

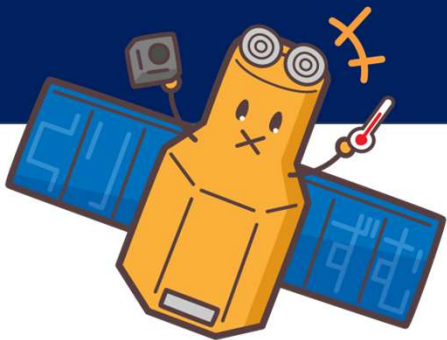


XRISM/Xtend calibration status: contamination, energy resolution, and pile-up

**YONEYAMA Tomokage (ISAS/JAXA)
on behalf of the XRISM Xtend team**



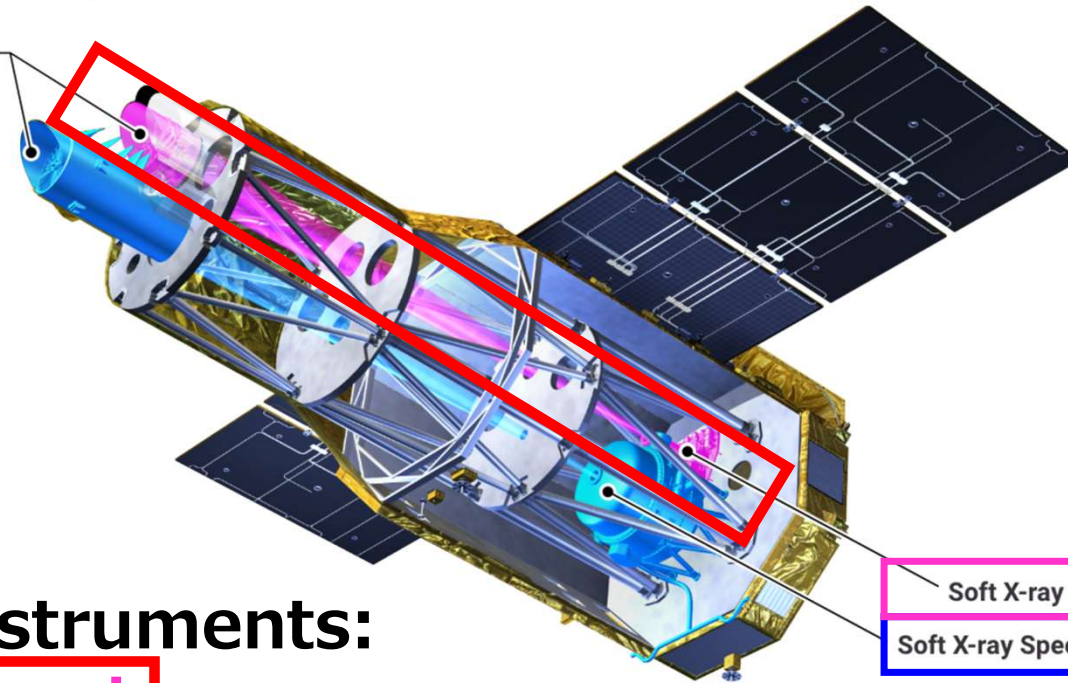
18th IACHEC Meeting
2026-04-21@Seeblick Pelham, Deutschland



X-ray **Imaging** and **Spectroscopy** Mission

X-ray telescope to collect X-rays

X-ray Mirror Assembly
(XMA)



Two mission instruments:
Resolve and **Xtend**

Soft X-ray Imager(Xtend)

