



AI-Assisted Super-Resolution and De-Noising for XMM-Newton EPIC-pn

Sam Sweere

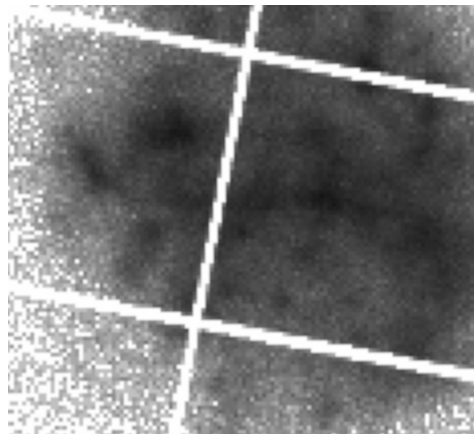
Remote trainee at the XMM SOC, ESAC, ESA

Supervisors:

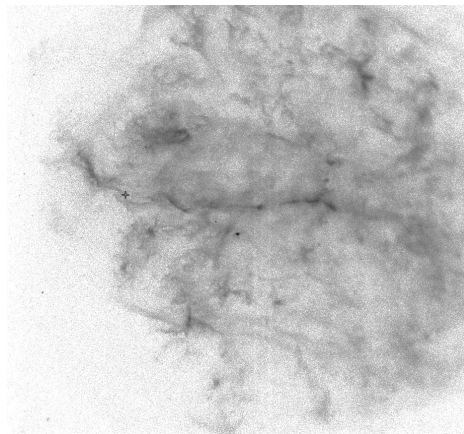
*Ivan Valtchanov, Maggie Lieu, Eva Verdugo, Antónia Vojtekova,
Maria Santos-Lieo*

Research Objective

- Increase the scientific exploratory value of XMM-Newton data
 - Decrease noise
 - Improve spatial resolution



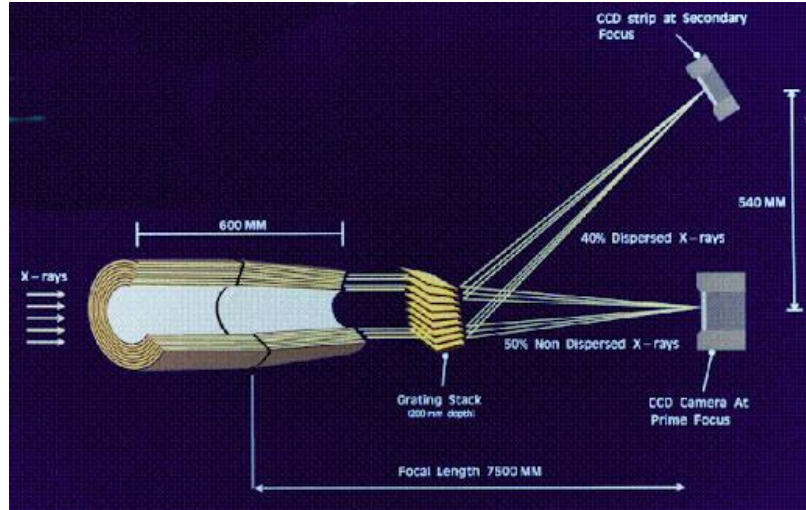
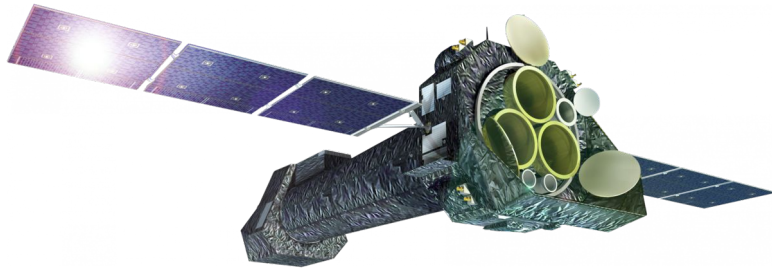
XMM
(6 arcsec FWHM PSF)



Chandra
(0.5 arcsec FWHM PSF)

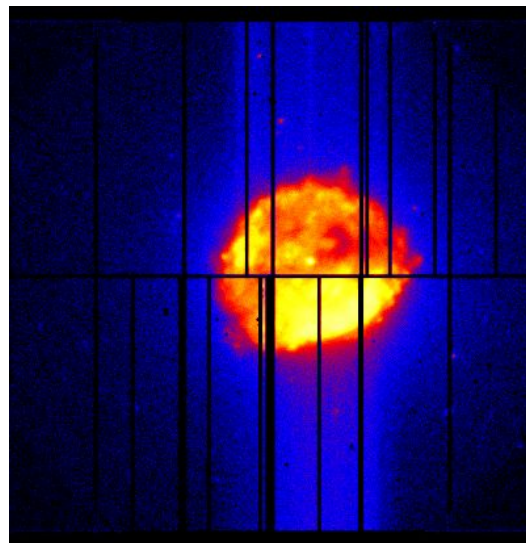
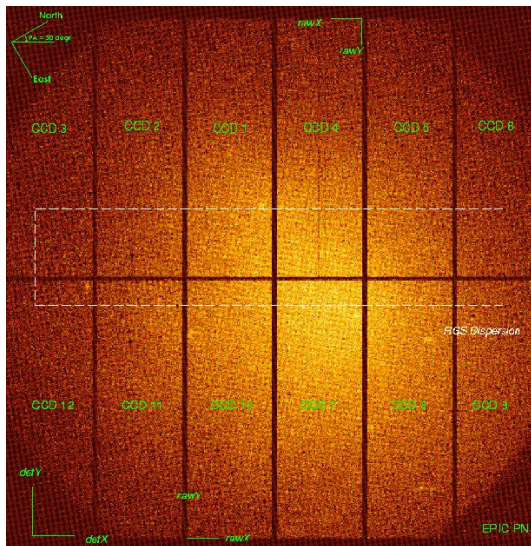
XMM-Newton

- Over 20 years in operation, vast amount of data
- X-ray telescope
- Multi-shell grazing incidence mirrors



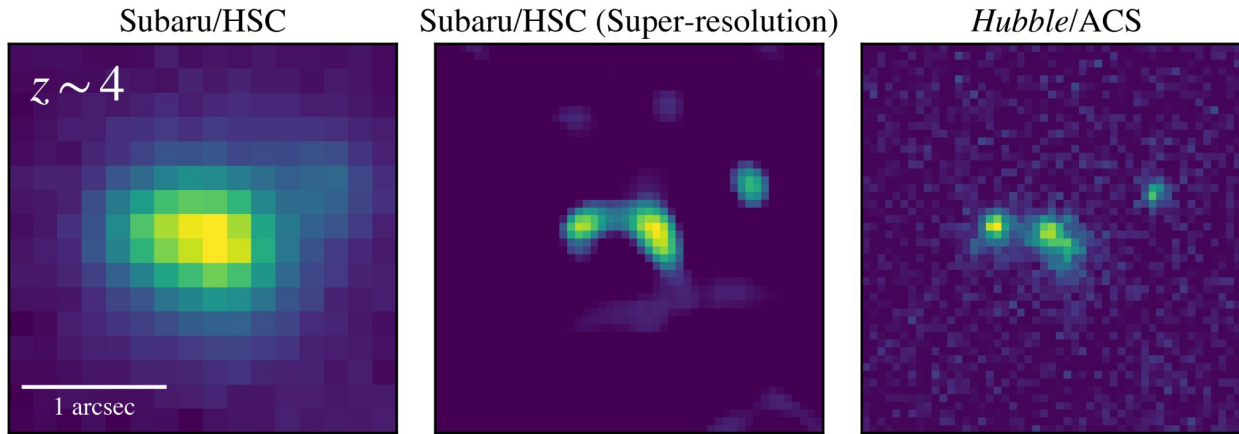
XMM EPIC-PN

- Most sensitive sensor
- 0.5 - 2.0 keV energy range

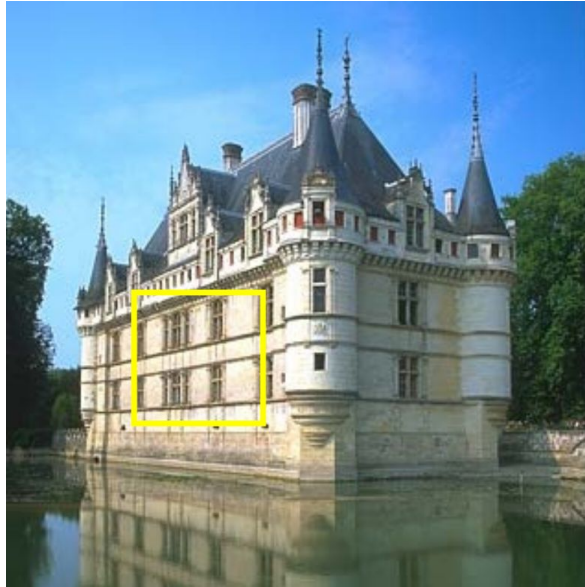


Traditional Approaches

- Richardson–Lucy deconvolution
- Not applicable to XMM because of changing PSF
- Machine Learning (AI)



AI Image Super-Resolution

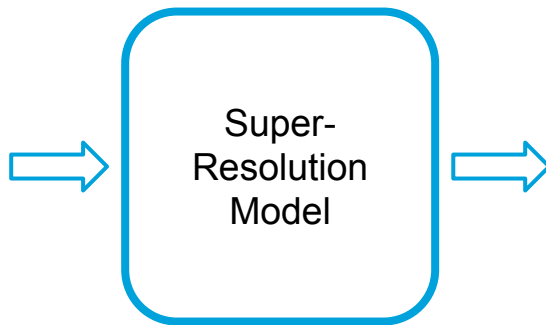


Low Resolution
Input

AI Image Super-Resolution



Low Resolution Input
(1x)

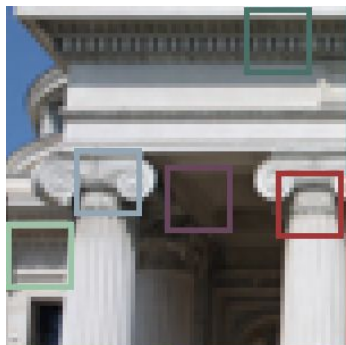


Generated
Super-Resolution
(4x)



Ground Truth

State of the Art Super-Resolution

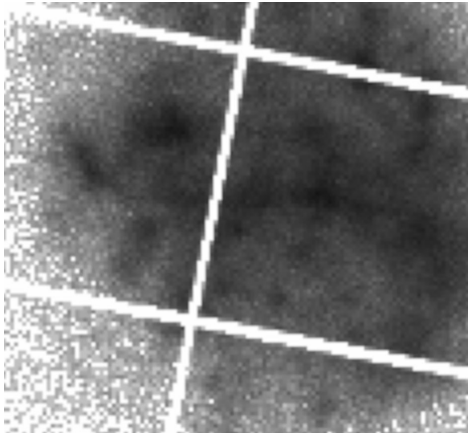


Low Resolution Input
(1x)

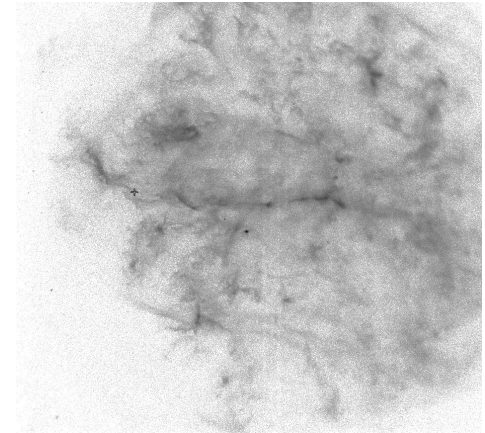
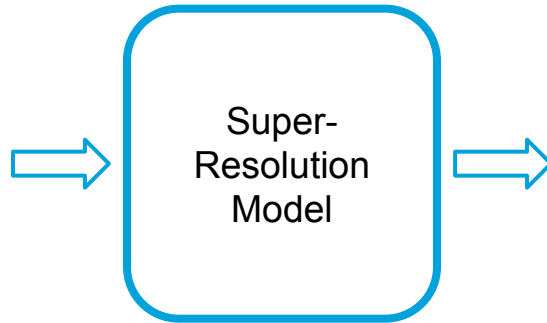
Generated
Super-Resolution
(4x)

Ground-Truth

XMM Super-Resolution



XMM
(6 arcsec FWHM PSF)



Chandra
(0.5 arcsec FWHM PSF)

How do you build a AI Super-Resolution model?

Intuition, how would a human do it?



Intuition, how would a human do it?



