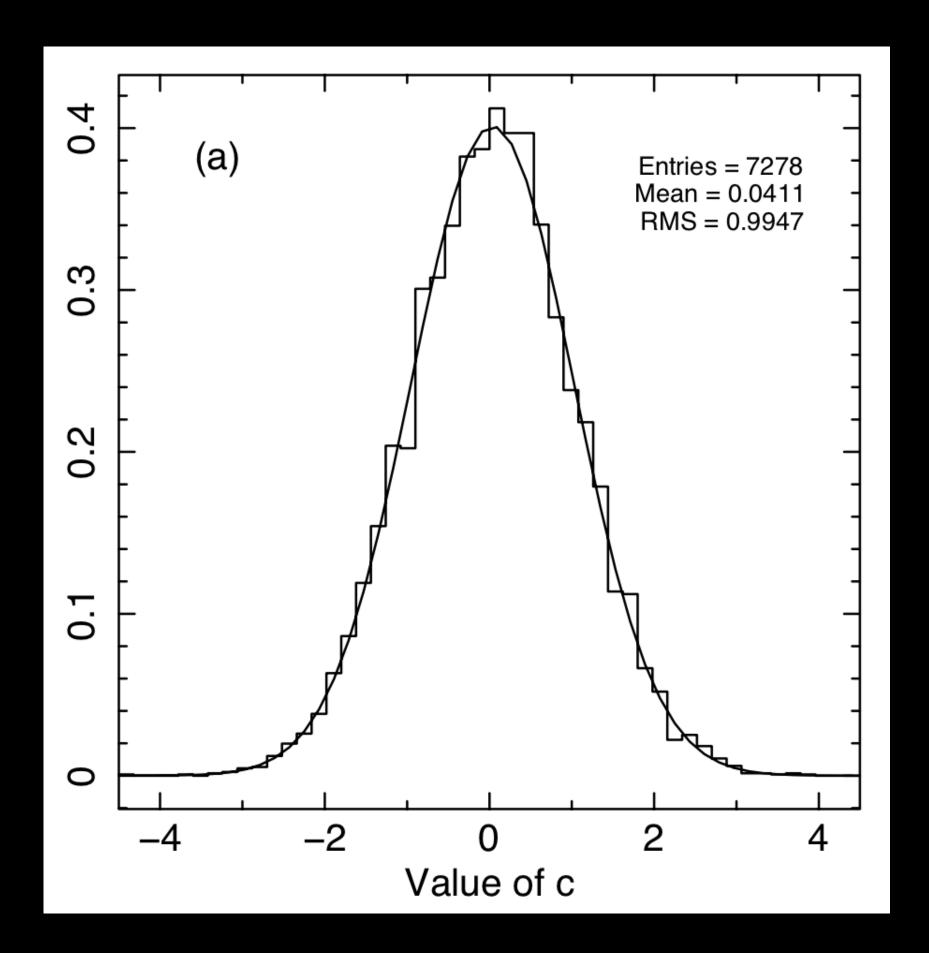
Treating the detection of transient phenomena as a statistical problem

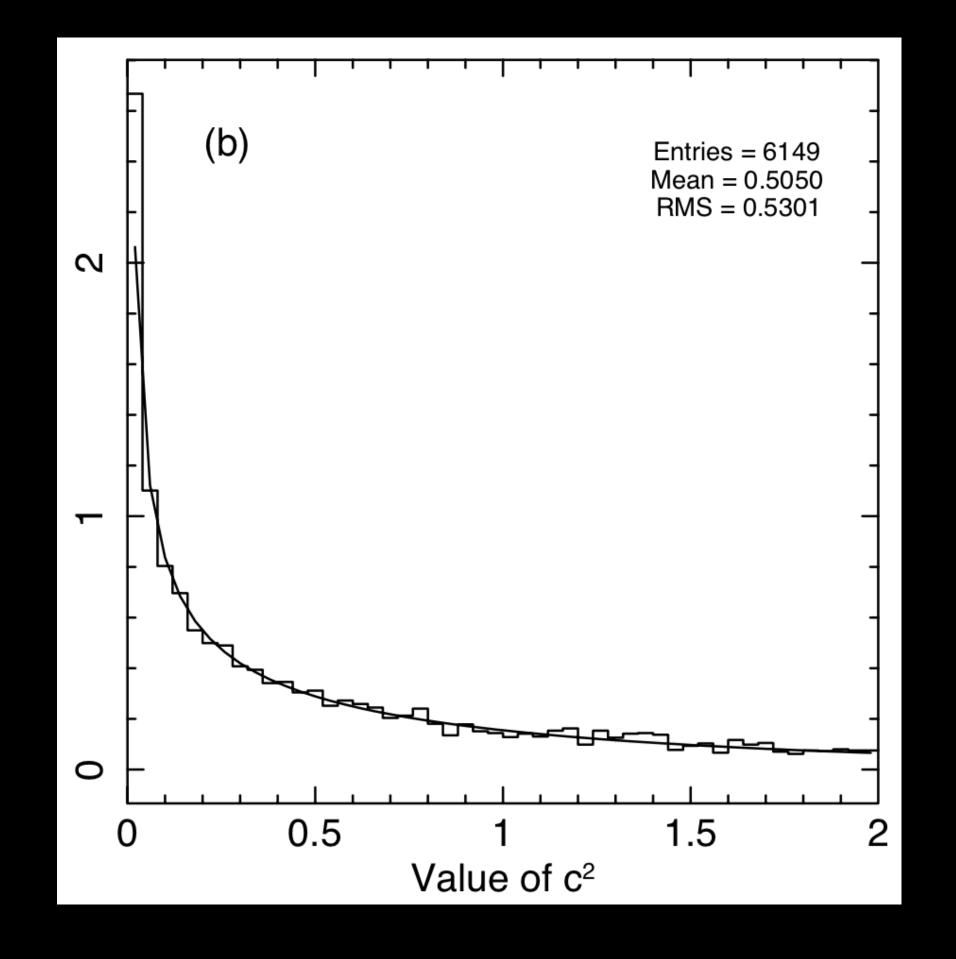
G. Belanger, ESA, ESAC

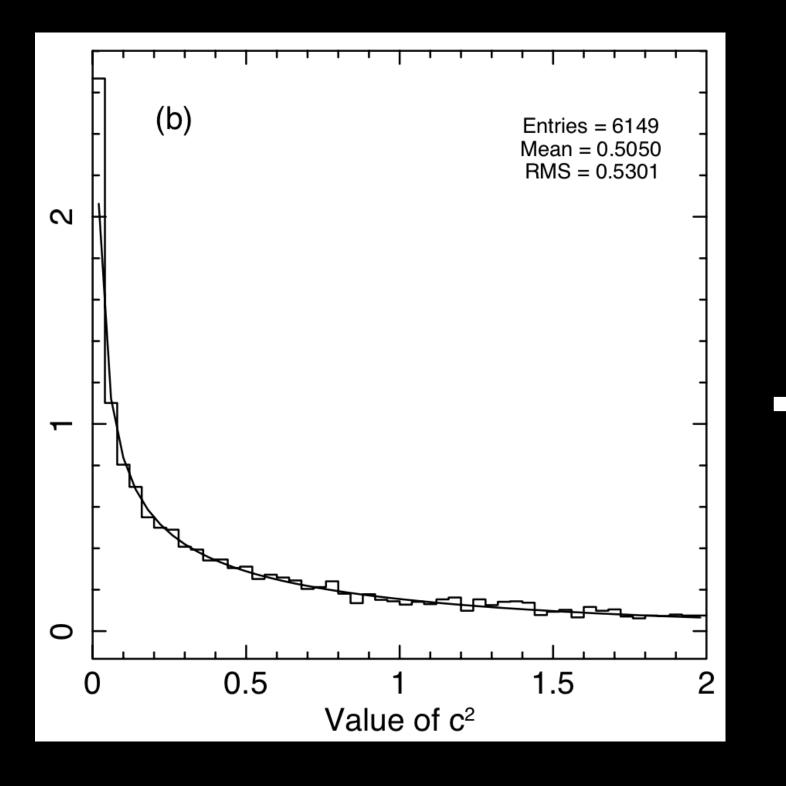
$$R^2 = 2N(C^2 + S^2)$$

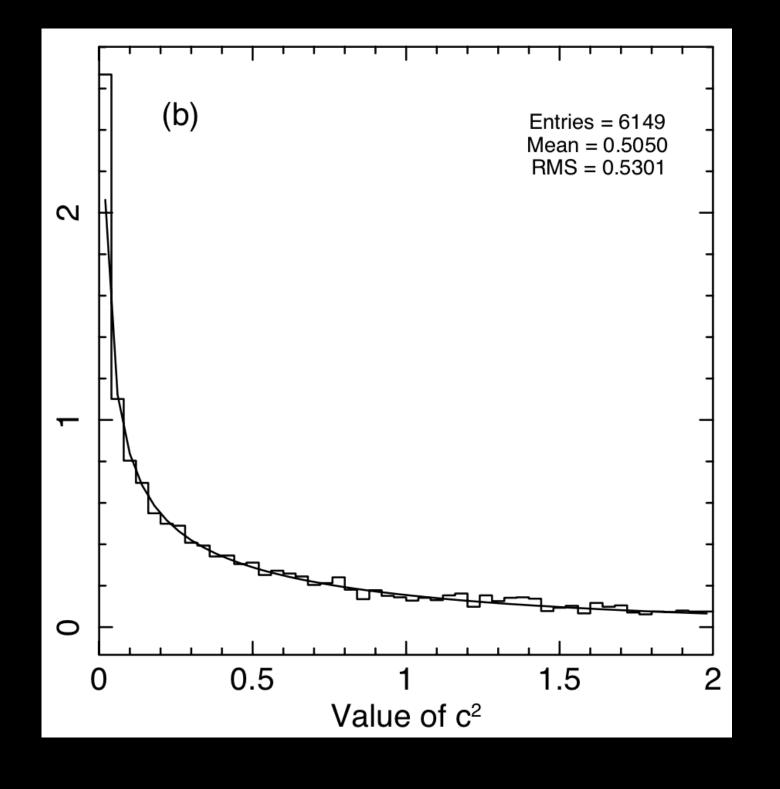
$$C = \frac{1}{N} \sum_{i=1}^{N} \cos \phi_i$$
 and $S = \frac{1}{N} \sum_{i=1}^{N} \sin \phi_i$.

