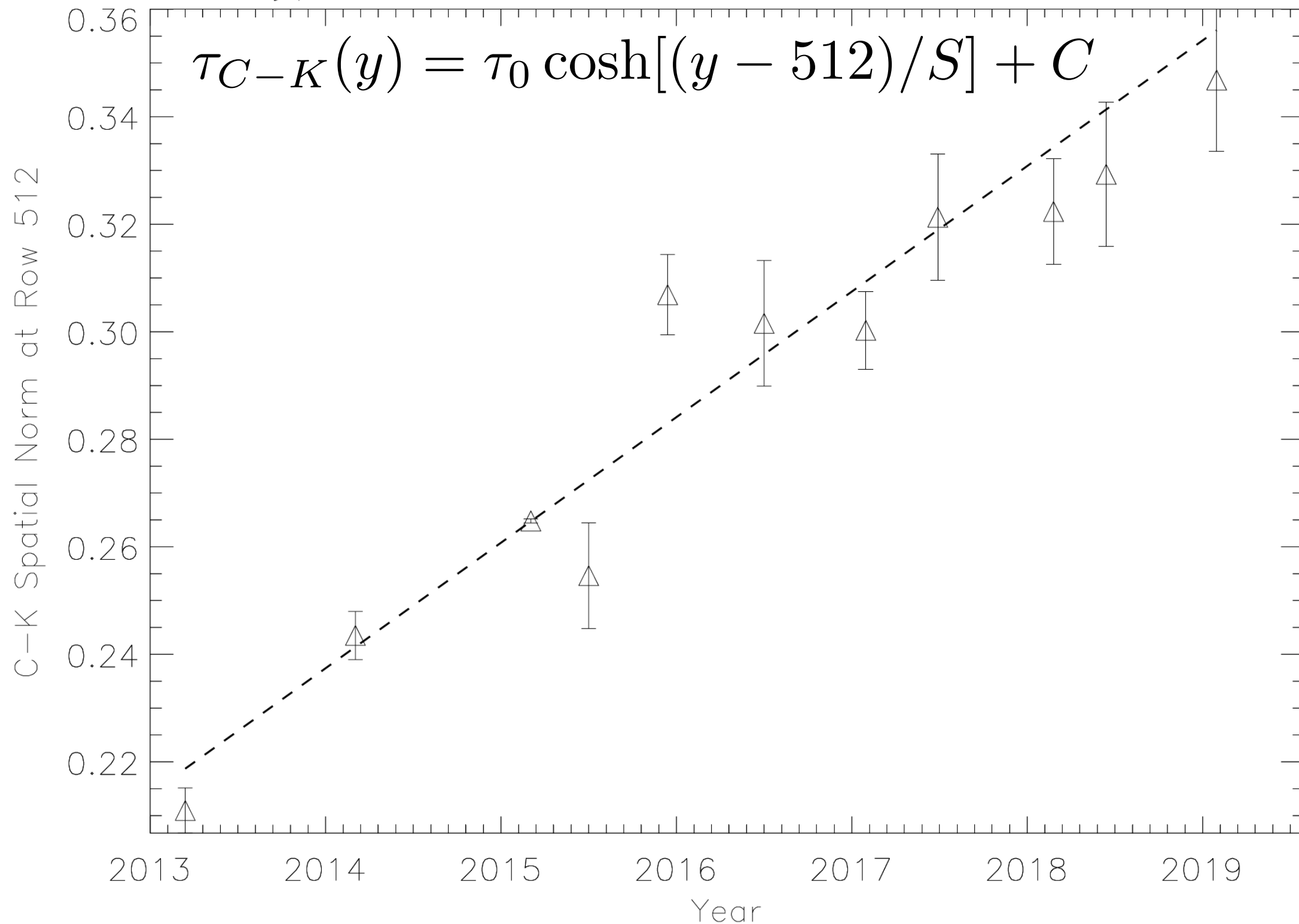


Big Dither 2014–8



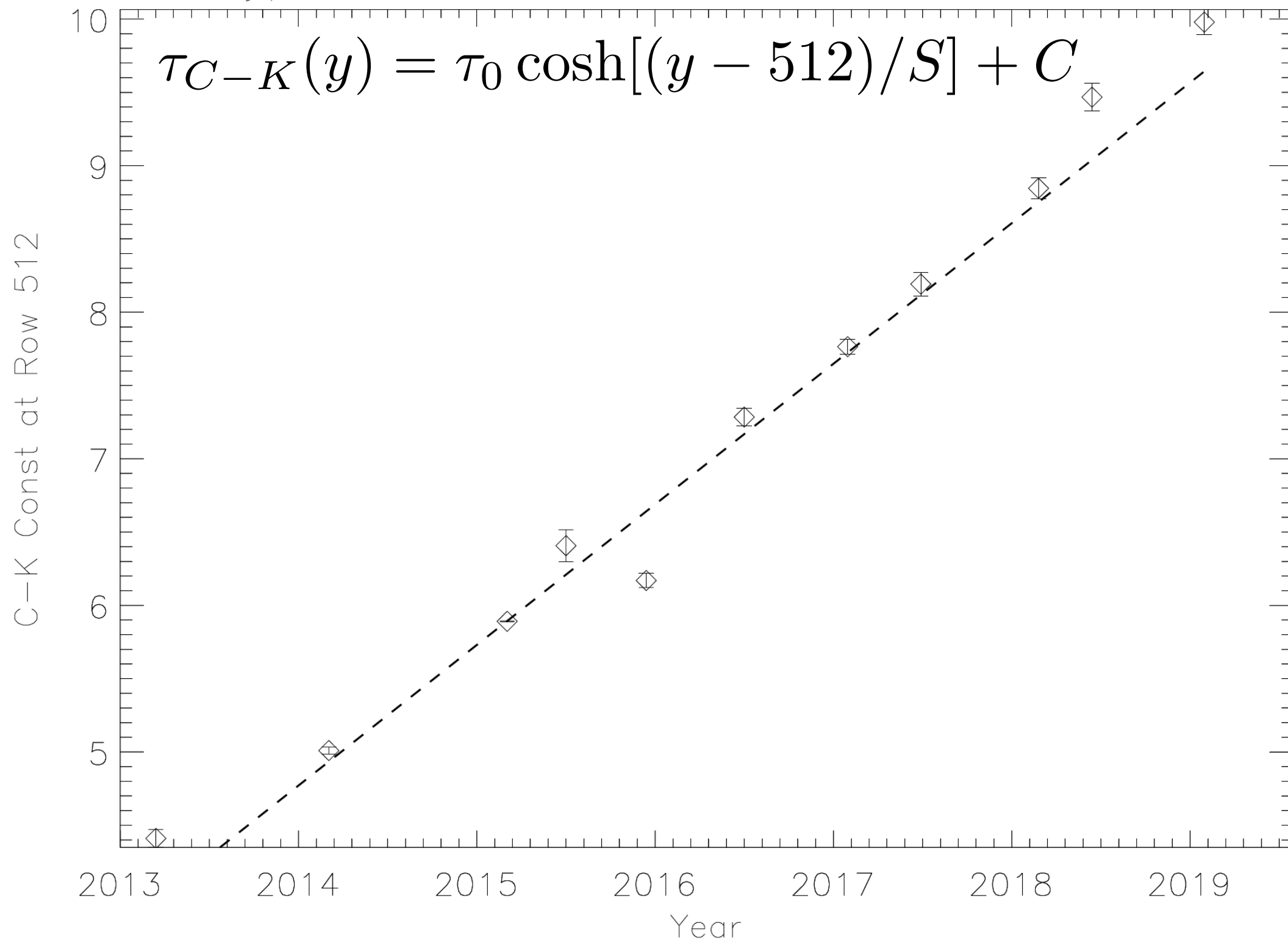
Spatial Part, C-K

Hyperbolic Cosine + Constant, scale=120 row