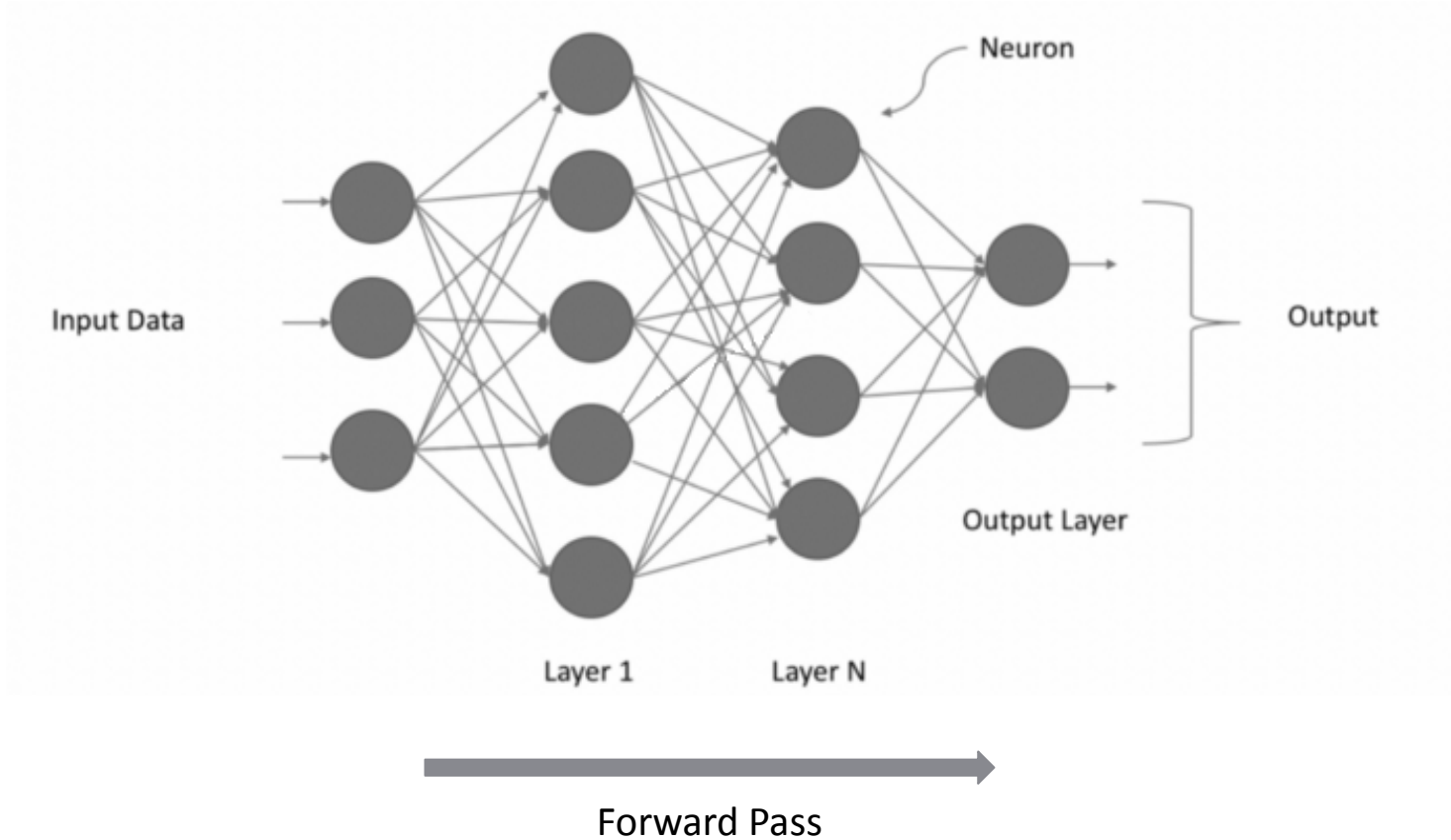
How to build a Denoising/SR AI Model

- Model Architecture
- Loss Functions
- Training Data
- Evaluation Metrics

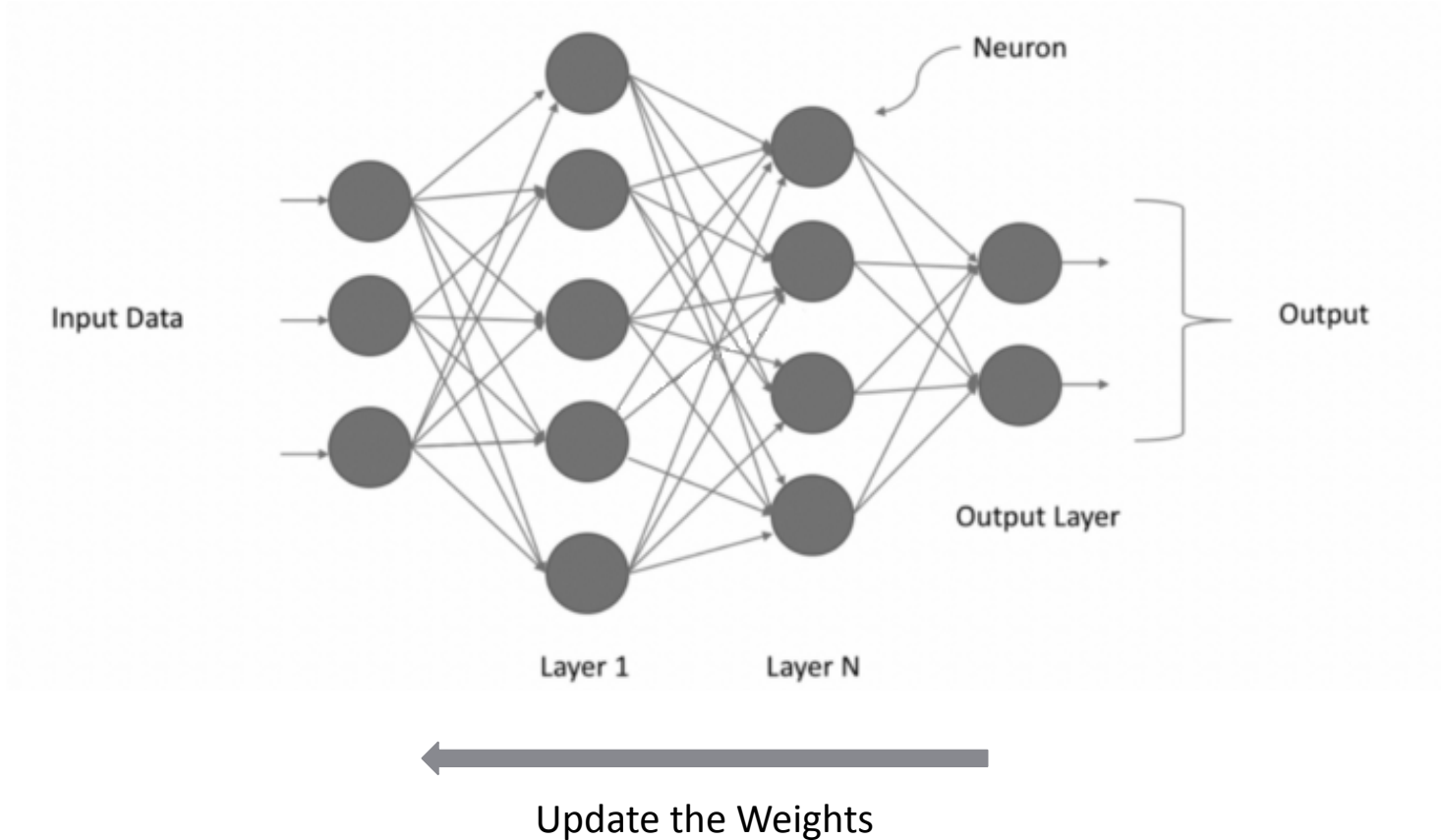
How to build a Denoising/SR AI Model

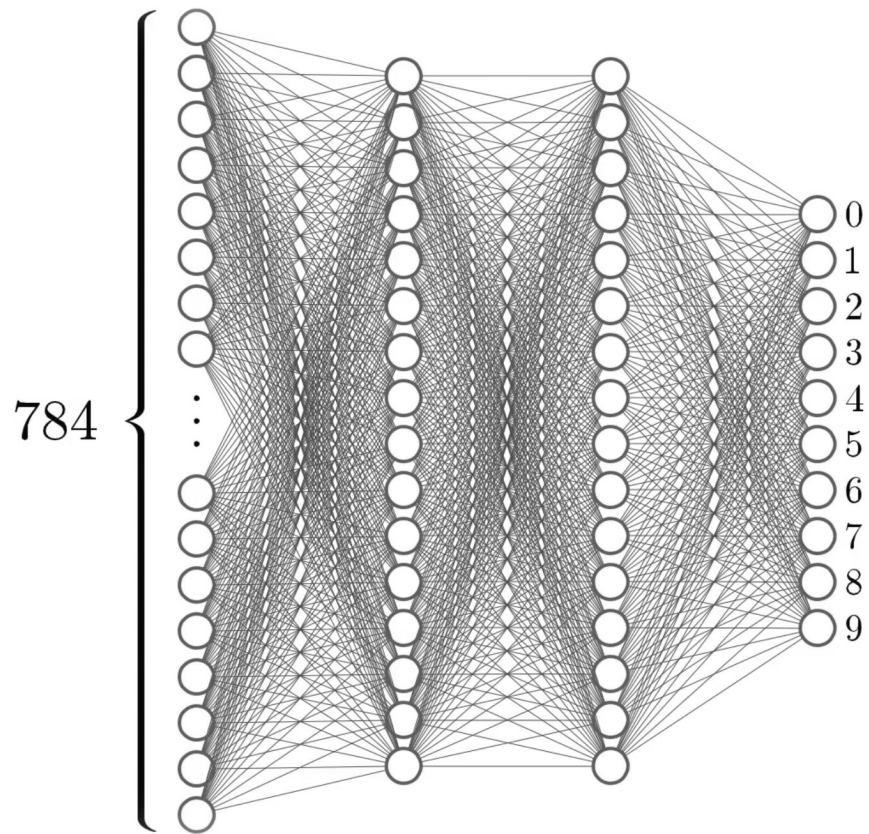
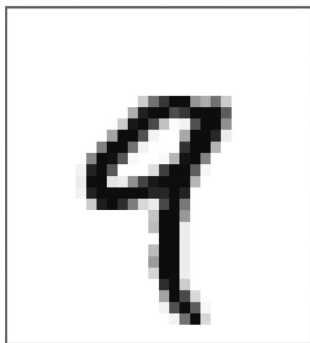
- **Model Architecture**
- Loss Functions
- Training Data
- Evaluation Metrics

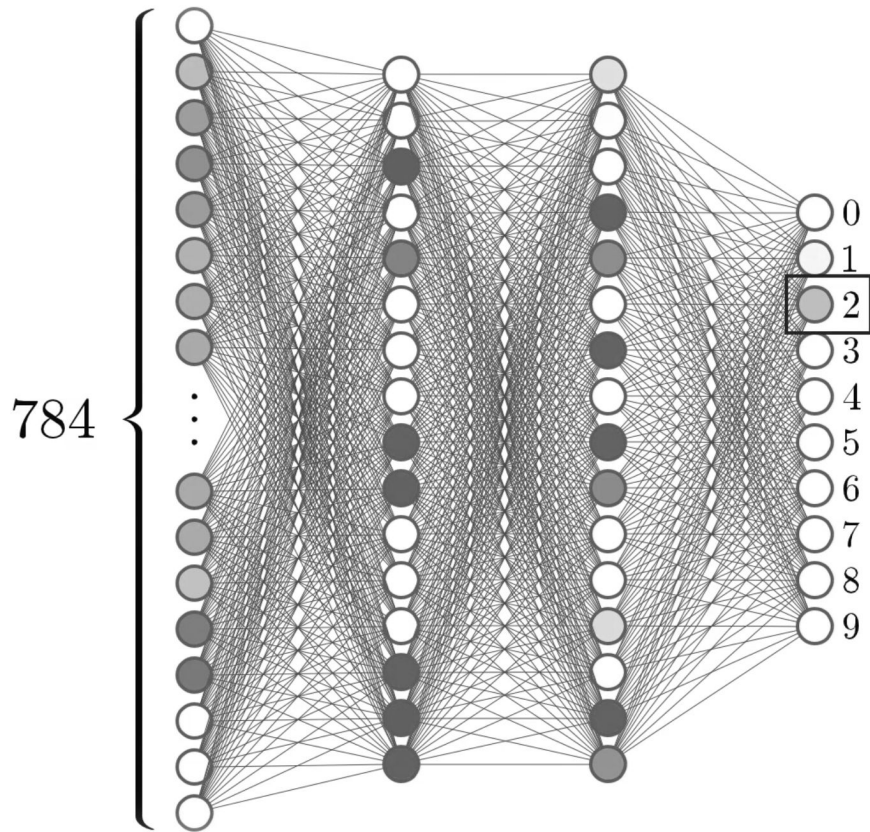
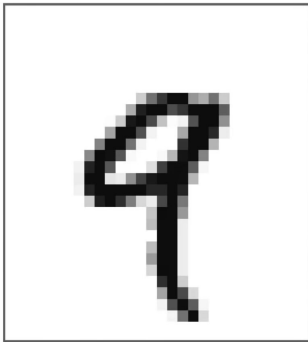
Deep Neural Networks

