



Thermal SNRs WG Report
at the IACHEC Meeting

Wednesday May 22 8:30 JST: Tuesday May 21 23:30 UTC



Attending:

XMM-Newton pn Konrad Dennerl, Michael Freyberg

XMM-Newton RGS & MOS Martin Stuhlinger

Suzaku XIS Eric Miller

Swift XRT Andy Beardmore

NuSTAR Brian Grefenstette

Chandra ACIS Paul Plucinsky

ASTROSAT Gulab Dewangan

NICER Kenji Yamaguchi

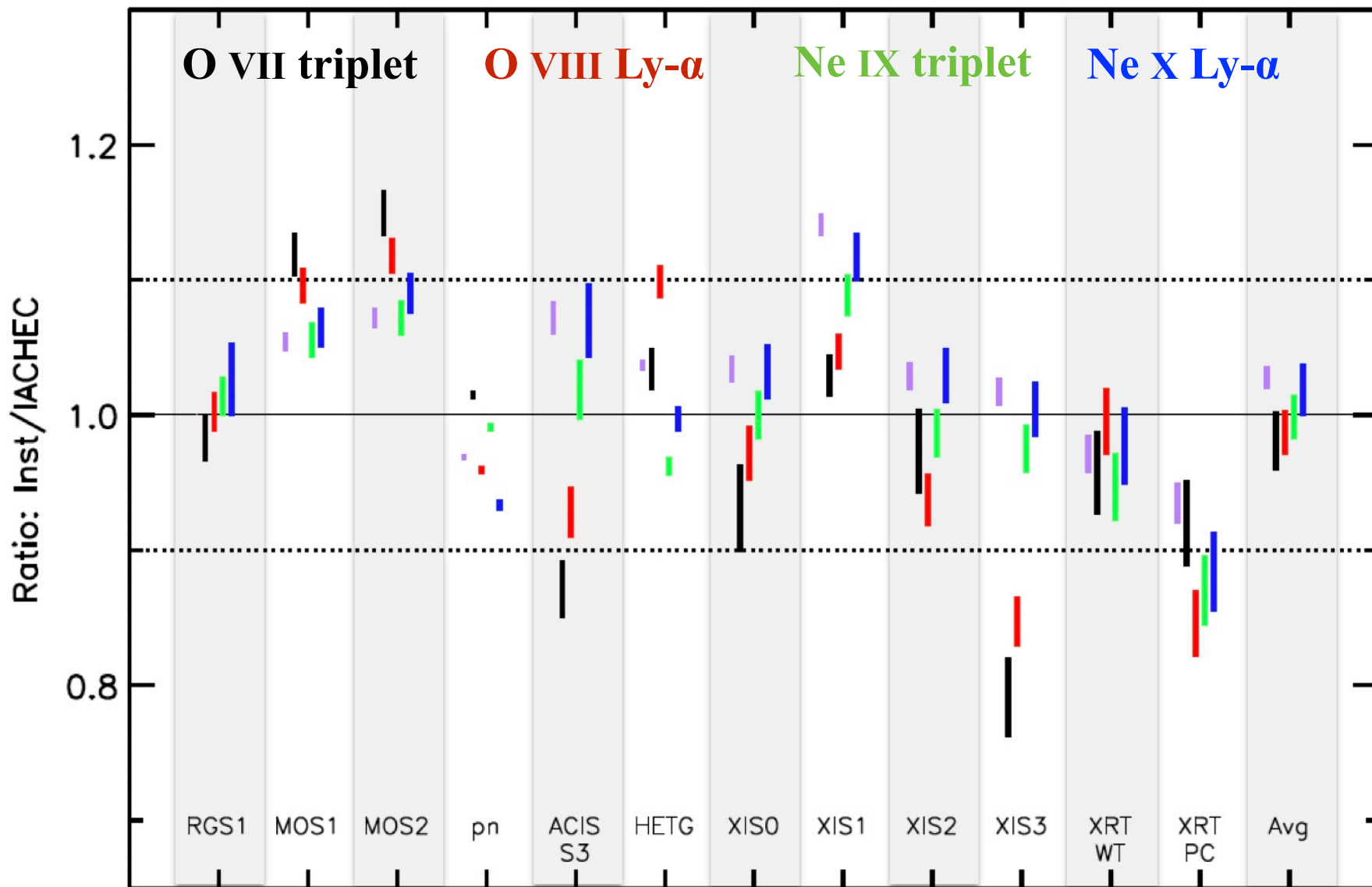


E0102 line norms
from IACHEC paper

1 E0102.2-7219 Results Dennerl (MPE)

Plucinsky et al. A&A 597, A35 (2017)

