

IACHEC

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# NICER

Neutron star Interior Composition Explorer

## Non-Thermal SNR Working Group

Craig Markwardt (NASA/GSFC) &  
Lorenzo Natalucci



MOOG

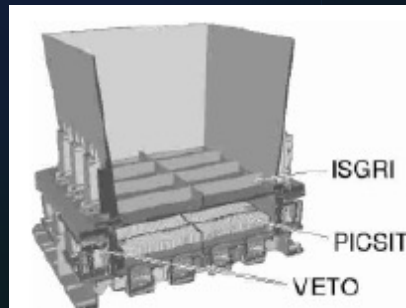


## *Summary*

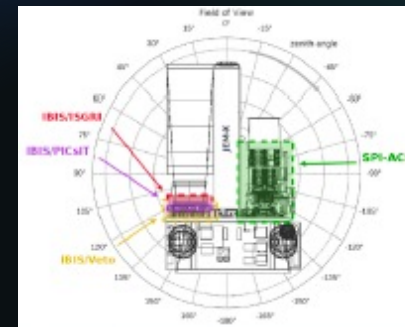
- INTEGRAL PICsIT analysis of Crab (Rodi)
- Crab Reference Model proposal (Markwardt)
- Crab Cross-Calibration Paper (Natalucci)
- Actions

# IBIS-PICsIT Instrument Review

- IBIS-PICsIT (Pixelated Caesium Iodide Telescope)
  - High-energy imaging plane of IBIS
  - Energy range ~175 keV - 14 MeV imaging
  - Spectral-imaging: ~1800 - 3600 sec



Bird et al.(2003)

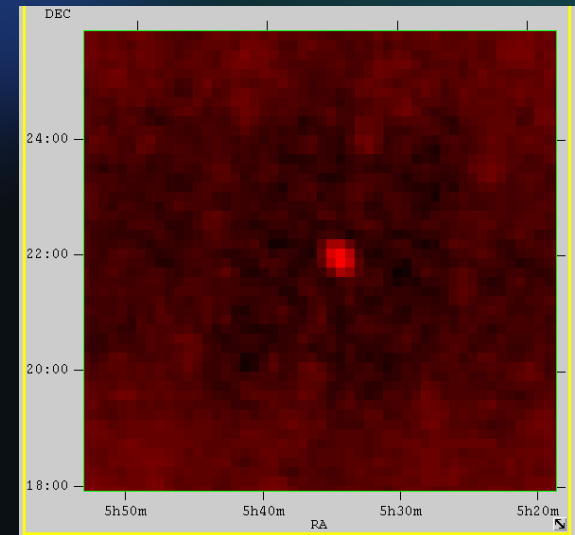


Savchenko et al.(2017)



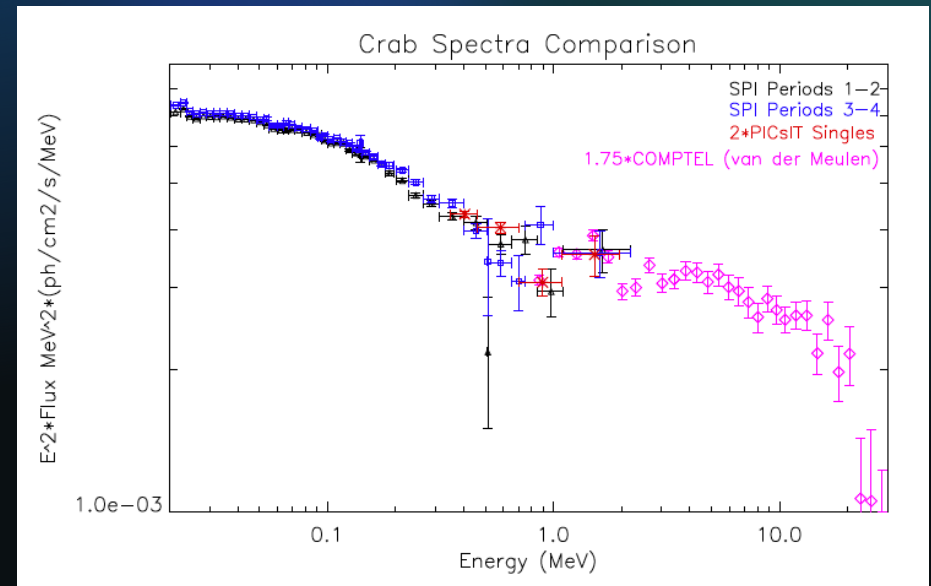
# Crab Analysis

- Selected Crab observations  $< 6$  deg of pointing dir.
- Ran standard OSA analysis until imaging step
- Stacked the images (varmosaic) for lc/spectra
- Fit count-rate and significance images with 2D Gaussian (MPFIT2DPEAK) for source count rates and errors