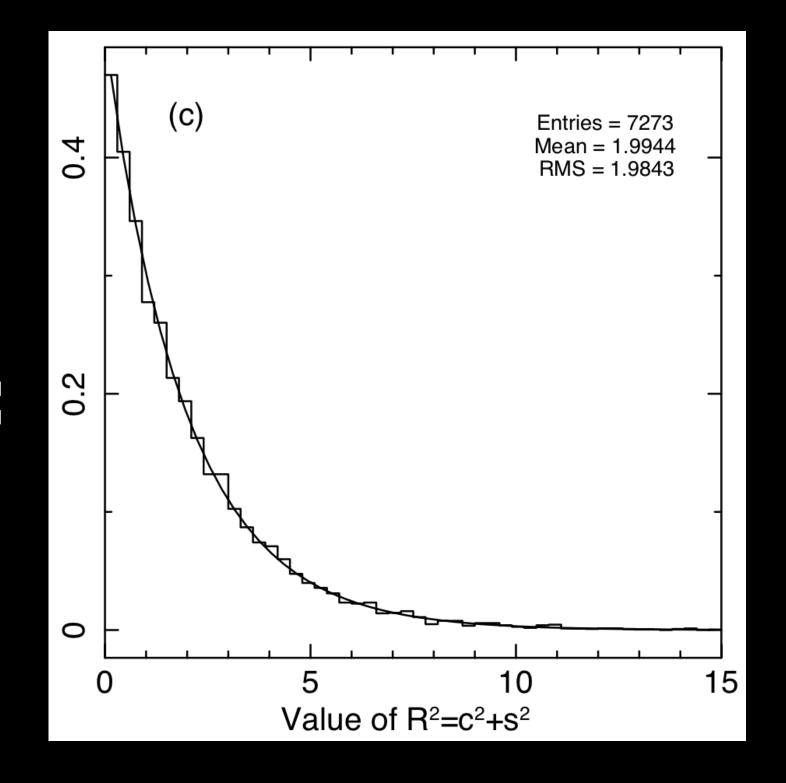


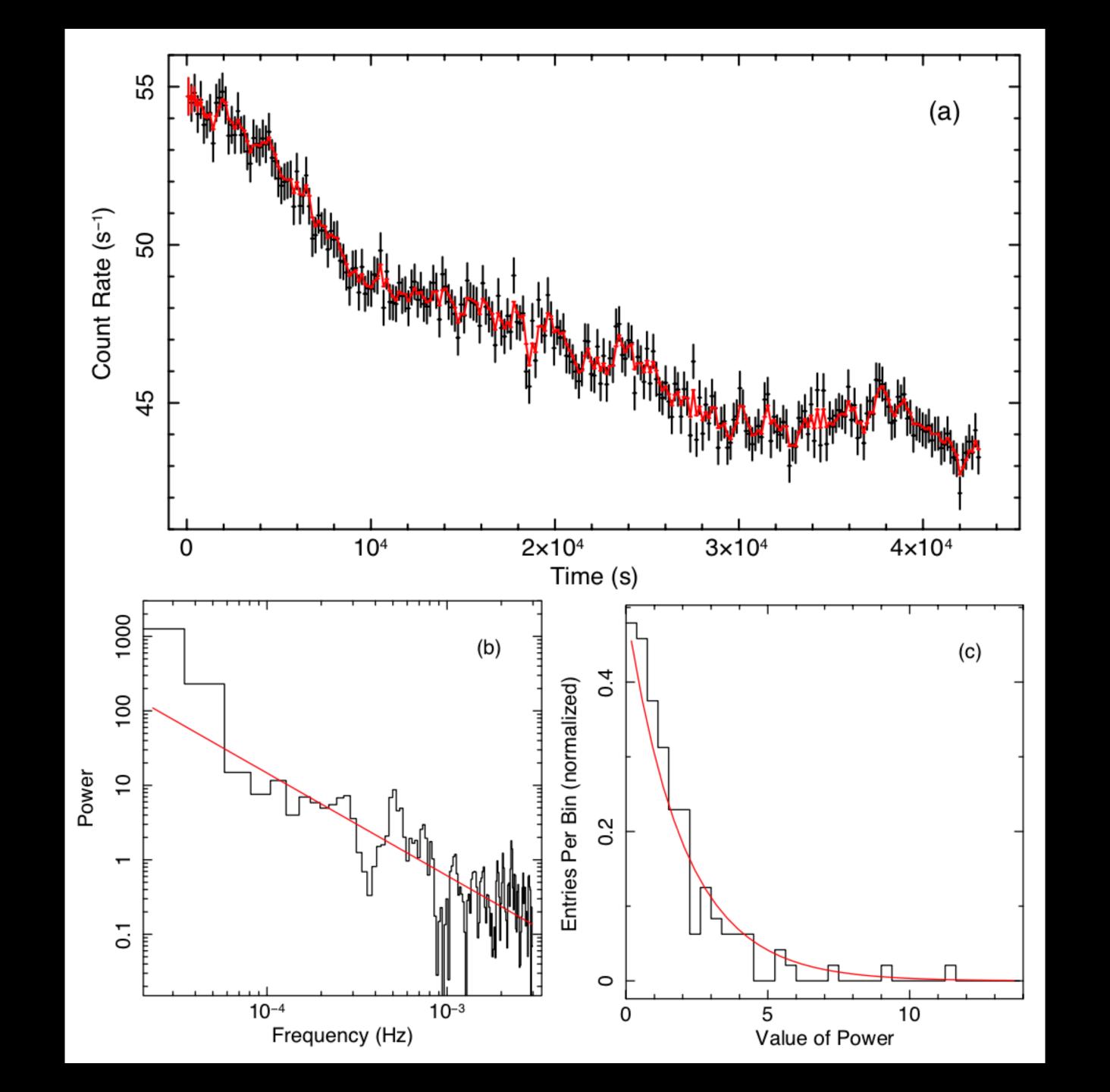


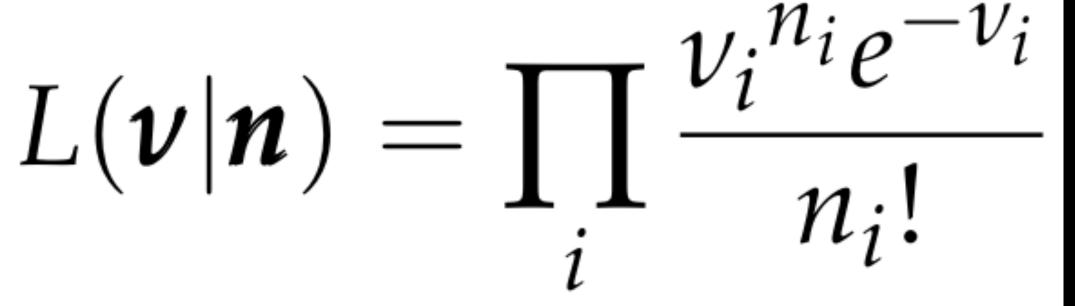


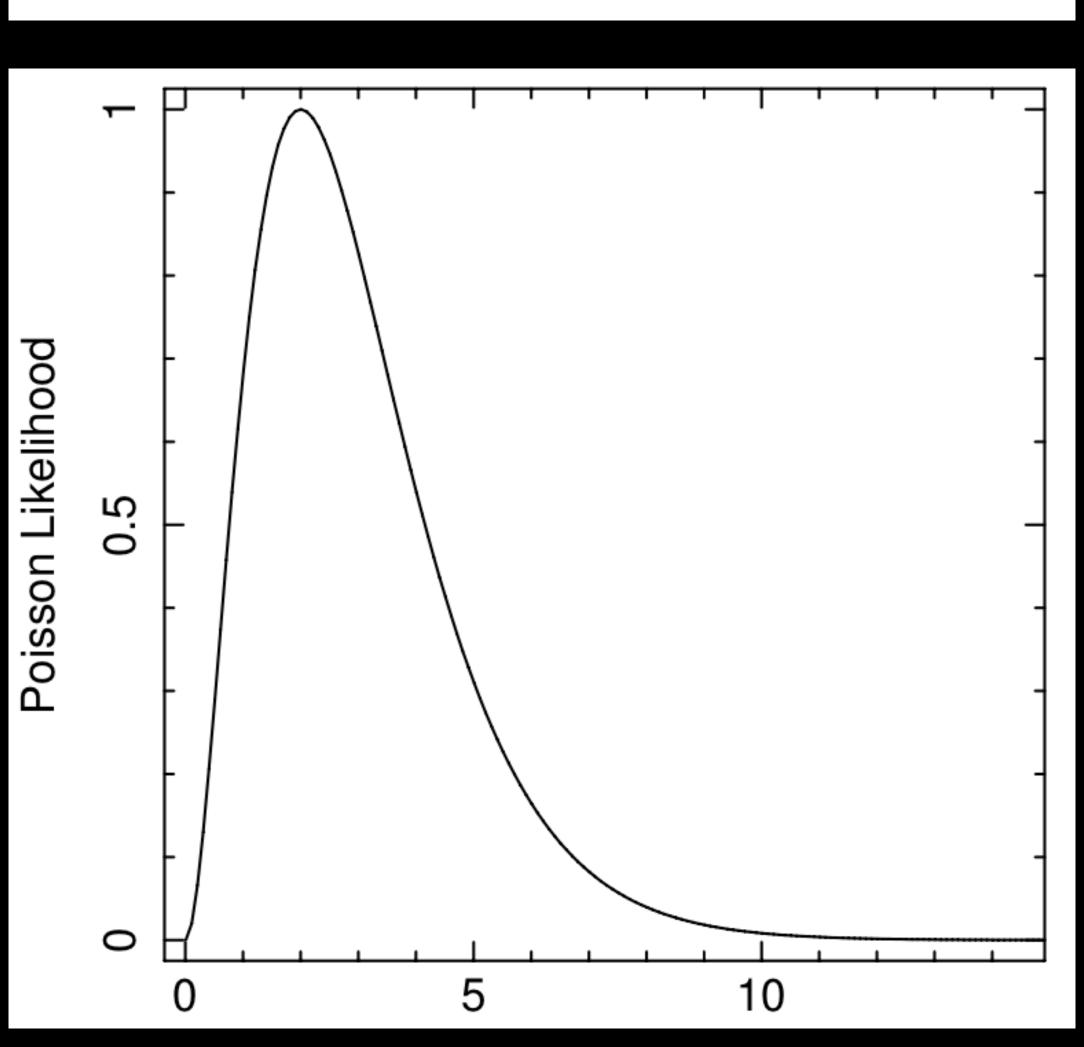




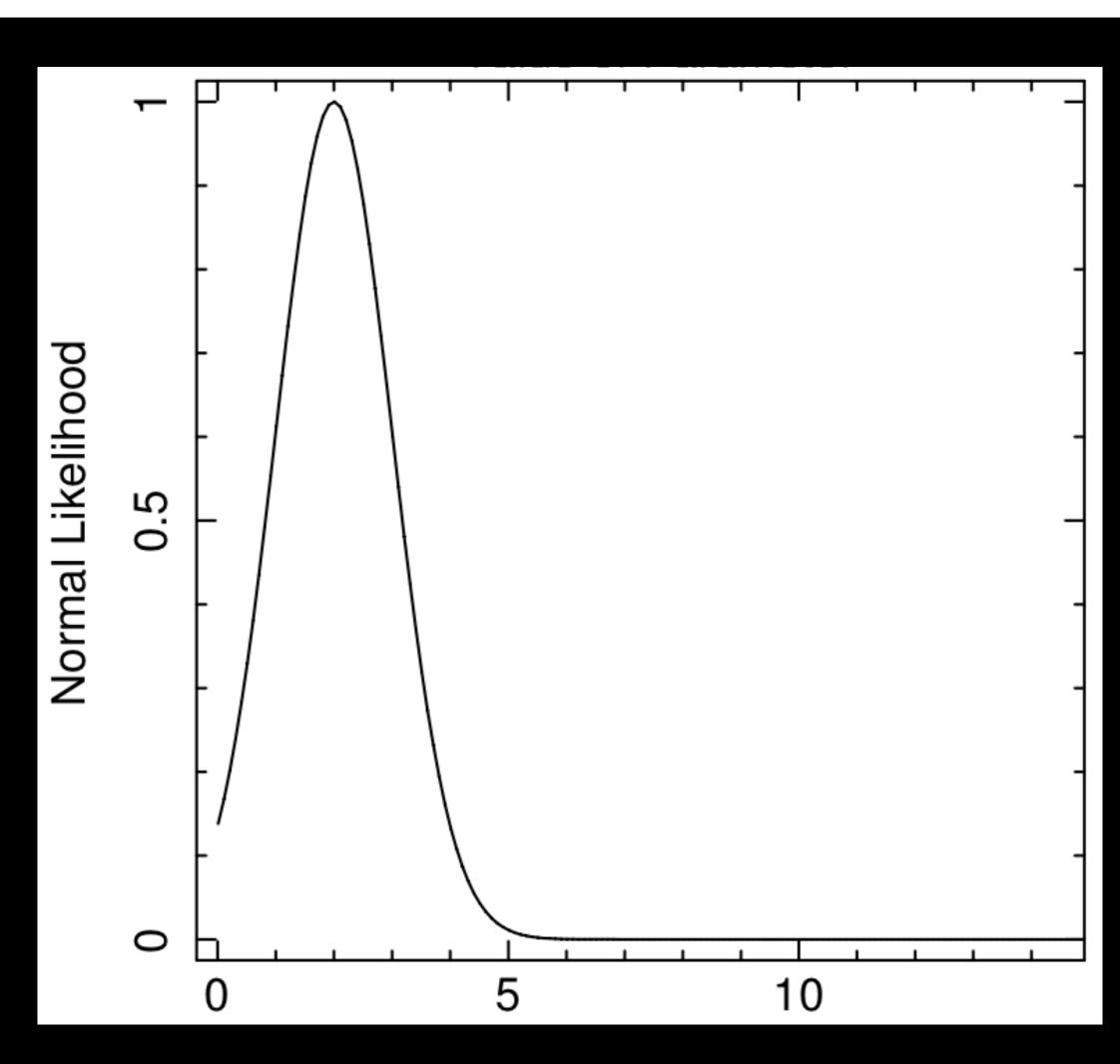




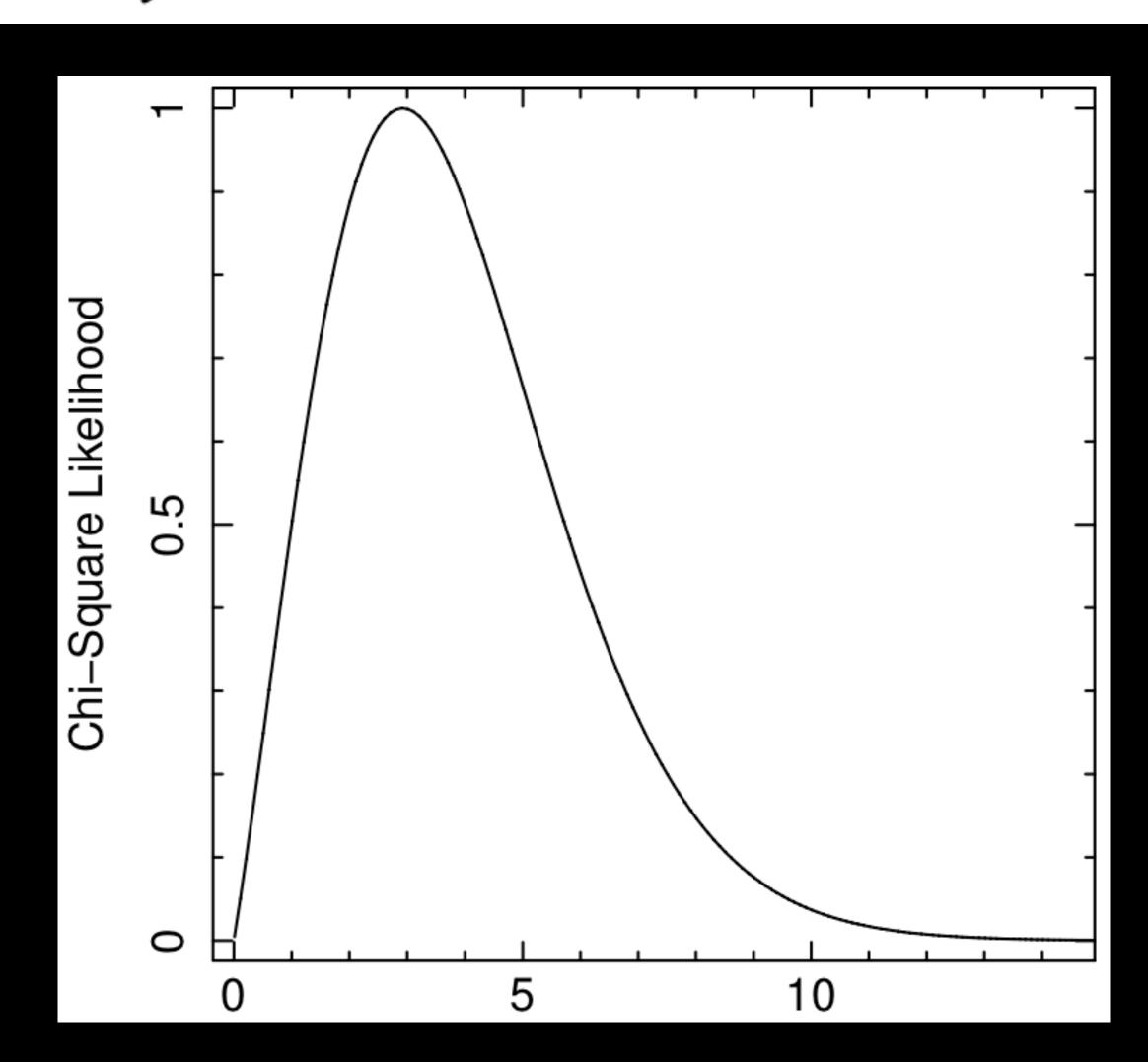