Results with release rmf

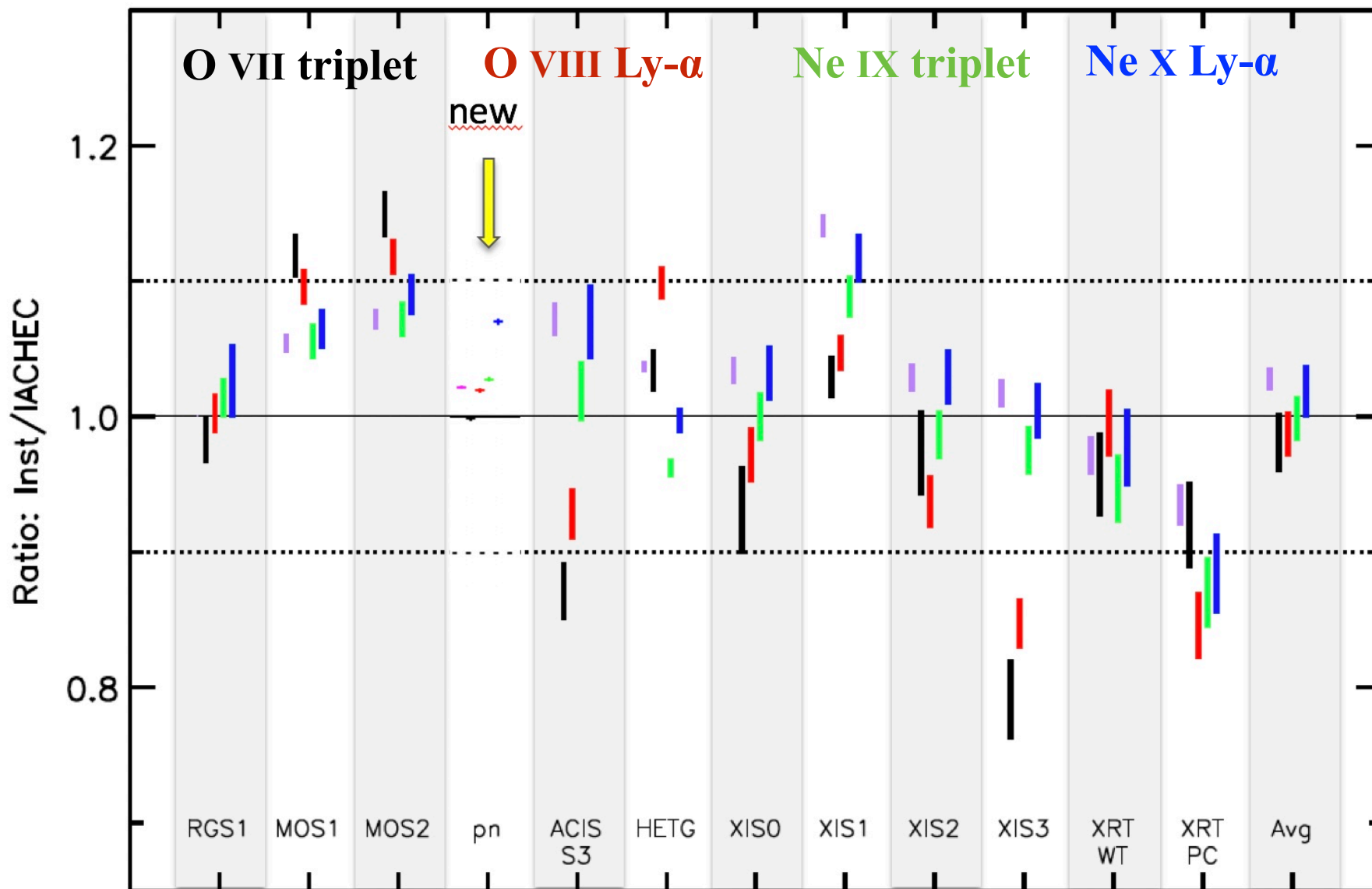




1 E0102.2-7219 Results Dennerl (MPE)

Plucinsky et al. A&A 597, A35 (2017)