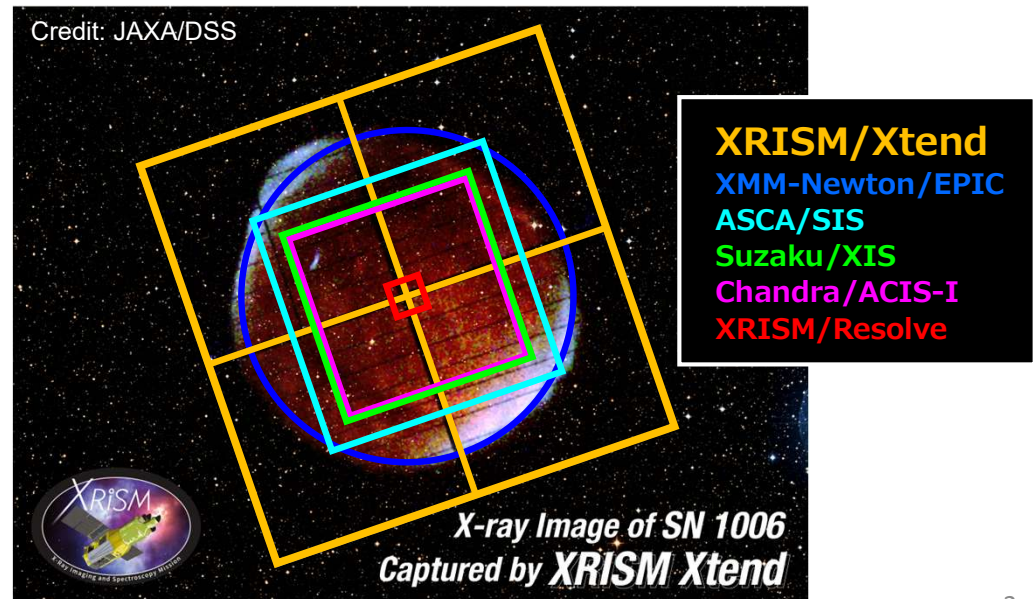
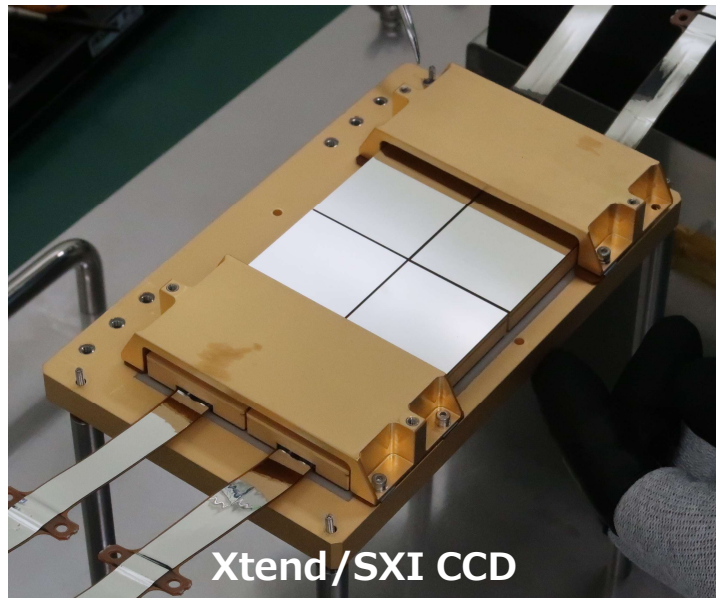
Soft X-ray Spectrometer(Resolve)

Mission Instruments

Xtend, the Soft X-ray Imager



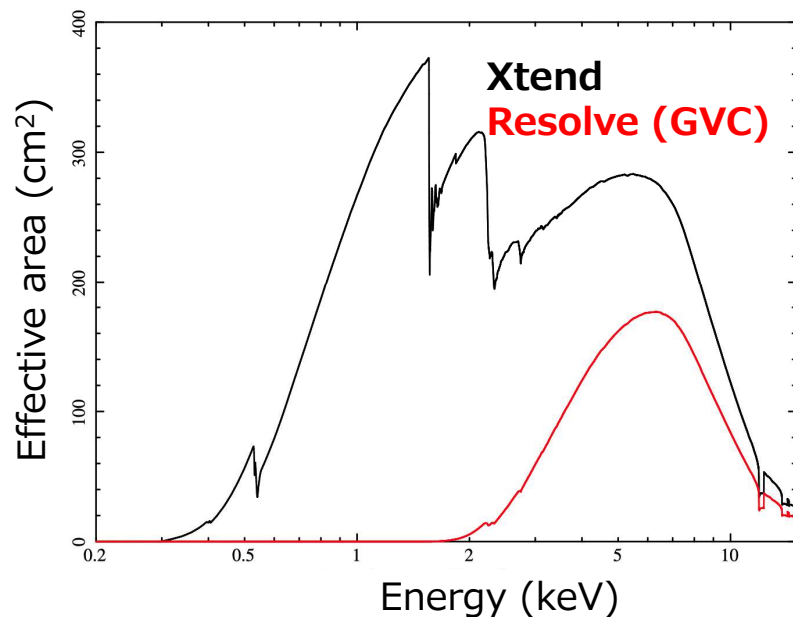
- X-ray telescope (XMA) + X-ray CCD array (SXI)
- Bandpass: 0.4 -- 13 keV
- Angular resolution: $\sim 1.4'$ (HPD)
- Timing resolution: 4 s – 0.05 s
- FOV: $38.5' \times 38.5'$