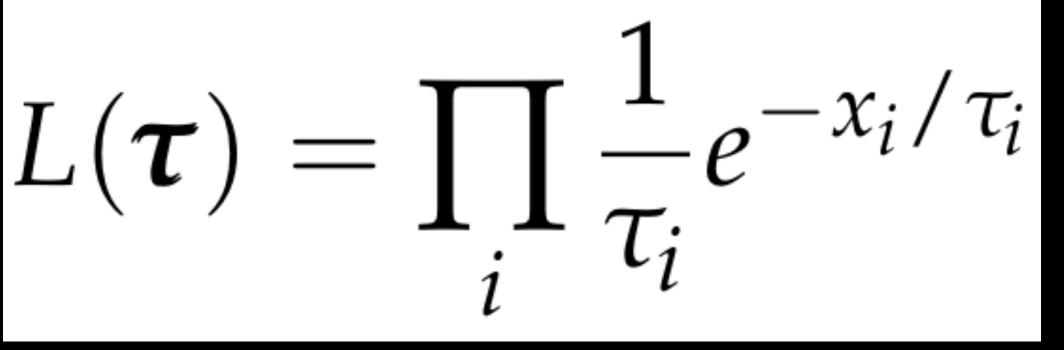


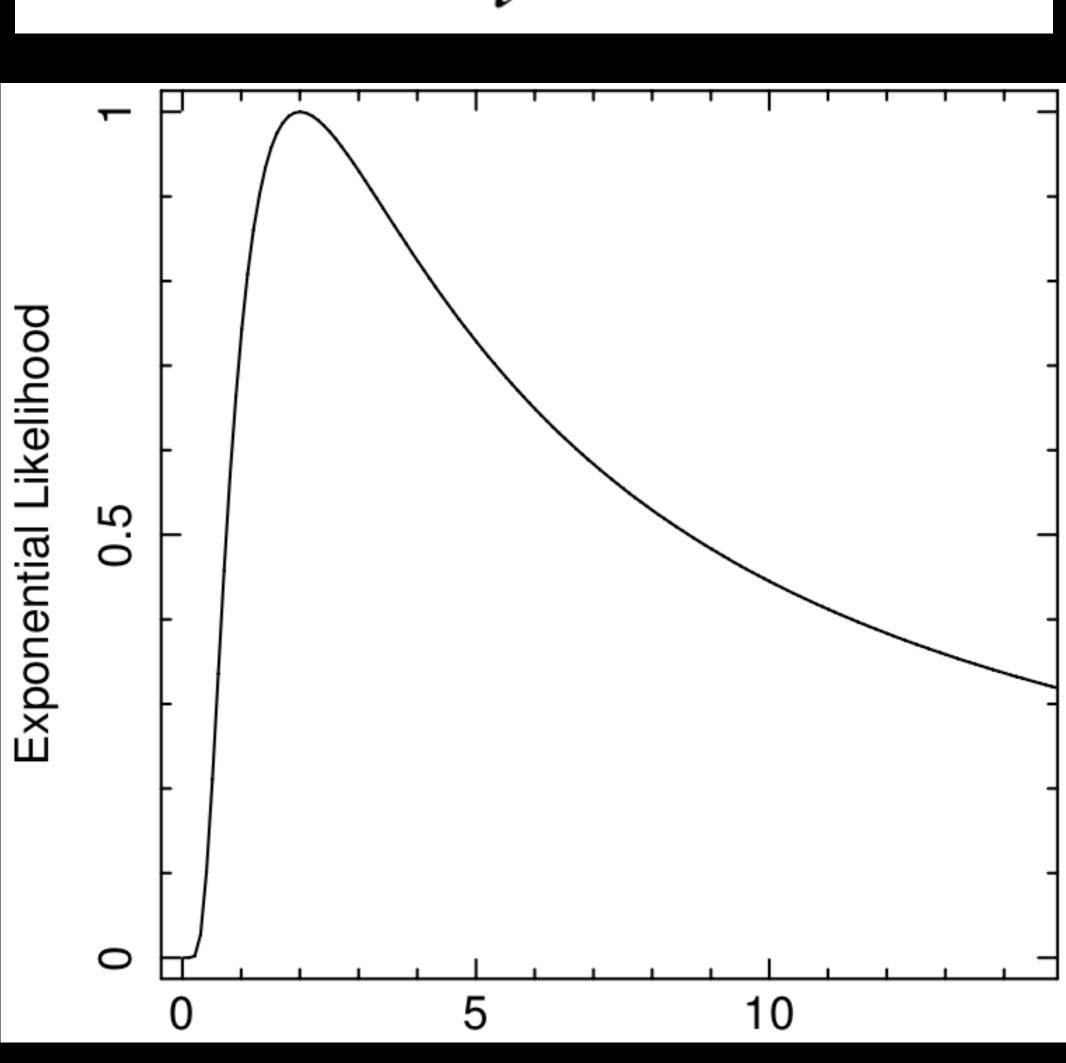
$L(\boldsymbol{\mu}, \boldsymbol{\sigma} | \boldsymbol{x}) = \prod_{i} \frac{1}{\sqrt{2\pi\sigma_i^2}} e^{-(x_i - \mu_i)^2/2\sigma_i^2}$



