

XMM-Newton — Chandra Cross-Calibration with Blazars

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- 25 XMM-Newton observations coordinated with Chandra:
 - 40 strictly simultaneous exposures for flux comparison
- Instruments being compared are:
 - **EPIC, RGS, ACISS-L/HETG, HRCS-LETG**
- Data reduction:
 - SAS 17 + CCFs as of Jan 2019
 - CIAO 4.10 + CALDB 4.8.1

➤ Energy bands:

- 0.15 – 0.33 keV (Lower EPIC - Lower RGS bound)
- 0.33 – 0.54 keV (Up to the O-edge)
- 0.54 – 0.85 keV
- 0.80 – 1.20 keV } O-VII/VIII , Ne-IX/X
- 1.20 – 1.50 keV }
- 1.50 – 1.82 keV (Up to the Si-edge)
- 1.82 – 2.20 keV (Up to the Au-edge)
- 2.20 – 3.50 keV
- 3.50 – 5.50 keV
- 5.50 – 10.0 keV

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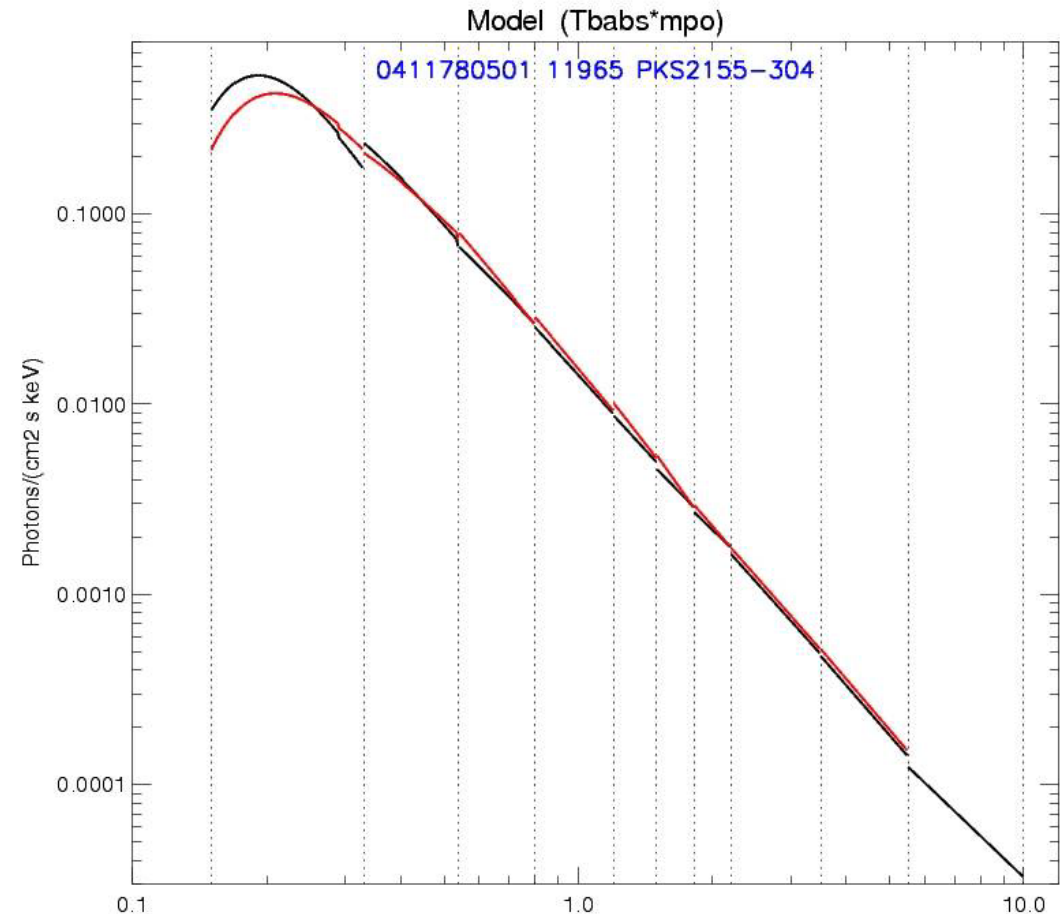
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➤ Spectral fitting: model consists of:

- multiple independent power laws
- absorption with nH fixed
 - PKS 2155-304: $1.42 \times 10^{20} \text{ cm}^{-2}$
 - 3C 273: $1.79 \times 10^{20} \text{ cm}^{-2}$
 - H 1426+428: $1.36 \times 10^{20} \text{ cm}^{-2}$