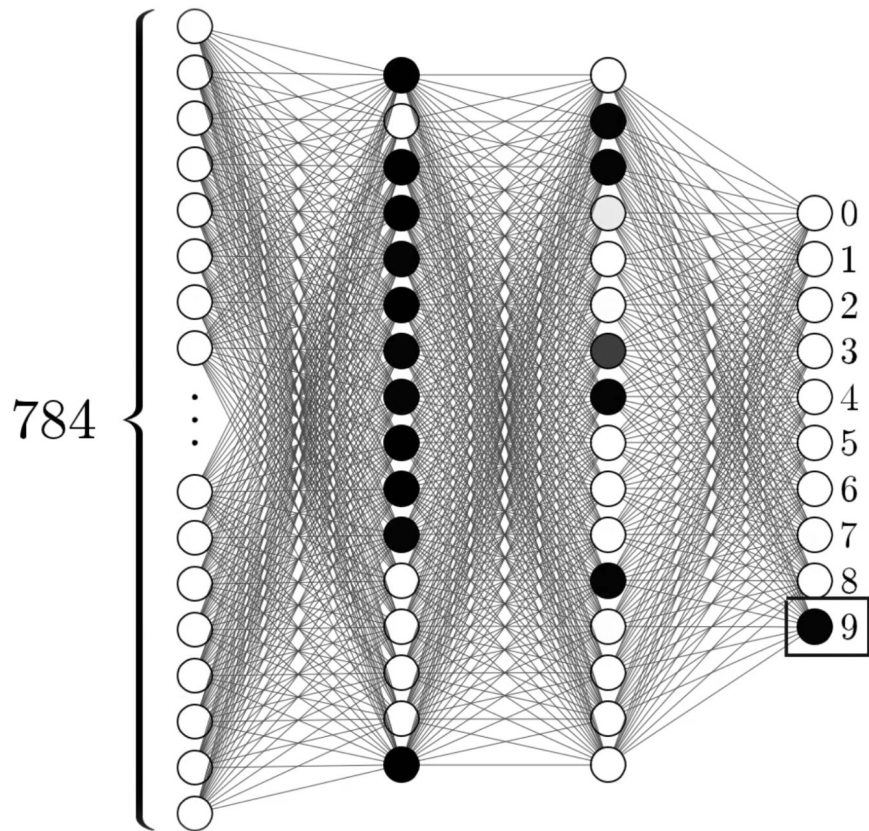
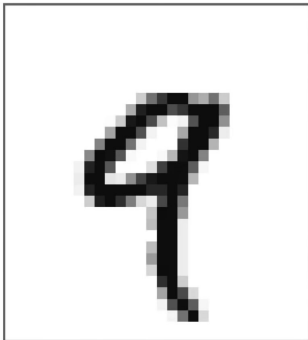


Deep Neural Networks





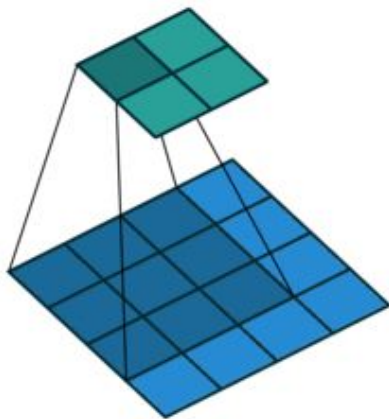




Convolutions



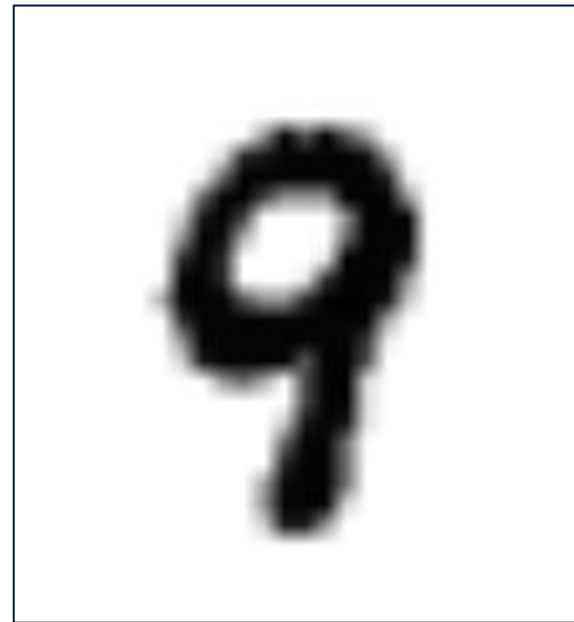
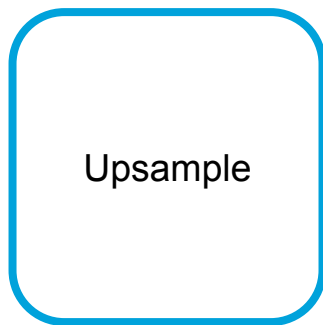
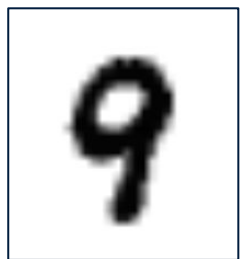
0	1	0
1	-4	1
0	1	0



Learned Convolutions



Upsampling



Deep Convolutional Neural Network

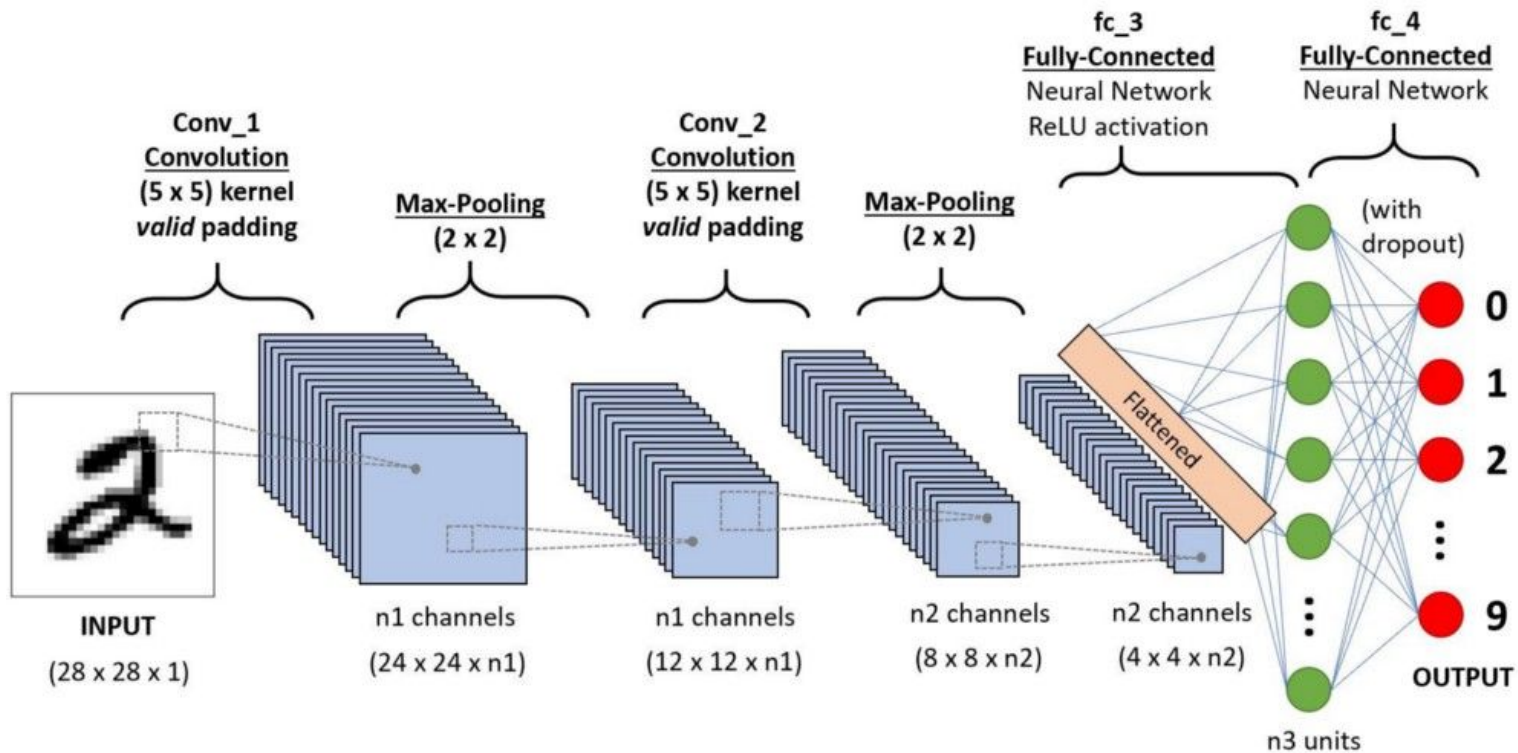
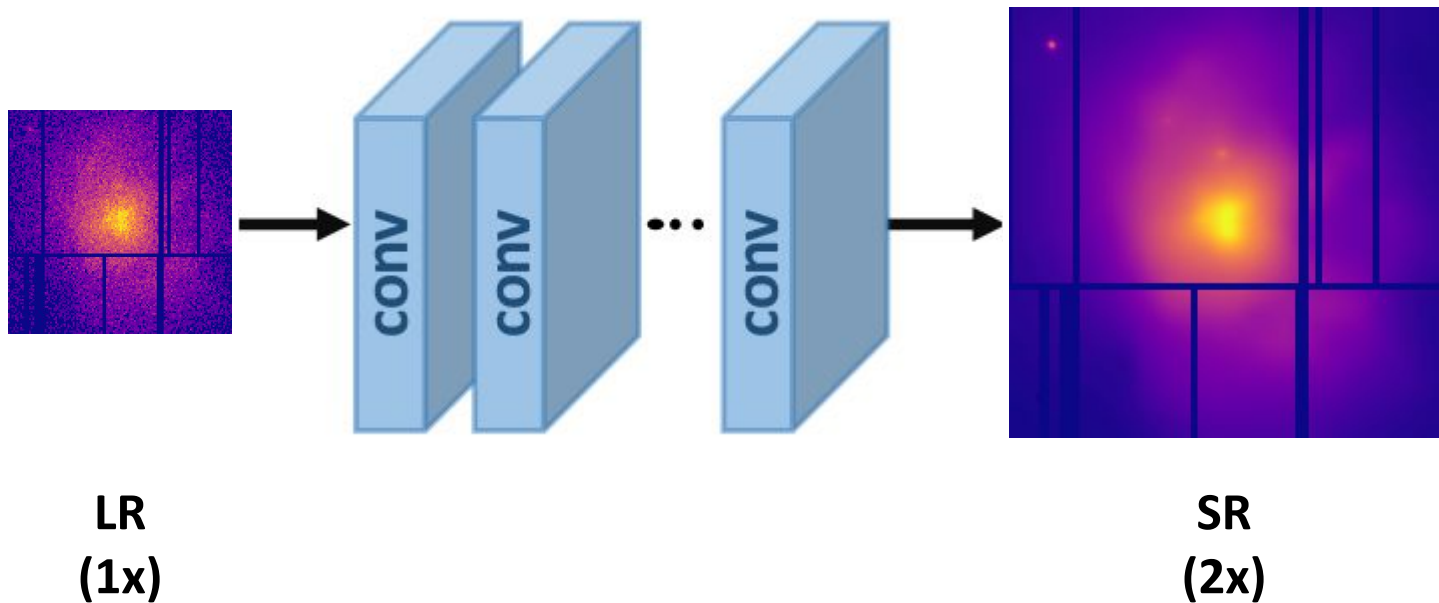
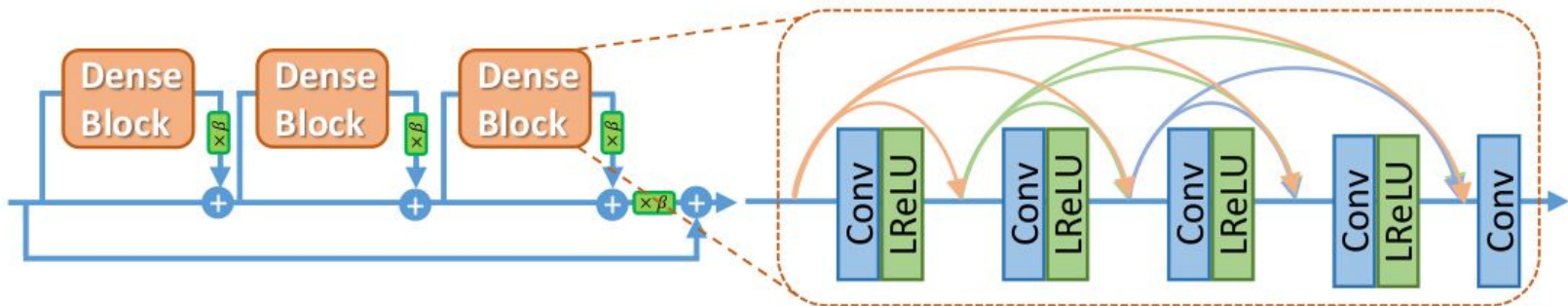


Image to Image Networks

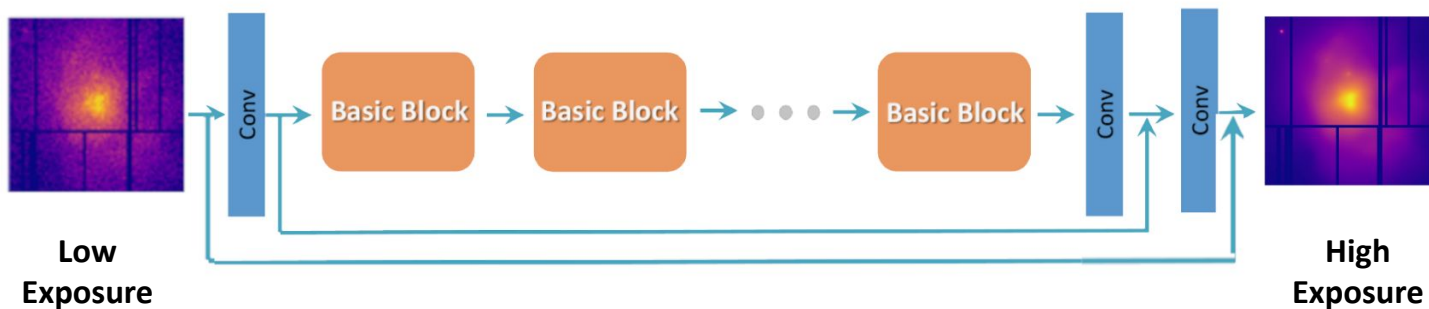


Residual in Residual Dense Block (RRDB)