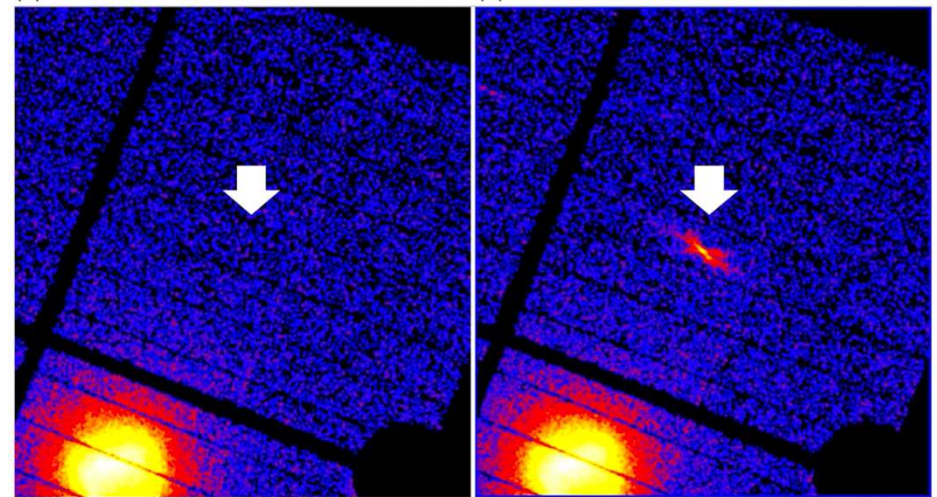
Xtend's Mission



- To observe diffuse sources (like SNRs, GCs) with the wide FOV
- To support Resolve's observation
 - by resolving point sources contaminating the target source
 - by determining the continuum with the wider bandpass
 - (due to closed gate valve of Resolve)
- To search for transient sources (Xtend Transient Search; XTS)



AX J1910.7+0917 (Atel #16607); Yoshimoto+ in prep.



The XRISM Xtend Team



Tohoku Gakuin University

H. Murakami

Rikkyo University

S. B. Kobayashi

Tokyo University of Science

T. Kohmura

Kanto Gakuin University

H. Nakajima (sub-PI)

ISAS/JAXA

Y. Kanemaru, T. Yoneyama, Y. Omiya

H. Tomida, Y. Maeda, M. Ishida

Meiji University

T. Sato

Shizuoka University

H. Uchiyama

Nagoya University

K. Yamaoka

Kyoto University

H. Uchida, T. G. Tsuru, T. Hayashi

Nara University of Education

M. Nobukawa

Nara University of Education

M. Nobukawa

Kindai University

K. K. Nobukawa

The University of Tokyo

K. Hagino

Tohoku University

H. Noda

University of Osaka

K. Odaka, H. Matsumoto, K. Hayashida

Konan University

T. Tanaka

University of Miyazaki

K. Mori (PI), M. Yamauchi

I. Hatsukade, H. Suzuki, T. Kayanoki

Kanazawa University

S. Ueda

Ehime University

M. Yoshimoto

Chuo University

D. Ishi

NASA's GSFC

T. Okajima, K. Tamura

University of Maryland

R. Boissay-Malaquini

Advisor

J. S. Hiraga, M. Ozaki, T. Dotani, H. Tsunemi

XRISM SOC (Joint work)

T. Mizuno (Hiroshima University)

Students

T. Hakamata, A. Miyamoto, A. Nagao,

S. Inoue, Y. Aoki, S. Takayama, M. Higuchi (graduated),

S. Itoh, S. Nakatake, R. Takamoto,

T. Matsushima, J. Kurashima, N. Asama,

Y. Tagi, M. Fujii, S. Morita





- Cross Cal. of the Effective Area: **INOUE Shun** (done)
- Xtend Cal. Status (1): **TY** (we are here)
- Xtend Cal. Status (2): **SUZUKI Hiromasa**
 - Response, background
- Xtend Transient Search: **YOSHIMOTO Marina**



Going well.

Current Status of Xtend



- All CCDs are in operation
- Maintains the energy resolution of 170 – 180 eV
 - Degradation of the charge transfer inefficiency (CTI) is as expected from the ground experiment
 - Meets the requirement of 250 eV at EOL
- No degradation of the effective area/quantum efficiency
 - Contamination is very well suppressed

Current Calibration Status of Xtend



- Xtend CalDB ver. 12: 20250915
- Updates from CalDB 11 based on the in-flight cals:
 - Gain-related files
 - Charge trail
 - CTI
 - Gain
 - RMF parameters

XRISM CALIBRATION DATA			
GEN Index Summary	Current Version: 20241115	Release Notes	Download
Resolve Index Summary	Current Version: 20250915	Release Notes	Download
XTEND Index Summary	Current Version: 20250915	Release Notes	Download
General Index Summary	Current Version: 20240822	Release Notes	Download



Routine cal. observation

- 2 observations per year
- Perseus cluster, Cygnus loop
 - Energy (gain, response...)
- 1E0102.2-7219
 - Contamination

Rframe dump

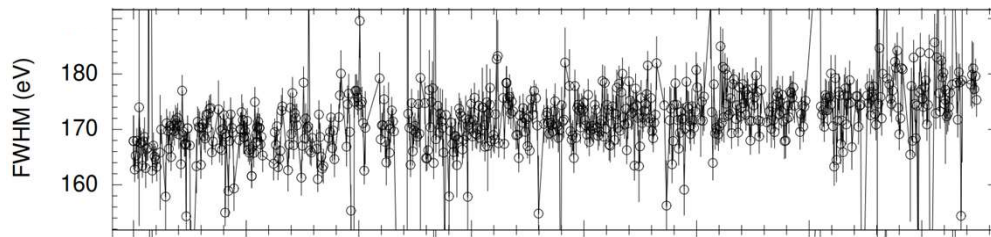
- Twice per year
- To check noise and light leak
- Day Earth and Night Earth