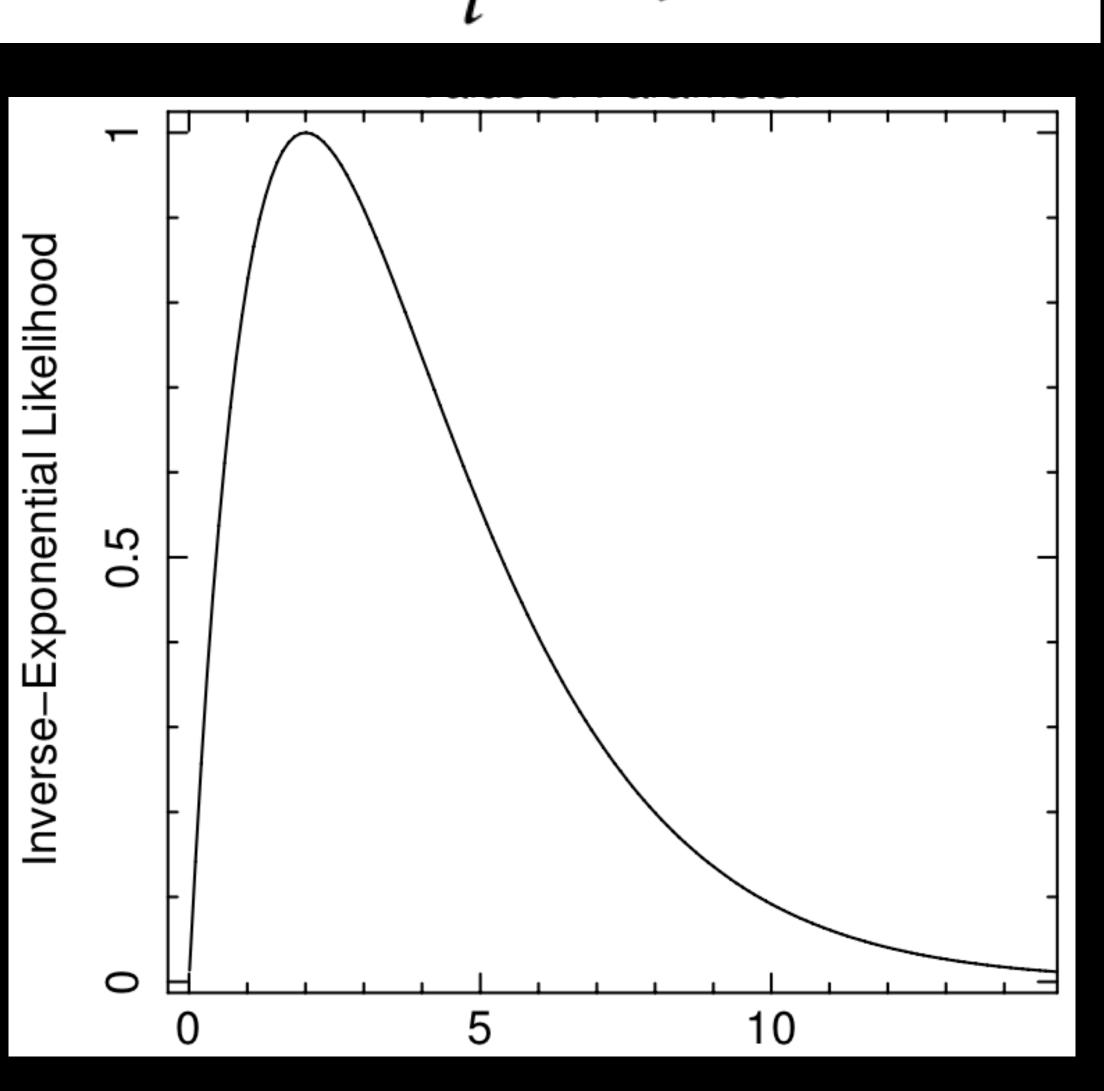
$L(\mathbf{k}|\mathbf{x}) = \prod_{i} \frac{1}{2^{k_i/2} \Gamma(k_i/2)} x_i^{k_i/2-1} e^{-x_i/2}$