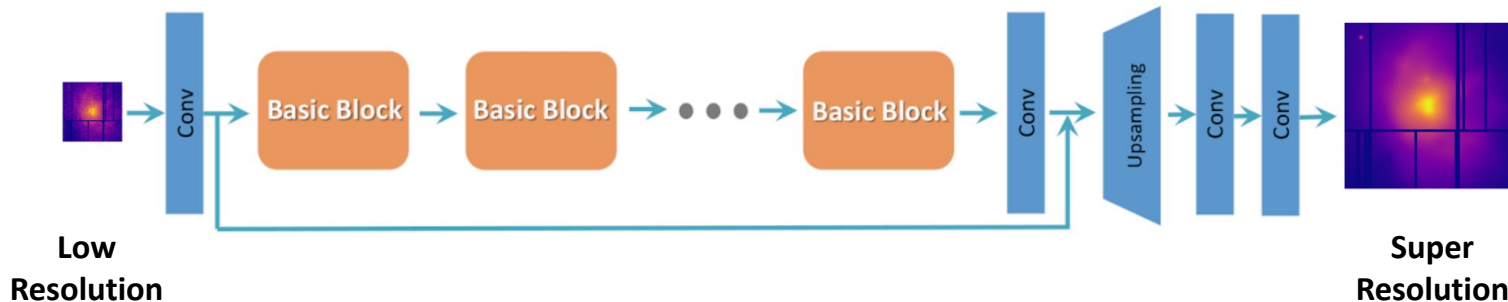


RRDB Denoise and Super-Resolution Models

Denoise Model



Super-Resolution Model



How to build a Denoising/SR AI Model

- Model Architecture
- **Loss Functions**
- Training Data
- Evaluation Metrics

Loss metrics

- Pixel loss
 - L1 (Least Absolute Deviations)
 - Poisson
- Adversarial loss

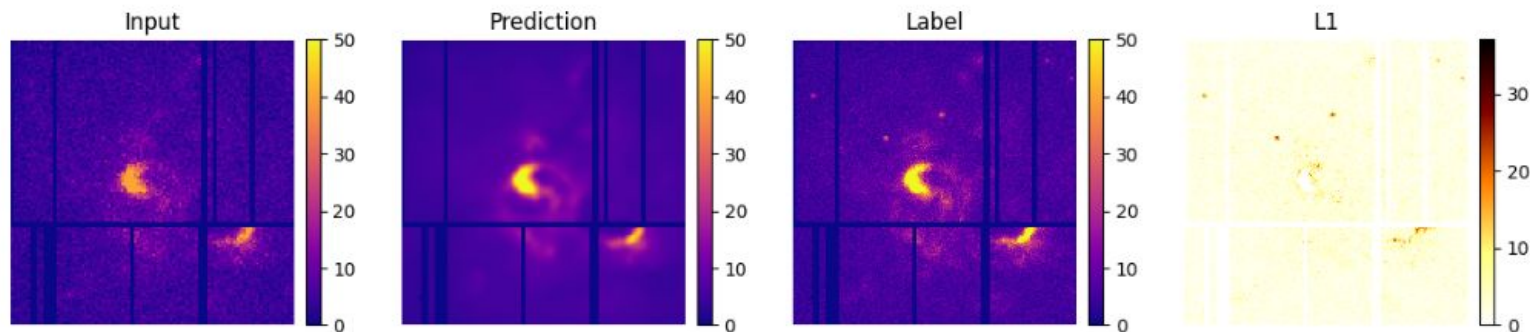
Pixel Loss

- L1 (Least Absolute Deviations):

$$L_1 = \frac{1}{N} \sum_{i=0}^N |y_i - \hat{y}_i|$$

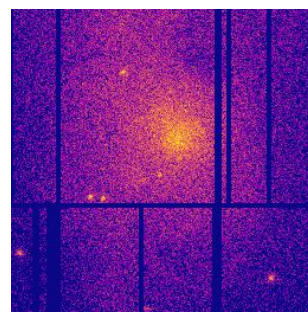
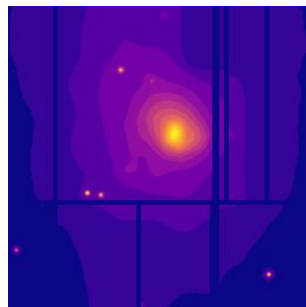
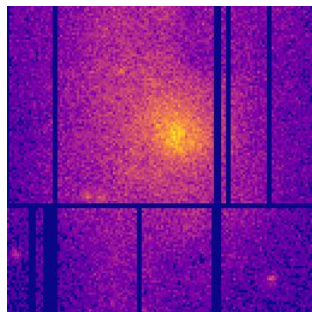
- Poisson Loss:

$$Poisson = \frac{1}{N} \sum_{i=0}^N (\hat{y}_i - y_i \log(\hat{y}_i))$$

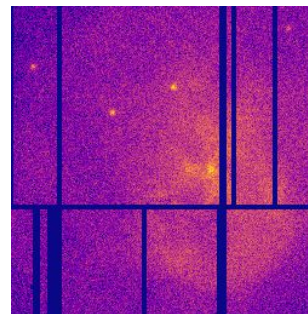
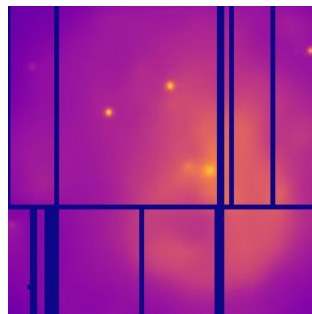
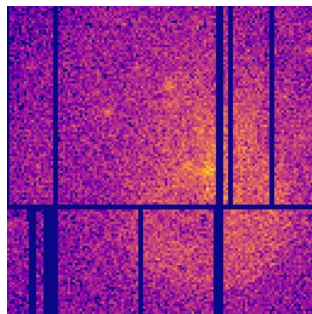


L1 Loss vs Poisson Loss

L1 Loss:



Poisson Loss:

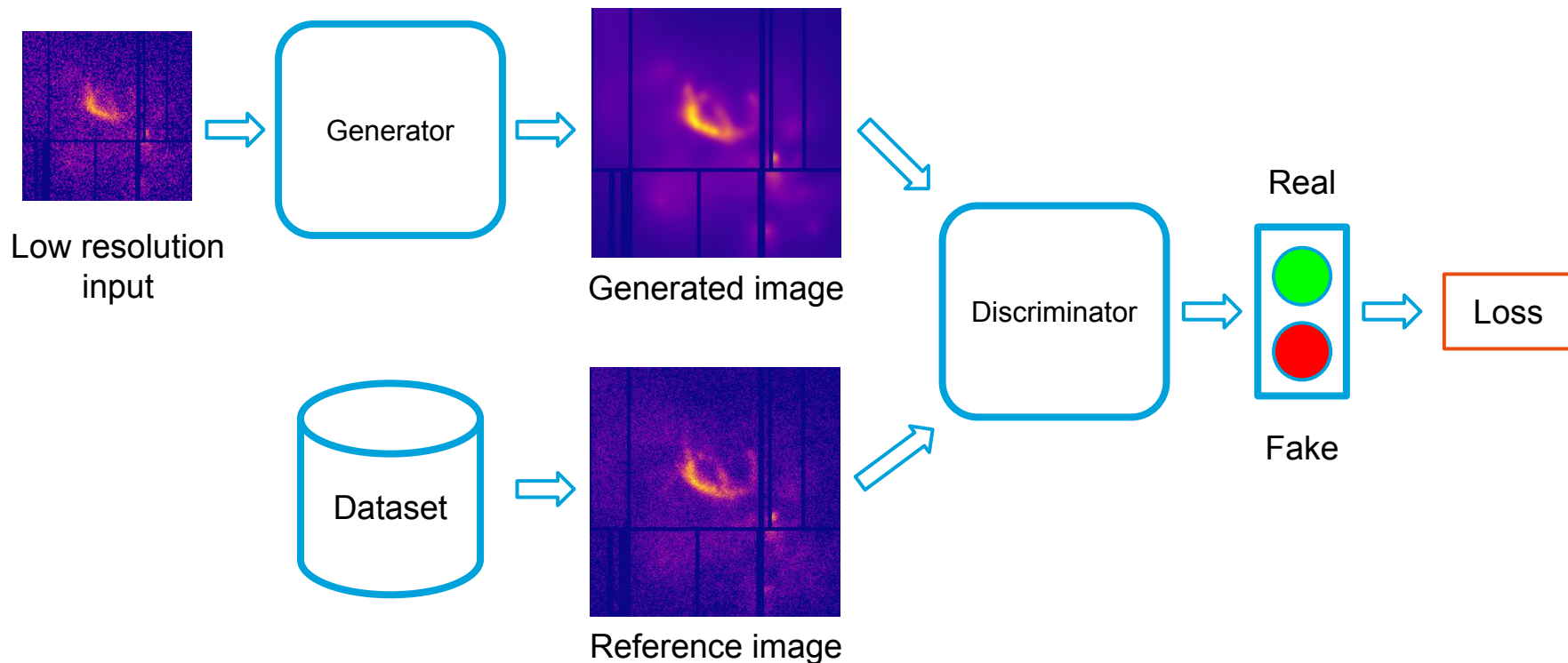


Low Resolution
(1x)

Generated
(2x)

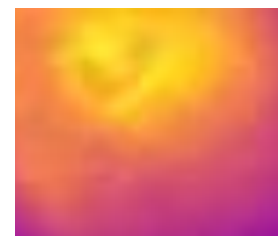
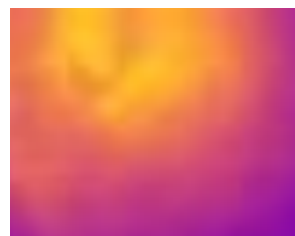
Target
(2x)

Generative Adversarial Network (GAN)

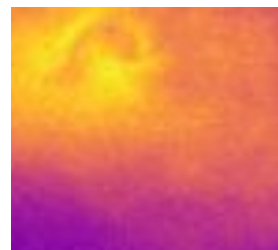
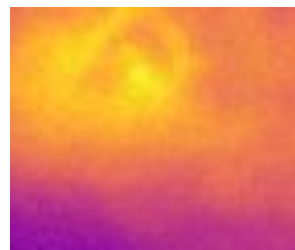
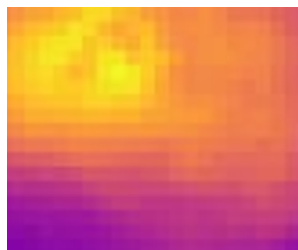


L1 Loss vs Adversarial Loss

L1 Loss:



L1 + Adversarial Loss:



Low Resolution
(1x)

Generated
(2x)

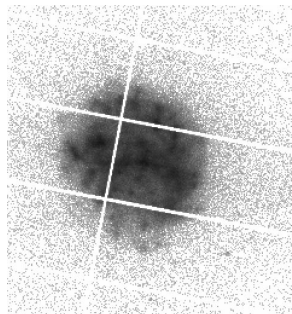
Target
(2x)

How to build a Denoising/SR AI Model

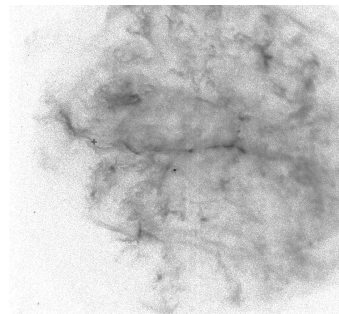
- Model Architecture
- Loss Functions
- **Training Data**
- Evaluation Metrics

Training Data

- XMM cannot zoom
- Chandra has a higher spatial resolution, however:
 - Different image characteristics
 - Only a few hundred good pairs