Energy Resolution (@5.9 keV; by ^{55}Fe Cal. Sources)

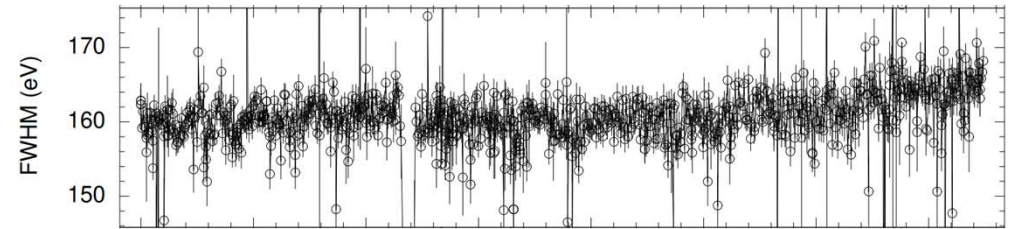


- Note: These plots are **BEFORE** applying corrections -> pessimistic case
- Gradually degrades by ~ 10 eV / 2 yr
- Slightly chip dependent
- 220 – 230 eV at 2033 (10 yr after launch)?

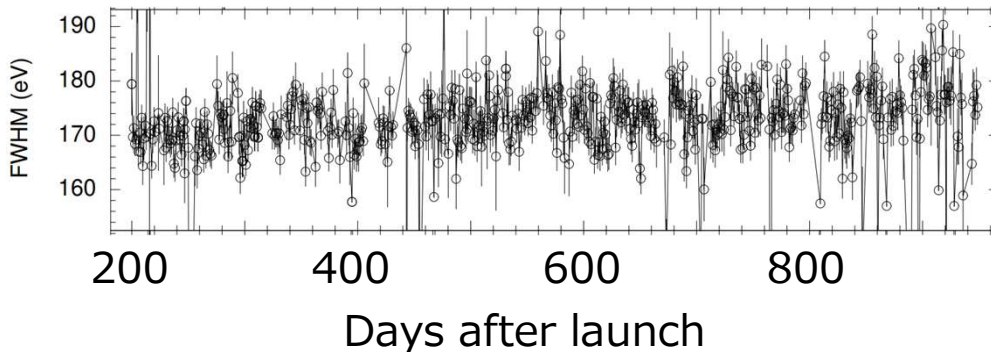
CCD1 (seg. CD)



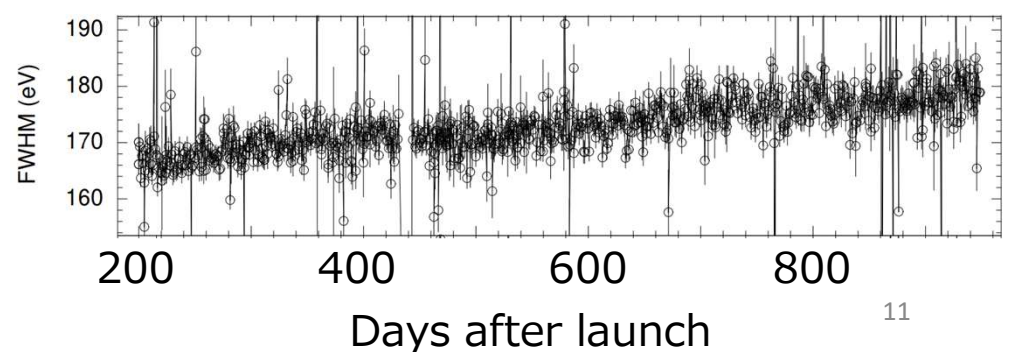
CCD3 (seg. AB)



CCD2 (seg. AB)



