







## Einstein Probe mission status

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on behalf of the Einstein Probe consortium

Image credit CAS/ESA

### Instruments & spacecraft





Wolter-1 + pn-CCD (eROSITA) FoV: ~1 deg Band: 0.3 -10keV **Resolution: 24**" (HPD, on-axis) Effe. area: ~300 cm<sup>2</sup> @1keV (x 2 units)







#### Spacecraft



On-board data processing Quick slew & autonomous follow-up

MPE



X/S-band (several hours) B (down/up-link; minutes) VHF (down-link; minutes)

Yuan, et al. 2022 Handbook of X-ray and Gamma-ray Astrophysics



#### Launch of EP on Jan. 9, 2024

6 yrs after mission adoption11 yrs after EP proposed13 yrs after beginning of MPO technology R&D @ NAOC/CAS

LM-2C @Xichang

A loong in Chinese mythology

# **Mission status**

- Orbit: 570 km (20 km lower since launch)
   Inclination: 29°
- Period: 95 min
  FXT and WXT calibrition :
- Spectrum,
- positioning



- WXT : ~ 20 micro- meteroid hitting
  - ~ 2 severe hitting (1/4 CMOS lost)
  - Effective area and gain do not show variations (huaqing's talk).

## **Data Real time performance**

- Data downlink:
- Alert : several minutesWhole data: < 10 hours</li>



~ 10 Orbits per day
 Including China and ESA's station.

Data uplink: Too observation command in 1 minute.



#### EP-WXT sources (5242)

Galactic coordinate



# WXT's All sky map



0.2

0.1

0.4

## Sensitivity



#### FXT localisation precision for bright sources



#### EP X-ray transients since launch



#### X-ray transients detected by EP-WXT



### EP240908a: example of quick onboard follow-up





- T0=2024-09-08T17:28:27 (UTC)
- peak flux: ~1e-9 erg/s/cm^2(EP team et al. GCN 37443)
- T0+7min: automated FXT follow-up (EP team et al. GCN 37432)
- T0+19hrs: FXT ToO
  - flux ~1.1x10^(-13) erg/s/cm^2
- optical afterglow candidate:
  - AB magnitude r ~ 24(Quirola-Vasquez et al. GCN 37438)









#### Follow-up of GW event: S250206dm

**S240422ed**: T0=2025-02-06 21:25:30.439 UTC; NSBH (55%), BNS (37%); 373 +/- 104 Mpc; 90% area=547 deg<sup>2</sup>

#### **EP follow-up observations:**

 $1^{st}$  round: 2×2400s WXT coverage + 100×300s FXT targeting galaxies (2025-02-07T02:55:12Z ~ Feb. 8<sup>th</sup>)  $2^{nd}$  round: 59×2000s FXT tiling (from Feb. 11<sup>th</sup> to 18<sup>th</sup>)

#### Search for transient or variable candidates in the FXT data (apparently associated with the galaxies):

found ~20 uncatalogued sources, two of which exhibit significant flux decay. Detailed information: GCN Circular 39545; <u>https://ep.bao.ac.cn/ep/cms/article/view?id=185</u>



 $1^{st}$ : WXT covered 302  $deg^2$  within (55% of) the 90% GW area WXT 0.5-4 keV flux limits:  $1 \times 10^{-11}$  erg s<sup>-1</sup>cm<sup>-2</sup> FXT 0.5-10 keV flux limits:  $3 \times 10^{-13}$  erg s<sup>-1</sup>cm<sup>-2</sup>



## WXT's CMOS detector response

#### Analysis on Tycho spectrum



Spectral analysis of (stacked) spectrum of Tycho (FM4, CMOS16) Energy resolution (compared with ground measurements, FM4, CMOS16)

No obvious variations in GAIN and energy resolution for most of the detectors after launch. See Huaqing's talk in detail.

# FXT calibration







- Science operations have been started since July 2024
- ~110 X-ray transients with high/SN detected
- A wide range of targets: FXT, GRB, TDE, WD+NS+BH, flaring stars, MMA .....
- Monitored the activity of large samples of known sources
- Great potential for more discoveries in X-ray time-domain and MM astronomy

<u>http://ep.bao.ac.cn</u> https://www.esa.int/Science\_Exploration/Space\_Science/Einstein\_Probe\_factsheet

Thanks to the Swift, NICER, NuSTAR, XMM, Chandra, etc. and ground-based optical and radio telescopes around the world for supports of follow-up observations