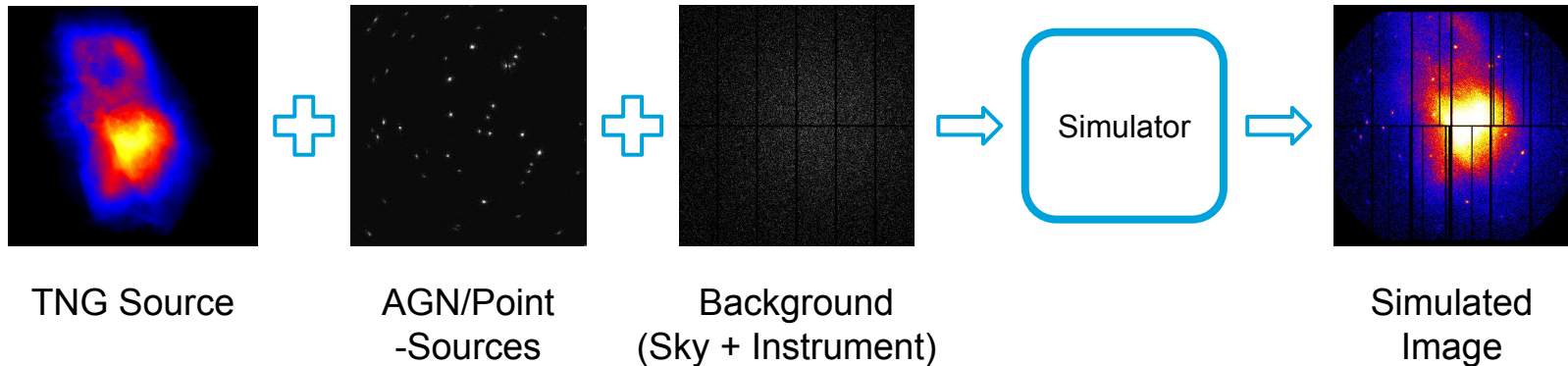
XMM



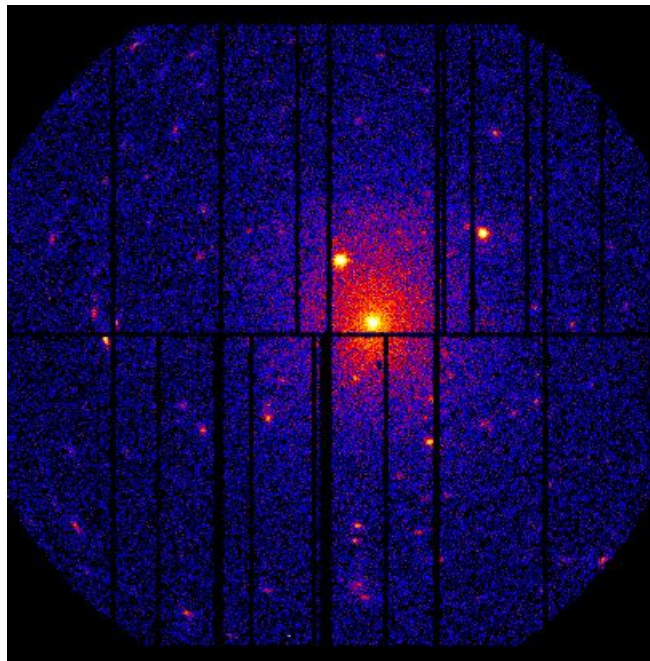
Chandra

XMM Simulation

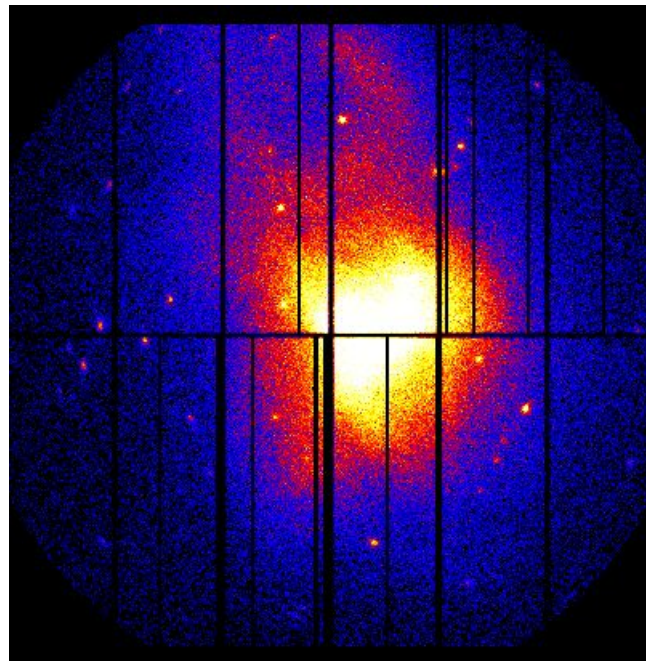
- Create training pairs with increased psf and spatial resolution
- Increase data on rare astronomical structures
- SIXTE X-ray simulation software (Bamberg University)
- IllustrisTNG simulation input



XMM Simulation: Comparison

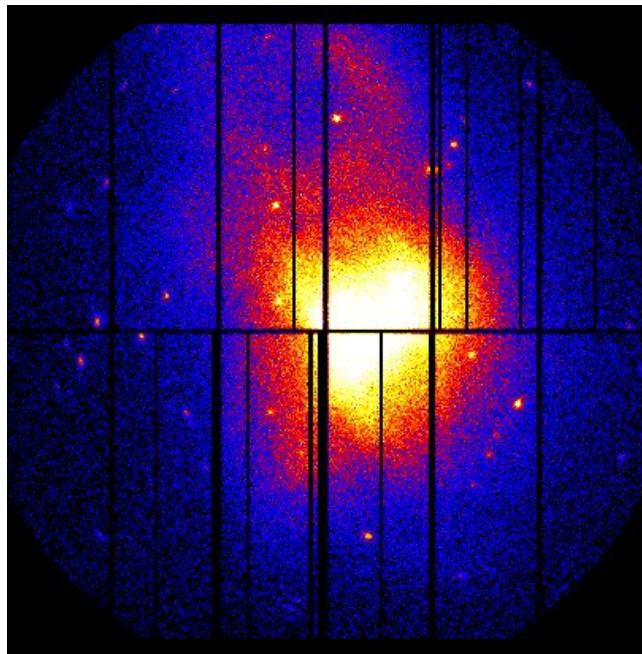


Real

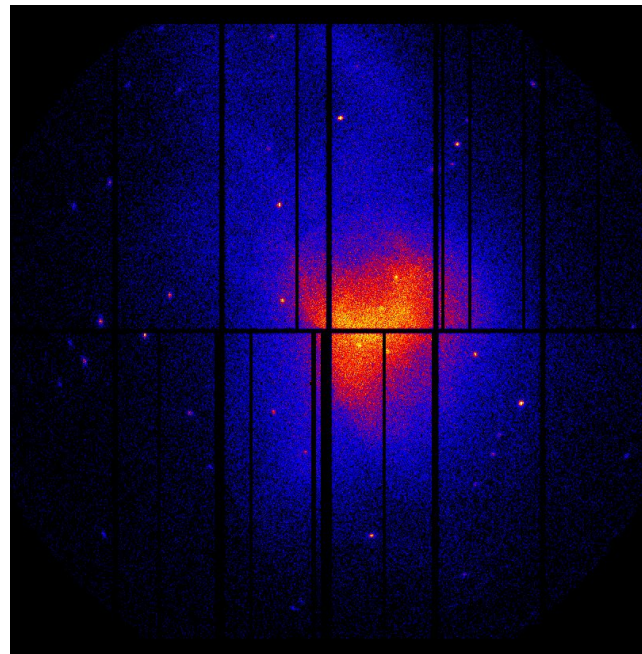


Simulated

XMM Simulation: PSF and Spatial Resolution



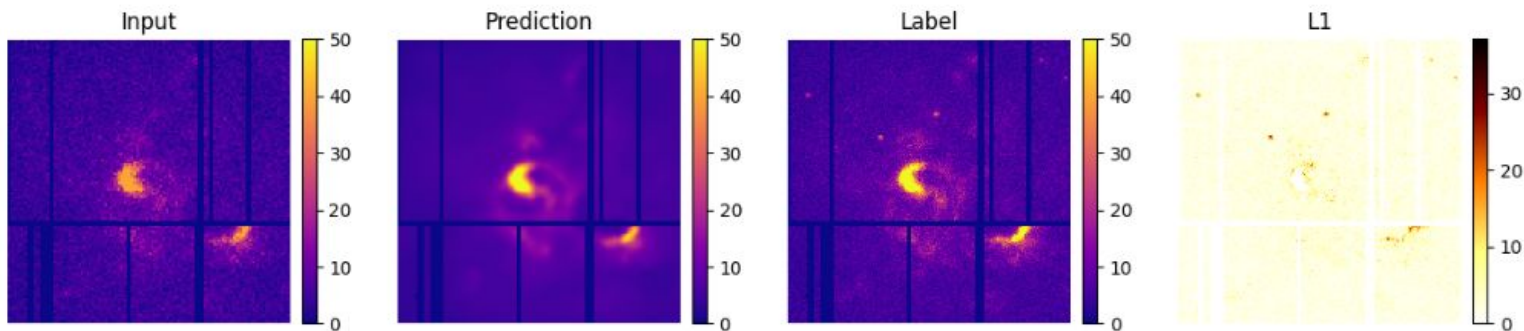
1x



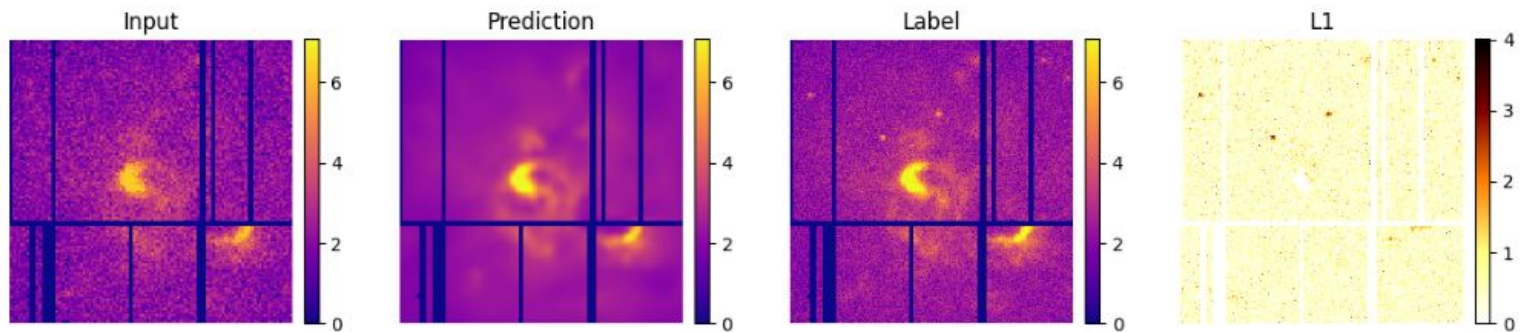
2x

Data Scaling

No scaling:



Sqrt:



How to build a Denoising/SR AI Model

- Model Architecture
- Loss Functions
- Training Data
- **Evaluation Metrics**

Evaluation Metrics

- Visually compare with Chandra
- Least Absolute Deviations (L1)
- Peak Signal to Noise Ratio (PSNR)
- Structural Similarity Index (SSIM)

Peak Signal to Noise Ratio (PSNR)

- Measure of ratio between maximum possible value and background noise

$$MSE = \frac{1}{N} \sum_{i=0}^N (y_i - \hat{y}_i)^2$$

$$PSNR = 10 \cdot \log_{10} \left(\frac{MAX_I^2}{MSE} \right)$$

Structural Similarity Index (SSIM)

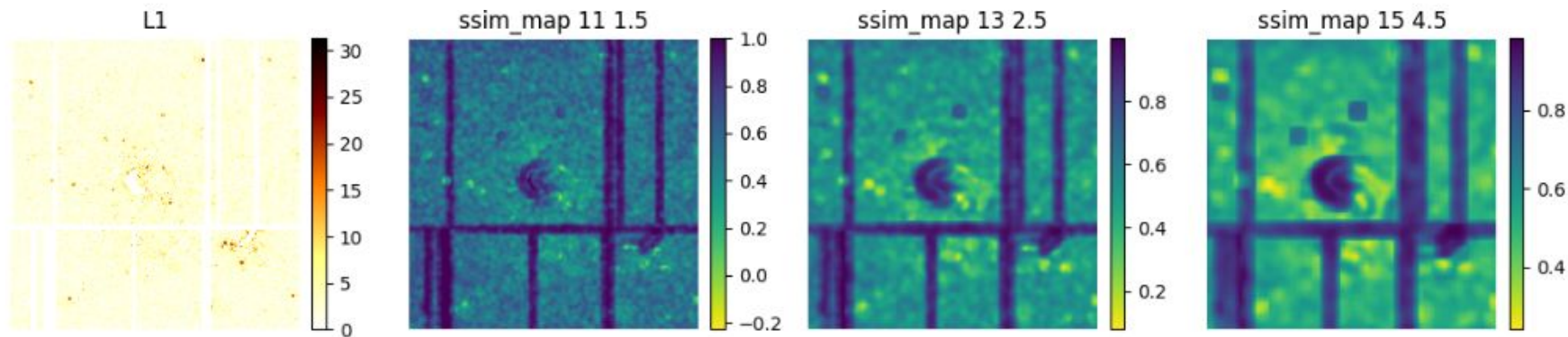
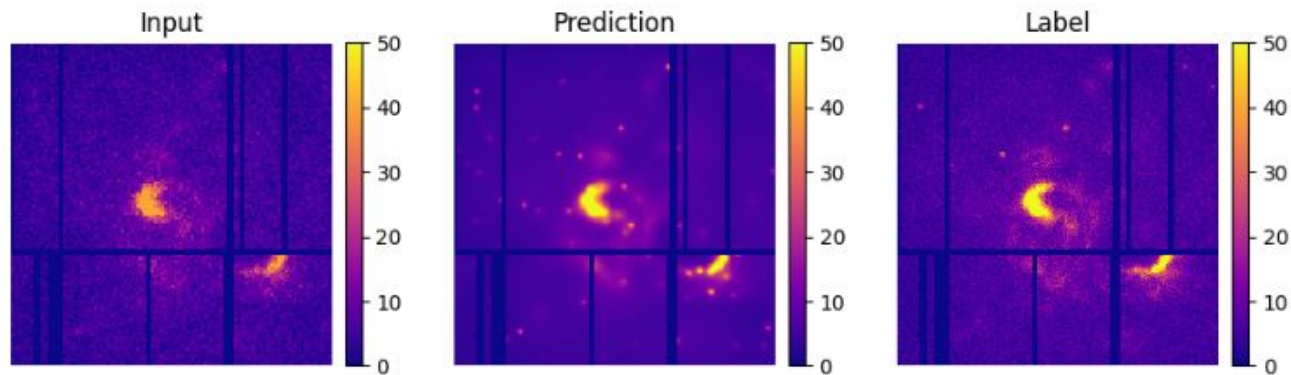
Luminance: $l(\mathbf{x}, \mathbf{y}) = \frac{2\mu_x\mu_y + C_1}{\mu_x^2 + \mu_y^2 + C_1}$

Contrast: $c(\mathbf{x}, \mathbf{y}) = \frac{2\sigma_x\sigma_y + C_2}{\sigma_x^2 + \sigma_y^2 + C_2}$

Structure: $s(\mathbf{x}, \mathbf{y}) = \frac{\sigma_{xy} + C_3}{\sigma_x\sigma_y + C_3}$