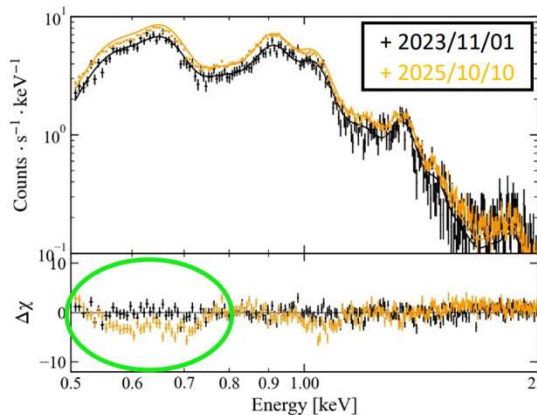
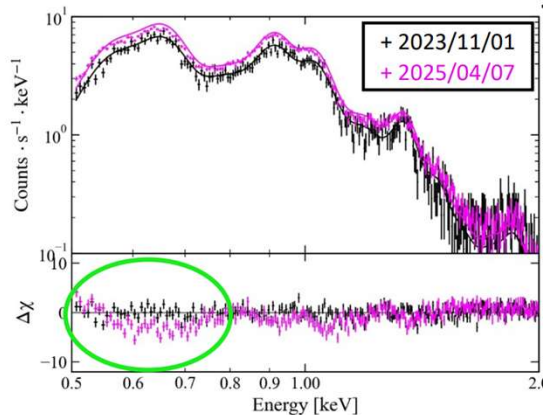
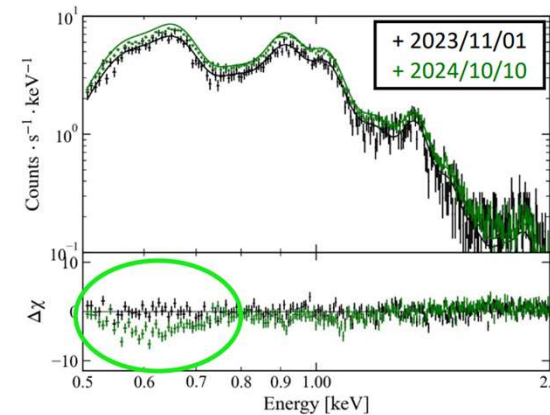
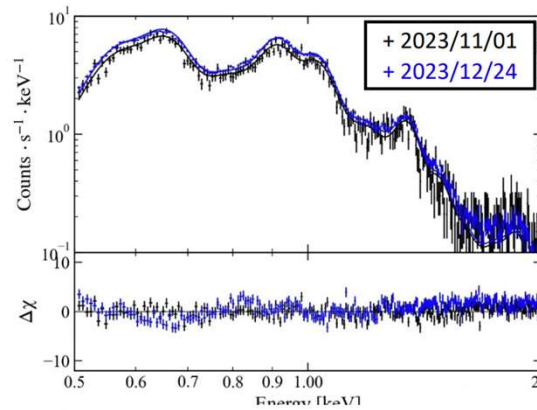
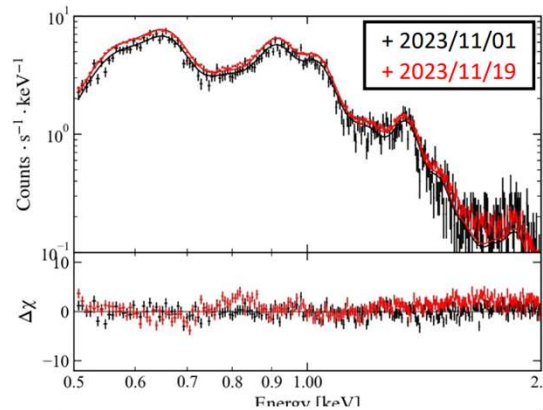
CCD4 (seg. CD)



Contamination



- Have been evaluated with E0102 fitted by the IACHEC model
- Model fitting method was improved; slight extra absorption appeared after 2024

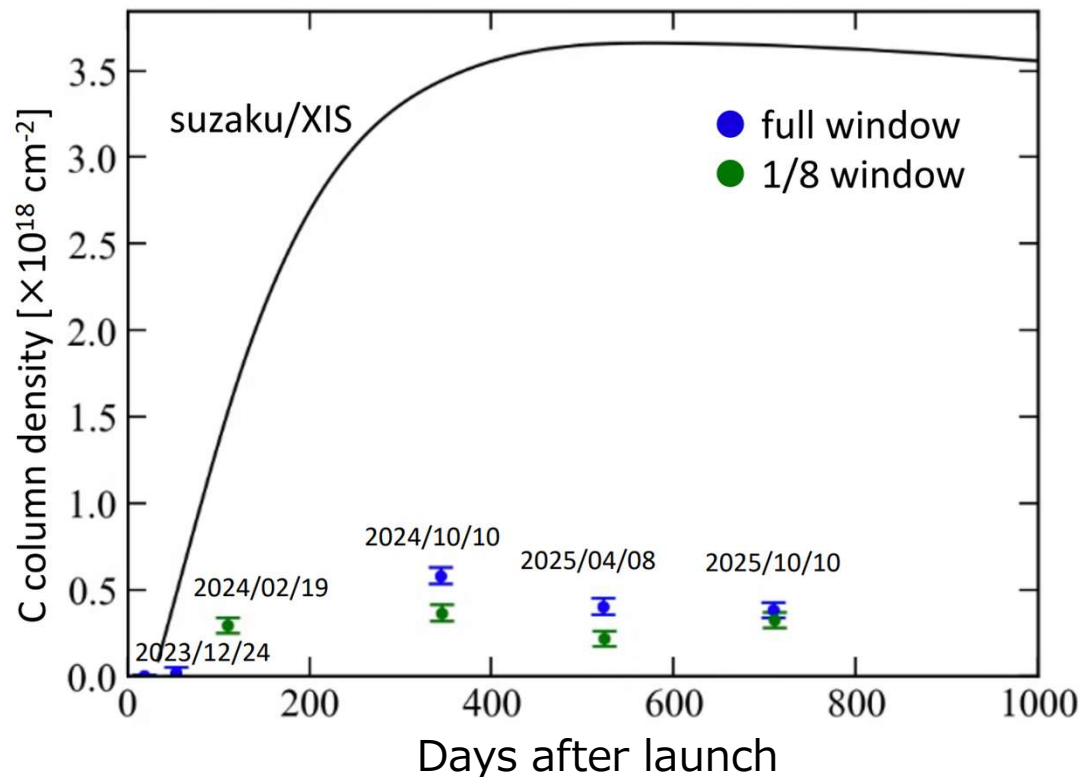


Full-window obs. of E0102
by M. Higuchi (TUS; graduated)