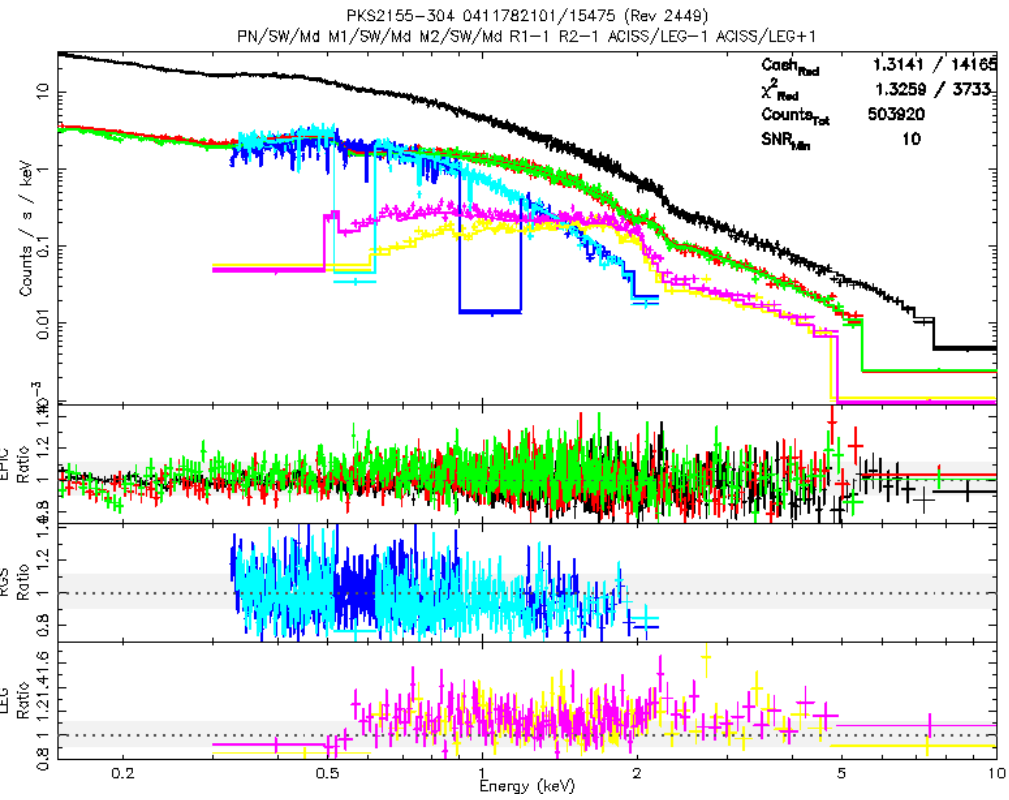
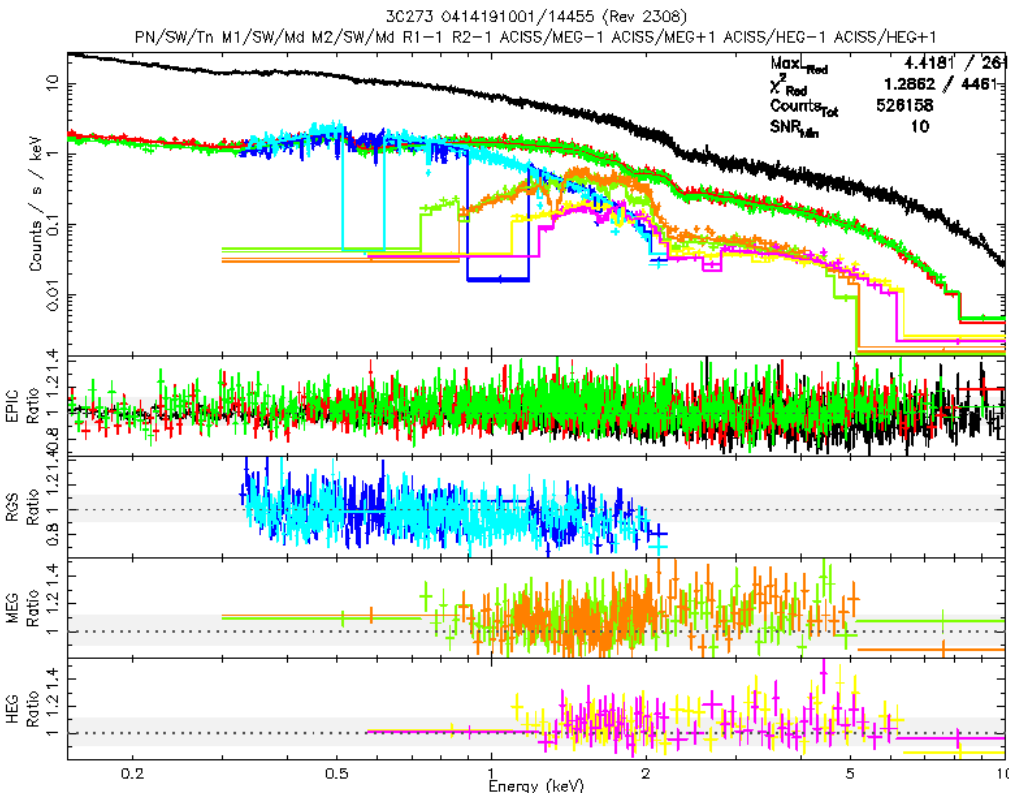
➤ Per simultaneous exposure:

- fit each instrument independently
- determine band fluxes from resulting best fits
- normalise to the PN flux