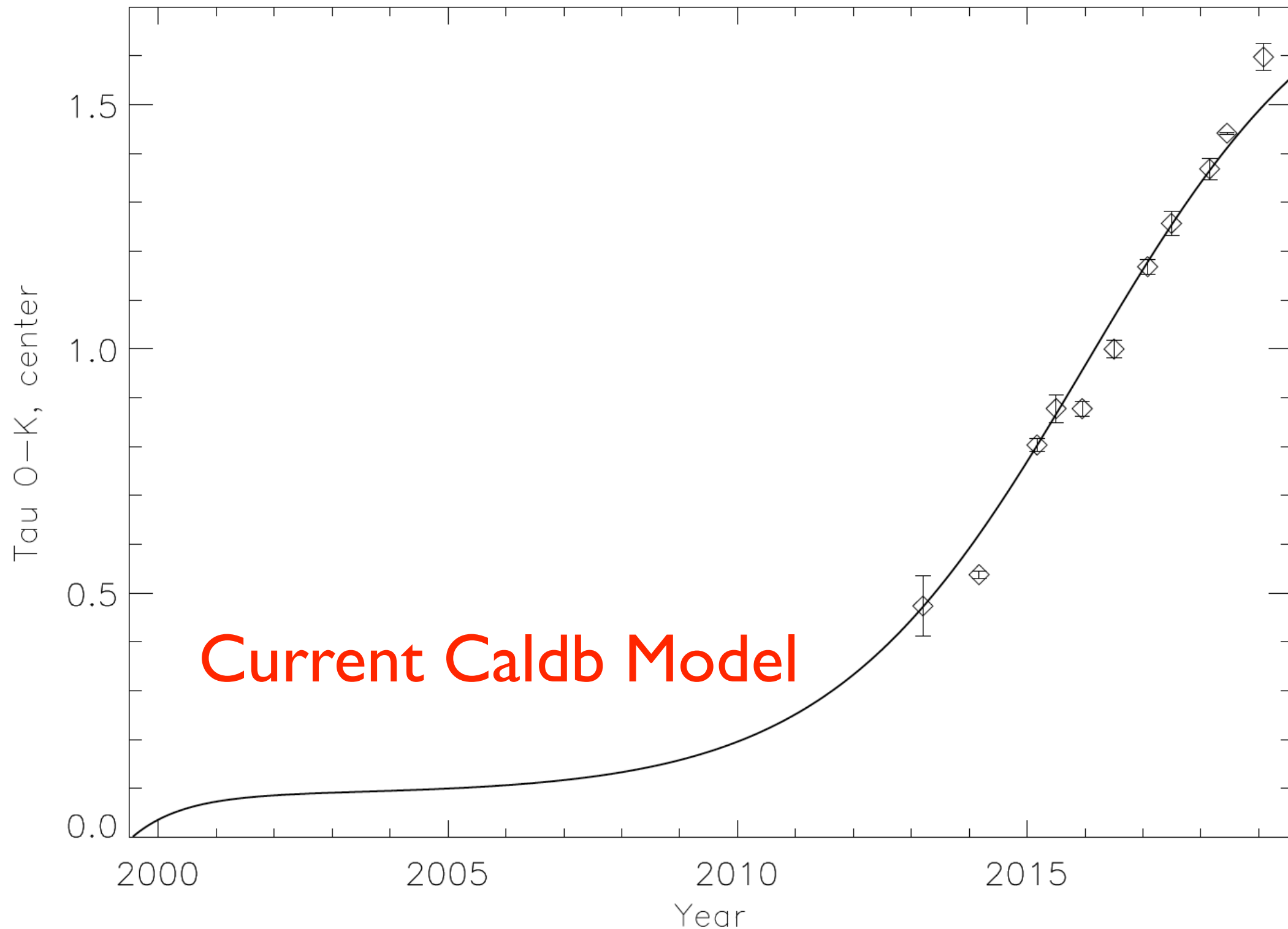
Uniform Part, C-K

Hyperbolic Cosine + Constant, scale=120 row



Uniform Part, O-K