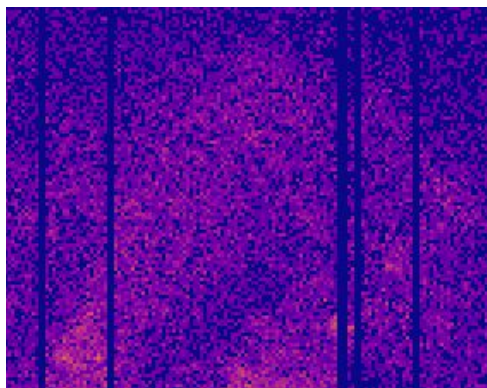
$$\text{SSIM}(\mathbf{x}, \mathbf{y}) = [l(\mathbf{x}, \mathbf{y})]^\alpha \cdot [c(\mathbf{x}, \mathbf{y})]^\beta \cdot [s(\mathbf{x}, \mathbf{y})]^\gamma$$

Structural Similarity Index (SSIM)

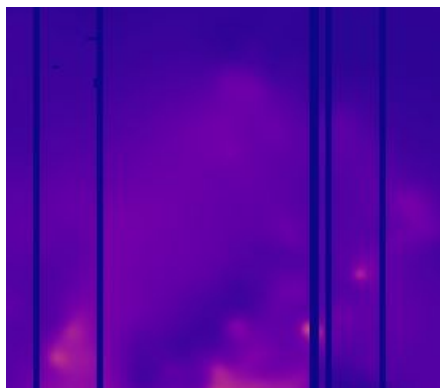


Challenges: Hallucinations

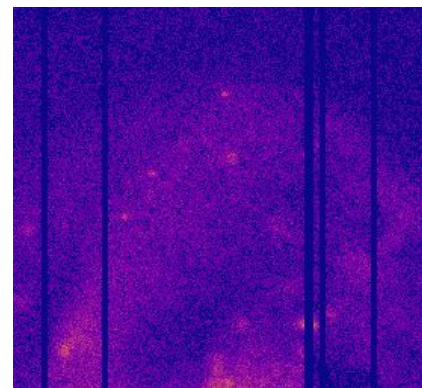
- Super-Resolution ill-posed problem
- Statistically validate the results
- Limited scope
- AI-Assisted



Low Resolution
(1x)



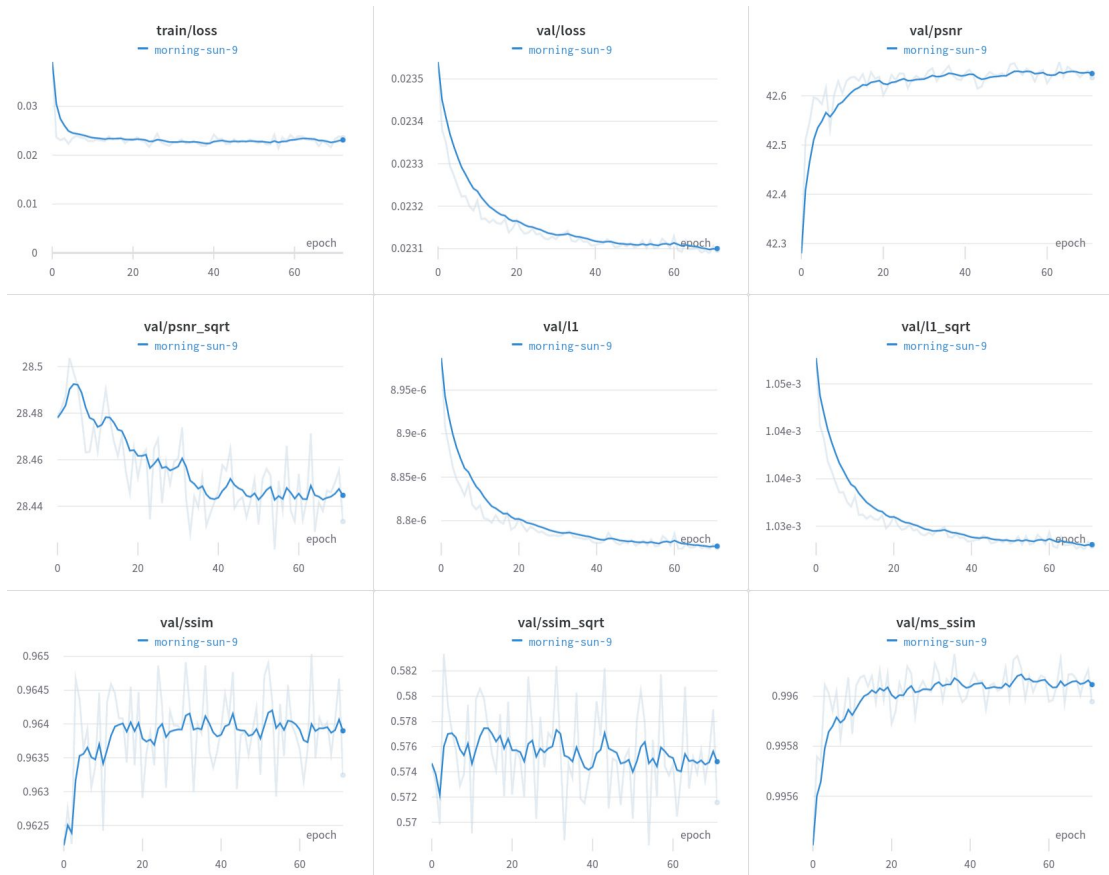
Generated
(2x)



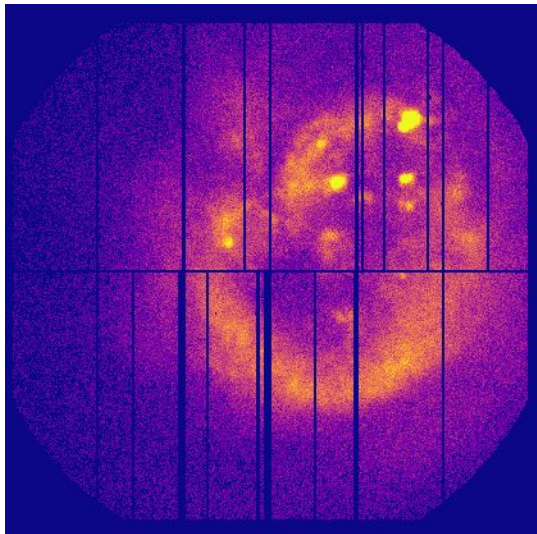
Target
(2x)

Preliminary Results

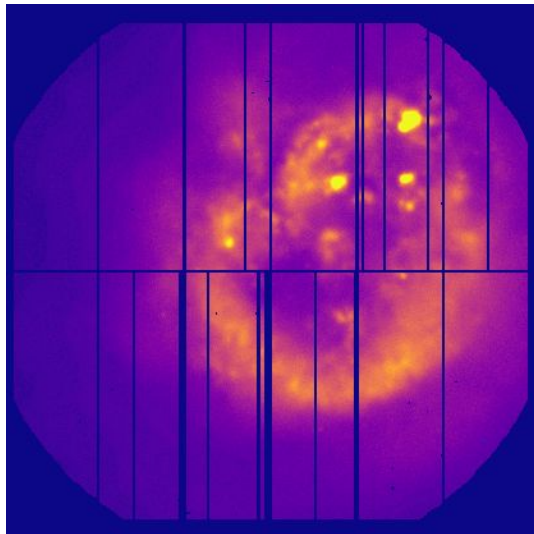
Preliminary Results Denoise: Simulated Data



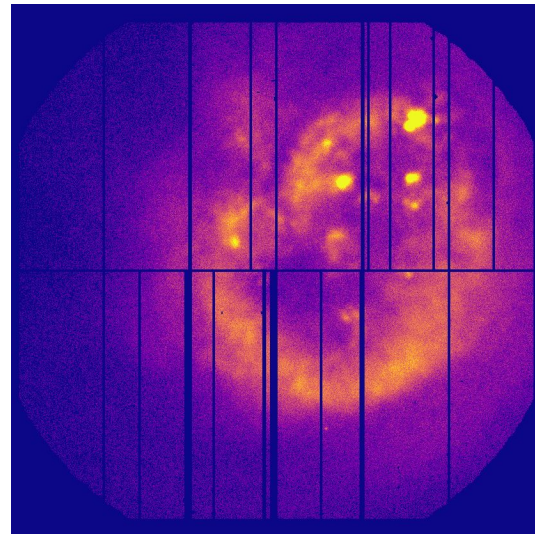
Preliminary Results Denoise: Simulated Data



Low Resolution Input
(1x, 20ks)

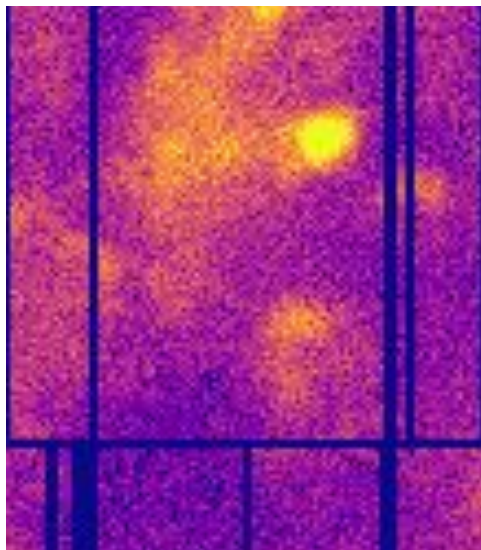


Generated
(1x, 100ks)

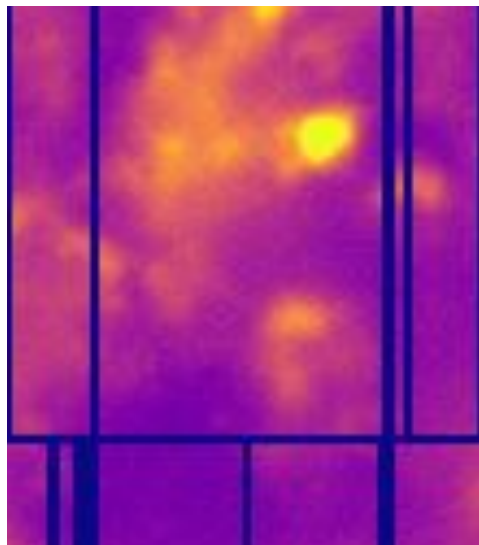


Target
(1x, 100ks)

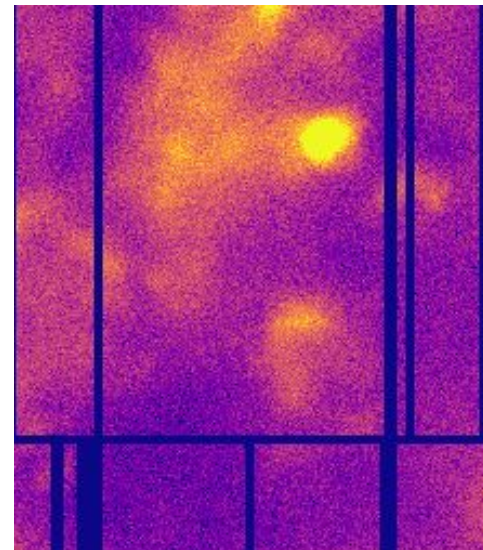
Preliminary Results Denoise: Simulated Data



Low Resolution Input
(1x, 20ks)

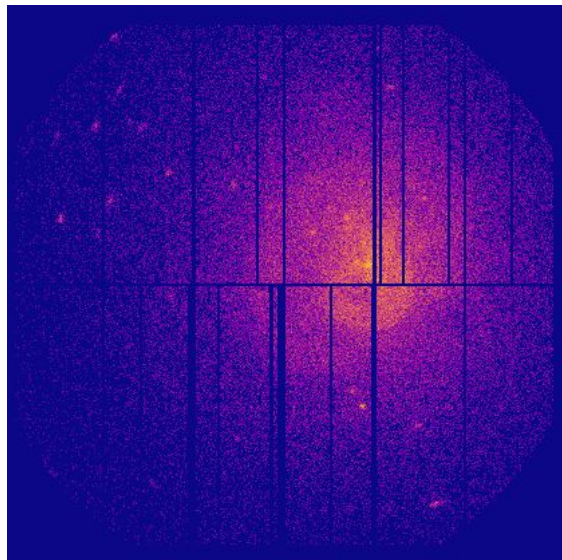


Generated
(1x, 100ks)

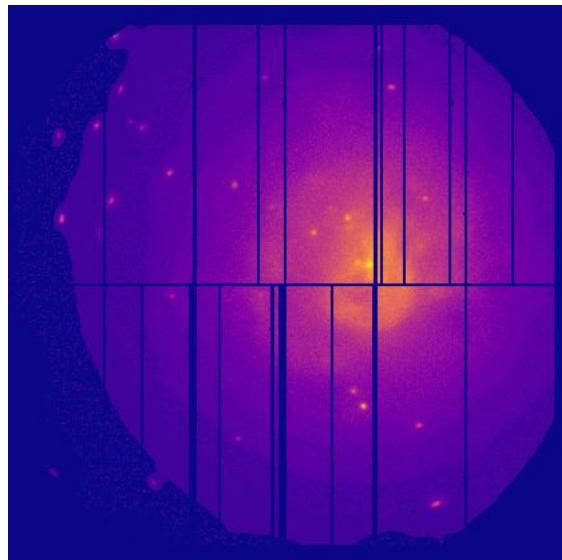


Target
(1x, 100ks)