3C 273

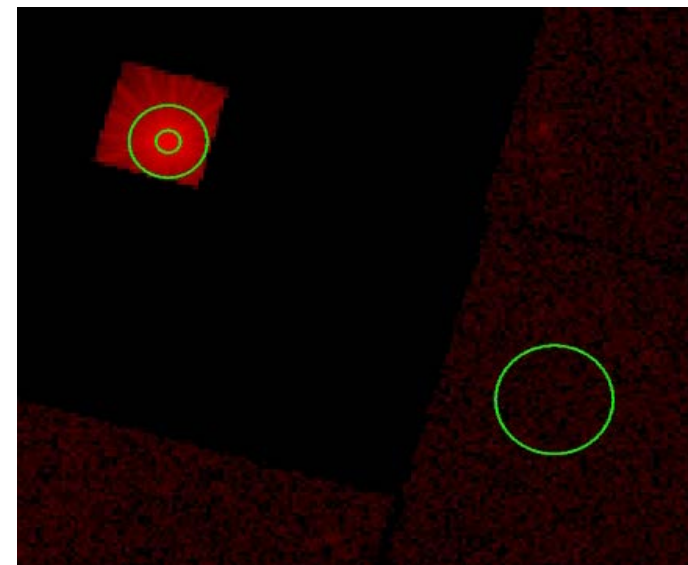
PKS 2155-304



Systematic uncertainties:

➤ Pile-up:

- EPIC requires excision of PSF core: use source extraction annuli.
- Per observation: for both MOSs use the largest common outer radius within window, and a common inner radius.
- However, radii vary from observation to observation, and are generally different from the PN radii.
- Differing annuli may introduce systematic uncertainties due to imperfect EE correction and RMF weighting.



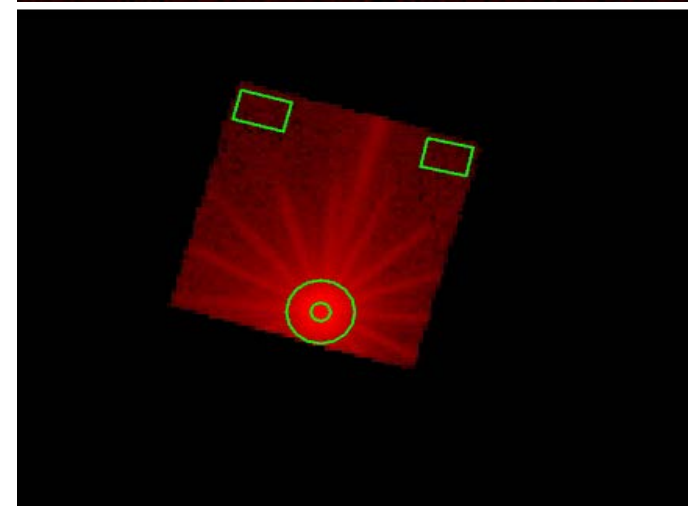
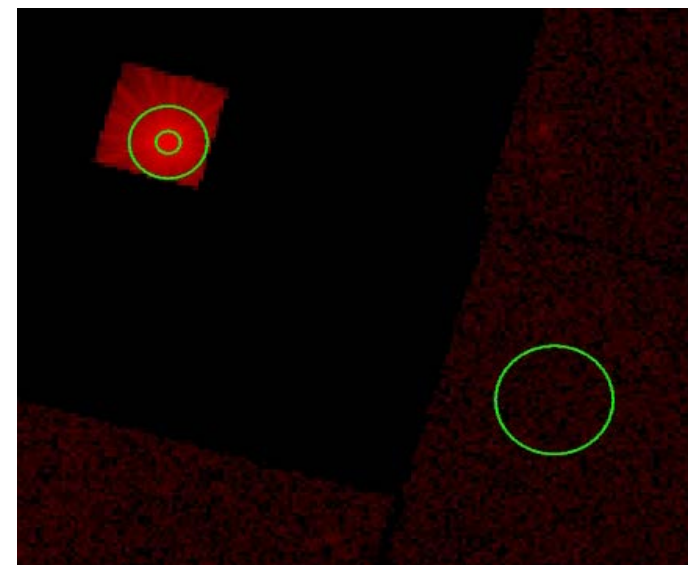
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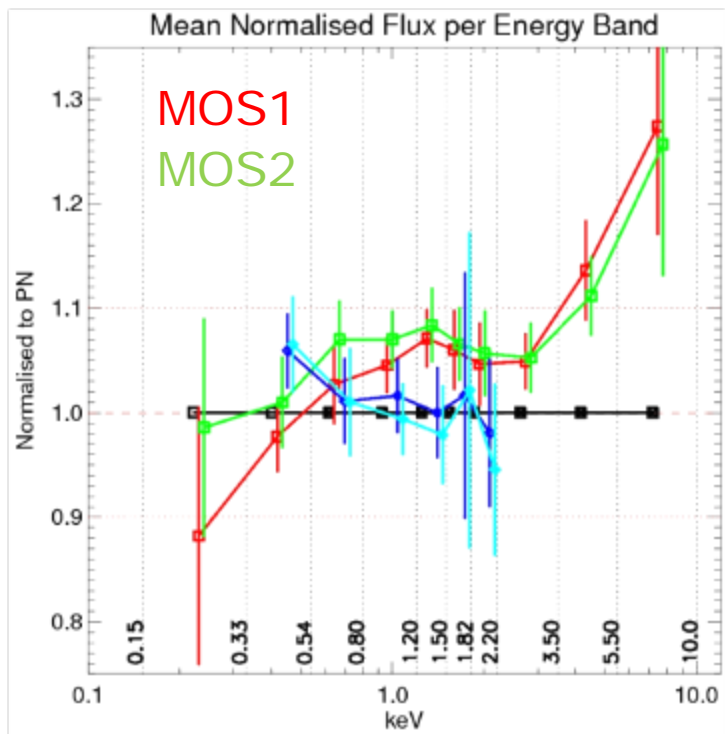
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➤ PN background:

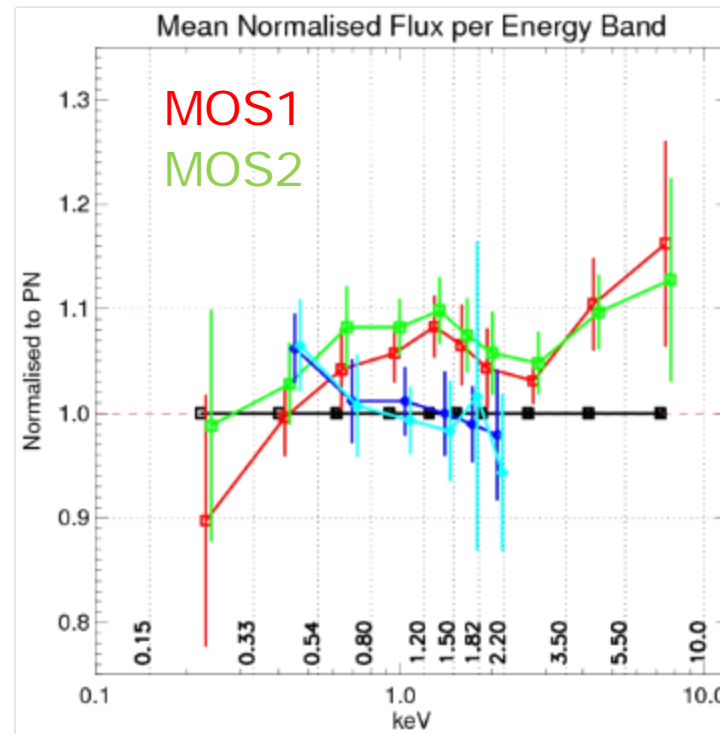
- Extracted from regions within the small window: some degree of source contamination.