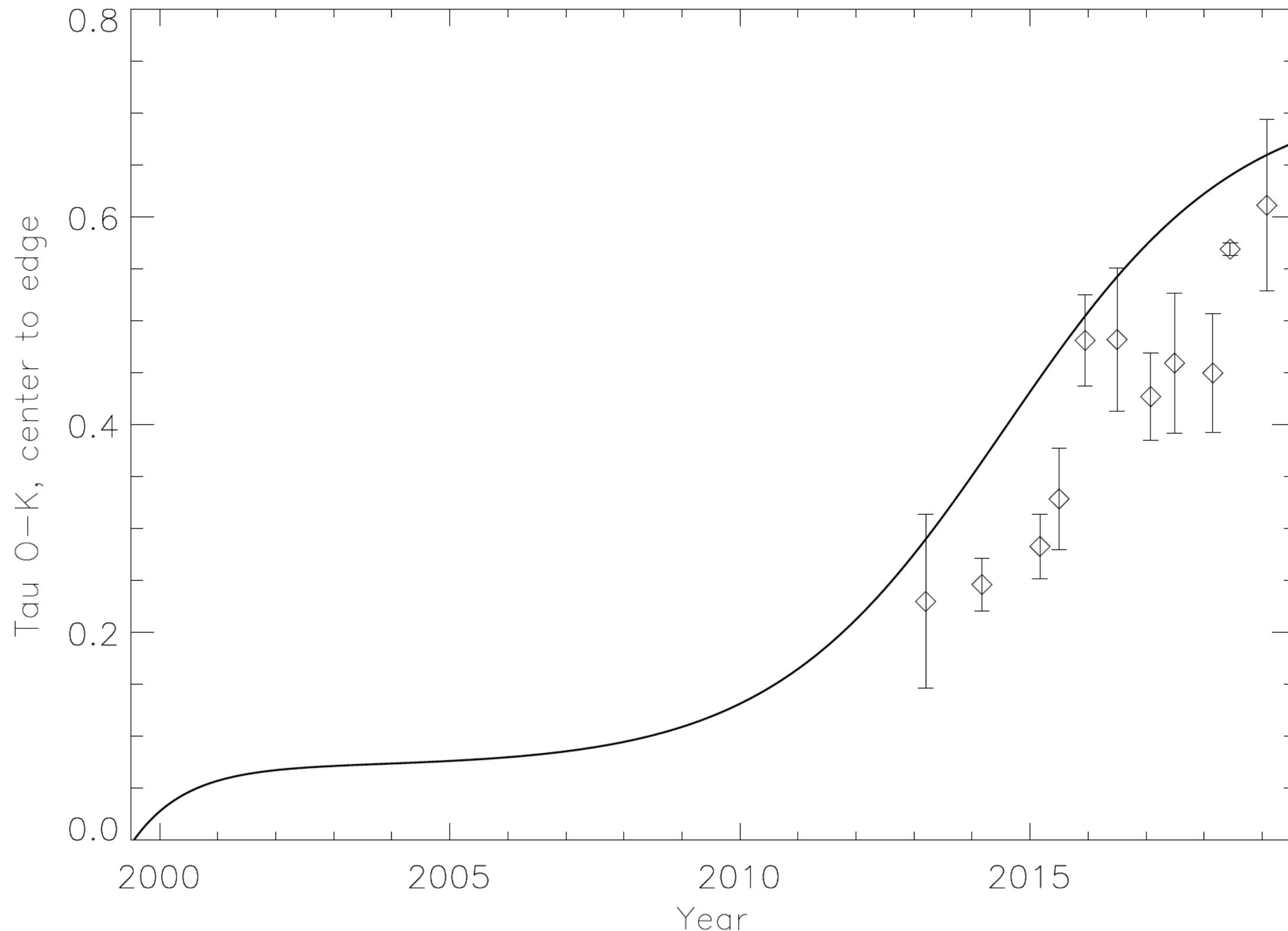
Scale = 120 row



Current Caldb Model

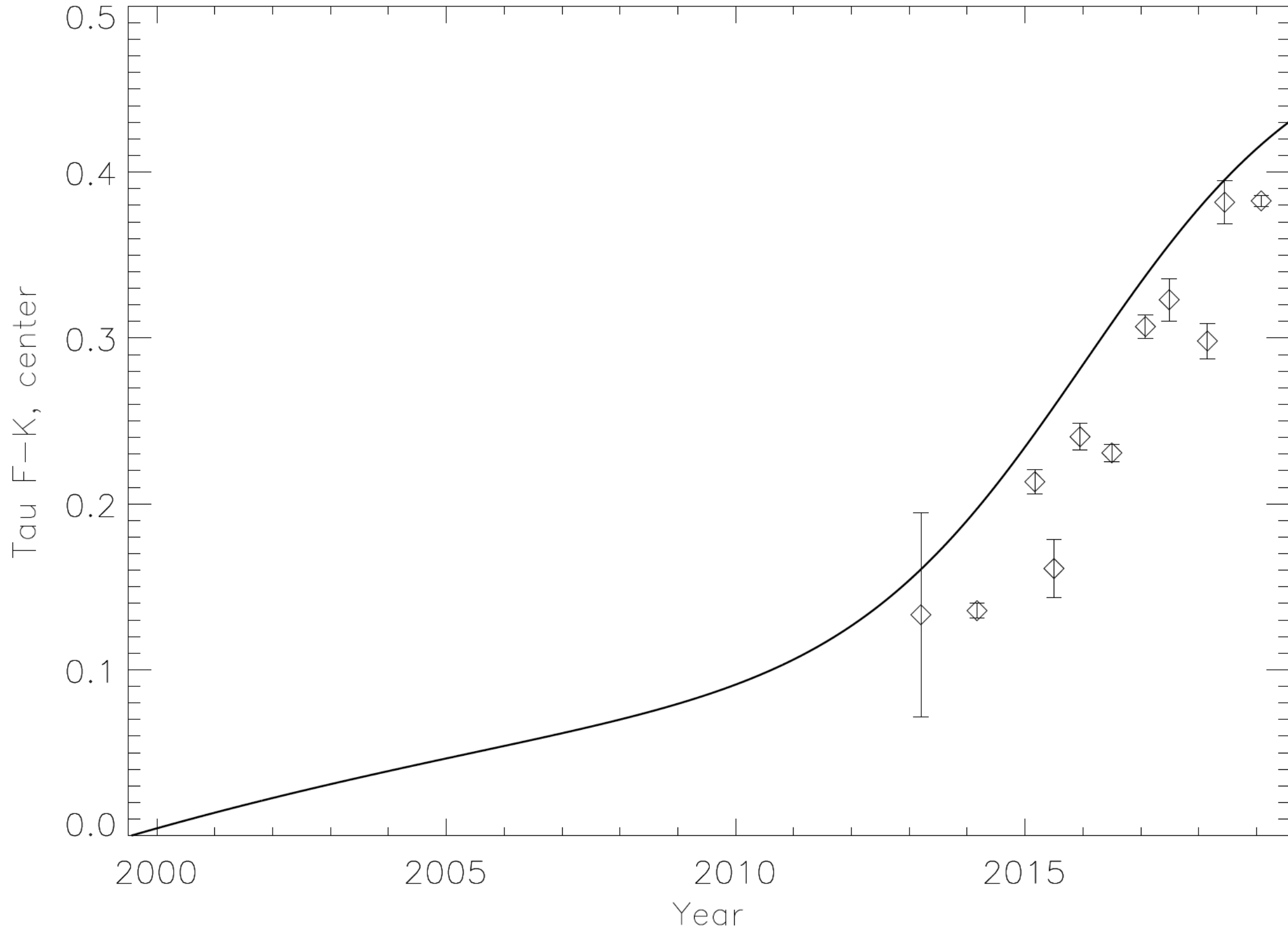
Spatial Part, O-K

Scale = 120 row