# Crab Spectrum

- SPI spectrum (Jourdain & Roque 2020) vs PICsIT single events vs COMPTEL (van der Meulen 1998)
- Spectrum:
  - COMPTEL has harder spectrum
  - Fit:
    - $\Gamma \sim 1.95$ ,  $E_{\text{cut}} \sim 27$  MeV (locally)
    - Knodlseder (2022):  $\Gamma \sim 2.0$ ,  $E_{\text{cut}} \sim 39$  MeV





## Crab Reference Model - Motivations

- Most observers use an absorbed power law for Crab nebula+pulsar observations

$tbabs * pow$  (and sometimes  $tbabs*logpar$ )

- Pro

- It's simple
- It does work well >2 keV

- Con

- Doesn't work as well at lower energies
- Not really physically motivated, inconsistent use in literature
- We know that the pulsar and nebula have distinct spectra (Kuiper 2000)
- Nebula has spatial variations (Mori et al 2001)

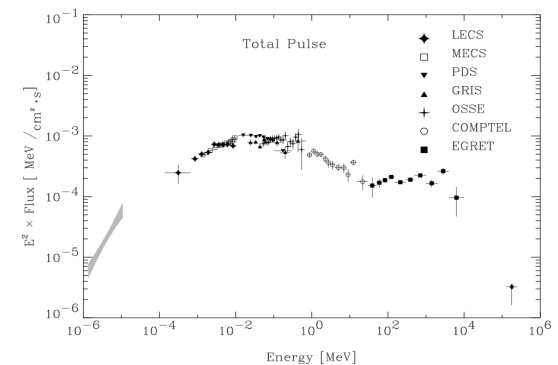
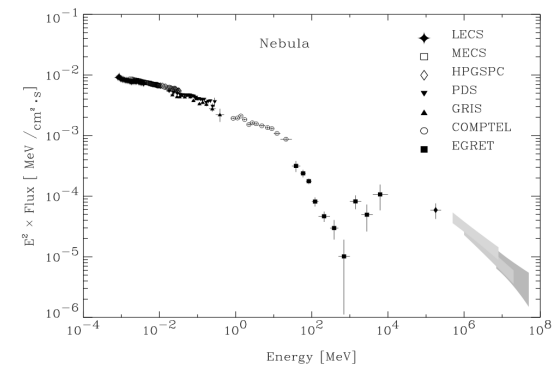
- Would be nice to have a Crab reference model that covers larger range of energies and can distinguish contributions from nebula and pulsar

- Goal: support X-ray observations in the 0.1-100 keV band
- Recognize that for energy band < 2 keV absorption model matters
- Recognize that dust scattering can be important
- Timing studies can separate pulsar from nebula



## Craig's Proposed Model

- XSPEC model  
 $\text{model TBvarabs} * \text{xscat} * (\text{curv} * \text{powerlaw} + \text{crab\_pulsar})$   
 Absorption Dust Nebula Pulsar
- Absorption (Kaastra et al 2009)
- Dust ("xscat" model Smith)
- Curvature correction (Kaastra)
- Crab pulsar
  - Kuiper 2000 & Kaastra 2009



Kuiper 2000



## *Reference Model Plans*

- Craig will post NICER model
- Model can be improved, made a bit more general and usable
- Maybe worth a quick paper





## *Crab Cross-Calibration Paper*

- Multi-mission cross-calibration paper using Nebula + Pulsar
  - ~10 instruments
- Lorenzo Natalucci coordinating author
- Status
  - Multiple mission data collected through 2017
  - Overleaf draft started
  - Open question of soliciting data through 2020?



## *Working Group Actions*

- Crab Cross-Calibration Paper (in work)
- G21.5 – delayed at moment
- Crab Reference Model (proposed)