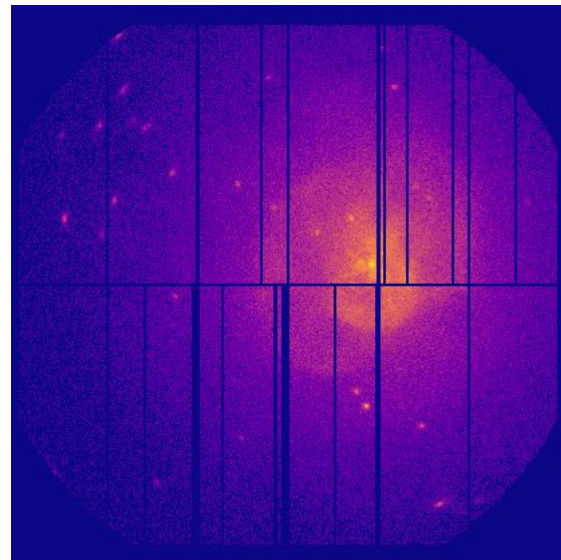
Preliminary Results Denoise: Simulated Data



Low Resolution Input
(1x, 20ks)

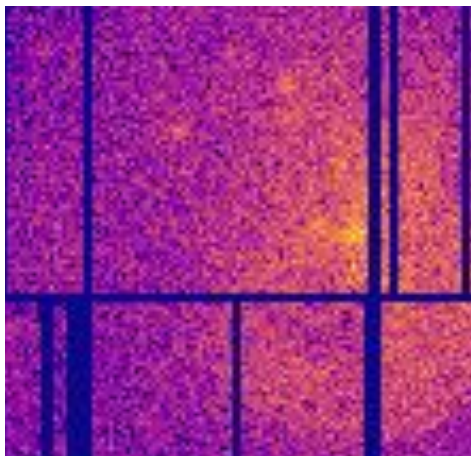


Generated
(1x, 100ks)

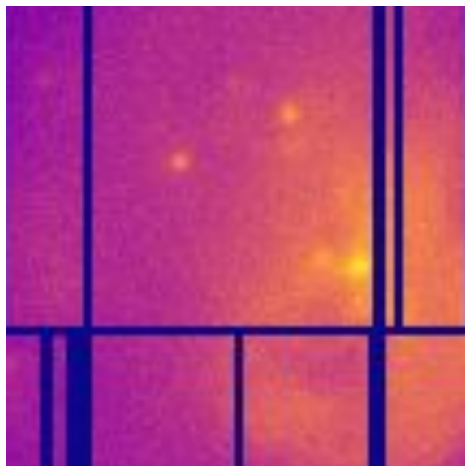


Target
(1x, 100ks)

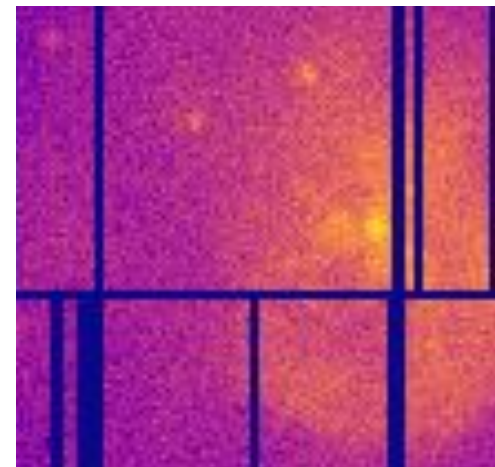
Preliminary Results Denoise: Simulated Data



Low Resolution Input
(1x, 20ks)

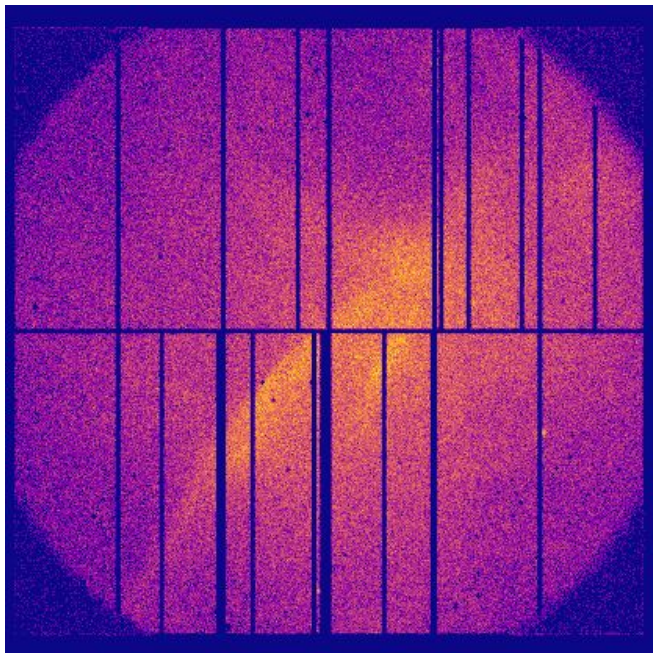


Generated
(1x, 100ks)

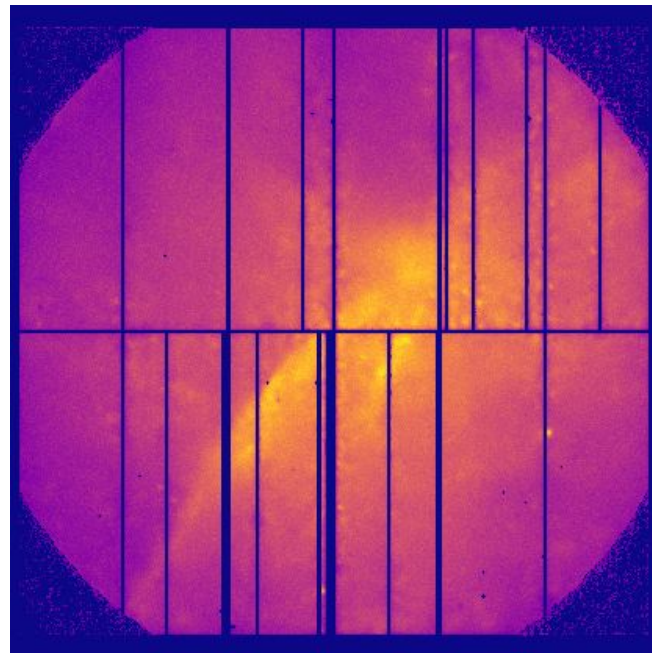


Target
(1x, 100ks)

Preliminary Results Denoise: Real Data



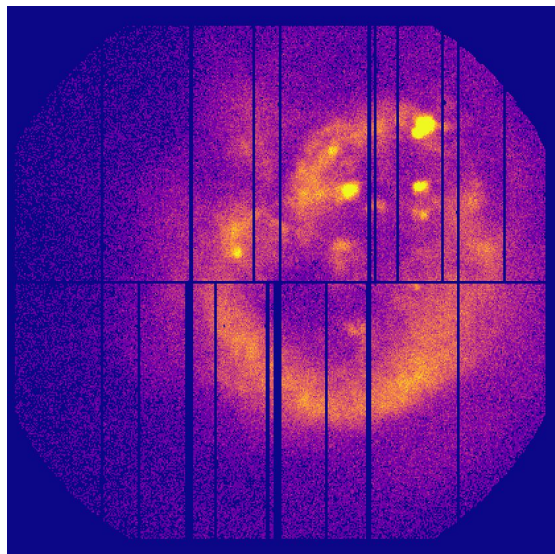
Low Resolution Input
(1x, 20ks)



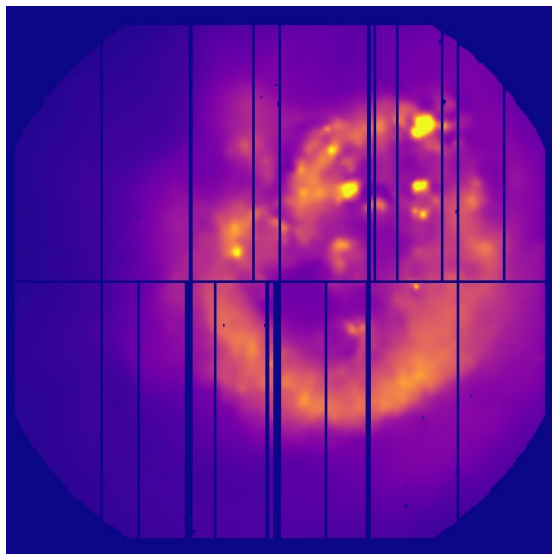
Generated
(1x, 100ks)

Source: Vela SNR

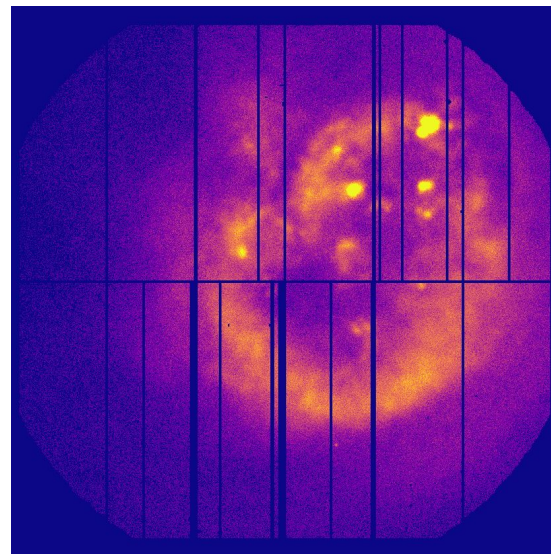
Preliminary Results SR: Simulated Data



Low Resolution Input
(1x, 20ks)

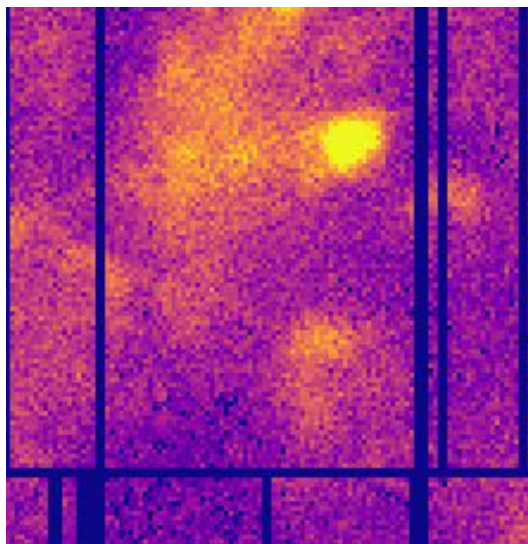


Generated
(2x, 100ks)

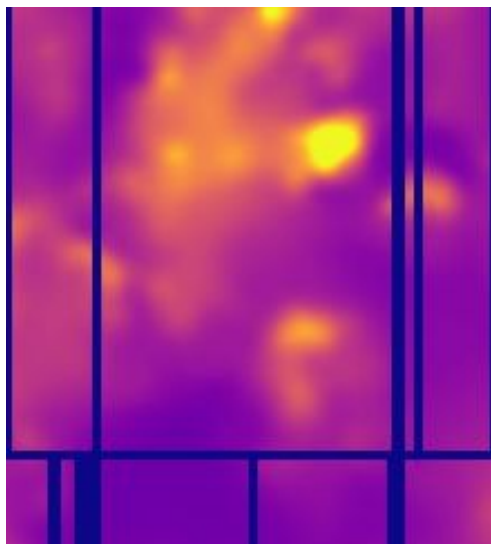


Target
(2x, 100ks)

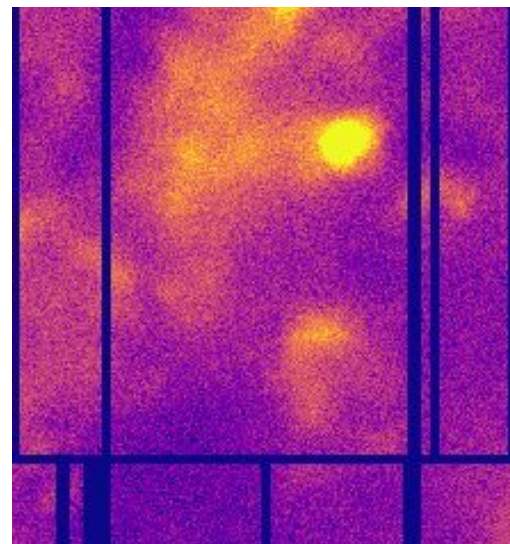
Preliminary Results SR: Simulated Data



Low Resolution Input
(1x, 20ks)



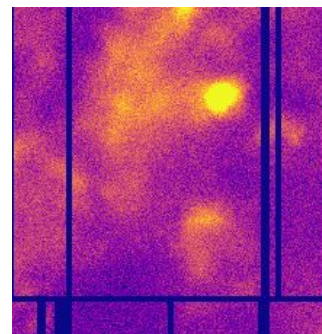
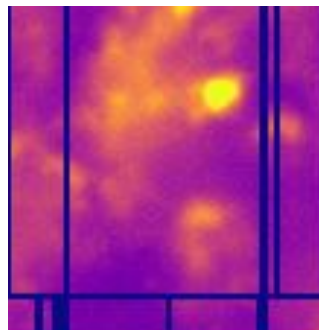
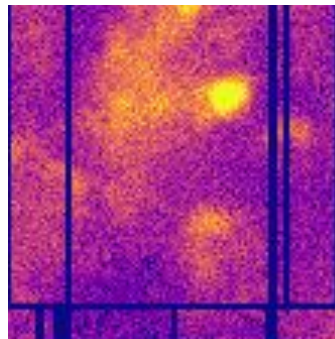
Generated
(2x, 100ks)