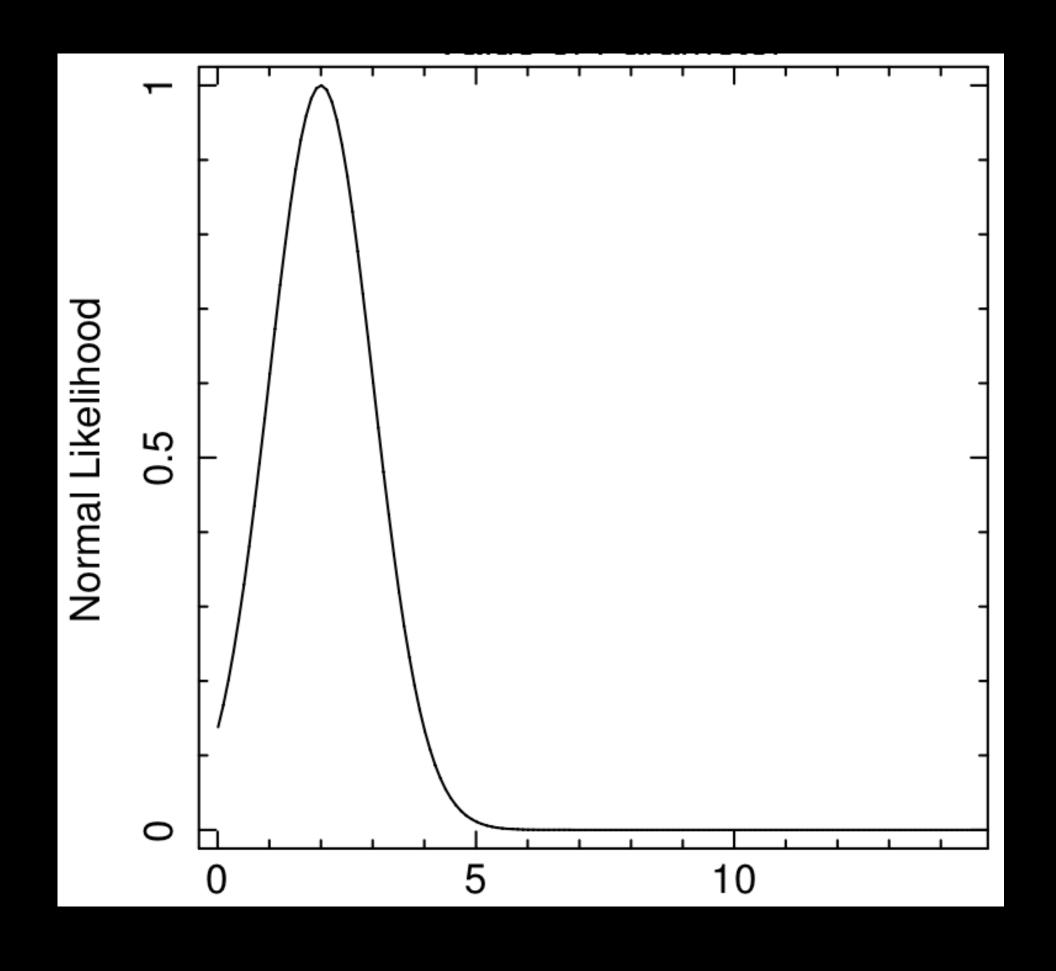


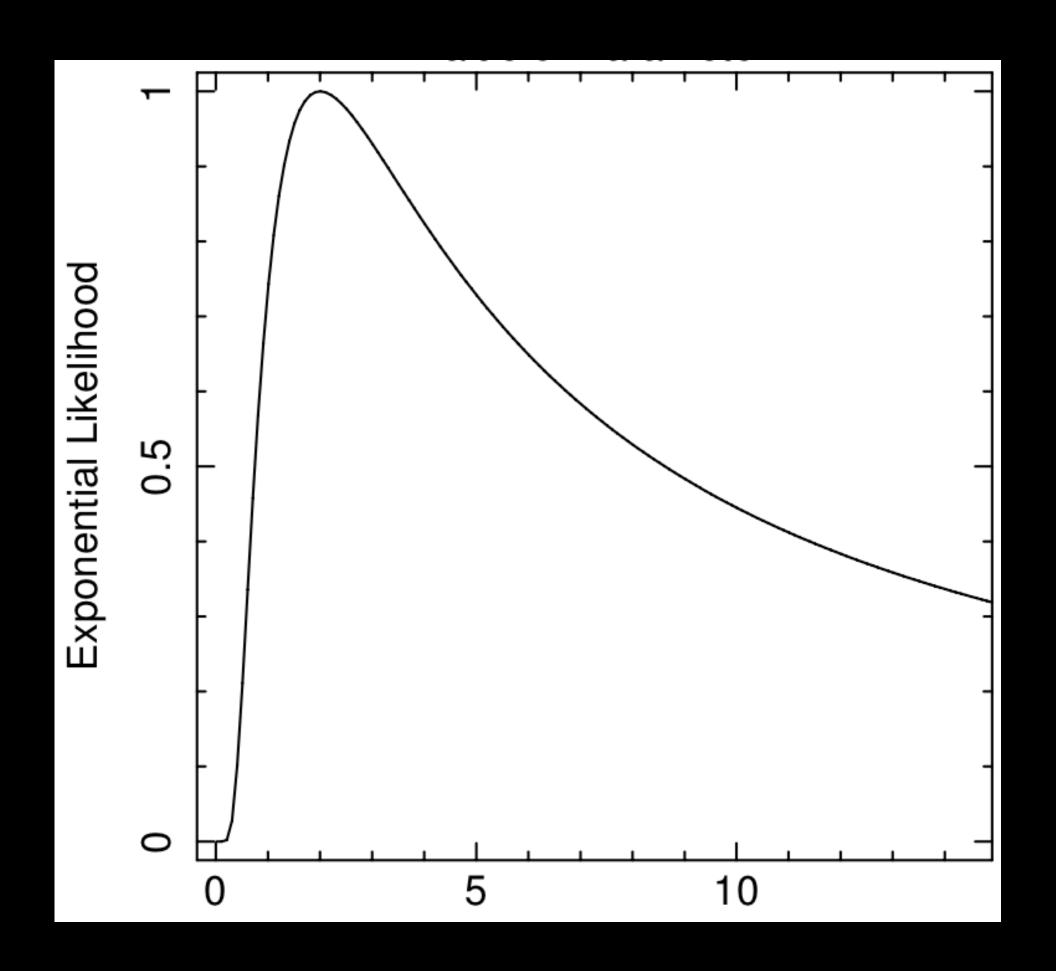


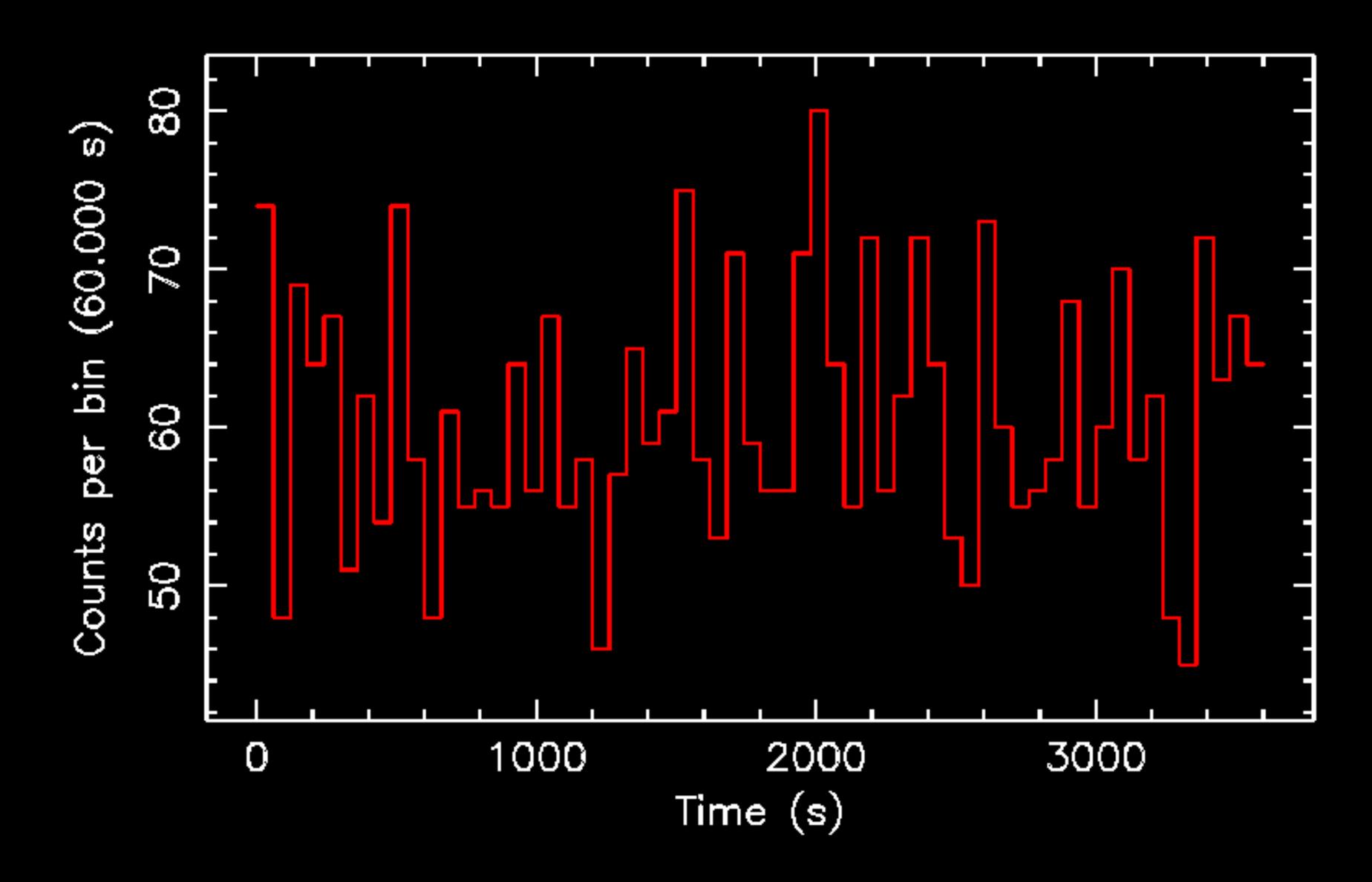


$L(t) = \prod_{i} \frac{t_i}{x_i^2} e^{-t_i/x_i}$