Revised results with new rmf

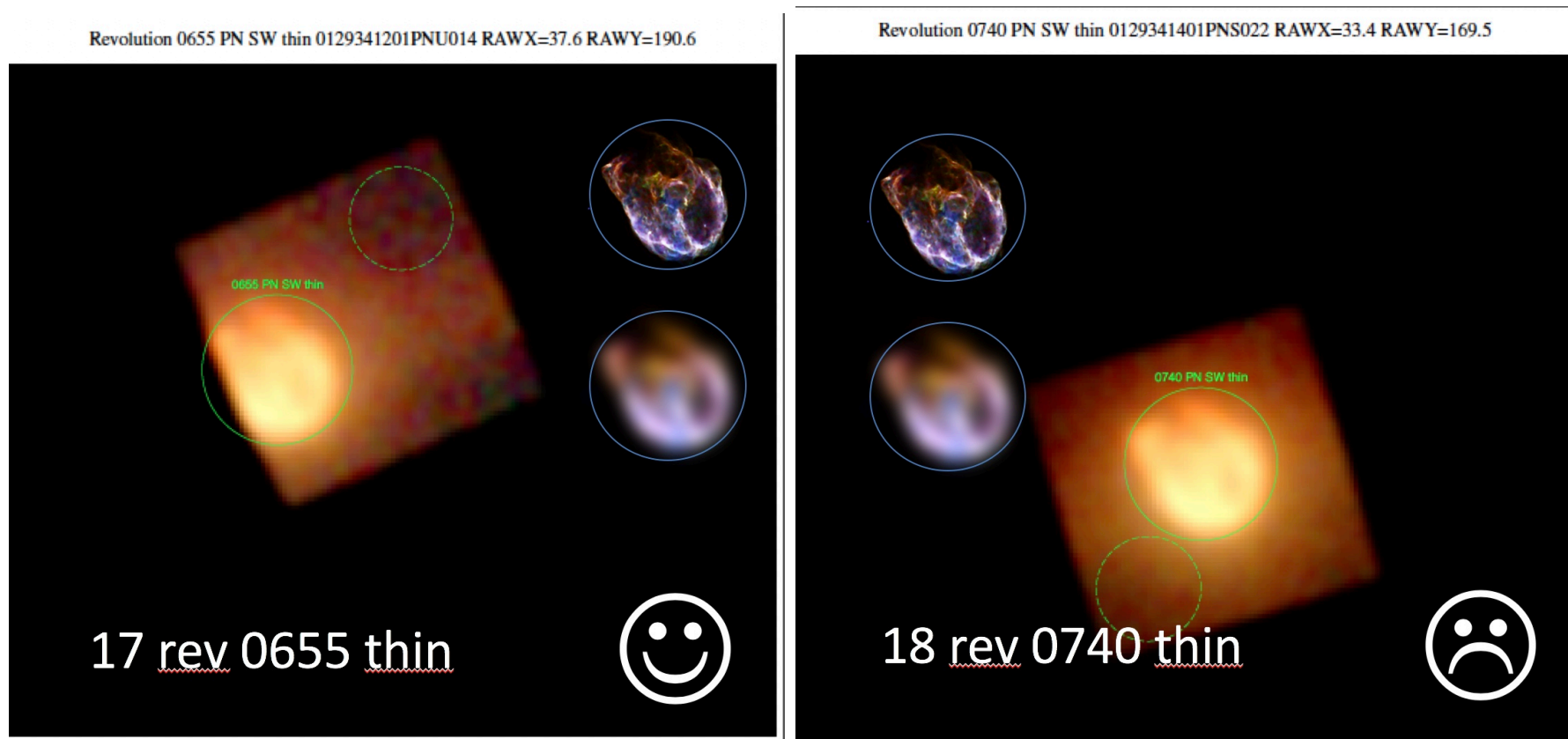




N132D Model Development

- IACHEC model is currently at version 2.11
- Dennerl showed fits to the pn data of N132D in small window mode. Better fits are achieved when N132D is positioned at the edge of the small window mode compared to fits when N132D is positioned in the center of small window. This difference is under investigation.
- Miller showed fits to the Fe-K region with Suzaku XIS with a new background model.
- Grefenstette showed NuSTAR fits and argued for the addition of a high temperature component with $kT \sim 6.0$ keV (consistent with Bamba et al. 2018 paper).