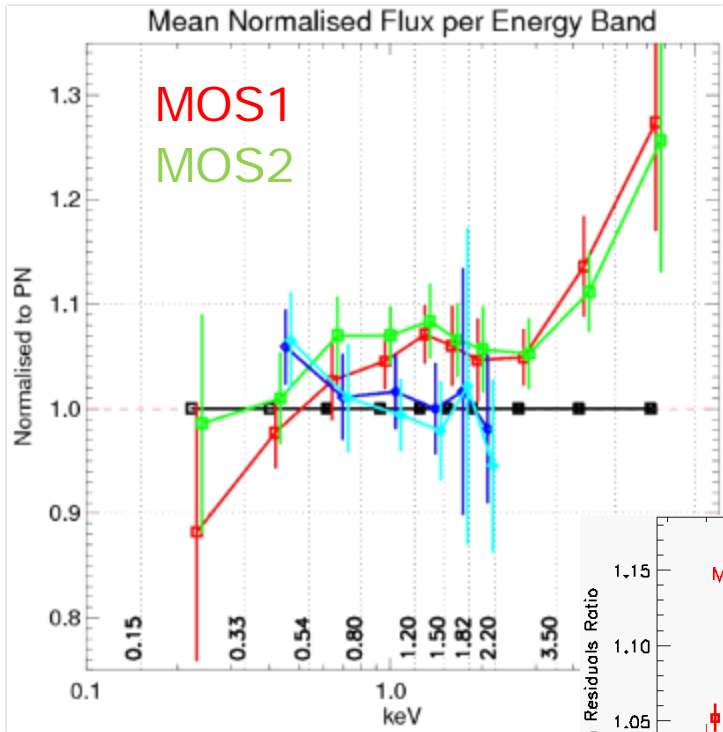
Previous PSF calibration



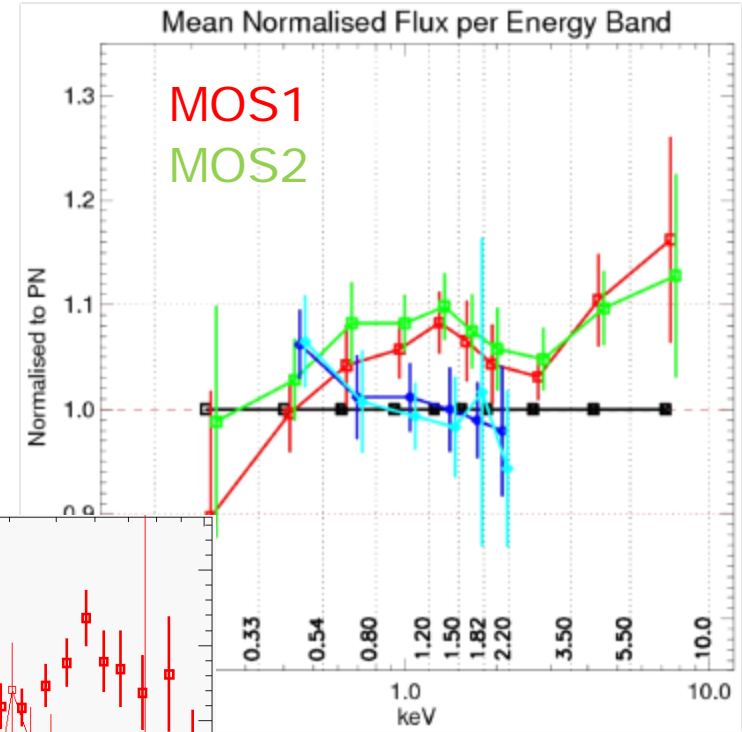
Current PSF calibration



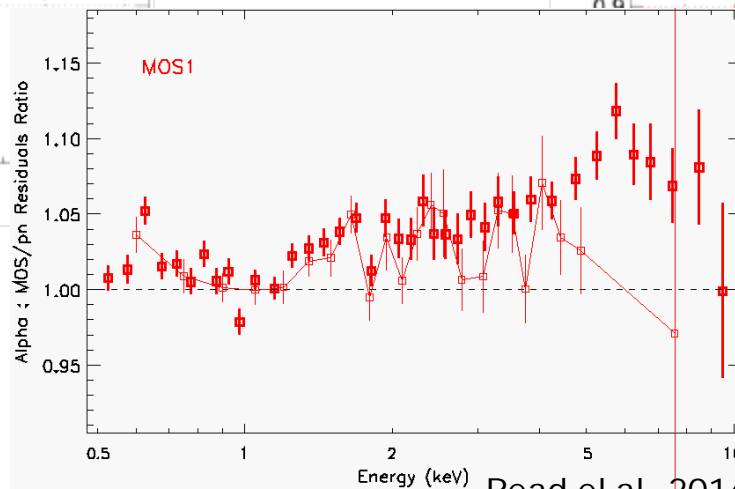
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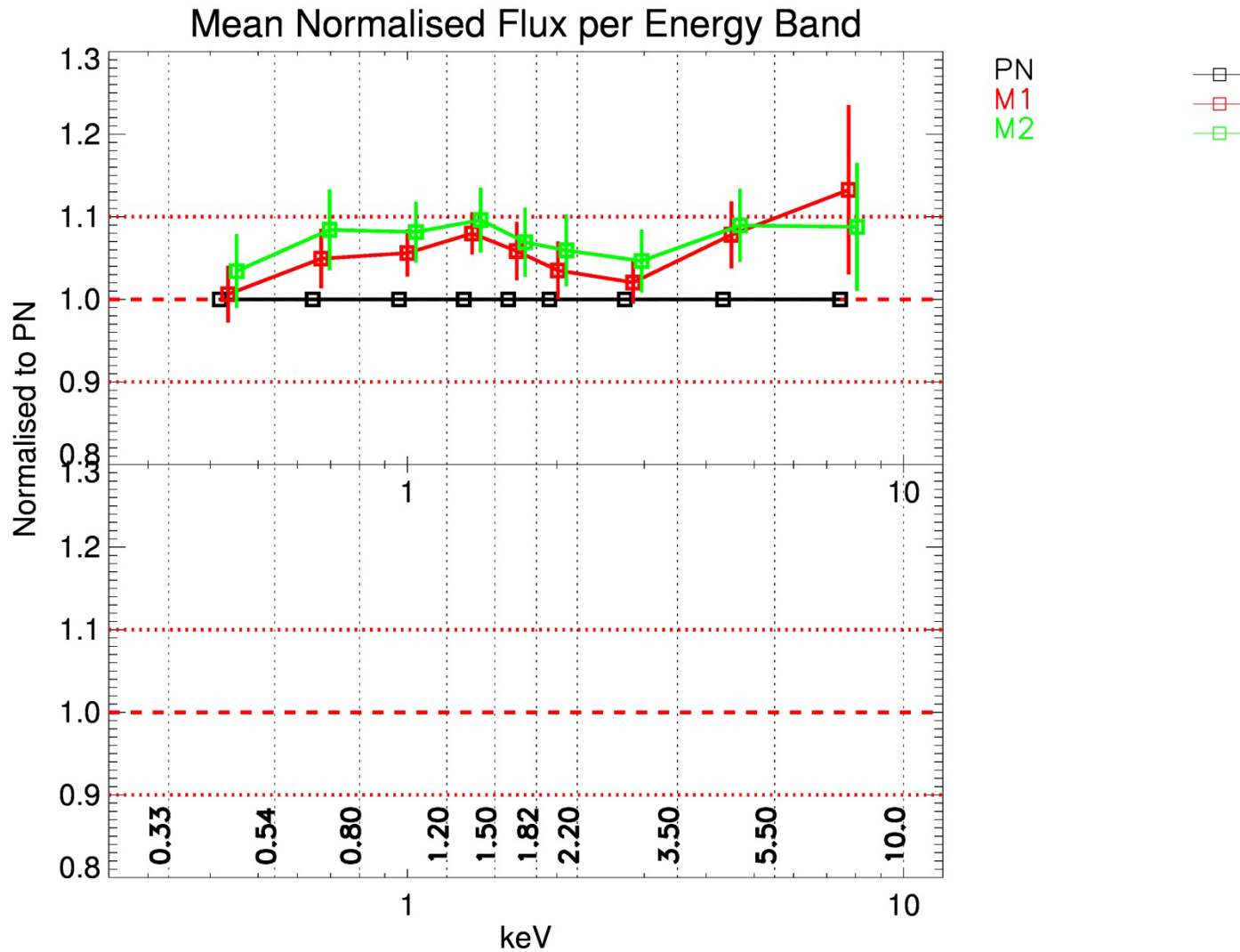