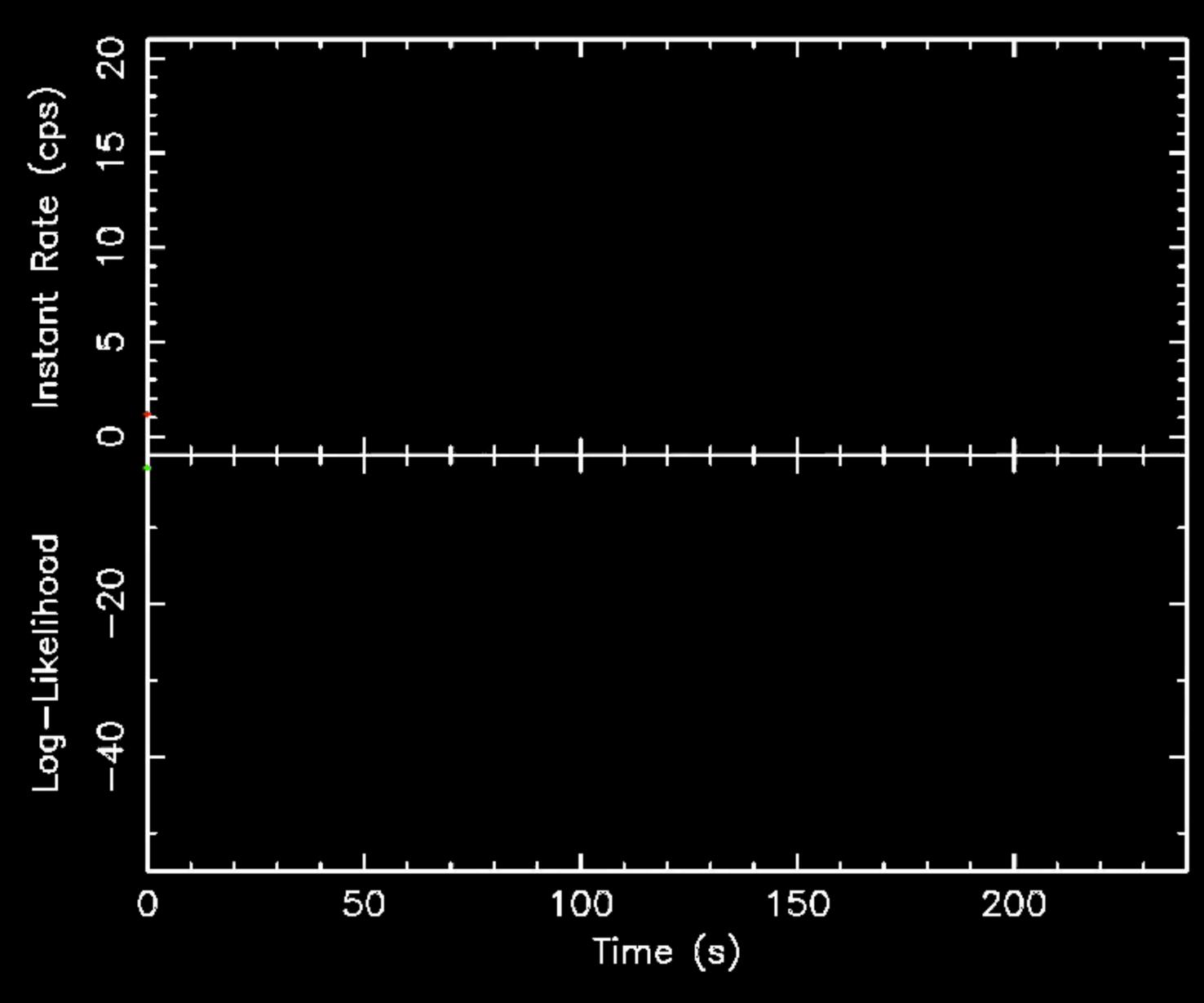




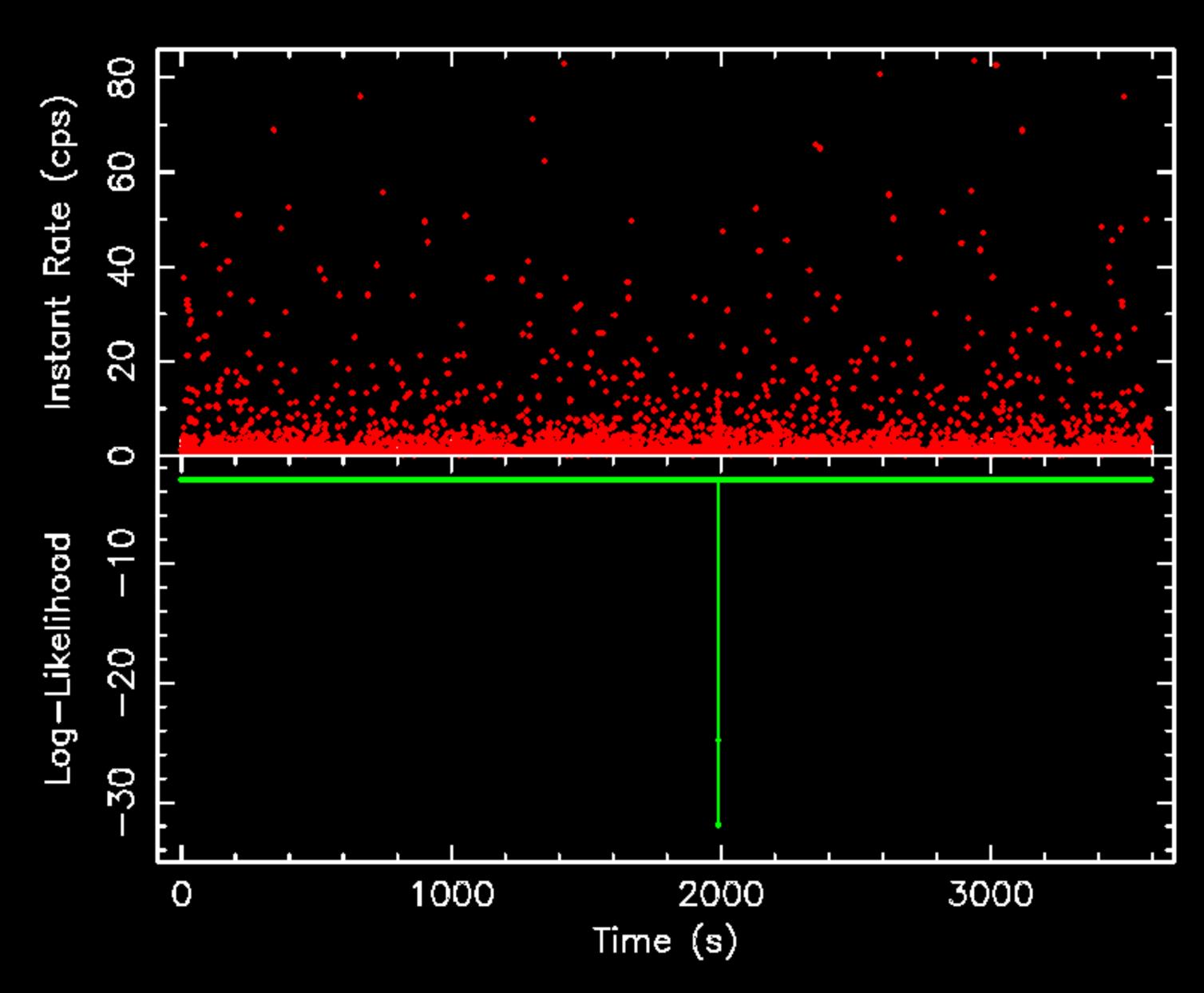




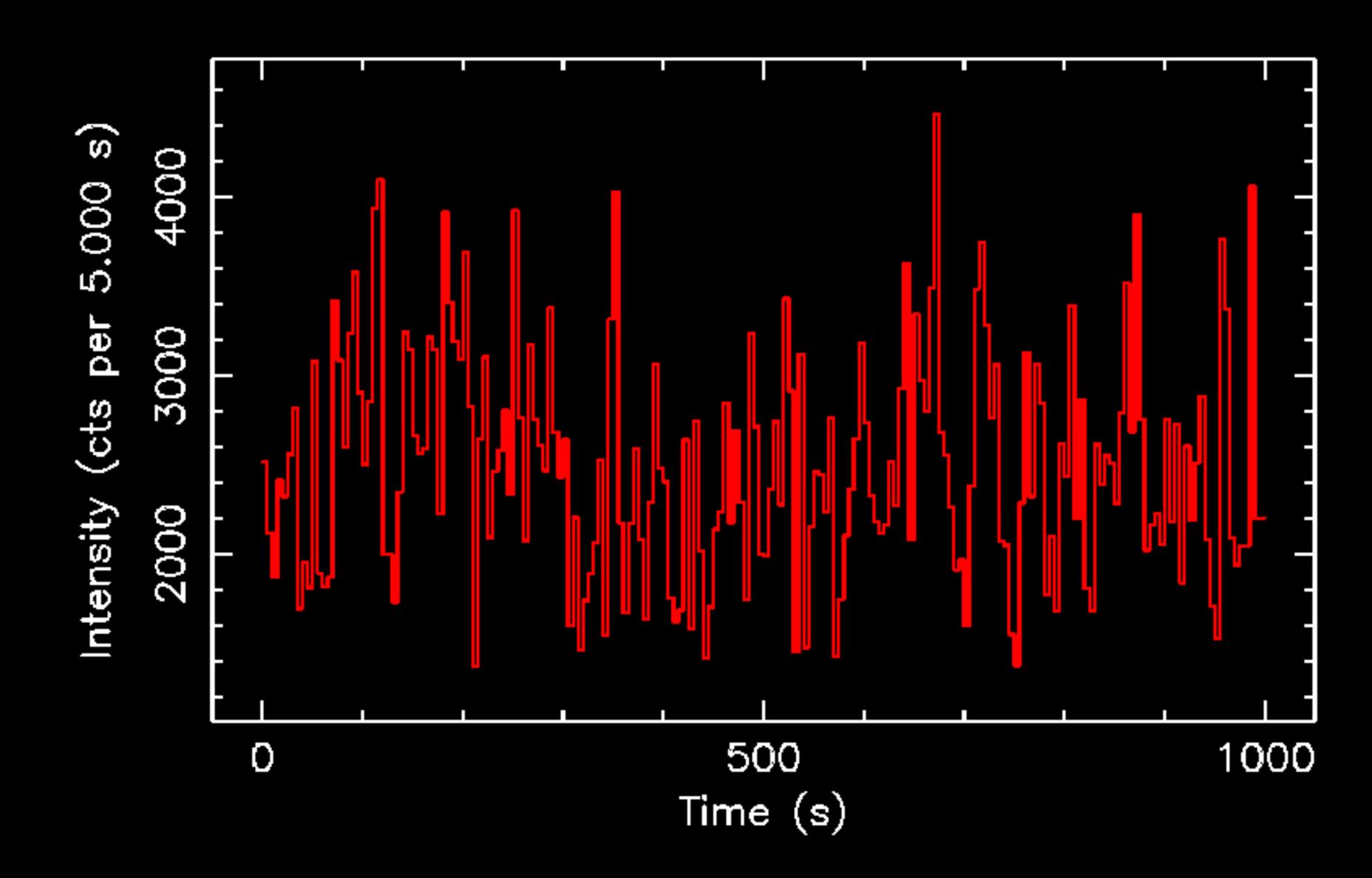
T=3600 s, mu=1 cps, dt=60 s



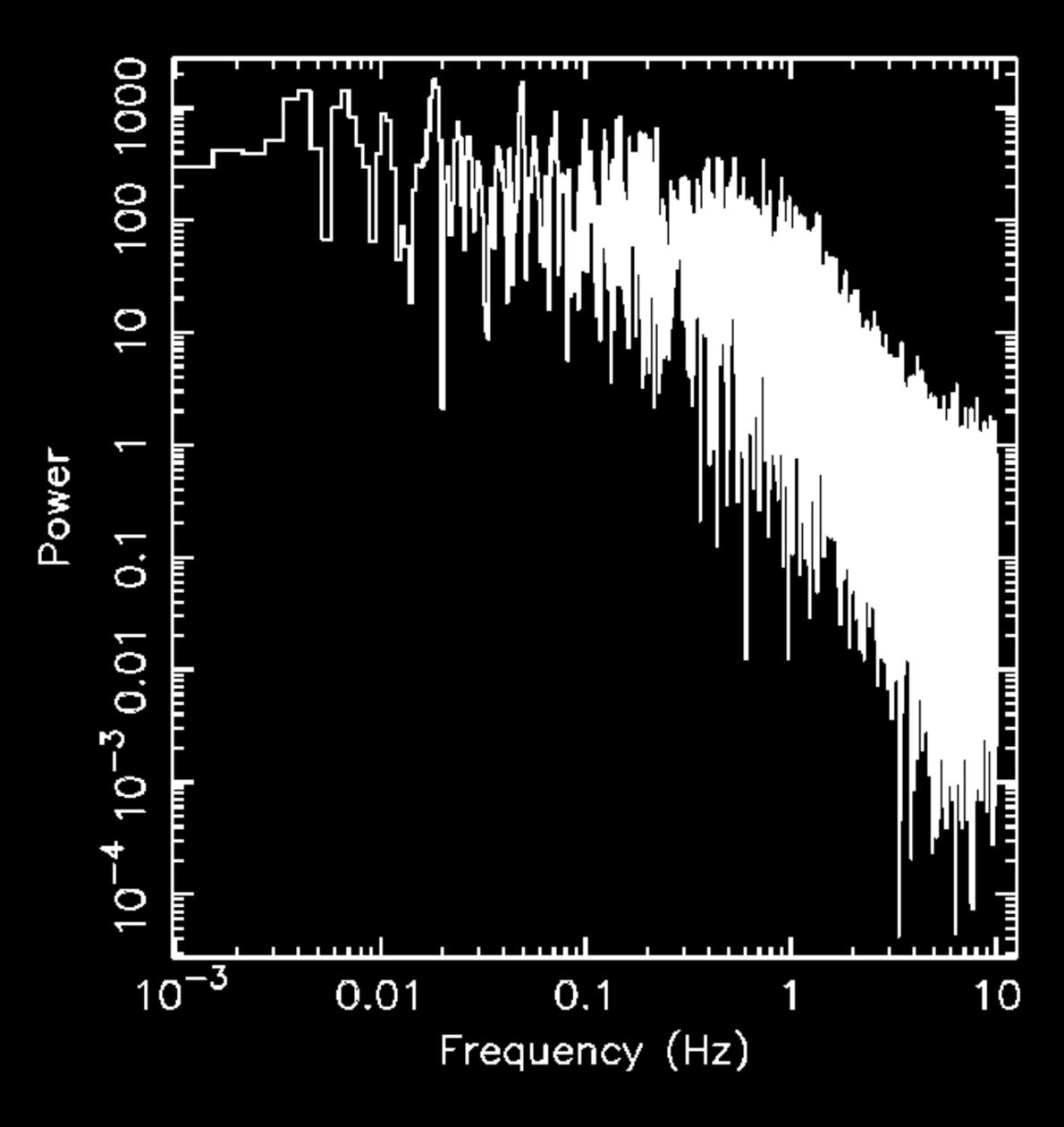