Dennerl
(MPE)
XMM pn



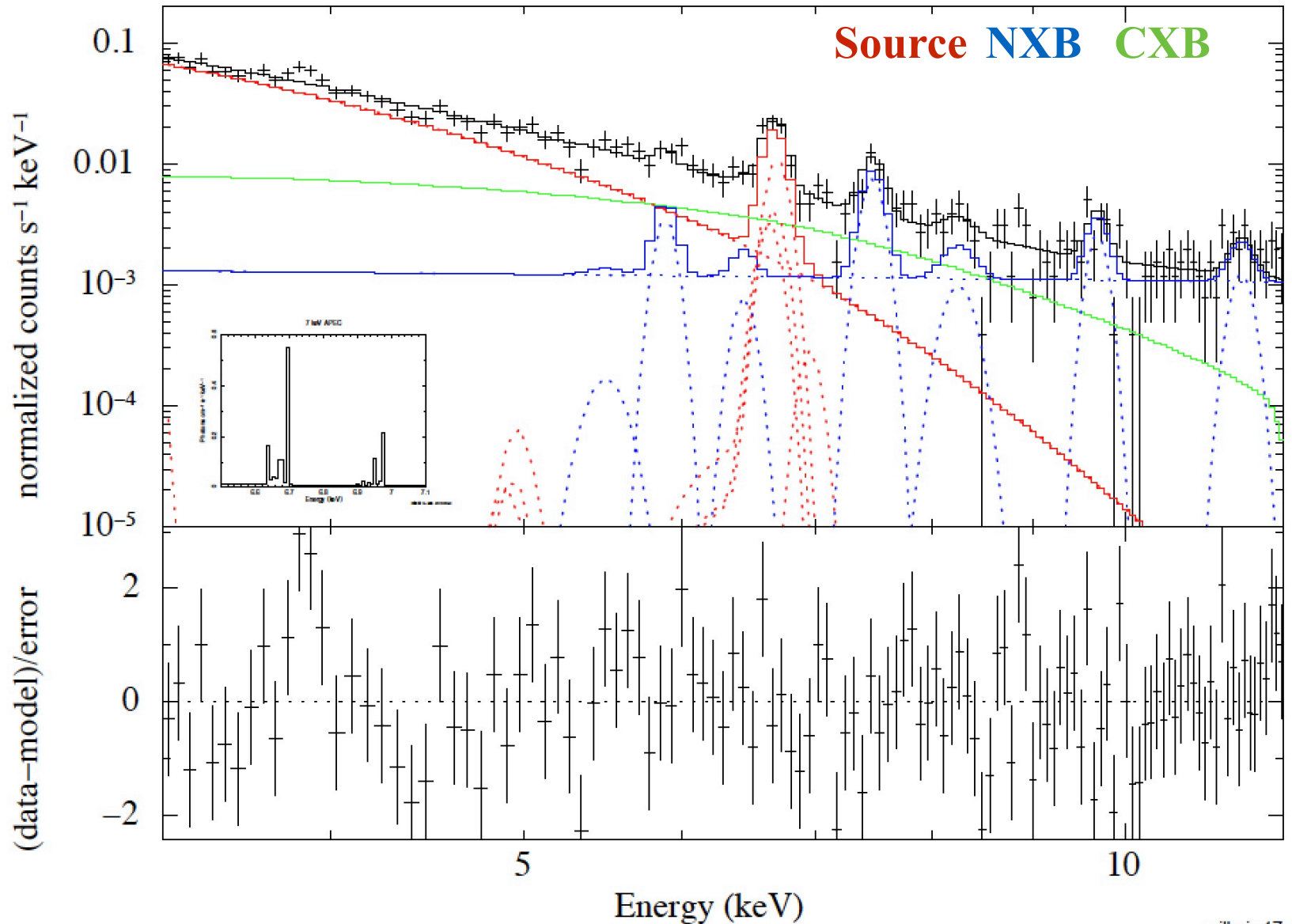


Miller
(MIT)
Suzaku XIS

Suzaku XIS fit to N132D data

N132D – XIS0 20100727, IACHEC model v2.11_20180406, contami_20140825

Background modeling is crucial to achieve reliable results in the Fe K region.