Contamination



- Improved method results in smaller error bars
- Conclusion not deviated; **contamination is low enough even in 2025**
- Cross-check with other sources is now in progress (at TUS)



- The inevitable fate of CCDs when observing a bright source
- A simulation-based study has been performed for Xtend (TY+25 JATIS)
- Spectral distortion and grade migration were reproduced

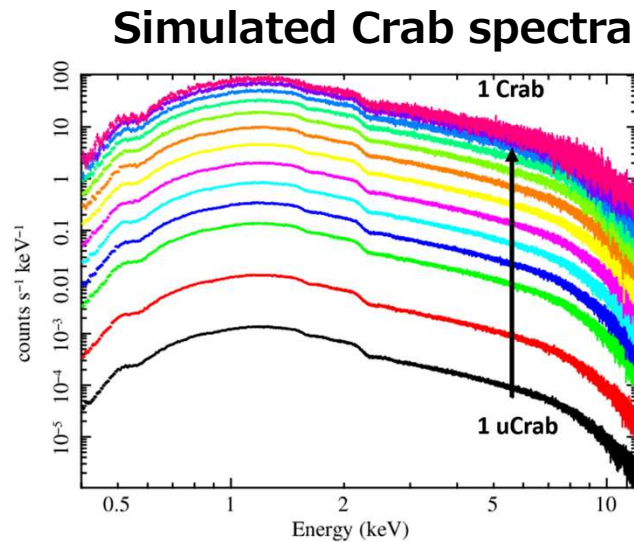
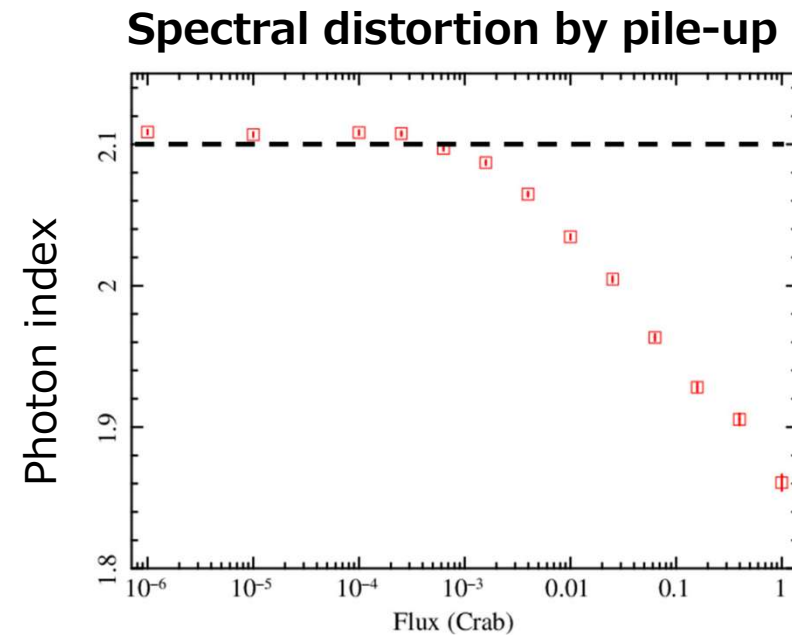