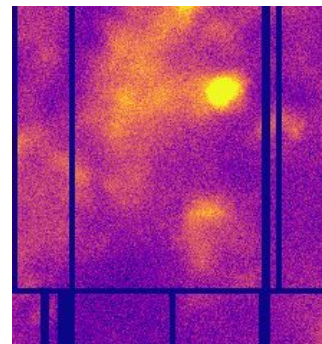
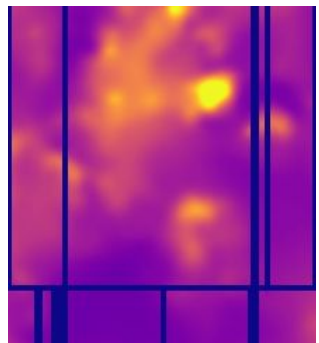
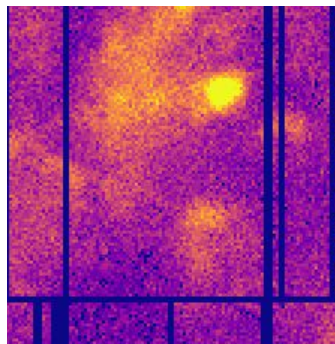
Target
(2x, 100ks)

Preliminary Results: Denoise vs SR

Denoised (1x):



Super-Resolution (2x):

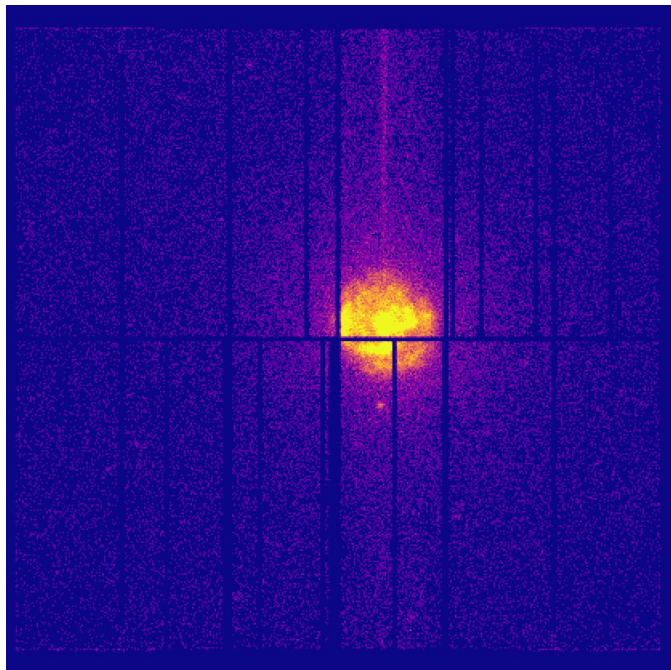


Low Resolution Input
(20ks)

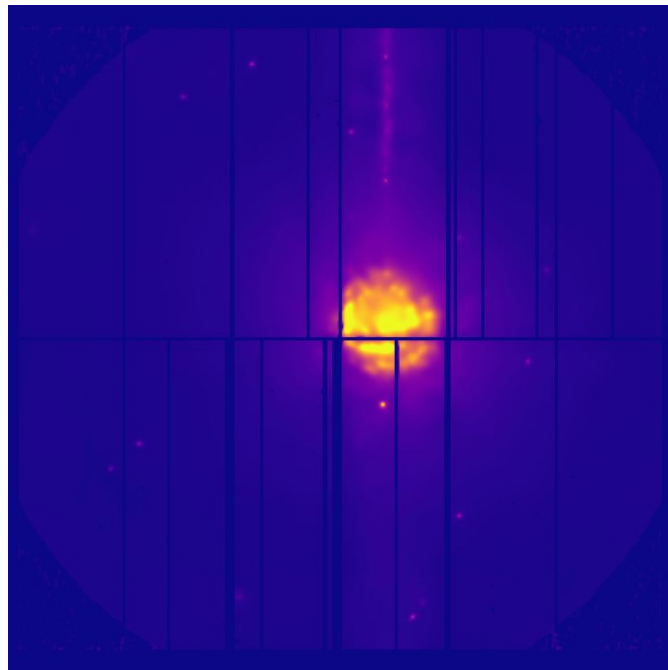
Generated
(100ks)

Target
(100ks)

Preliminary Results SR: Real Data



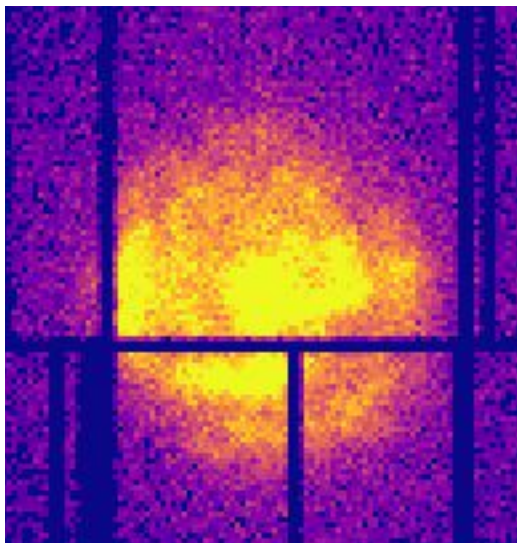
Low Resolution Input
(1x, 20ks)



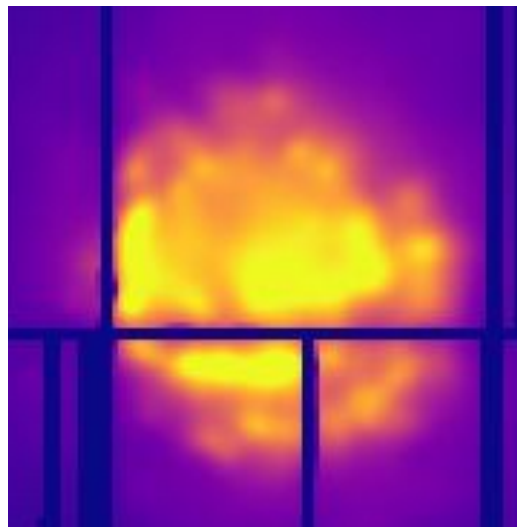
Generated
(2x, 100ks)

Source: 1E 1841-045

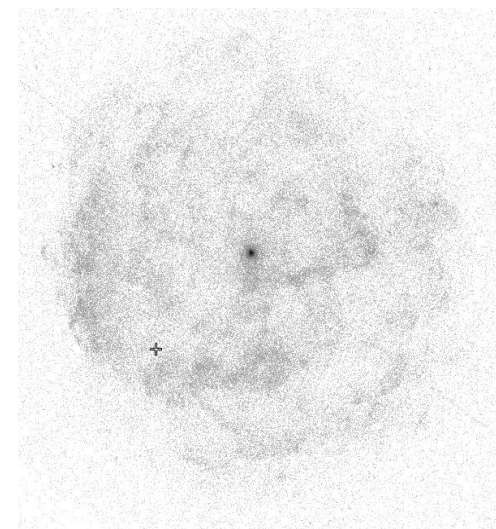
Preliminary Results SR: Real Data



Low Resolution Input
(1x, 20ks)

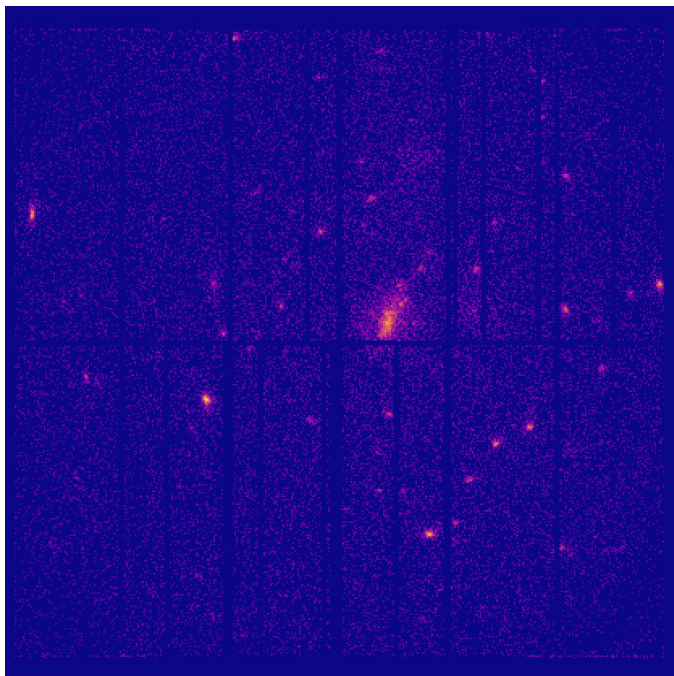


Generated
(2x, 100ks)

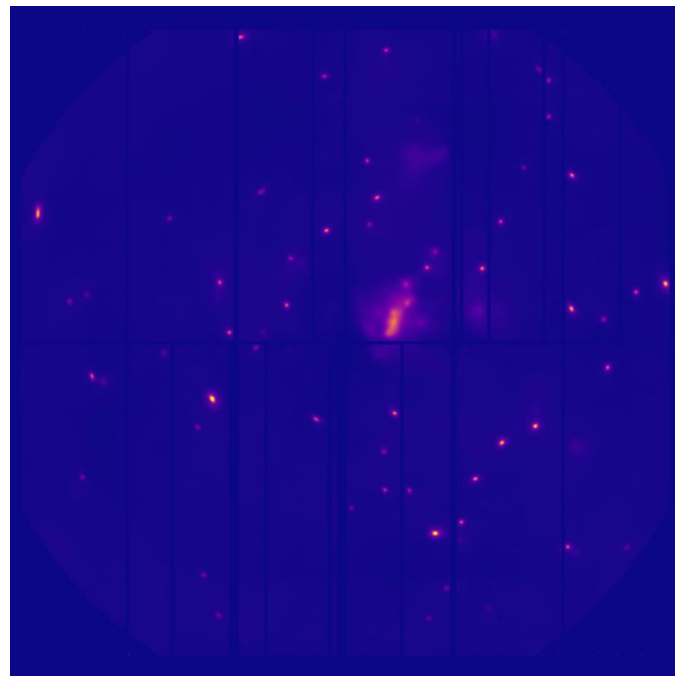


Chandra

Preliminary Results SR: Real Data



Low Resolution Input
(1x, 20ks)



Generated
(2x, 100ks)

Source: NGC 4666

Summary

- Denoising/Super-Resolution model for XMM-Newton Epic-PN
- Improved visual clarity
- Find contaminating agn sources
- Follow-up observations/further analysis
- Long training, short inference

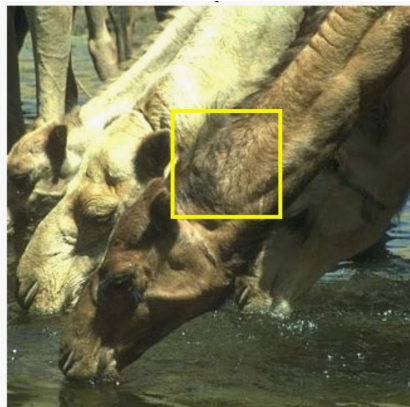
Next Steps

- Fine-Tuning on real XMM data
- Testing multiple loss functions
- More extensive model evaluation (on simulated and real data)

Any questions?

Potential Problems: Hallucinations

- Statistically validate the results
- Limited scope
- AI-Assisted



Low Resolution Input
(1x)

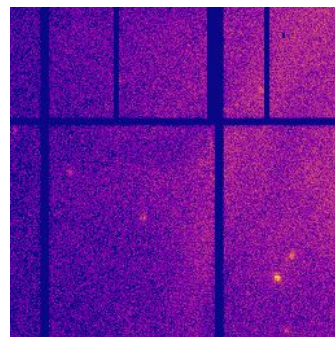
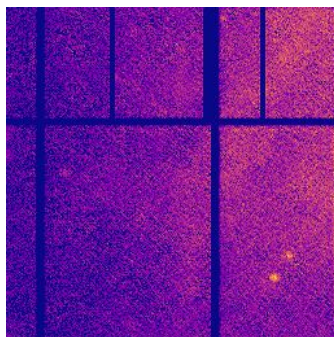
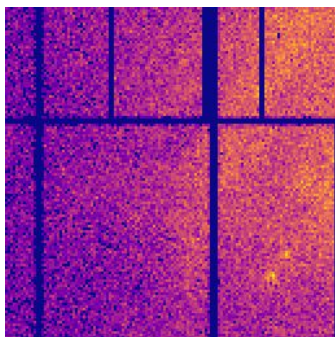
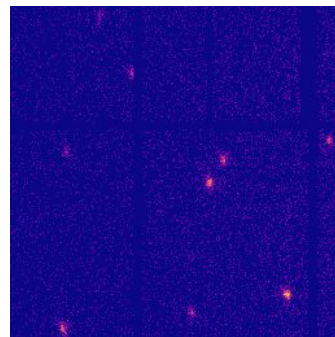
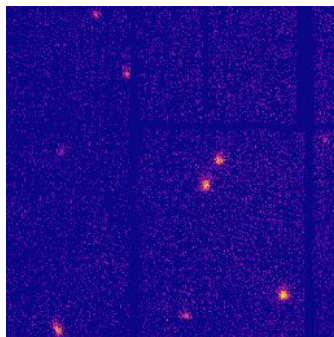
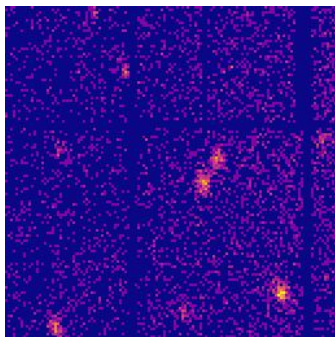


Generated
Super-Resolution
(4x)



Ground-Truth

AGN De-Blending



Low Resolution Input
(1x)

Generated
Super-Resolution
(2x)

Target