Neutron star: Burst at 2000 s



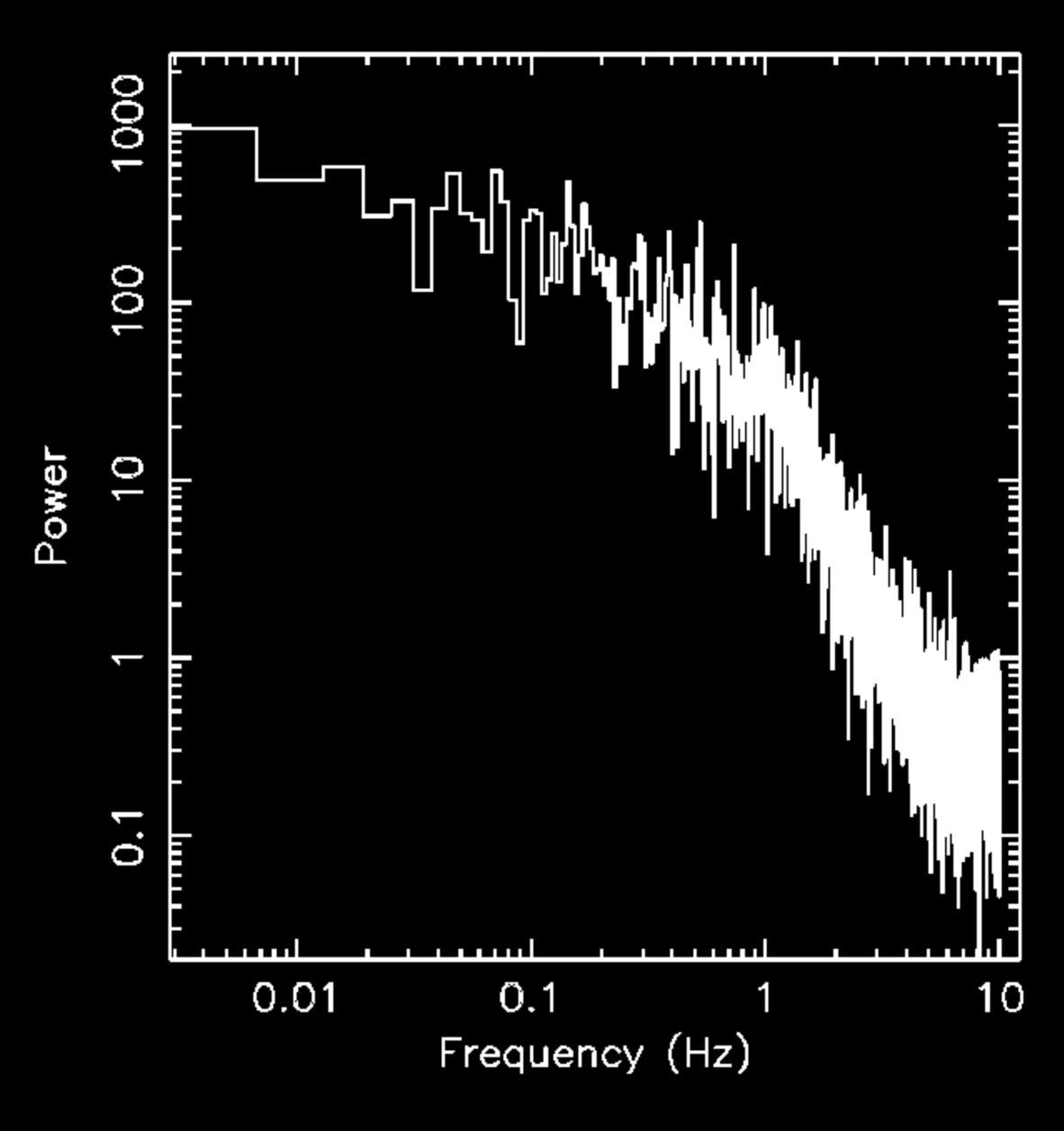
Burst: 30 s, 33 events



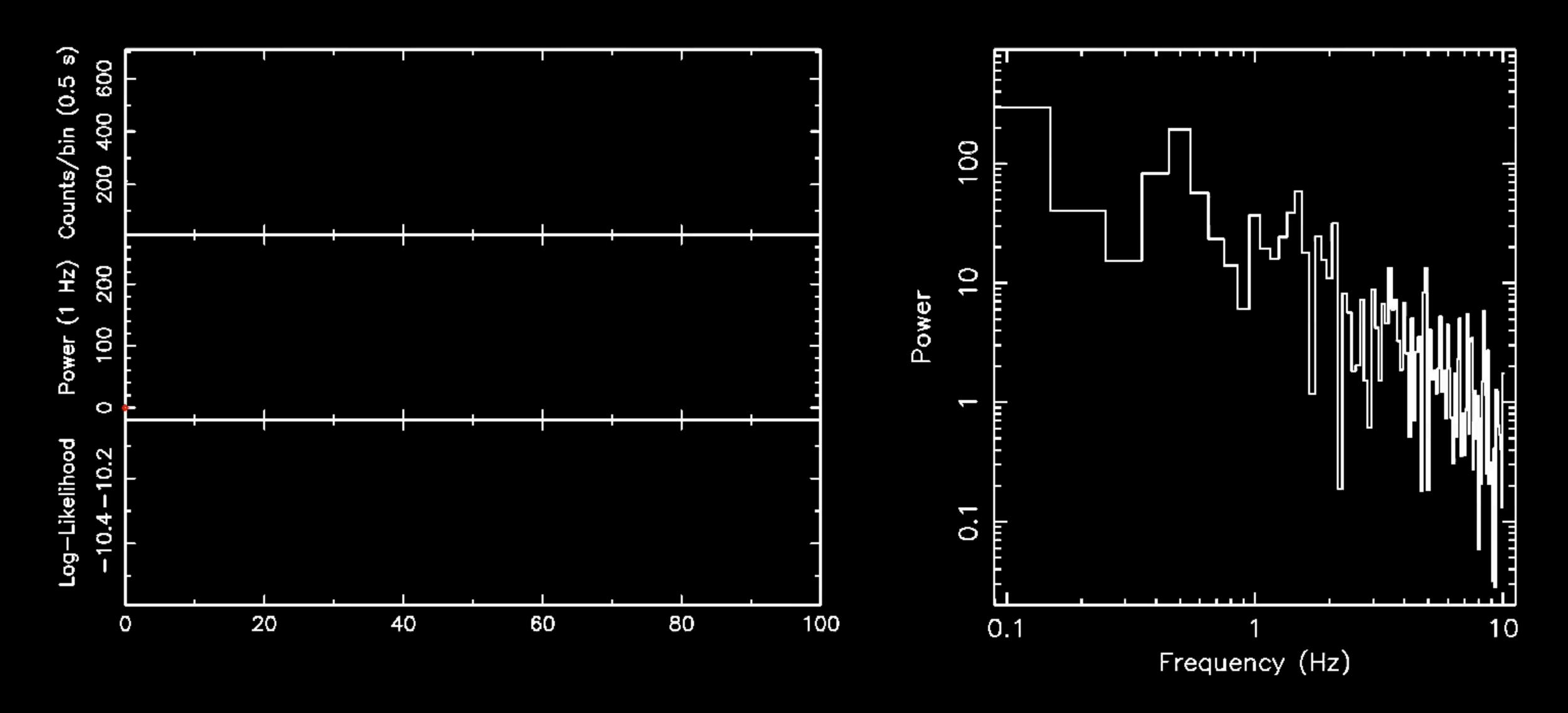
Black Hole: T=1000 s, mu=400 cps, dt=5 s