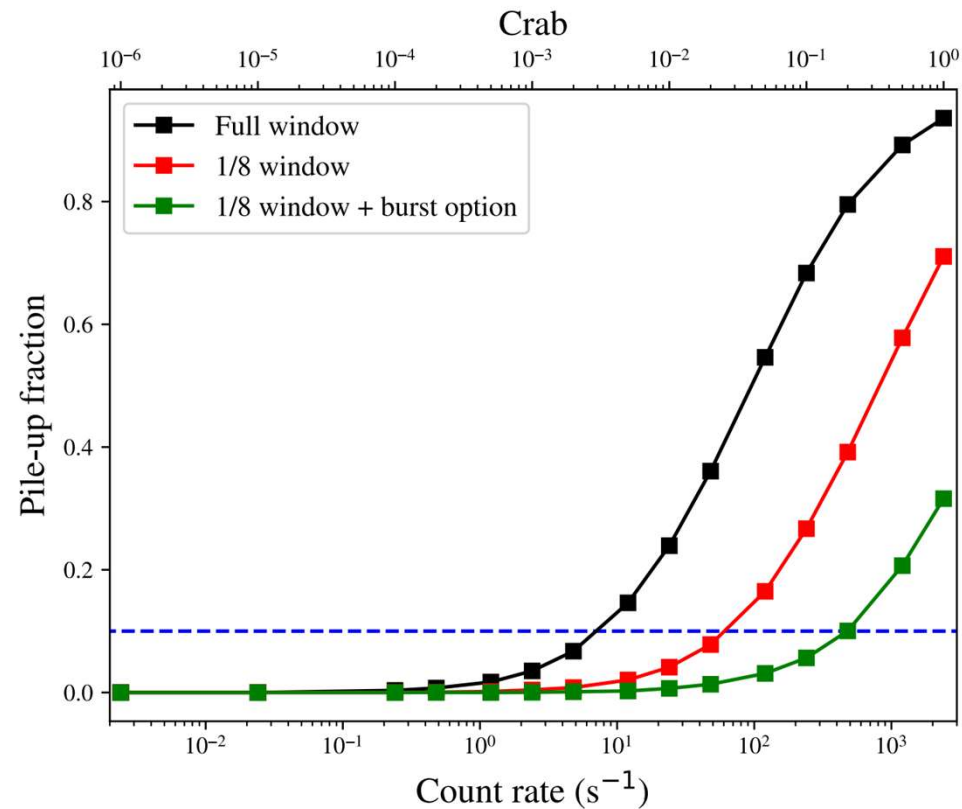
Fig. 1 Simulated spectra by the pile-up simulator by assuming a point source with the Crab spectrum with fluxes of 10^{-6} – 1 Crab. The full window mode is simulated in this figure. High input fluxes result in the high pile-up effect, by which fractions of photons are lost in the simulated spectra (see Fig. 2 for the quantitative evaluation of the pile-up).



Pile-up



- Simple suggestion:
Use 1/8 win (+ burst) for > 8 (66) cps for a point source

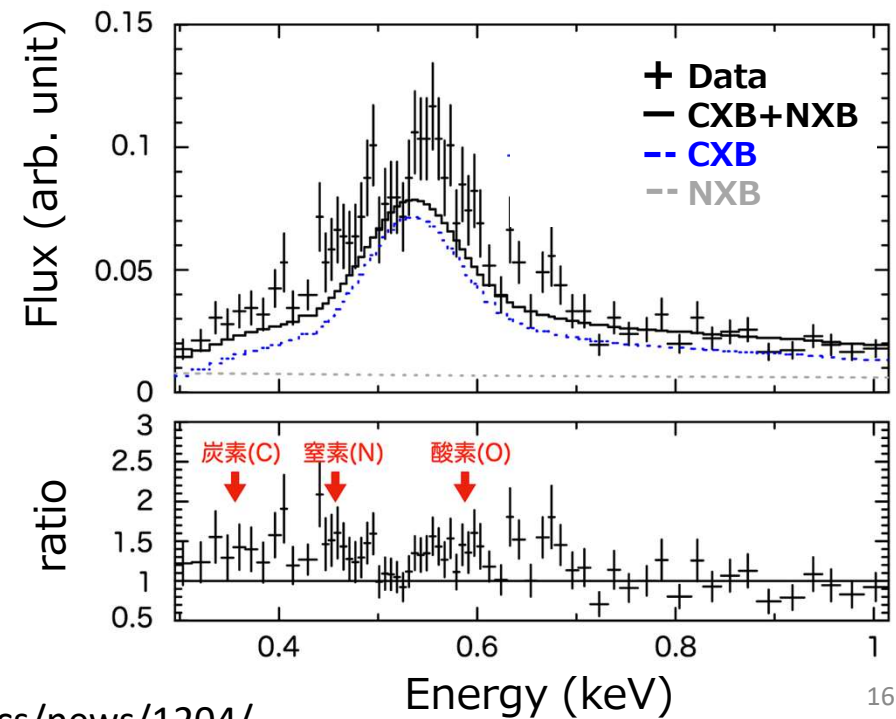
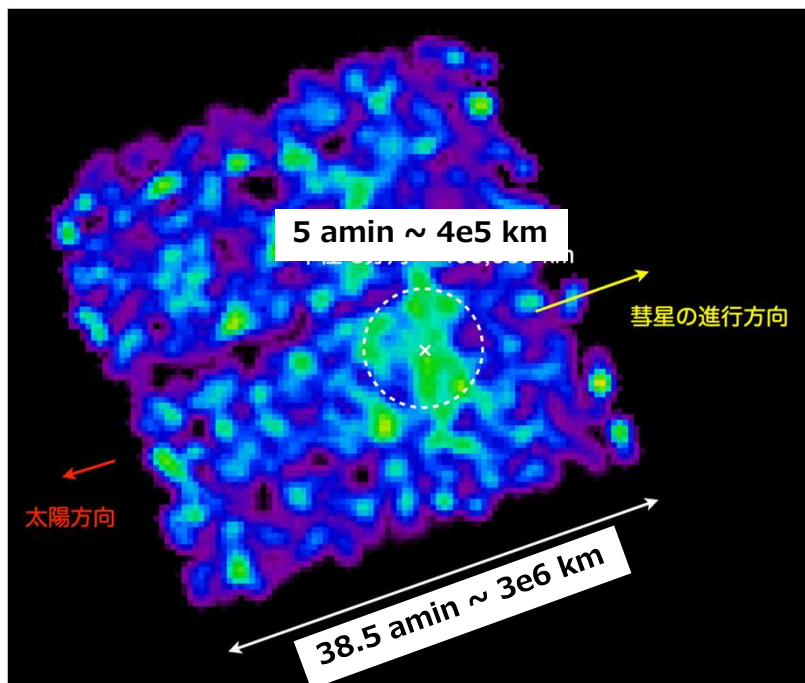