Sample of non-piled-up
on-axis sources

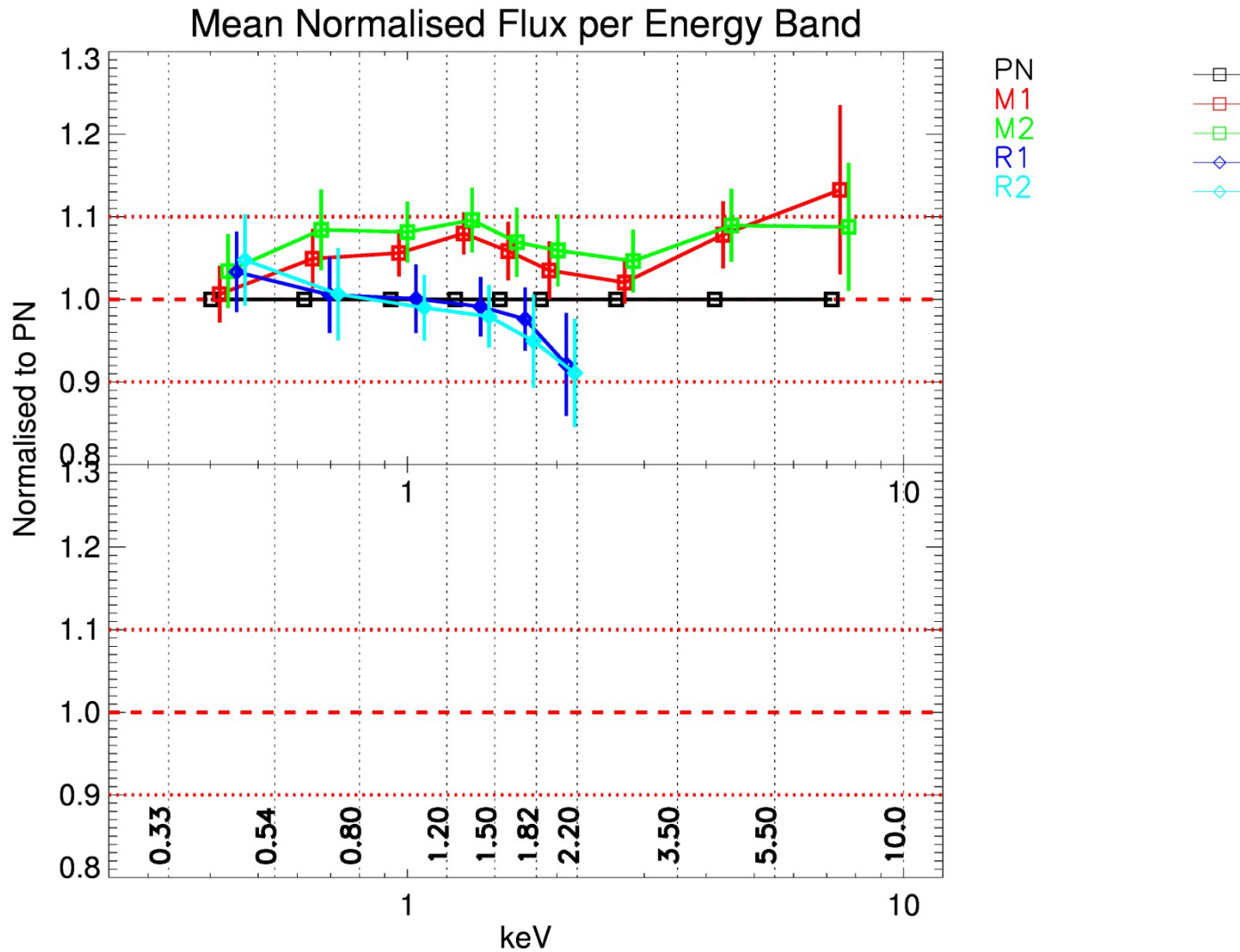


Read et al. 2014

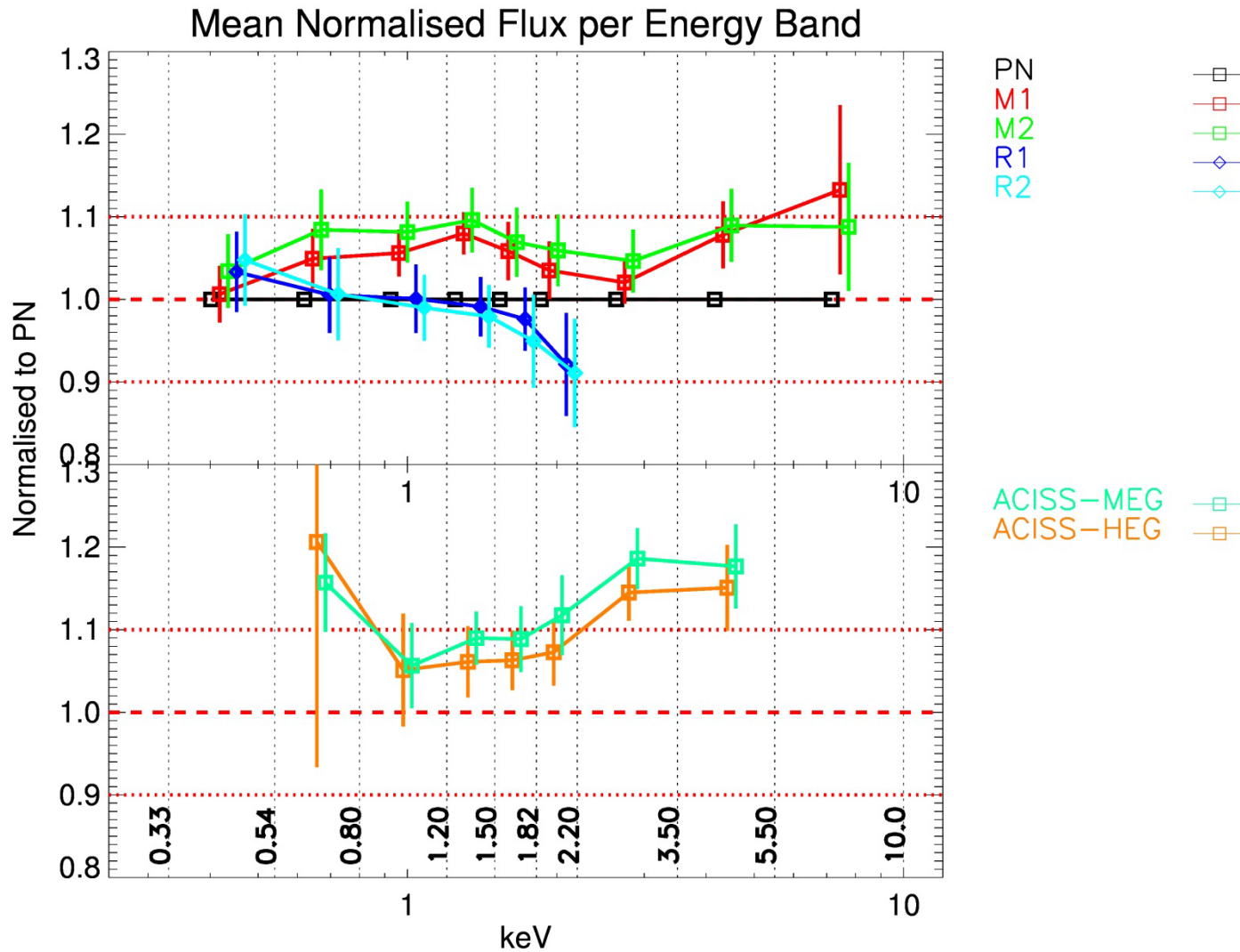