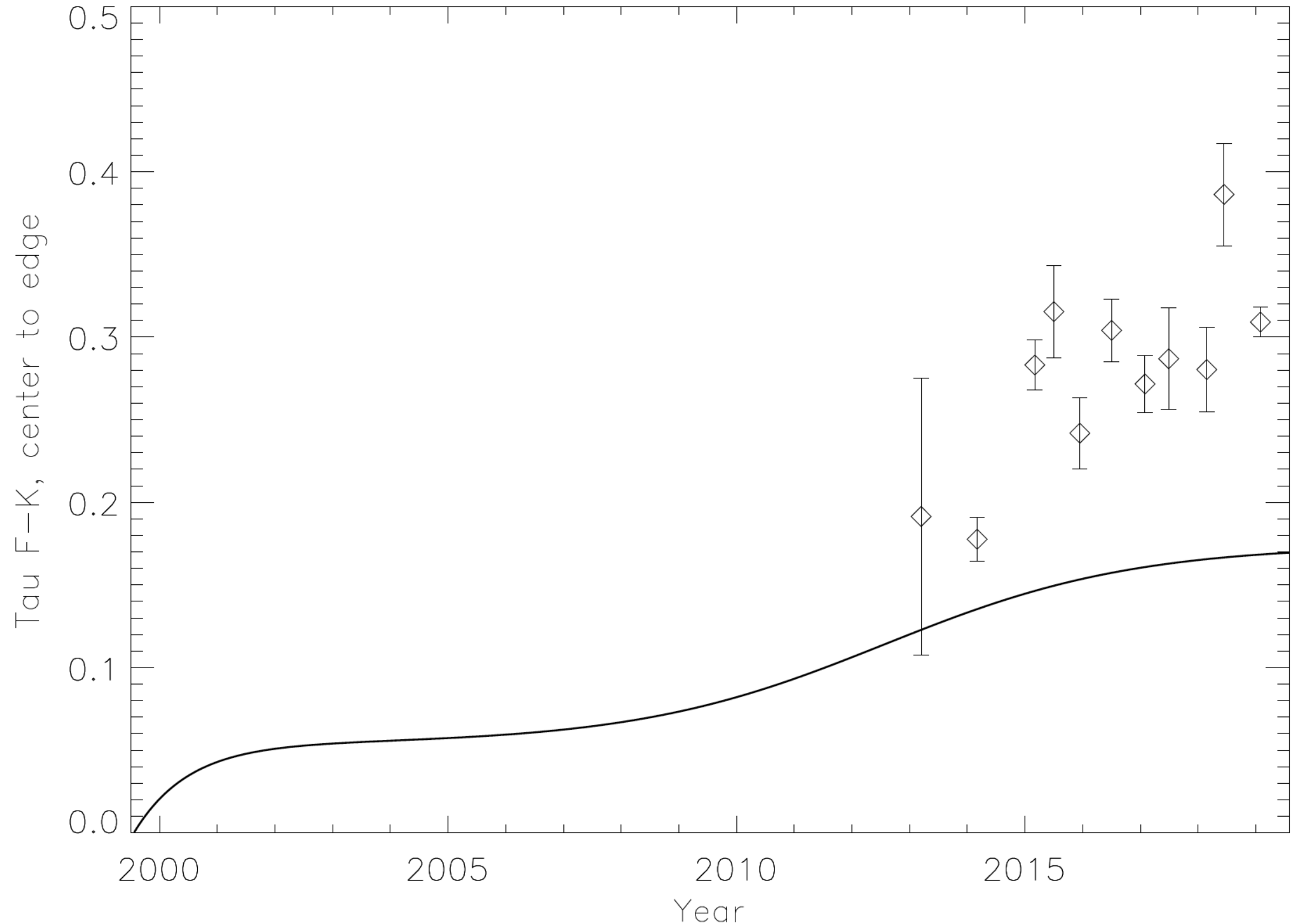
Uniform Part, F-K

Scale = 120 row



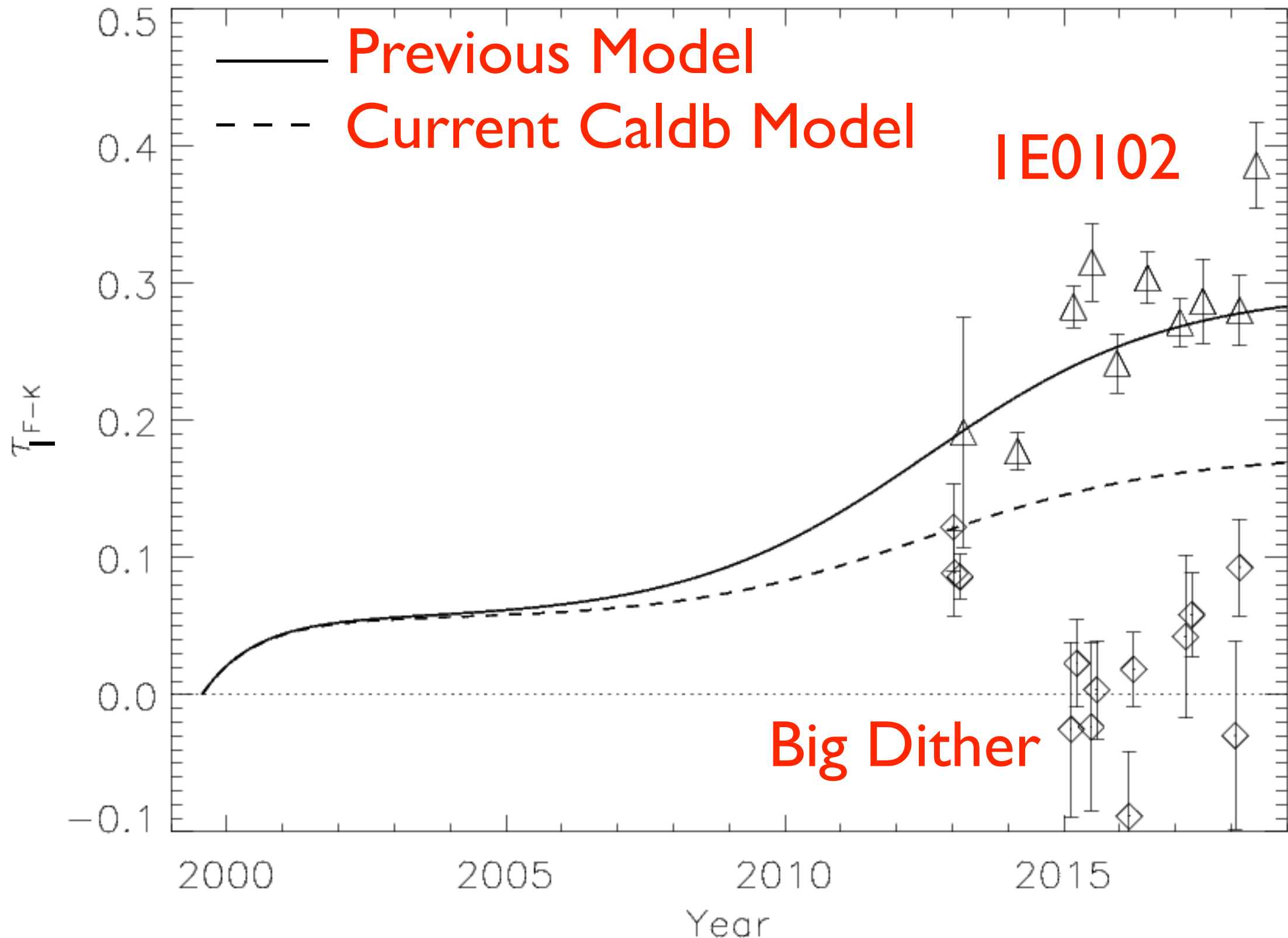
Spatial Part, F-K