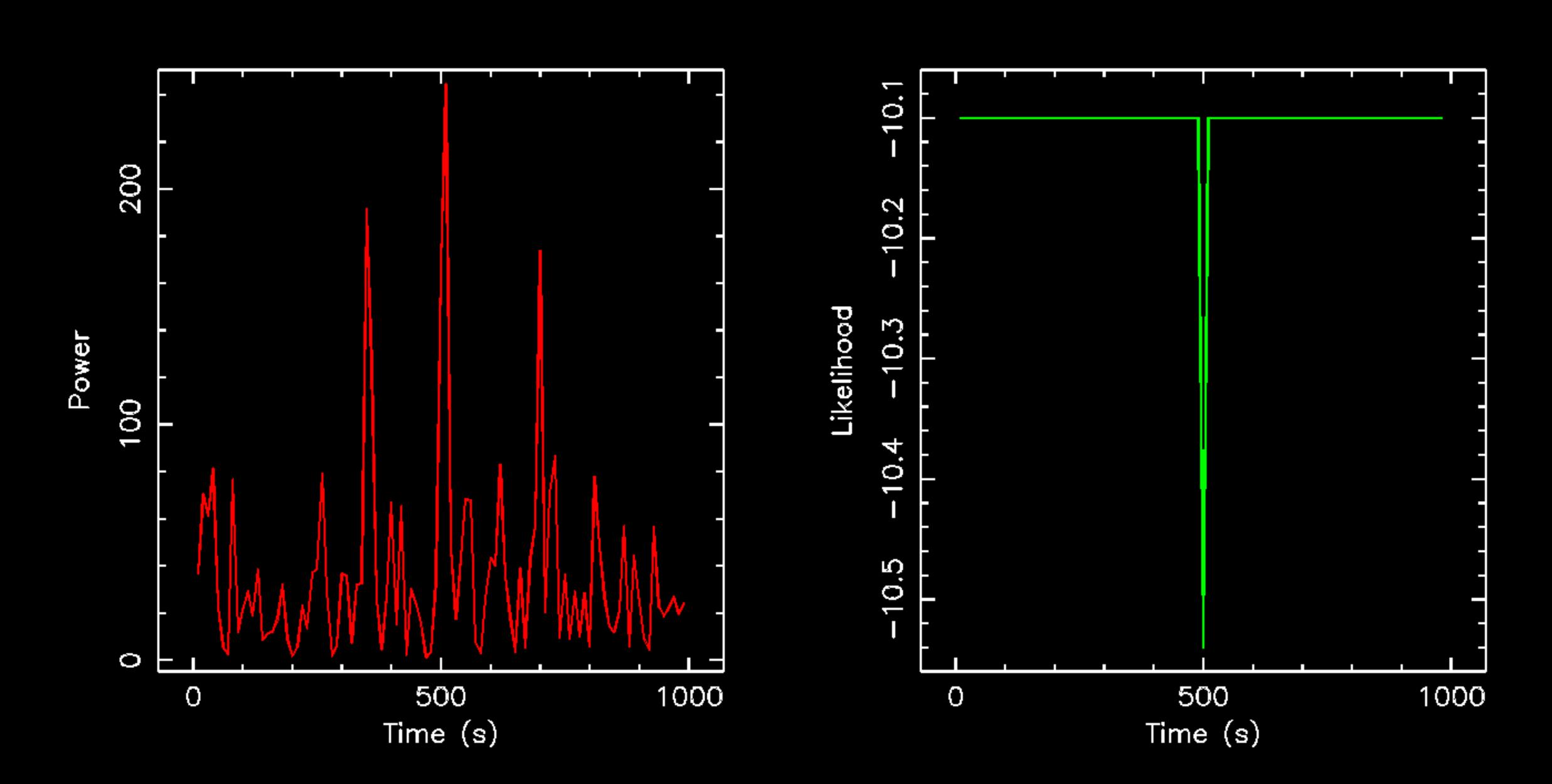
Two-zones with transition @ 1 Hz

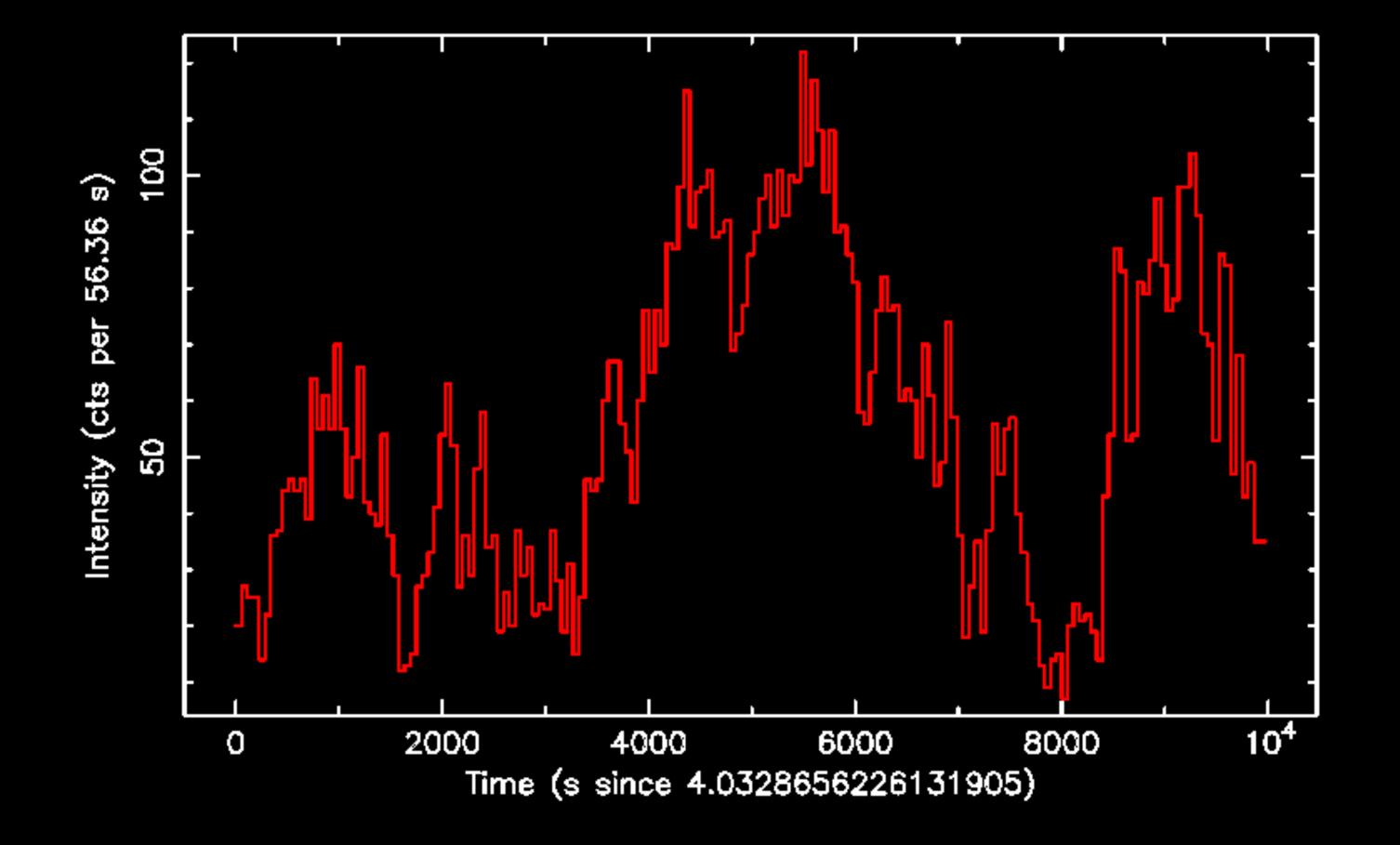


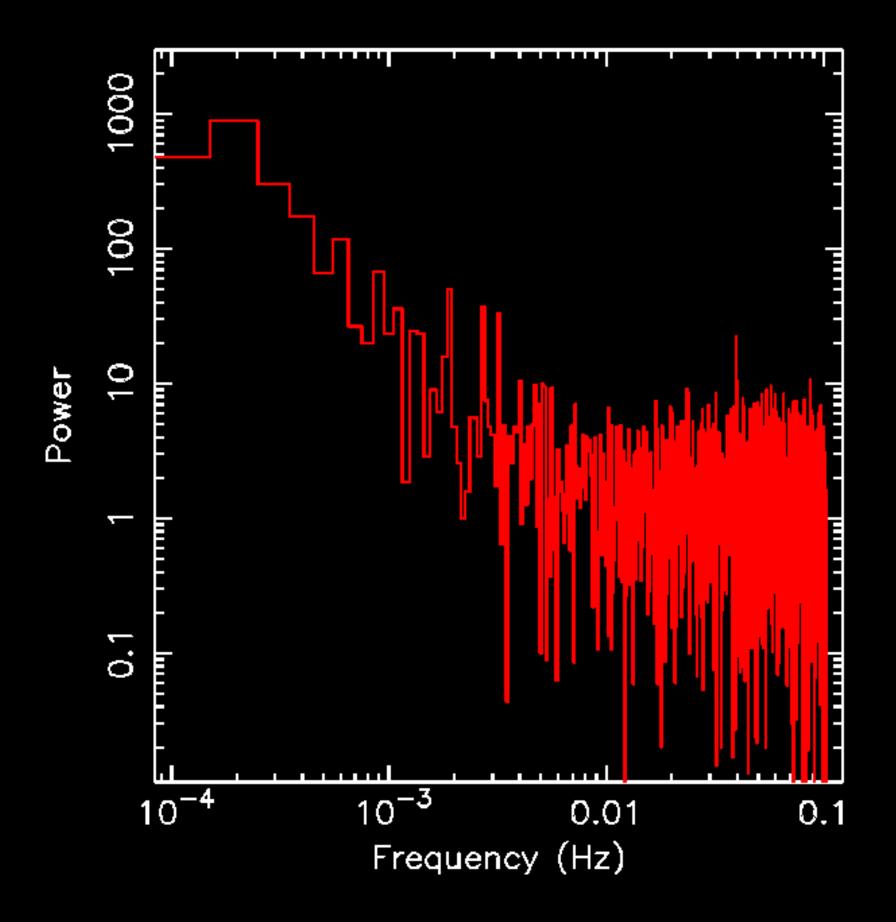
Kalman filtered data

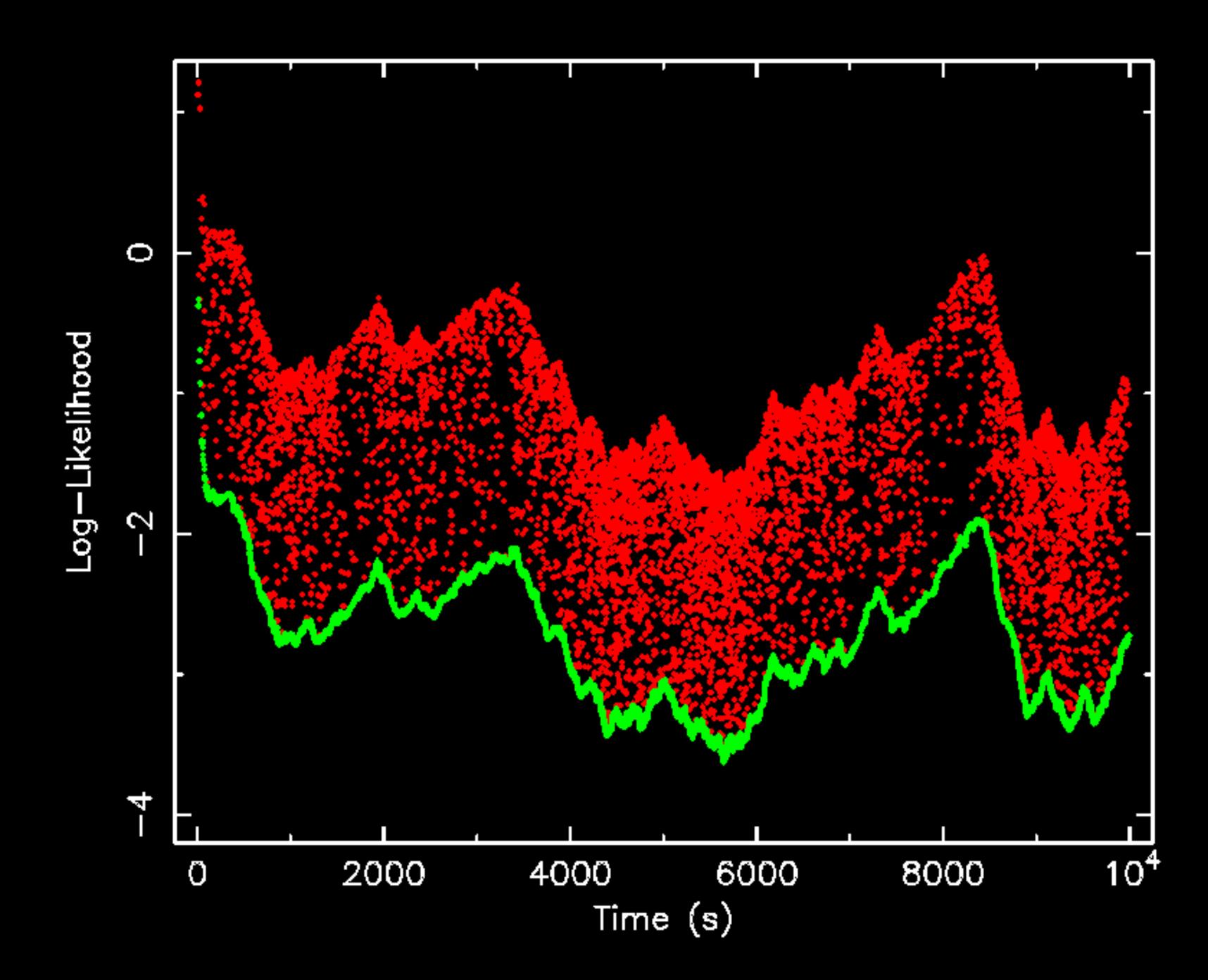


QPO @ 1 Hz: t = 500 s, dt = 30 s (3 points)









Thank you for listening