Mean Normalised Instrumental Fluxes



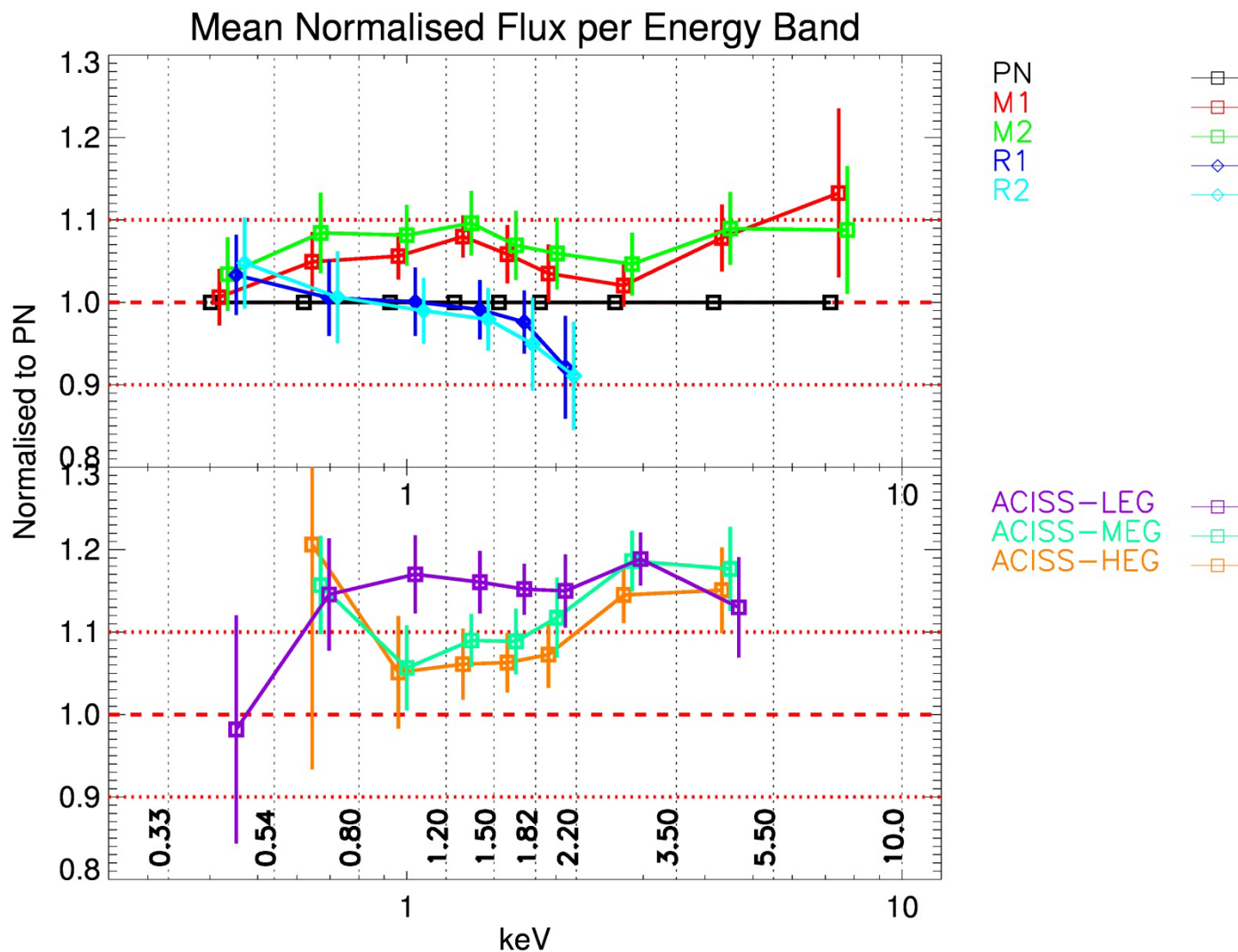
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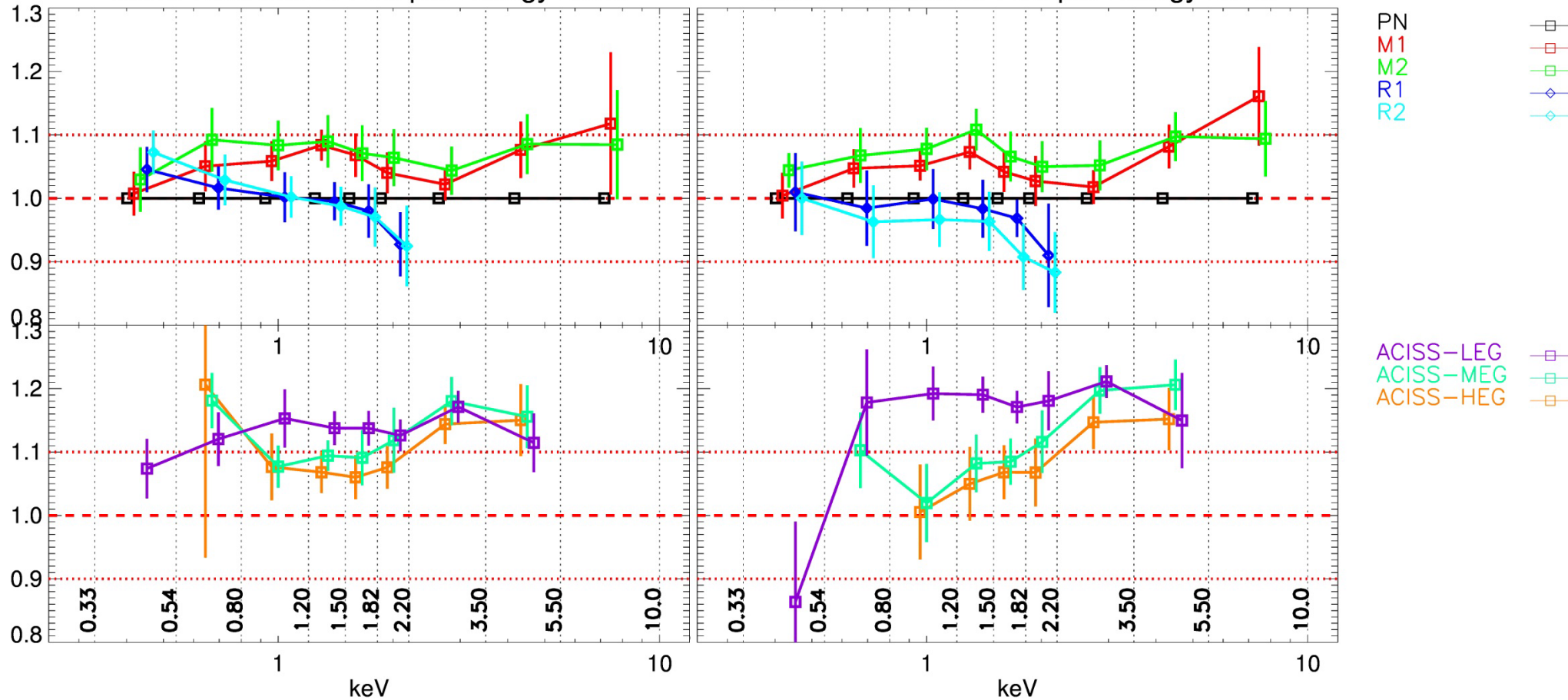


Before 2008

After 2008

Mean Normalised Flux per Energy Band

Mean Normalised Flux per Energy Band



chisqr v. cash

ml v. cash