Scale = 120 row



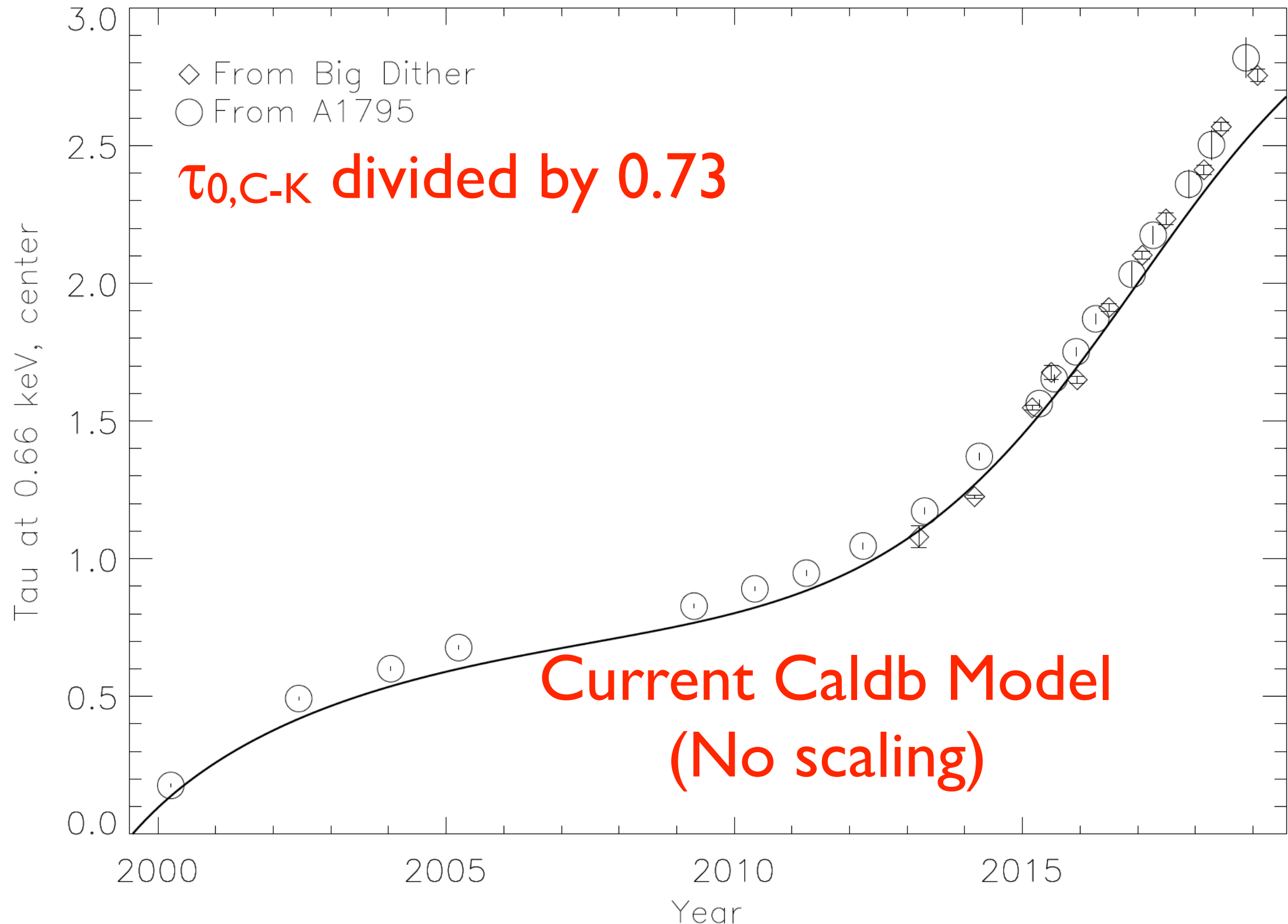
Compromise Model

Data from 1E0102 Fits



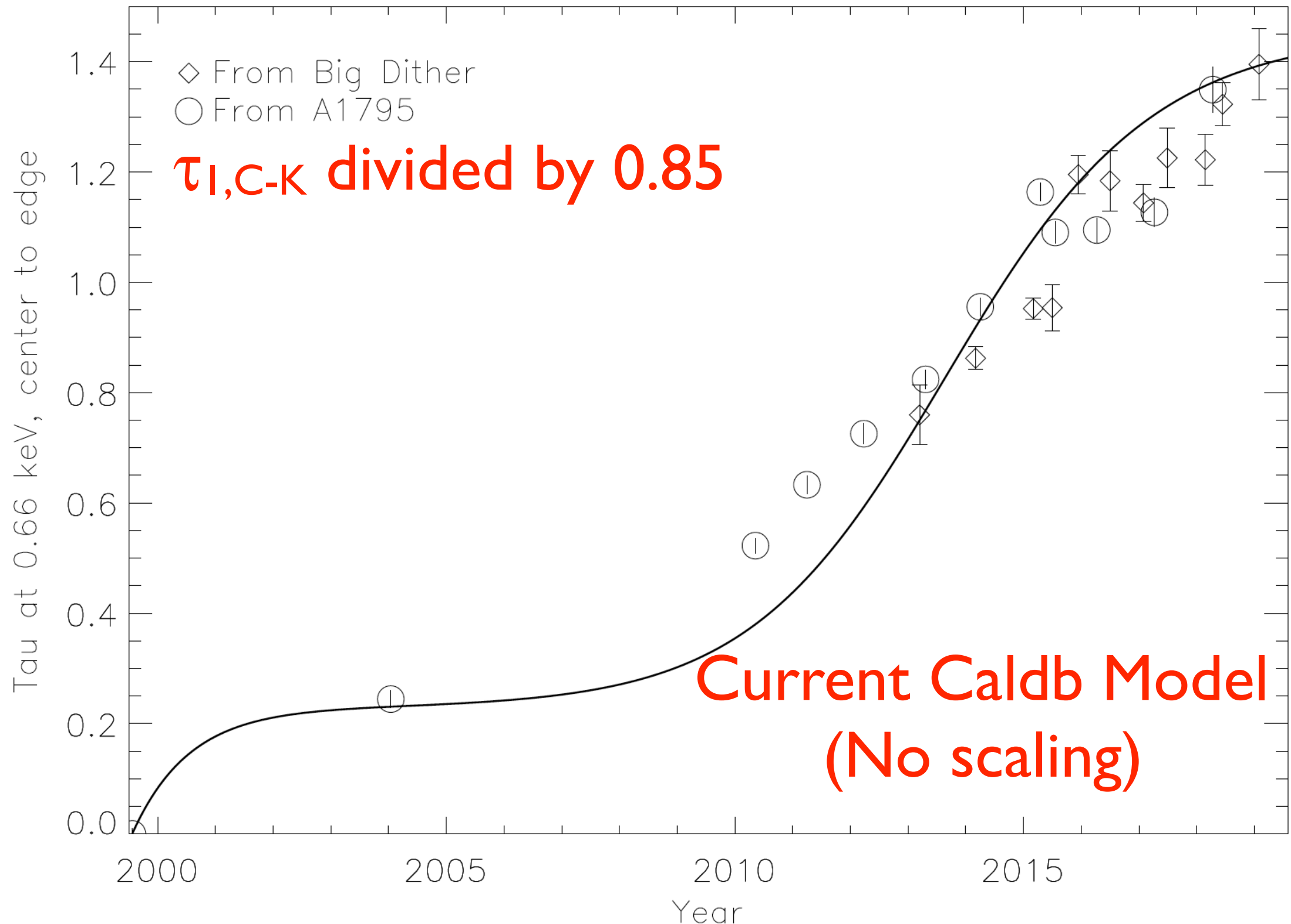
Compare to A1795

Scale = 120 row



Compare to A1795

Scale = 120 row



Summary

- Contaminant deposition continues
 - rate hasn't changed in 3-5 years
 - ACIS-I measurements are irregular, more like ACIS-S
- New material is different than in 2001-5
 - O/C is higher
 - 285 eV feature is present — it's aliphatic
- New material is more uniform than old
- Spatial dependence is still asymmetric, with more at high rows
- Model is OK now; to be updated this year