※Diffuse source is also evaluated in the paper

Xtend Observation of the Interstellar Object 3I ATLAS



- Performed at the end of Nov. 2025 (2025-11-26 – 11-28)
- First X-ray detection from an interstellar object
- CNO charge exchange lines...?
- Further analysis in progress



<https://www.xrism.jaxa.jp/topics/news/1204/>

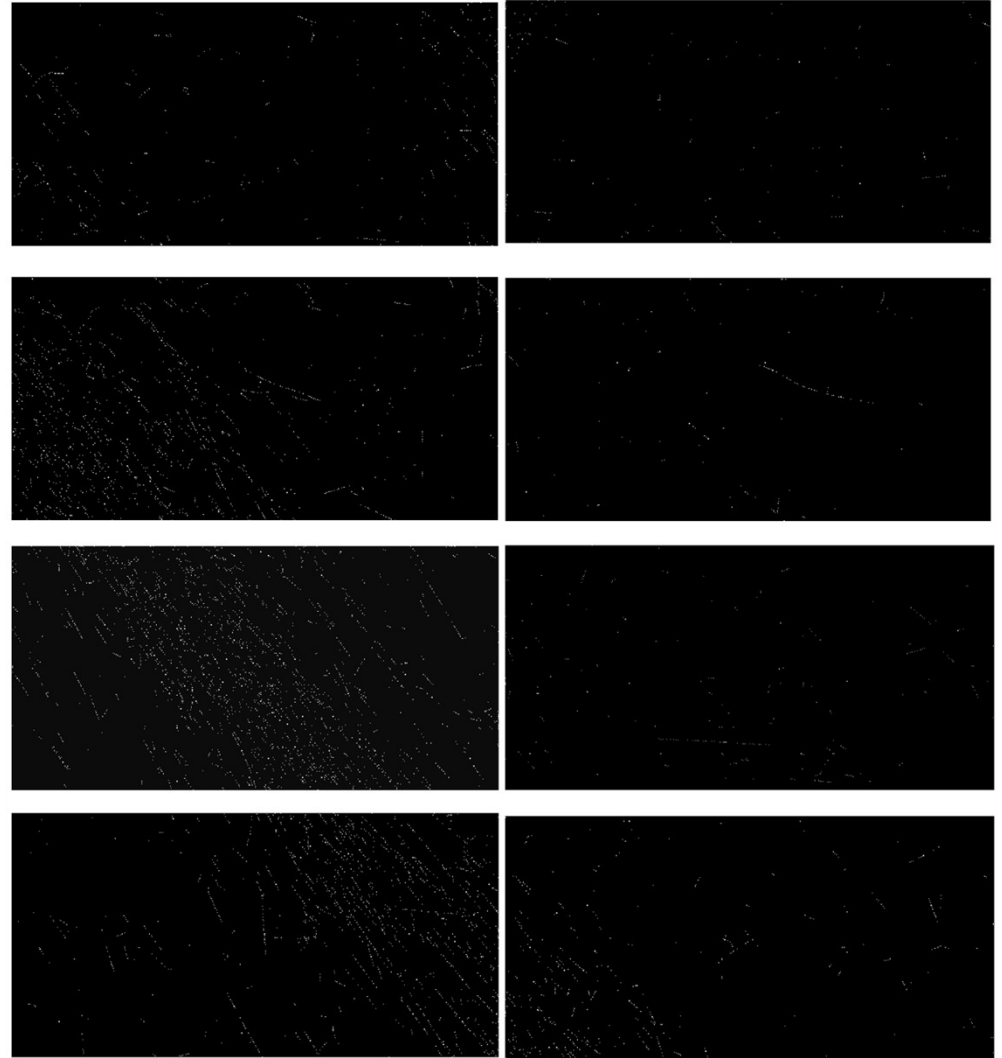
Particle Event in 2026-03-29



- A prominent event occurred at 03/29 15:57:13
- Trajectory of an energetic particle was recorded as > 6000 events at a single exposure frame
- Impressive image just for share
- Any science...?

Note:

- Extracted from QL events (not final products)
- CCDs got no damage from this event





- **Xtend is going well**
- Calibration activity is being continued
- Energy resolution is degraded as expected from the ground experiments
- Contamination is well suppressed, while a new method is applied to estimate
- Pile-up is evaluated with a simulation
- Xtend worked as a cloud chamber