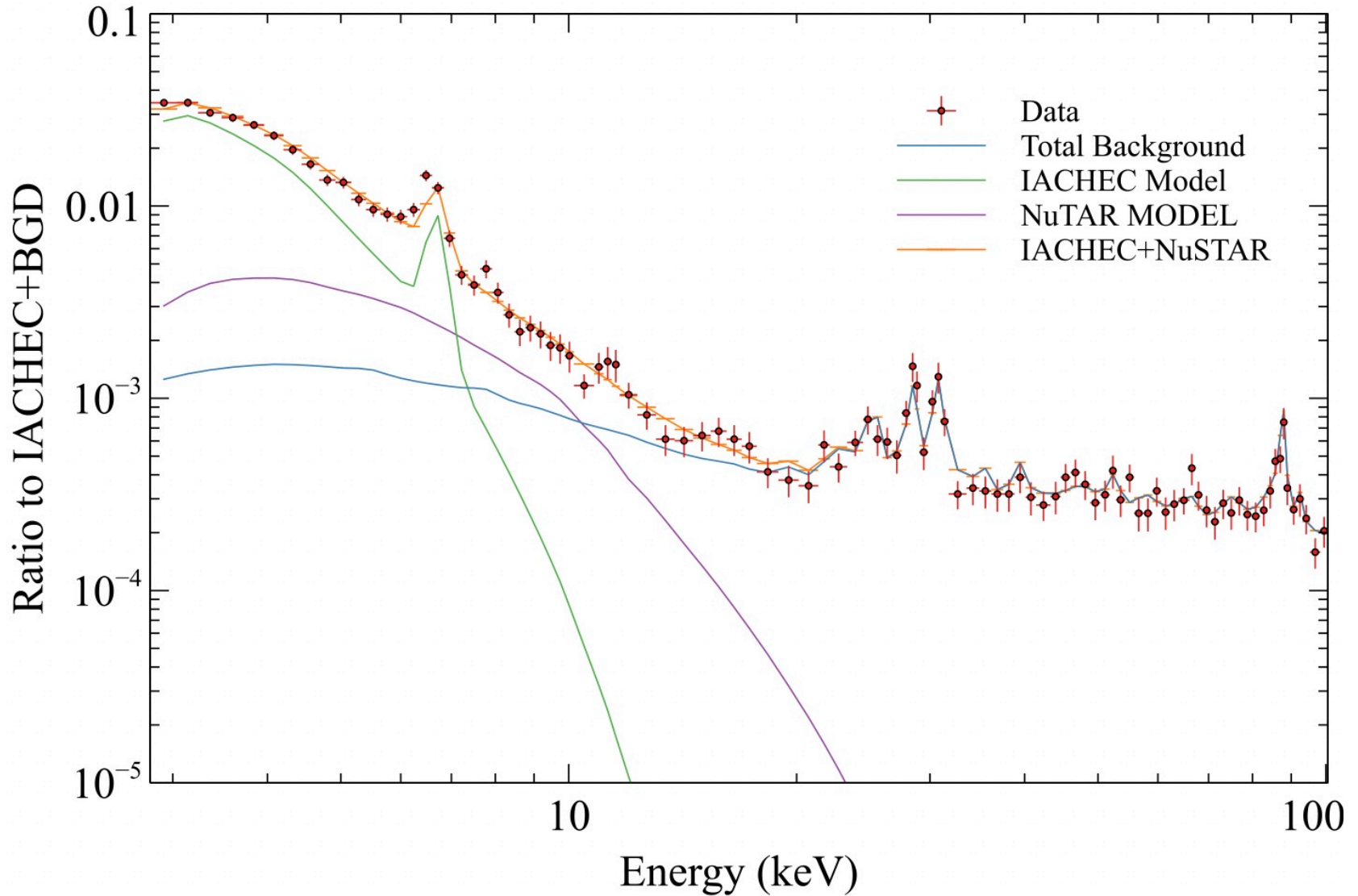




Grefenstette
(Caltech)
NuSTAR

NuSTAR fit to N132D data

NuSTAR spectrum shows the need for a hard component with $kT \sim 6.0$ keV





N132D Model Development ACTIONS

- **Dennerl will explore background subtraction for pn small window mode data**
- **Miller will complete XIS fits and send new model components to Grefenstette and Plucinsky**
- **Stuhlinger will fit RGS data with IACHEC model and identify issues**
- **Grefenstette will fit NuSTAR data with new components suggested by Miller**
- **Plucinsky will update the model with new components and distribute next version**
- **Plucinsky will schedule telecon in a month, ~14-18 June**



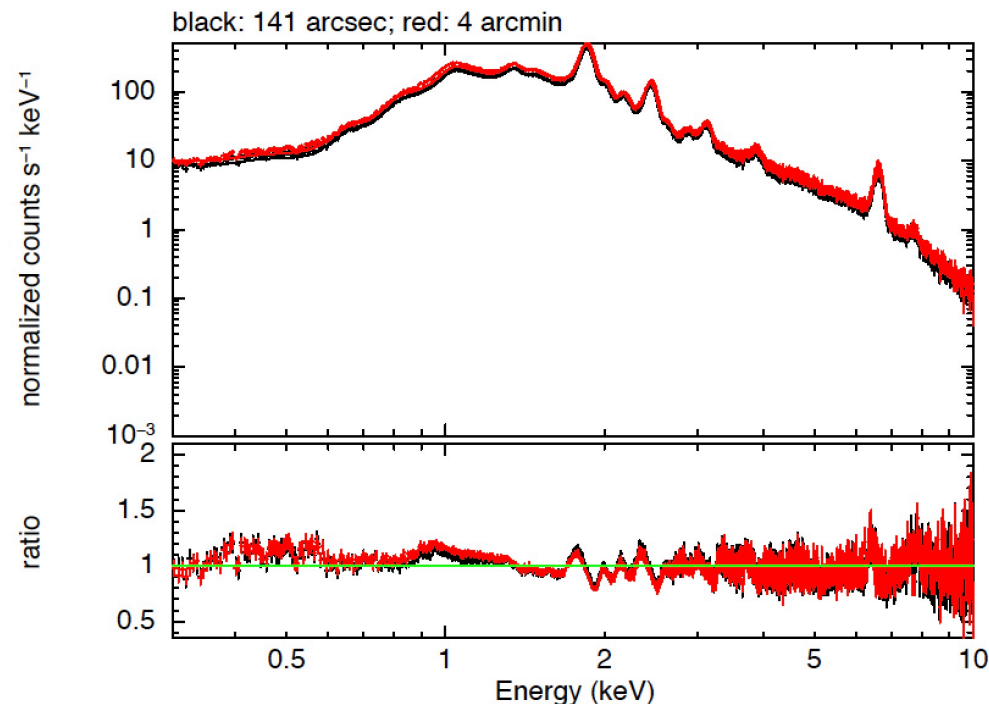
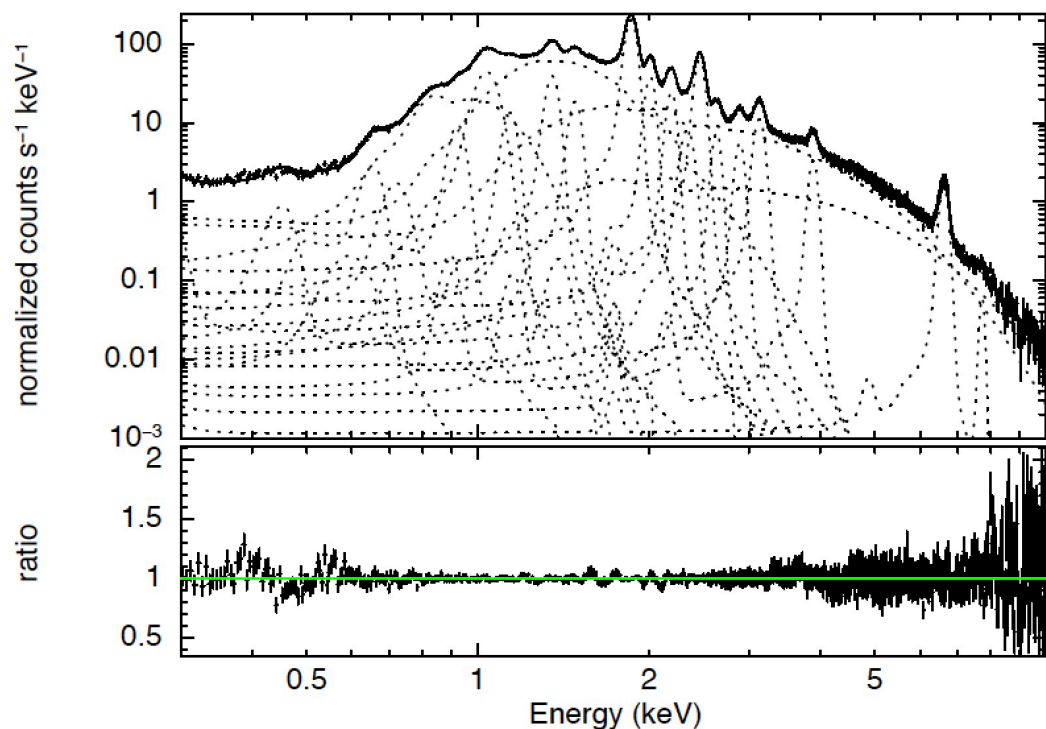
Cas A Model

- Beardmore (Leicester) showed his Cas A model that is used for Swift XRT calibration fit to the XMM MOS1 and pn data
- Dewangan (IUCAA) uses this model for ASTROSAT calibration
- we recommend that we should release this model as an IACHEC model but make clear that it is based on CCD-resolution data, not high-resolution data

Beardmore (Leicester)

XMM MOS1

XMM pn





Organization for the Coming Year

- **We have made little progress on the N132D model. We need someone to take responsibility for the RGS data. Stuhlinger has volunteered.**
- **We need to meet regularly. We should meet once a month by WebEx or zoom.**
- **We need to make progress this year to take advantage of the information from Suzaku and the new information from NuSTAR**